# Analytical and Numerical Methods for Solving Partial Differential Equations and Integral Equations Arising in Physical Models 

Guest Editors: Santanu Saha Ray, Om P. Agrawal, R. K. Bera, Shantanu Das, and T. Raja Sekhar

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## Abstract and Applied Analysis

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## Editorial

# Analytical and Numerical Methods for Solving Partial Differential Equations and Integral Equations Arising in Physical Models 

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Received 15 December 2013; Accepted 15 December 2013; Published 9 January 2014
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Mathematical modelling of real-life problems usually results in functional equations, like ordinary or partial differential equations, integral and integrodifferential equations, and stochastic equations. Many mathematical formulations of physical phenomena contain integrodifferential equations; these equations arise in many fields like fluid dynamics, biological models, and chemical kinetics. Partial differential equations (PDEs) have become a useful tool for describing the natural phenomena of science and engineering models. In addition, most physical phenomena of fluid dynamics, quantum mechanics, electricity, ecological systems, and many other models are controlled within their domain of validity by PDEs. Therefore, it becomes increasingly important to be familiar with all traditional and recently developed methods for solving PDEs and the implementations of these methods. Leaving aside quantum mechanics, which remains to date an inherently linear theory, most real-world physical systems, including gas dynamics, fluid mechanics, elasticity, relativity, ecology, neurology, and thermodynamics, are modelled by nonlinear partial differential equations.

The aim of this special issue is to bring together the leading researchers of dynamics, quantum mechanics, ecology, and neurology area including applied mathematicians and allow them to share their original research work. Analytical and numerical methods with advanced mathematical and real physical modelling, recent developments of PDEs, and
integral equations in physical systems are included in the main focus of the issue.

Accordingly, various papers on partial differential equations and integral equations have been included in this special issue after completing a heedful, rigorous, and peer-review process. In particular, the nonlinear hydroelastic waves propagating beneath an infinite ice sheet floating on an inviscid fluid of finite depth are investigated analytically in one of the papers. In this paper, the approximate series solutions for the velocity potential and the wave surface elevation are derived, respectively, by an analytic approximation technique named homotopy analysis method (HAM) and are presented for the second-order components.

In another paper, a domain decomposition method is proposed for the coupled stationary Navier-Stokes and Darcy equations with the Beavers-Joseph-Saffman interface condition in order to improve the efficiency of the finite element method. The physical interface conditions are directly utilized to construct the boundary conditions on the interface and then decouple the Navier-Stokes and Darcy equations. Newton iteration is used to deal with the nonlinear systems.

Another paper proposes a pressure-stabilized LagrangeGalerkin method in a parallel domain decomposition system in which the new stabilization strategy is proved to be effective for large Reynolds number and Rayleigh number simulations. The symmetry of the stiffness matrix enables
the interface problems of the linear system to be solved by the preconditioned conjugate method, and an incomplete balanced domain preconditioner is applied to the flowthermal coupled problems.

One of the papers is of use of Sumudu transform on fractional derivatives for solving some interesting nonhomogeneous fractional ordinary differential equations. Then spectral and spectral element methods have been discussed with Legendre-Gauss-Lobatto nodal basis for general 2ndorder elliptic eigenvalue problems. A priori and a posteriori error estimates for spectral and spectral element methods have been proposed. In the another paper, a generalized double sinh-Gordon equation has many more applications in various fields such as fluid dynamics, integrable quantum field theory, and kink dynamics has been solved by Expfunction method to obtain new exact solutions for this generalized double sinh-Gordon equation. A semianalytical method called the optimal homotopy asymptotic method has been also applied for solving the linear Fredholm integral equations of the first kind in another paper. In one of the papers, two strategies for inverting the open boundary conditions with adjoint method are compared by carrying out semi-idealized numerical experiments. In the first strategy, the open boundary curves are assumed to be partly space varying and are generated by linearly interpolating the values at feature points and, in the second strategy, the open boundary conditions are assumed to be fully space varying and the values at every open boundary points are taken as control variables. Another paper contains the use of a relatively new analytical method like homotopy decomposition method to solve the 2 D and 3 D Poisson equations and biharmonic equations. The method does not require the linearization or assumptions of weak nonlinearity, the solutions are generated in the form of general solution, and it is more realistic compared to the method of simplifying the physical problems.

One of the papers has shown that a strong solution of the Degasperis-Procesi equation possesses persistence property in the sense that the solution with algebraically decaying initial data and its spatial derivative must retain this property. In another paper, the fractional complex transformation has been used to transform nonlinear partial differential equations to nonlinear ordinary differential equations. The improved $\left(G^{\prime} / G\right)$-expansion method has suggested solving the space and time fractional foam drainage and KdV equations. Integral equation has been one of the essential tools for various areas of applied mathematics. For solving nonlinear Fredholm integrodifferential equations, the method based on hybrid functions approximate has been proposed in one of the papers. The properties of hybrid of block pulse functions and orthonormal Bernstein polynomials have been presented and utilized to reduce the problem to the solution of nonlinear algebraic equations. Another paper contains many numerical methods, namely, B-Spline wavelet method, Wavelet Galerkin method, and quadrature method, for solving Fredholm integral equations of second kind. A peer-review of different numerical methods for solving both linear and nonlinear Fredholm integral equations of second kind has been presented. This paper has more emphasized on
the importance of interdisciplinary effort for advancing the study on numerical methods for solving integral equations. Also one of the papers has used a numerical method like function approximation to determine the numerical solution of system of linear Volterra integrodifferential equations using Bezier curves. Two-dimensional Volterra integral equations have also been solved using more recent semianalytic method like the reduced differential transform method and also compared with the differential transform method. One of the papers has presented a numerical method to achieve the approximate solutions in a generalized expansion form of two-dimensional fractional-order Legendre functions (2DFLFs). The operational matrices of integration and derivative for 2D-FLFs have been derived.

Then a mixed finite element method has been introduced for an elliptic equation modelling of Darcy flow in porous media. In present mixed finite element, the approximate velocity is continuous and the conservation law holds locally. In order to assess the rotational potential vorticity-conserved equation with topography effect and dissipation effect, the multiple-scale method has been studied to describe the Rossby solitary waves in deep rotational fluids. A one step optimal homotopy analysis method has been applied numerically to harmonic wave propagation in a nonlinear thermoelasticity under influence of rotation, thermal relaxation times, and magnetic field. The problem has been solved in one-dimensional elastic half-space model subjected initially to a prescribed harmonic displacement and the temperature of the medium. In one of the papers, the analytical and multishaped solitary wave solutions have been presented for extended reduced Ostrovsky equation. The exact solitary (traveling) wave solutions are also expressed by three types of functions which are hyperbolic function solution, trigonometric function solution, and rational solution. In order to classify the exact solutions, including solitons and elliptic solutions, of the generalized $K(m, n)$ equation by the complete discrimination system a polynomial method has been obtained. To examine the possible approximate solutions of both integer and noninteger systems of nonlinear differential equations which describe tuberculosis disease population dynamics, the relatively new analytical technique like homotopy decomposition method has been proposed. In one of the papers, a relatively new operator called the triple Laplace transform has been introduced and to make use of the operator some kind of third-order differential equation called Mboctara equations has been solved.

Another paper investigates the effect of boundary slip on the transient pulsatile fluid flow through a vessel with body acceleration. To describe the non-Newtonian behavior, the modified second-grade fluid model has been analyzed in which the viscosity and the normal stresses have been represented in terms of the shear rate. One of the papers proves the existence of global solutions for nonlinear wave equations with damping and source terms by constructing a stable set and also obtaining the asymptotic stability of global solutions through the use of a difference inequality. In order to assess the spatial dynamical behavior of a predatorprey system with Allee effect, the bifurcation analyses have been used in which the exact Turing domain has been found
in the parameters space. According to the operator theory, the temperature dependence of the solution to the BCS gap equation has been connected with superconductivity. When the potential is a positive constant, the BCS gap equation reduces to the simple gap equation. The solution to the BCS gap equation has been indeed continuous with respect to both the temperature and the energy under a certain condition when the potential is not a constant. This study represents that there is a unique nonnegative solution to the simple gap equation, which is continuous and strictly decreasing and is of class $C^{2}$ with respect to the temperature.

At present, the use of partial differential equation and integral equation in real physical systems is commonly encountered in the fields of science and engineering. Analysis and numerical approximate of such physical models are required for efficient computational tools. The present issue has addressed recent trends and developments regarding the analytical and numerical methods that may be used in the dynamical system. Eventually, it may be expected that the present special issue would certainly helpful to explore the researchers with their new arising problems and elevate the efficiency and accuracy of the solution methods in use nowadays.

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## Research Article

# Numerical Methods for Solving Fredholm Integral Equations of Second Kind 

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Received 3 September 2013; Accepted 3 October 2013
Academic Editor: Rasajit Bera
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Integral equation has been one of the essential tools for various areas of applied mathematics. In this paper, we review different numerical methods for solving both linear and nonlinear Fredholm integral equations of second kind. The goal is to categorize the selected methods and assess their accuracy and efficiency. We discuss challenges faced by researchers in this field, and we emphasize the importance of interdisciplinary effort for advancing the study on numerical methods for solving integral equations.

## 1. Introduction

Integral equations occur naturally in many fields of science and engineering [1]. A computational approach to solve integral equation is an essential work in scientific research.

Integral equation is encountered in a variety of applications in many fields including continuum mechanics, potential theory, geophysics, electricity and magnetism, kinetic theory of gases, hereditary phenomena in physics and biology, renewal theory, quantum mechanics, radiation, optimization, optimal control systems, communication theory, mathematical economics, population genetics, queuing theory, medicine, mathematical problems of radiative equilibrium, the particle transport problems of astrophysics and reactor theory, acoustics, fluid mechanics, steady state heat conduction, fracture mechanics, and radiative heat transfer problems. Fredholm integral equation is one of the most important integral equations.

Integral equations can be viewed as equations which are results of transformation of points in a given vector spaces of integrable functions by the use of certain specific integral operators to points in the same space. If, in particular, one is concerned with function spaces spanned by polynomials for which the kernel of the corresponding transforming integral operator is separable being comprised of polynomial
functions only, then several approximate methods of solution of integral equations can be developed.

A computational approach to solving integral equation is an essential work in scientific research. Some methods for solving second kind Fredholm integral equation are available in the open literature. The $B$-spline wavelet method, the method of moments based on $B$-spline wavelets by Maleknejad and Sahlan [2], and variational iteration method (VIM) by He [3-5] have been applied to solve second kind Fredholm linear integral equations. The learned researchers Maleknejad et al. proposed some numerical methods for solving linear Fredholm integral equations system of second kind using Rationalized Haar functions method, Block-Pulse functions, and Taylor series expansion method [6-8]. Haar wavelet method with operational matrices of integration [9] has been applied to solve system of linear Fredholm integral equations of second kind. Quadrature method [10], $B$-spline wavelet method [11], wavelet Galerkin method [12], and also VIM [13] can be applied to solve nonlinear Fredholm integral equation of second kind. Some iterative methods like Homotopy perturbation method (HPM) [14-16] and Adomian decomposition method (ADM) [16-18] have been applied to solve nonlinear Fredholm integral equation of second kind.

## 2. Fredholm Integral Equation

The general form of linear Fredholm integral equation is defined as follows:

$$
\begin{equation*}
g(x) y(x)=f(x)+\lambda \int_{a}^{b} K(x, t) y(t) d t \tag{1}
\end{equation*}
$$

where $a$ and $b$ are both constants. $f(x), g(x)$, and $K(x, t)$ are known functions while $y(x)$ is unknown function. $\lambda$ (nonzero parameter) is called eigenvalue of the integral equation. The function $K(x, t)$ is known as kernel of the integral equation.
2.1. Fredholm Integral Equation of First Kind. The linear integral equation is of form (by setting $g(x)=0$ in (1))

$$
\begin{equation*}
f(x)+\lambda \int_{a}^{b} K(x, t) y(t) d t=0 \tag{2}
\end{equation*}
$$

Equation (2) is known as Fredholm integral equation of first kind.
2.2. Fredholm Integral Equation of Second Kind. The linear integral equation is of form (by setting $g(x)=1$ in (1))

$$
\begin{equation*}
y(x)=f(x)+\int_{a}^{b} K(x, t) y(t) d t \tag{3}
\end{equation*}
$$

Equation (3) is known as Fredholm integral equation of second kind.
2.3. System of Linear Fredholm Integral Equations. The general form of system of linear Fredholm integral equations of second kind is defined as follows:

$$
\begin{array}{r}
\sum_{j=1}^{n} g_{i, j} y_{j}(x)=f_{i}(x)+\sum_{j=1}^{n} \int_{a}^{b} K_{i, j}(x, t) y_{j}(t) d t  \tag{4}\\
i=1,2, \ldots, n
\end{array}
$$

where $f_{i}(x)$ and $K_{i, j}(x, t)$ are known functions and $y_{j}(x)$ are the unknown functions for $i, j=1,2, \ldots, n$.
2.4. Nonlinear Fredholm-Hammerstein Integral Equation of Second Kind. Nonlinear Fredholm-Hammerstein integral equation of second kind is defined as follows:

$$
\begin{equation*}
y(x)=f(x)+\int_{a}^{b} K(x, t) F(y(t)) d t \tag{5}
\end{equation*}
$$

where $K(x, t)$ is the kernel of the integral equation, $f(x)$ and $K(x, t)$ are known functions, and $y(x)$ is the unknown function that is to be determined.
2.5. System of Nonlinear Fredholm Integral Equations. System of nonlinear Fredholm integral equations of second kind is defined as follows:

$$
\begin{array}{r}
\sum_{j=1}^{n} g_{i, j} y_{j}(x)=f_{i}(x)+\sum_{j=1}^{n} \int_{a}^{b} K_{i, j}(x, t) F_{i, j}\left(t, y_{j}(t)\right) d t \\
i=1,2, \ldots, n \tag{6}
\end{array}
$$

where $f_{i}(x)$ and $K_{i, j}(x, t)$ are known functions and $y_{j}(x)$ are the unknown functions for $i, j=1,2, \ldots, n$.

## 3. Numerical Methods for Linear Fredholm Integral Equation of Second Kind

Consider the following Fredholm integral equation of second kind defined in (3)

$$
\begin{equation*}
y(x)=f(x)+\int_{a}^{b} K(x, t) y(t) d t, \quad a \leq x \leq b \tag{7}
\end{equation*}
$$

where $K(x, t)$ and $g(x)$ are known functions and $y(x)$ is unknown function to be determined.

### 3.1. B-Spline Wavelet Method

3.1.1. B-Spline Scaling and Wavelet Functions on the Interval $[0,1]$. Semiorthogonal wavelets using $B$-spline are specially constructed for the bounded interval and this wavelet can be represented in a closed form. This provides a compact support. Semiorthogonal wavelets form the basis in the space $L^{2}(R)$.

Using this basis, an arbitrary function in $L^{2}(R)$ can be expressed as the wavelet series. For the finite interval $[0,1]$, the wavelet series cannot be completely presented by using this basis. This is because supports of some basis are truncated at the left or right end points of the interval. Hence, a special basis has to be introduced into the wavelet expansion on the finite interval. These functions are referred to as the boundary scaling functions and boundary wavelet functions.

Let $m$ and $n$ be two positive integers and let

$$
\begin{align*}
a & =x_{-m+1}=\cdots=x_{0}<x_{1} \\
& <\cdots<x_{n}=x_{n+1}  \tag{8}\\
& =\cdots=x_{n+m-1}=b
\end{align*}
$$

be an equally spaced knots sequence. The functions

$$
\begin{align*}
& B_{m, j, X}(x)= \frac{x-x_{j}}{x_{j+m-1}-x_{j}} B_{m-1, j, X}(x) \\
&+\frac{x_{j+m}-x}{x_{j+m}-x_{j+1}} B_{m-1, j+1, X}(x),  \tag{9}\\
& j=-m+1, \ldots, n-1, \\
& B_{1, j, X}(x)= \begin{cases}1, & x \in\left[x_{j}, x_{j+1}\right), \\
0, & \text { otherwise },\end{cases}
\end{align*}
$$

are called cardinal $B$-spline functions of order $m \geq 2$ for the knot sequence $X=\left\{x_{i}\right\}_{i=-m+1}^{n+m-1}$ and $\operatorname{Supp} B_{m, j, X}(x)=$ $\left[x_{j}, x_{j+m}\right] \cap[a, b]$.

By considering the interval $[a, b]=[0,1]$, at any level $j \in$ $\mathrm{Z}^{+}$, the discretization step is $2^{-j}$, and this generates $n=2^{j}$ number of segments in $[0,1]$ with knot sequence

$$
X^{(j)}=\left\{\begin{array}{l}
x_{-m+1}^{(j)}=\cdots=x_{0}^{(j)}=0  \tag{10}\\
x_{k}^{(j)}=\frac{k}{2^{j}}, \quad k=1, \ldots, n-1 \\
x_{n}^{(j)}=\cdots=x_{n+m-1}^{(j)}=1
\end{array}\right.
$$

Let $j_{0}$ be the level for which $2^{j_{0}} \geq 2 m-1$; for each level, $j \geq j_{0}$, the scaling functions of order $m$ can be defined as follows in [2]:

$$
\begin{align*}
& \varphi_{m, j, i}(x) \\
& \quad= \begin{cases}B_{m, j_{0}, i}\left(2^{j-j_{0}} x\right) & i=-m+1, \ldots,-1, \\
B_{m, j_{0}, 2^{j}-m-i}\left(1-2^{j-j_{0}} x\right) & i=2^{j}-m+1, \ldots, 2^{j}-1, \\
B_{m, j_{0}, 0}\left(2^{j-j_{0}} x-2^{-j_{0}} i\right) & i=0, \ldots, 2^{j}-m .\end{cases} \tag{11}
\end{align*}
$$

And the two scale relations for the $m$-order semiorthogonal compactly supported $B$-wavelet functions are defined as follows:

$$
\begin{gather*}
\psi_{m, j, i-m}=\sum_{k=i}^{2 i+2 m-2} q_{i, k} B_{m, j, k-m}, \quad i=1, \ldots, m-1, \\
\psi_{m, j, i-m}=\sum_{k=2 i-m}^{2 i+2 m-2} q_{i, k} B_{m, j, k-m}, \quad i=m, \ldots, n-m+1, \\
\psi_{m, j, i-m}=\sum_{k=2 i-m}^{n+i+m-1} q_{i, k} B_{m, j, k-m}, \quad i=n-m+2, \ldots, n, \tag{12}
\end{gather*}
$$

where $q_{i, k}=q_{k-2 i}$.
Hence, there are $2(m-1)$ boundary wavelets and $(n-$ $2 m+2$ ) inner wavelets in the bounded interval $[a, b]$. Finally, by considering the level $j$ with $j \geq j_{0}$, the $B$-wavelet functions in $[0,1]$ can be expressed as follows:

$$
\begin{align*}
& \psi_{m, j, i}(x) \\
& = \begin{cases}\psi_{m, j_{0}, i}\left(2^{j-j_{0}} x\right) & i=-m+1, \ldots,-1, \\
\psi_{m, 2^{j}-2 m+1-i, i}\left(1-2^{j-j_{0}} x\right) & i=2^{j}-2 m+2, \ldots, 2^{j}-m, \\
\psi_{m, j_{0}, 0}\left(2^{j-j_{0}} x-2^{-j_{0}} i\right) & i=0, \ldots, 2^{j}-2 m+1 .\end{cases} \tag{13}
\end{align*}
$$

The scaling functions $\varphi_{m, j, i}(x)$ occupy $m$ segments and the wavelet functions $\psi_{m, j, i}(x)$ occupy $2 m-1$ segments.

When the semiorthogonal wavelets are constructed from $B$-spline of order $m$, the lowest octave level $j=j_{0}$ is determined in $[19,20]$ by

$$
\begin{equation*}
2^{j_{0}} \geq 2 m-1 \tag{14}
\end{equation*}
$$

so as to have a minimum of one complete wavelet on the interval $[0,1]$.
3.1.2. Function Approximation. A function $f(x)$ defined over [ 0,1 ] may be approximated by $B$-spline wavelets as $[21,22$ ]

$$
\begin{align*}
f(x)= & \sum_{k=1-m}^{2^{j_{0}}-1} c_{j_{0}, k} \varphi_{j_{0}, k}(x) \\
& +\sum_{j=j_{0}}^{\infty} \sum_{k=1-m}^{2^{j}-m} d_{j, k} \psi_{j, k}(x) . \tag{15}
\end{align*}
$$

If the infinite series in (15) is truncated at $M$, then (15) can be written as [2]

$$
\begin{align*}
f(x) \cong & \sum_{k=1-m}^{2^{j_{0}}-1} c_{j_{0}, k} \varphi_{j_{0}, k}(x) \\
& +\sum_{j=j_{0}}^{M} \sum_{k=1-m}^{2^{j}-m} d_{j, k} \psi_{j, k}(x) \tag{16}
\end{align*}
$$

where $\varphi_{2, k}$ and $\psi_{j, k}$ are scaling and wavelets functions, respectively, and $C$ and $\Psi$ are $\left(2^{M+1}+m-1\right) \times 1$ vectors given by

$$
\begin{align*}
& C=\left[c_{j_{0}, 1-m}, \ldots, c_{j_{0}, 2^{j_{0}-1}}, d_{j_{0}, 1-m}, \ldots\right. \\
&  \tag{17}\\
& \left.\quad d_{j_{0}, 2^{j_{0}-m}}, \ldots, d_{M, 1-m}, \ldots, d_{M, 2^{M}-m}\right]^{T} \\
& \Psi=\left[\varphi_{j_{0}, 1-m}, \ldots, \varphi_{j_{0}, 2^{j_{0}-1}}, \psi_{j_{0}, 1-m}, \ldots\right.  \tag{18}\\
& \\
& \left.\quad \psi_{j_{0}, 2^{j_{0}-m}}, \ldots, \psi_{M, 1-m}, \ldots, \psi_{M, 2^{M}-m}\right]^{T}
\end{align*}
$$

with

$$
\begin{gather*}
c_{j_{0}, k}=\int_{0}^{1} f(x) \widetilde{\varphi}_{j_{0}, k}(x) d x, \quad k=1-m, \ldots, 2^{j_{0}}-1, \\
d_{j, k}=\int_{0}^{1} f(x) \widetilde{\psi}_{j, k}(x) d x,  \tag{19}\\
j=j_{0}, \ldots, M, \quad k=1-m, \ldots, 2^{M}-m,
\end{gather*}
$$

where $\widetilde{\varphi}_{j_{0}, k}(x)$ and $\widetilde{\psi}_{j, k}(x)$ are dual functions of $\varphi_{j_{0}, k}$ and $\psi_{j, k}$, respectively. These can be obtained by linear combinations of $\varphi_{j_{0}, k}, k=1-m, \ldots, 2^{j_{0}}-1$, and $\psi_{j, k}, j=j_{0}, \ldots, M, k=$ $1-m, \ldots, 2^{M}-m$, as follows. Let

$$
\begin{equation*}
\Phi=\left[\varphi_{j_{0}, 1-m}, \ldots, \varphi_{j_{0}, 2^{j_{0}-1}}\right]^{T} \tag{20}
\end{equation*}
$$

$$
\begin{equation*}
\bar{\Psi}=\left[\psi_{j_{0}, 1-m}, \ldots, \psi_{j_{0}, 2^{j_{0}-m}}, \ldots, \psi_{M, 1-m}, \ldots, \psi_{M, 2^{M}-m}\right]^{T} \tag{21}
\end{equation*}
$$

Using (11), (20), (12)-(13), and (21), we get

$$
\begin{align*}
& \int_{0}^{1} \Phi \Phi^{T} d x=P_{1}  \tag{22}\\
& \int_{0}^{1} \bar{\Psi} \bar{\Psi}^{T} d x=P_{2}
\end{align*}
$$

Suppose that $\widetilde{\Phi}$ and $\widetilde{\bar{\Psi}}$ are the dual functions of $\Phi$ and $\bar{\Psi}$, respectively; then

$$
\begin{gather*}
\int_{0}^{1} \widetilde{\Phi} \Phi^{T} d x=I_{1}, \\
\int_{0}^{1} \widetilde{\Psi} \bar{\Psi}^{T} d x=I_{2},  \tag{23}\\
\widetilde{\Phi}=P_{1}^{-1} \Phi, \\
\widetilde{\bar{\Psi}}=P_{2}^{-1} \bar{\Psi} . \tag{24}
\end{gather*}
$$

3.1.3. Application of B-Spline Wavelet Method. In this section, linear Fredholm integral equation of the second kind of form (7) has been solved by using $B$-spline wavelets. For this, we use (16) to approximate $y(x)$ as

$$
\begin{equation*}
y(x)=C^{T} \Psi(x) \tag{25}
\end{equation*}
$$

where $\Psi(x)$ is defined in (18) and $C$ is $\left(2^{M+1}+m-1\right) \times 1$ unknown vector defined similarly as in (17). We also expand $f(x)$ and $K(x, t)$ by $B$-spline dual wavelets $\widetilde{\Psi}$ defined in (24) as

$$
\begin{gather*}
f(x)=C_{1}^{T} \widetilde{\Psi}(x), \\
K(x, t)=\widetilde{\Psi}^{T}(t) \Theta \widetilde{\Psi}(x), \tag{26}
\end{gather*}
$$

where

$$
\begin{equation*}
\Theta_{i, j}=\int_{0}^{1}\left[\int_{0}^{1} K(x, t) \Psi_{i}(t) d t\right] \Psi_{j}(x) d x \tag{27}
\end{equation*}
$$

From (26) and (25), we get

$$
\begin{align*}
\int_{0}^{1} K(x, t) y(t) d t & =\int_{0}^{1} C^{T} \Psi(t) \widetilde{\Psi}^{T}(t) \Theta \widetilde{\Psi}(x) d t  \tag{28}\\
& =C^{T} \Theta \widetilde{\Psi}(x)
\end{align*}
$$

since

$$
\begin{equation*}
\int_{0}^{1} \Psi(t) \widetilde{\Psi}^{T}(t) d t=I \tag{29}
\end{equation*}
$$

By applying (25)-(28) in (7) we have

$$
\begin{equation*}
C^{T} \Psi(x)-C^{T} \Theta \widetilde{\Psi}(x)=C_{1}^{T} \widetilde{\Psi}(x) \tag{30}
\end{equation*}
$$

By multiplying both sides of (30) with $\Psi^{T}(x)$ from the right and integrating both sides with respect to $x$ from 0 to 1 , we get

$$
\begin{equation*}
C^{T} P-C^{T} \Theta=C_{1}^{T} \tag{31}
\end{equation*}
$$

since

$$
\begin{equation*}
\int_{0}^{1} \widetilde{\Psi}(x) \Psi^{T}(x) d x=I \tag{32}
\end{equation*}
$$

and $P$ is a $\left(2^{M+1}+m-1\right) \times\left(2^{M+1}+m-1\right)$ square matrix given by

$$
P=\int_{0}^{1} \Psi(x) \Psi^{T}(x) d x=\left(\begin{array}{rr}
P_{1} & 0  \tag{33}\\
0 & P_{2}
\end{array}\right) .
$$

Consequently, from (31), we get $C^{T}=C_{1}^{T}(P-\Theta)^{-1}$. Hence, we can calculate the solution for $y(x)=C^{T} \Psi(x)$.

### 3.2. Method of Moments

3.2.1. Multiresolution Analysis (MRA) and Wavelets [2]. A set of subspaces $\left\{V_{j}\right\}_{j \in Z}$ is said to be MRA of $L^{2}(R)$ if it possesses the following properties:

$$
\begin{gather*}
V_{j} \subset V_{j+1}, \quad \forall j \in \mathrm{Z}  \tag{34}\\
\bigcup_{j \in \mathrm{Z}} V_{j} \text { is dense in } L^{2}(R),  \tag{35}\\
\bigcap_{j \in \mathrm{Z}} V_{j}=\phi,  \tag{36}\\
f(x) \in V_{j} \Longleftrightarrow f(2 x) \in V_{j+1}, \quad \forall j \in \mathrm{Z} \tag{37}
\end{gather*}
$$

where $Z$ denotes the set of integers. Properties (34)-(36) state that $\left\{V_{j}\right\}_{j \in Z}$ is a nested sequence of subspaces that effectively covers $L^{2}(R)$. That is, every square integrable function can be approximated as closely as desired by a function that belongs to at least one of the subspaces $V_{j}$. A function $\varphi \in L^{2}(R)$ is called a scaling function if it generates the nested sequence of subspaces $V_{j}$ and satisfies the dilation equation; namely,

$$
\begin{equation*}
\varphi(x)=\sum_{k} p_{k} \varphi(a x-k), \tag{38}
\end{equation*}
$$

with $p_{k} \in l^{2}$ and $a$ being any rational number.
For each scale $j$, since $V_{j} \subset V_{j+1}$, there exists a unique orthogonal complementary subspace $W_{j}$ of $V_{j}$ in $V_{j+1}$. This subspace $W_{j}$ is called wavelet subspace and is generated by $\psi_{j, k}=\psi\left(2^{j} x-k\right)$, where $\psi \in L^{2}$ is called the wavelet. From the above discussion, these results follow easily:

$$
\begin{array}{cl}
V_{j_{1}} \cap V_{j_{2}}=V_{j_{2}}, & j_{1}>j_{2} \\
W_{j_{1}} \cap W_{j_{2}}=0, & j_{1} \neq j_{2},  \tag{39}\\
V_{j_{1}} \cap W_{j_{2}}=0, & j_{1} \leq j_{2} .
\end{array}
$$

Some of the important properties relevant to the present analysis are given below [2, 19].
(1) Vanishing Moment. A wavelet is said to have a vanishing moment of order $m$ if

$$
\begin{equation*}
\int_{-\infty}^{\infty} x^{p} \psi(x) d x=0 ; \quad p=0, \ldots, m-1 \tag{40}
\end{equation*}
$$

All wavelets must satisfy the previously mentioned condition for $p=0$.
(2) Semiorthogonality. The wavelets $\psi_{j, k}$ form a semiorthogonal basis if

$$
\begin{equation*}
\left\langle\psi_{j, k}, \psi_{s, i}\right\rangle=0 ; \quad j \neq s ; \quad \forall j, k, s, i \in \mathrm{Z} . \tag{41}
\end{equation*}
$$

3.2.2. Method of Moments for the Solution of Fredholm Integral Equation. In this section, we solve the integral equation of form (7) in interval $[0,1]$ by using linear $B$-spline wavelets [2]. The unknown function in (7) can be expanded in terms of the scaling and wavelet functions as follows:

$$
\begin{align*}
y(x) \approx & \sum_{k=-1}^{2^{j_{0}}-1} c_{k} \varphi_{j_{0}, k}(x) \\
& +\sum_{j=j_{0}}^{M} \sum_{k=-1}^{2^{j}-2} d_{j, k} \psi_{j, k}(x)  \tag{42}\\
= & C^{T} \Psi(x) .
\end{align*}
$$

By substituting this expression into (7) and employing the Galerkin method, the following set of linear system of order $\left(2^{M}+1\right)$ is generated. The scaling and wavelet functions are used as testing and weighting functions:

$$
\left(\begin{array}{cc}
\langle\varphi, \varphi\rangle-\langle K \varphi, \varphi\rangle & \langle\psi, \varphi\rangle-\langle K \psi, \varphi\rangle  \tag{43}\\
\langle\varphi, \psi\rangle-\langle K \varphi, \psi\rangle & \langle\psi, \psi\rangle-\langle K \psi, \psi\rangle
\end{array}\right)\binom{C}{D}=\binom{F_{1}}{F_{2}},
$$

where

$$
\begin{gathered}
C=\left[c_{-1}, c_{0}, \ldots, c_{3}\right]^{T}, \\
D=\left[d_{2,-1}, \ldots, d_{2,2}, d_{3,-1}, \ldots, d_{3,6}, \ldots,\right. \\
\left.d_{M,-1}, \ldots, d_{M, 2^{M}-2}\right]^{T}, \\
\langle\varphi, \varphi\rangle-\langle K \varphi, \varphi\rangle \\
=\left(\int_{0}^{1} \varphi_{j_{0}, r}(x) \varphi_{j_{0}, i}(x) d x\right. \\
\langle\psi, \varphi\rangle-\langle K \psi, \varphi\rangle \\
=\left(\int_{0}^{1} \varphi_{j_{0}, r}(x) \int_{0}^{1} K(x, t) \varphi_{j_{0}, i}(t) d t d x\right)_{i, r}, \\
\langle\varphi, \psi\rangle-\langle K \varphi, \psi\rangle \\
\left.\quad-\int_{0}^{1} \varphi_{j_{0}, r}(x) \int_{0}^{1} K(x, t) \psi_{k, j}(t) d t d x\right)_{r, k, j}, j \\
=\left(\int_{0}^{1} \psi_{s, l}(x) \varphi_{j_{0}, i}(x) d x\right. \\
\left.\quad-\int_{0}^{1} \psi_{s, l}(x) \int_{0}^{1} K(x, t) \varphi_{j_{0}, i}(t) d t d x\right)_{i, l, s},
\end{gathered}
$$

$$
\begin{gather*}
\langle\psi, \psi\rangle-\langle K \psi, \psi\rangle \\
=\left(\int_{0}^{1} \psi_{s, l}(x) \psi_{k, j}(x) d x\right. \\
\left.-\int_{0}^{1} \psi_{s, l}(x) \int_{0}^{1} K(x, t) \psi_{k, j}(t) d t d x\right)_{l, s, k, j}, \\
F_{1}=\int_{0}^{1} f(x) \varphi_{j_{0}, r}(x) d x \\
F_{2}=\int_{0}^{1} f(x) \psi_{s, l}(x) d x \tag{44}
\end{gather*}
$$

and the subscripts $i, r, k, j, l$, and $s$ assume values as given below:

$$
\begin{gather*}
i, r=-1, \ldots, 2^{j_{0}}-1 \\
l, k=j_{0}, \ldots, M  \tag{45}\\
s, j=-1, \ldots, 2^{M}-2
\end{gather*}
$$

In fact, the entries with significant magnitude are in the $\langle K \varphi, \varphi\rangle-\langle\varphi, \varphi\rangle$ and $\langle K \psi, \psi\rangle-\langle\psi, \psi\rangle$ submatrices which are of order $\left(2^{j_{0}}+1\right)$ and $\left(2^{M+1}+1\right)$, respectively.
3.3. Variational Iteration Method [3-5]. In this section, Fredholm integral equation of second kind given in (7) has been considered for solving (7) by variational iteration method. First, we have to take the partial derivative of (7) with respect to $x$ yielding

$$
\begin{equation*}
Y^{\prime}(x)=f^{\prime}(x)+\int_{0}^{1} K^{\prime}(x, t) y(t) d t \tag{46}
\end{equation*}
$$

We apply variation iteration method for (46). According to this method, correction functional can be defined as

$$
\begin{align*}
& y_{n+1}(x) \\
& \quad=y_{n}(x) \\
& \quad+\int_{0}^{x} \lambda(\xi)\left(y_{n}^{\prime}(\xi)-f^{\prime}(\xi)-\int_{a}^{b} K^{\prime}(\xi, t) \tilde{y}_{n}(t) d t\right) d \xi \tag{47}
\end{align*}
$$

where $\lambda(\xi)$ is a general Lagrange multiplier which can be identified optimally by the variational theory, the subscript $n$ denotes the $n$th order approximation, and $\widetilde{y}_{n}$ is considered as a restricted variation; that is, $\delta \tilde{y}_{n}=0$. The successive approximations $y_{n}(x), n \geq 1$ for the solution $y(x)$ can be readily obtained after determining the Lagrange multiplier and selecting an appropriate initial function $y_{0}(x)$. Consequently the approximate solution may be obtained by using

$$
\begin{equation*}
y(x)=\lim _{n \rightarrow \infty} y_{n}(x) \tag{48}
\end{equation*}
$$

To make the above correction functional stationary, we have

$$
\begin{align*}
\delta y_{n+1}(x)= & \delta y_{n}(x) \\
& +\delta \int_{0}^{x} \lambda(\xi)\left(y_{n}^{\prime}(\xi)-f^{\prime}(\xi)\right. \\
& \left.\quad-\int_{a}^{b} K^{\prime}(\xi, t) \tilde{y}_{n}(t) d t\right) d \xi \\
= & \delta y_{n}(x)+\int_{0}^{x} \lambda(\xi) \delta\left(y_{n}^{\prime}(\xi)\right) d \xi \\
= & \delta y_{n}(x)+\left.\lambda \delta y_{n}\right|_{\xi=x}-\int_{0}^{x} \lambda^{\prime}(\xi) \delta y_{n}(\xi) d \xi . \tag{49}
\end{align*}
$$

Under stationary condition,

$$
\begin{equation*}
\delta y_{n+1}=0 \tag{50}
\end{equation*}
$$

implies the following Euler Lagrange equation:

$$
\begin{equation*}
\lambda^{\prime}(\xi)=0, \tag{51}
\end{equation*}
$$

with the following natural boundary condition:

$$
\begin{equation*}
1+\left.\lambda(\xi)\right|_{\xi=x}=0 . \tag{52}
\end{equation*}
$$

Solving (51), along with boundary condition (52), we get the general Lagrange multiplier

$$
\begin{equation*}
\lambda=-1 . \tag{53}
\end{equation*}
$$

Substituting the identified Lagrange multiplier into (47) results in the following iterative scheme:

$$
\begin{align*}
y_{n+1}(x)= & y_{n}(x) \\
& -\int_{0}^{x}\left(y_{n}^{\prime}(\xi)-f^{\prime}(\xi)-\int_{a}^{b} K^{\prime}(\xi, t) \widetilde{y}_{n}(t) d t\right) d \xi \\
n & \geq 0 \tag{54}
\end{align*}
$$

By starting with initial approximate function $y_{0}(x)=f(x)$ (say), we can determine the approximate solution $y(x)$ of (7).

## 4. Numerical Methods for System of Linear Fredholm Integral Equations of Second Kind

Consider the system of linear Fredholm integral equations of second kind of the following form:

$$
\begin{array}{r}
\sum_{j=1}^{n} y_{j}(x)=f_{i}(x)+\sum_{j=1}^{n} \int_{0}^{1} K_{i, j}(x, t) y_{j}(t) d t,  \tag{55}\\
i=1,2, \ldots, n,
\end{array}
$$

where $f_{i}(x)$ and $K_{i, j}(x, t)$ are known functions and $y_{j}(x)$ are the unknown functions for $i, j=1,2, \ldots, n$.
4.1. Application of Haar Wavelet Method [9]. In this section, an efficient algorithm for solving Fredholm integral equations with Haar wavelets is analyzed. The present algorithm takes the following essential strategy. The Haar wavelet is first used to decompose integral equations into algebraic systems of linear equations, which are then solved by collocation methods.
4.1.1. Haar Wavelets. The compact set of scale functions is chosen as

$$
h_{0}= \begin{cases}1, & 0 \leq x<1  \tag{56}\\ 0, & \text { others }\end{cases}
$$

The mother wavelet function is defined as

$$
h_{1}(x)= \begin{cases}1, & 0 \leq x<\frac{1}{2}  \tag{57}\\ -1, & \frac{1}{2} \leq x<1 \\ 0, & \text { others }\end{cases}
$$

The family of wavelet functions generated by translation and dilation of $h_{1}(x)$ are given by

$$
\begin{equation*}
h_{n}(x)=h_{1}\left(2^{j} x-k\right) \tag{58}
\end{equation*}
$$

where $n=2^{j}+k, j \geq 0,0 \leq k<2^{j}$.
Mutual orthogonalities of all Haar wavelets can be expressed as

$$
\int_{0}^{1} h_{m}(x) h_{n}(x) d x=2^{-j} \delta_{m n}= \begin{cases}2^{-j}, & m=n=2^{j}+k  \tag{59}\\ 0, & m \neq n\end{cases}
$$

4.1.2. Function Approximation. An arbitrary function $y(x) \in$
$L^{2}[0,1)$ can be expanded into the following Haar series:

$$
\begin{equation*}
y(x)=\sum_{n=0}^{+\infty} c_{n} h_{n}(x), \tag{60}
\end{equation*}
$$

where the coefficients $c_{n}$ are given by

$$
\begin{array}{r}
c_{n}=2^{j} \int_{0}^{1} y(x) h_{n}(x) d x,  \tag{61}\\
n=2^{j}+k, \quad j \geq 0, \quad 0 \leq k<2^{j} .
\end{array}
$$

In particular, $c_{0}=\int_{0}^{1} y(x) d x$.
The previously mentioned expression in (60) can be approximately represented with finite terms as follows:

$$
\begin{equation*}
y(x) \approx \sum_{n=0}^{m-1} c_{n} h_{n}(x)=C_{(m)}^{T} h_{(m)}(x), \tag{62}
\end{equation*}
$$

where the coefficient vector $C_{(m)}^{T}$ and the Haar function vector $h_{(m)}(x)$ are, respectively, defined as

$$
\begin{gather*}
C_{(m)}^{T}=\left[c_{0}, c_{1}, \ldots, c_{m-1}\right], \quad m=2^{j}, \\
h_{(m)}(x)=\left[h_{0}(x), h_{1}(x), \ldots, h_{m-1}(x)\right]^{T}, \quad m=2^{j} \tag{63}
\end{gather*}
$$

The Haar expansion for function $K(x, t)$ of order $m$ is defined as follows:

$$
\begin{equation*}
K(x, t) \approx \sum_{u=0}^{m-1} \sum_{v=0}^{m-1} a_{u v} h_{v}(x) h_{u}(t) \tag{64}
\end{equation*}
$$

where $a_{u v}=2^{i+q} \iint_{0}^{1} K(x, t) h_{v}(x) h_{u}(t) d x d t, u=2^{i}+j, v=$ $2^{q}+r, i, q \geq 0$.

From (62) and (64), we obtain

$$
\begin{equation*}
K(x, t) \approx h_{(m)}^{T}(t) K h_{(m)}(x), \tag{65}
\end{equation*}
$$

where

$$
\begin{equation*}
K=\left(a_{u v}\right)_{m \times m}^{T} \tag{66}
\end{equation*}
$$

4.1.3. Operational Matrices of Integration. We define

$$
\begin{equation*}
H_{(m)}=\left[h_{(m)}\left(\frac{1}{2 m}\right), h_{(m)}\left(\frac{3}{2 m}\right), \ldots, h_{(m)}\left(\frac{2 m-1}{2 m}\right)\right] \tag{67}
\end{equation*}
$$

where $H_{(1)}=[1], H_{(2)}=\left[\begin{array}{cc}1 & 1 \\ 1 & -1\end{array}\right]$.
Then, for $m=4$, the corresponding matrix can be represented as

$$
\begin{align*}
H_{(4)} & =\left[h_{(4)}\left(\frac{1}{8}\right), h_{(4)}\left(\frac{3}{8}\right), \ldots, h_{(4)}\left(\frac{7}{8}\right)\right] \\
& =\left[\begin{array}{cccc}
1 & 1 & 1 & 1 \\
1 & 1 & -1 & -1 \\
1 & -1 & 0 & 0 \\
0 & 0 & 1 & -1
\end{array}\right] . \tag{68}
\end{align*}
$$

The integration of the Haar function vector $h_{(m)}(t)$ is

$$
\begin{array}{r}
\int_{0}^{x} h_{(m)}(t) d t=P_{(m)} h_{(m)}(x),  \tag{69}\\
x \in[0,1),
\end{array}
$$

where $P_{(m)}$ is the operational matrix of order $m$, and

$$
\begin{gather*}
P_{(1)}=\left[\frac{1}{2}\right] \\
P_{(m)}=\frac{1}{2 m}\left[\begin{array}{cc}
2 m P_{(m / 2)} & -H_{(m / 2)} \\
H_{(m / 2)}^{-1} & 0
\end{array}\right] . \tag{70}
\end{gather*}
$$

By recursion of the above formula, we obtain

$$
\begin{gather*}
P_{(2)}=\frac{1}{4}\left[\begin{array}{cc}
2 & -1 \\
1 & 0
\end{array}\right], \\
P_{(4)}=\frac{1}{16}\left[\begin{array}{cccc}
8 & -4 & -2 & -2 \\
4 & 0 & -2 & 2 \\
1 & 1 & 0 & 0 \\
1 & -1 & 0 & 0
\end{array}\right], \\
P_{(8)}=\frac{1}{64}\left[\begin{array}{cccccccc}
32 & -16 & -8 & -8 & -4 & -4 & -4 & -4 \\
16 & 0 & -8 & 8 & -4 & -4 & 4 & 4 \\
4 & 4 & 0 & 0 & -4 & 4 & 0 & 0 \\
4 & 4 & 0 & 0 & 0 & 0 & -4 & 4 \\
1 & 1 & 2 & 0 & 0 & 0 & 0 & 0 \\
1 & 1 & -2 & 0 & 0 & 0 & 0 & 0 \\
1 & -1 & 0 & 2 & 0 & 0 & 0 & 0 \\
1 & -1 & 0 & -2 & 0 & 0 & 0 & 0
\end{array}\right] . \tag{71}
\end{gather*}
$$

Therefore, we get

$$
\begin{aligned}
H_{(m)}^{-1}= & \left(\frac{1}{m}\right) H_{(m)}^{T} \\
& \times \operatorname{diag}(1,1,2,2, \underbrace{2^{2}, \ldots, 2^{2}}_{2^{2}}, \ldots, \underbrace{2^{\alpha-1}, \ldots, 2^{\alpha-1}}_{2^{\alpha-1}})
\end{aligned}
$$

where $m=2^{\alpha}$ and $\alpha$ is a positive integer.
The inner product of two Haar functions can be represented as

$$
\begin{equation*}
\int_{0}^{1} h_{(m)}(t) h_{(m)}^{T}(t) d t=D \tag{73}
\end{equation*}
$$

where

$$
\begin{gather*}
D=\operatorname{diag}(1,1,1 / 2,1 / 2, \underbrace{1 / 2^{2}, \ldots, 1 / 2^{2}}_{2^{2}}, \ldots,  \tag{74}\\
\\
\left.\frac{1 / 2^{\alpha-1}, \ldots, 1 / 2^{\alpha-1}}{2^{\alpha-1}}\right) .
\end{gather*}
$$

4.1.4. Haar Wavelet Solution for Fredholm Integral Equations System [9]. Consider the following Fredholm integral equations system defined in (55):

$$
\begin{array}{r}
\sum_{j=1}^{m} y_{j}(x)=f_{i}(x)+\sum_{j=1}^{m} \int_{0}^{1} K_{i, j}(x, t) y_{j}(t) d t  \tag{75}\\
i=1,2, \ldots, m
\end{array}
$$

The Haar series of $y_{j}(x)$ and $K_{i, j}(x, t), i=1,2, \ldots, m ; j=$ $1,2, \ldots, m$ are, respectively, expanded as

$$
\begin{gather*}
y_{j}(x) \approx Y_{j}^{T} h_{(m)}(x), \quad j=1,2, \ldots, m \\
K_{i, j}(x, t) \approx h_{(m)}^{T}(t) K_{i, j} h_{(m)}(x)  \tag{76}\\
i, j=1,2, \ldots, m
\end{gather*}
$$

Substituting (76) into (75), we get

$$
\begin{align*}
& \sum_{j=1}^{m} Y_{j}^{T} h_{(m)}(x) \\
& =f_{i}(x)+\sum_{j=1}^{m} \int_{0}^{1} Y_{j}^{T} h_{(m)}(t) h_{(m)}^{T}(t) K_{i, j} h_{(m)}(x) d t \\
&  \tag{77}\\
& i=1,2, \ldots, m .
\end{align*}
$$

From (77) and (73), we get

$$
\begin{array}{r}
\sum_{j=1}^{m} Y_{j}^{T} h_{(m)}(x)=f_{i}(x)+\sum_{j=1}^{m} Y_{j}^{T} D K_{i, j} h_{(m)}(x),  \tag{78}\\
i=1,2, \ldots, m .
\end{array}
$$

Interpolating $m$ collocation points, that is, $\left\{x_{i}\right\}_{i=1}^{m}$, in the interval $[0,1]$ leads to the following algebraic system of equations:

$$
\begin{array}{r}
\sum_{j=1}^{m} Y_{j}^{T} h_{(m)}\left(x_{i}\right)=f_{i}\left(x_{i}\right)+\sum_{j=1}^{m} Y_{j}^{T} D K_{i, j} h_{(m)}\left(x_{i}\right)  \tag{79}\\
i=1,2, \ldots, m
\end{array}
$$

Hence, $Y_{j}, j=1,2, \ldots, m$ can be computed by solving the above algebraic system of equations and consequently the solutions $y_{j}(x) \approx Y_{j}^{T} h_{(m)}(x), j=1,2, \ldots, m$.
4.2. Taylor Series Expansion Method. In this section, we present Taylor series expansion method for solving Fredholm integral equations system of second kind [7]. This method reduces the system of integral equations to a linear system of ordinary differential equation. After including boundary conditions, this system reduces to a system of equations that can be solved easily by any usual methods.

Consider the second kind Fredholm integral equations system defined in (55) as follows:

$$
\begin{gather*}
y_{i}(x)=f_{i}(x)+\sum_{j=1}^{n} \int_{0}^{1} K_{i, j}(x, t) y_{j}(t) d t  \tag{80}\\
i=1,2, \ldots, n, \quad 0 \leq x \leq 1 .
\end{gather*}
$$

A Taylor series expansion can be made for the solution of $y_{j}(t)$ in the integral equation (80):

$$
\begin{align*}
y_{j}(t)= & y_{j}(x)+y_{j}^{\prime}(x)(t-x)+\cdots \\
& +\frac{1}{m!} y_{j}^{(m)}(x)(t-x)^{m}+E(t) \tag{81}
\end{align*}
$$

where $E(t)$ denotes the error between $y_{j}(t)$ and its Taylor series expansion in (81).

If we use the first $m$ term of Taylor series expansion and neglect the term containing $E(t)$, that is,
$\int_{0}^{1} \sum_{j=1}^{n} K_{i, j}(x, t) E(t) d t$, then, substituting (81) for $y_{j}(t)$ into the integral in (80), we have

$$
\begin{align*}
& y_{i}(x) \approx f_{i}(x) \\
& \quad+\sum_{j=1}^{n} \int_{0}^{1} K_{i, j}(x, t) \sum_{r=0}^{m} \frac{1}{r!}(t-x)^{r} y_{j}^{(r)}(x) d t, \\
& \quad i=1,2, \ldots, n, \\
& y_{i}(x) \approx f_{i}(x)  \tag{82}\\
& +\sum_{j=1}^{n} \sum_{r=0}^{m} \frac{1}{r!} y_{j}^{(r)}(x) \int_{0}^{1} K_{i, j}(x, t)(t-x)^{r} d t, \\
& \quad i=1,2, \ldots, n, \\
& y_{i}(x)-\sum_{j=1}^{n} \sum_{r=0}^{m} \frac{1}{r!} y_{j}^{(r)}(x)\left[\int_{0}^{1} K_{i, j}(x, t)(t-x)^{r} d t\right] \\
& \approx f_{i}(x), \quad i=1,2, \ldots, n . \tag{83}
\end{align*}
$$

Equation (83) becomes a linear system of ordinary differential equations that we have to solve. For solving the linear system of ordinary differential equations (83), we require an appropriate number of boundary conditions.

In order to construct boundary conditions, we first differentiate $s$ times both sides of (80) with respect to $x$; that is,

$$
\begin{gather*}
y_{i}^{(s)}(x)=f_{i}^{(s)}(x)+\sum_{j=1}^{n} \int_{0}^{1} K_{i, j}^{(s)}(x, t) y_{j}(t) d t  \tag{84}\\
i=1,2, \ldots, n, \quad s=1,2, \ldots, m
\end{gather*}
$$

where $K_{i, j}^{(s)}(x, t)=\partial^{(s)} K_{i, j}(x, t) / \partial x^{(s)}, s=1,2, \ldots, m$.
Applying the mean value theorem for integral in (84), we have

$$
\begin{gather*}
y_{i}^{(s)}(x)-\left[\sum_{j=1}^{n} \int_{0}^{1} K_{i, j}^{(s)}(x, t) d t\right] y_{j}(x) \approx f_{i}^{(s)}(x)  \tag{85}\\
i=1,2, \ldots, n, \quad s=1,2, \ldots, m
\end{gather*}
$$

Now (83) combined with (85) becomes a linear system of algebraic equations that can be solved analytically or numerically.

### 4.3. Block-Pulse Functions for the Solution of Fredholm Integral

 Equation. In this section, Block-Pulse functions (BPF) have been utilized for the solution of system of Fredholm integral equations [6].An $m$-set of BPF is defined as follows:

$$
\Phi_{i}(t)= \begin{cases}1, & (i-1) \frac{T}{m} \leq t<i \frac{T}{m},  \tag{86}\\ 0, & \text { otherwise }\end{cases}
$$

with $t \in[0, T), T / m=h$ and $i=1,2, \ldots, m$.

### 4.3.1. Properties of BPF

(1) Disjointness. One has

$$
\Phi_{i}(t) \Phi_{j}(t)= \begin{cases}\Phi_{i}(t), & i=j  \tag{87}\\ 0, & i \neq j\end{cases}
$$

$, i, j=1,2, \ldots, m$. This property is obtained from definition of BPF.
(2) Orthogonality. One has

$$
\int_{0}^{T} \Phi_{i}(t) \Phi_{j}(t) d t= \begin{cases}h, & i=j  \tag{88}\\ 0, & i \neq j\end{cases}
$$

$t \in[0, T), i, j=1,2, \ldots, m$. This property is obtained from the disjointness property.
(3) Completeness. For every $f \in L^{2},\{\Phi\}$ is complete; if $\int \Phi f=$ 0 then $f=0$ almost everywhere. Because of completeness of $\{\Phi\}$, we have

$$
\begin{equation*}
\int_{0}^{T} f^{2}(t) d t=\sum_{i=1}^{\infty} f_{i}^{2}\left\|\Phi_{i}(t)\right\|^{2} \tag{89}
\end{equation*}
$$

for every real bounded function $f(t)$ which is square integrable in the interval $t \in[0, T)$ and $f_{i}=(1 / h) f(t) \Phi_{i}(t) d t$.
4.3.2. Function Approximation. The orthogonality property of BPF is the basis of expanding functions into their BlockPulse series. For every $f(t) \in L^{2}(R)$,

$$
\begin{equation*}
f(t)=\sum_{i=1}^{m} f_{i} \Phi_{i}(t) \tag{90}
\end{equation*}
$$

where $f_{i}$ is the coefficient of Block-Pulse function, with respect to $i$ th Block-Pulse function $\Phi_{i}(t)$.

The criterion of this approximation is that mean square error between $f(t)$ and its expansion is minimum

$$
\begin{equation*}
\varepsilon=\frac{1}{T} \int_{0}^{T}\left(f(t)-\sum_{j=1}^{m} f_{j} \Phi_{j}(t)\right)^{2} d t \tag{91}
\end{equation*}
$$

so that we can evaluate Block-Pulse coefficients.

$$
\begin{align*}
& \text { Now } \frac{\partial \varepsilon}{\partial f_{i}}=-\frac{2}{T} \int_{0}^{T}\left(f(t)-\sum_{j=1}^{m} f_{j} \Phi_{j}(t)\right) \Phi_{i}(t) d t=0, \\
& \Longrightarrow f_{i}=\frac{1}{h} \int_{0}^{T} f(t) \Phi_{i}(t) d t \quad \text { (using orthogonal property). } \tag{92}
\end{align*}
$$

In the matrix form, we obtain the following from (90) as follow:

$$
\begin{gather*}
f(t)=\sum_{i=1}^{m} f_{i} \Phi_{i}(t)=F^{T} \Phi(t)=\Phi^{T} F \\
\text { where } F=\left[f_{1}, f_{2}, \ldots, f_{m}\right]^{T} \tag{93}
\end{gather*}
$$

Now let $K(t, s)$ be two-variable function defined on $t \in[0, T)$ and $s \in[0,1)$; then $K(t, s)$ can be expanded to BPF as

$$
\begin{equation*}
K(t, s)=\Phi^{T}(t) K \Psi(s) \tag{94}
\end{equation*}
$$

where $\Phi(t)$ and $\Psi(s)$ are $m_{1}$ and $m_{2}$ dimensional Block-Pulse function vectors and $k$ is a $m_{1} \times m_{2}$ Block-Pulse coefficient matrix.

There are two different cases of multiplication of two BPF. The first case is

$$
\Phi(t) \Phi^{T}(t)=\left(\begin{array}{cccc}
\Phi_{1}(t) & 0 & \cdots & 0  \tag{95}\\
0 & \Phi_{2}(t) & \cdots & 0 \\
\vdots & \vdots & \ddots & \vdots \\
0 & 0 & \cdots & \Phi_{m}(t)
\end{array}\right)
$$

It is obtained from disjointness property of BPF. It is a diagonal matrix with $m$ Block-Pulse functions.

The second case is

$$
\begin{equation*}
\Phi^{T}(t) \Phi(t)=1 \tag{96}
\end{equation*}
$$

because $\sum_{i=1}^{m}\left(\Phi_{i}(t)\right)^{2}=\sum_{i=1}^{m} \Phi_{i}(t)=1$.
Operational Matrix of Integration. BPF integration property can be expressed by an operational equation as

$$
\begin{equation*}
\int_{0}^{T} \Phi(t) d t=P \Phi(t) \tag{97}
\end{equation*}
$$

where

$$
\begin{equation*}
\Phi(t)=\left[\Phi_{1}(t), \Phi_{2}(t), \ldots, \Phi_{m}(t)\right]^{T} \tag{98}
\end{equation*}
$$

A general formula for $P_{m \times m}$ can be written as

$$
P=\frac{1}{2}\left(\begin{array}{ccccc}
1 & 2 & 2 & \cdots & 2  \tag{99}\\
0 & 1 & 2 & \cdots & 2 \\
0 & 0 & 1 & \cdots & 2 \\
\vdots & \vdots & \vdots & \ddots & \vdots \\
0 & 0 & 0 & \cdots & 1
\end{array}\right) \text {. }
$$

By using this matrix, we can express the integral of a function $f(t)$ into its Block-Pulse series

$$
\begin{equation*}
\int_{0}^{t} f(t) d t=\int_{0}^{t} F^{T} \Phi(t) d t=F^{T} P \Phi(t) \tag{100}
\end{equation*}
$$

4.3.3. Solution for Linear Integral Equations System. Consider the integral equations system from (55) as follows:

$$
\begin{array}{r}
\sum_{j=1}^{n} y_{j}(x)=f_{i}(x)+\sum_{j=1}^{n} \int_{\alpha}^{\beta} K_{i, j}(x, t) y_{j}(t) d t  \tag{101}\\
i=1,2, \ldots, n
\end{array}
$$

Block-Pulse coefficients of $y_{j}(x), j=1,2, \ldots, n$ in the interval $x \in[\alpha, \beta)$ can be determined from the known functions $f_{i}(x), i=1,2, \ldots, n$ and the kernels $K_{i, j}(x, t), i, j=1,2, \ldots n$. Usually, we consider $\alpha=0$ to facilitie the use of Block-Pulse
functions. In case $\alpha \neq 0$, we set $X=((x-\alpha) /(\beta-\alpha)) T$, where $T=m h$.

We approximate $f_{i}(x), y_{j}(x), K_{i, j}(x, t)$ by its BPF as follows:

$$
\begin{align*}
f_{i}(x) & \approx F_{i}^{T} \Phi(x), \\
y_{j}(x) & \approx Y_{j}^{T} \Phi(x),  \tag{102}\\
K_{i, j}(x, t) & \approx \Phi^{T}(t) K_{i, j} \Phi(x),
\end{align*}
$$

where $F_{i}, Y_{j}$, and $K_{i, j}$ are defined in Section 4.3.2, and substituting (102) into (101), we have

$$
\begin{align*}
& \sum_{j=1}^{n} Y_{j}^{T} \Phi(x)=F_{i}^{T} \Phi(x) \\
& +\sum_{j=1}^{n} \int_{0}^{m h} Y_{j}^{T} \Phi(t) \Phi^{T}(t) K_{i, j} \Phi(x) d t  \tag{103}\\
& \\
& i=1,2, \ldots, n  \tag{104}\\
& \sum_{j=1}^{n} Y_{j}^{T} \Phi(x)=F_{i}^{T} \Phi(x)+\sum_{j=1}^{n} Y_{j}^{T} h I K_{i, j} \Phi(x) \\
& \\
& i=1,2, \ldots, n
\end{align*}
$$

since

$$
\begin{equation*}
\int_{0}^{m h} \Phi(t) \Phi^{T}(t) d t=h I \tag{105}
\end{equation*}
$$

From (104), we get

$$
\begin{equation*}
\sum_{j=1}^{n}\left(I-h K_{i, j}^{T}\right) Y_{j}=F_{i}, \quad i=1,2, \ldots, n . \tag{106}
\end{equation*}
$$

Set $A_{i, j}=I-h K_{i, j}^{T}$; then we have from (106)

$$
\begin{equation*}
\sum_{j=1}^{n} A_{i, j} Y_{j}=F_{i}, \quad i=1,2, \ldots, n \tag{107}
\end{equation*}
$$

which is a linear system

$$
\left(\begin{array}{ccccc}
A_{11} & A_{12} & \cdot & \cdot & A_{1 n}  \tag{108}\\
A_{21} & A_{22} & \cdot & \cdot & A_{2 n} \\
\cdot & \cdot & \cdot & & \cdot \\
\cdot & \cdot & \cdot & \cdot \\
\cdot & \cdot & & \cdot & \cdot \\
A_{n 1} & A_{n 2} & \cdot & \cdot & A_{n n}
\end{array}\right)\left(\begin{array}{c}
Y_{1} \\
Y_{2} \\
\cdot \\
\cdot \\
\cdot \\
Y_{n}
\end{array}\right)=\left(\begin{array}{c}
F_{1} \\
F_{2} \\
\cdot \\
\cdot \\
\cdot \\
F_{n}
\end{array}\right)
$$

After solving the above system we can find $Y_{j}, j=1,2, \ldots, n$ and hence obtain the solutions $y_{j}=\Phi^{T} Y_{j}, j=1,2, \ldots, n$.

## 5. Numerical Methods for Nonlinear Fredholm-Hammerstein Integral Equation

We consider the second kind nonlinear Fredholm integral equation of the following form:

$$
\begin{array}{r}
u(x)=f(x)+\int_{0}^{1} K(x, t) F(t, u(t)) d t  \tag{109}\\
0 \leq x \leq 1
\end{array}
$$

where $K(x, t)$ is the kernel of the integral equation, $f(x)$ and $K(x, t)$ are known functions, and $u(x)$ is the unknown function that is to be determined.
5.1. B-Spline Wavelet Method. In this section, nonlinear Fredholm integral equation of second kind of the form given in (109) has been solved by using $B$-spline wavelets [11].
$B$-spline scaling and wavelet functions in the interval $[0,1]$ and function approximation have been defined in Sections 3.1.1 and 3.1.2, respectively.

First, we assume that

$$
\begin{gather*}
y(x)=F(x, u(x)),  \tag{110}\\
0 \leq x \leq 1 .
\end{gather*}
$$

Now, from (16), we can approximate the functions $u(x)$ and $y(x)$ as

$$
\begin{align*}
& u(x)=A^{T} \Psi(x) \\
& y(x)=B^{T} \Psi(x) \tag{111}
\end{align*}
$$

where $A$ and $B$ are $\left(2^{M+1}+m-1\right) \times 1$ column vectors similar to $C$ defined in (17).

Again, by using dual of the wavelet functions, we can approximate the functions $f(x)$ and $K(x, t)$ as follows:

$$
\begin{gather*}
F(x)=D^{T} \widetilde{\Psi}(x), \\
K(x, t)=\widetilde{\Psi}^{T}(t) \Theta \widetilde{\Psi}(x), \tag{112}
\end{gather*}
$$

where

$$
\begin{equation*}
\Theta_{(i, j)}=\int_{0}^{1}\left[\int_{0}^{1} K(x, t) \Psi_{i}(t) d t\right] \Psi_{j}(x) d x \tag{113}
\end{equation*}
$$

From (110)-(112), we get

$$
\begin{align*}
\int_{0}^{1} K & (x, t) F(t, u(t)) d t \\
& =\int_{0}^{1} B^{T} \Psi(t) \widetilde{\Psi}^{T}(t) \Theta \widetilde{\Psi}(x) d t \\
& =B^{T}\left[\int_{0}^{1} \Psi(t) \widetilde{\Psi}^{T}(t) d t\right] \Theta \widetilde{\Psi}(x)  \tag{114}\\
& =B^{T} \Theta \widetilde{\Psi}(x), \quad \text { since } \int_{0}^{1} \Psi(t) \widetilde{\Psi}^{T}(t) d t=I
\end{align*}
$$

Applying (110)-(114) in (109), we get

$$
\begin{equation*}
A^{T} \Psi(x)=D^{T} \widetilde{\Psi}(x)+B^{T} \Theta \widetilde{\Psi}(x) \tag{115}
\end{equation*}
$$

Multiplying (115) by $\Psi^{T}(x)$ both sides from the right and integrating both sides with respect to $x$ from 0 to 1 , we have

$$
\begin{gather*}
A^{T} P=D^{T}+B^{T} \Theta, \\
A^{T} P-D^{T}-B^{T} \Theta=0, \tag{116}
\end{gather*}
$$

where $P$ is a $\left(2^{M+1}+m-1\right) \times\left(2^{M+1}+m-1\right)$ square matrix given by

$$
\begin{gather*}
P=\int_{0}^{1} \Psi(x) \Psi^{T}(x) d x=\left[\begin{array}{ll}
P_{1} & \\
& P_{2}
\end{array}\right]  \tag{117}\\
\quad \int_{0}^{1} \widetilde{\Psi}(x) \Psi^{T}(x) d x=I
\end{gather*}
$$

Equation (116) gives a system of $\left(2^{M+1}+m-1\right)$ algebraic equations with $2\left(2^{M+1}+m-1\right)$ unknowns for $A$ and $B$ vectors given in (111).

To find the solution $u(x)$ in (111), we first utilize the following equation:

$$
\begin{equation*}
F\left(x, A^{T} \Psi(x)\right)=B^{T} \Psi(x) \tag{118}
\end{equation*}
$$

with the collocation points $x_{i}=(i-1) /\left(2^{M+1}+m-2\right)$, where $i=1,2, \ldots, 2^{M+1}+m-1$.

Equation (118) gives a system of $\left(2^{M+1}+m-1\right)$ algebraic equations with $2\left(2^{M+1}+m-1\right)$ unknowns, for $A$ and $B$ vectors given in (111).

Combining (116) and (118), we have a total of $2\left(2^{M+1}+\right.$ $m-1)$ system of algebraic equations with $2\left(2^{M+1}+m-\right.$ 1) unknowns for $A$ and $B$. Solving those equations for the unknown coefficients in the vectors $A$ and $B$, we can obtain the solution $u(x)=A^{T} \Psi(x)$.

### 5.2. Quadrature Method Applied to Fredholm Integral Equa-

 tion. In this section, Quadrature method has been applied to solve nonlinear Fredholm-Hammerstein integral equation [10].The quadrature methods like Simpson rule and modified trapezoid method are applied for solving a definite integral as follows.

### 5.2.1. Simpson's Rule. One has

$$
\begin{aligned}
\int_{a}^{b} f(x) d x & =\sum_{i=1}^{n-1} \int_{x_{i-1}}^{x_{i+1}} f(x) d x \\
& =\frac{h}{3} f(a)+\frac{4 h}{3} \sum_{i=1}^{n / 2} f\left(x_{2 i-1}\right)
\end{aligned}
$$

$$
\begin{align*}
& +\frac{2 h}{3} \sum_{i=1}^{(n-1) / 2} f\left(x_{2 i}\right) \\
& +\frac{h}{3} f(b) \\
& -\frac{(b-a)}{180} h^{4} f^{(4)}(\eta) \tag{119}
\end{align*}
$$

5.2.2. Modified Trapezoid Rule. One has

$$
\begin{align*}
\int_{a}^{b} f(x) d x= & \sum_{i=1}^{n} \int_{x_{i-1}}^{x_{i}} f(x) d x \\
= & \frac{h}{2} f(a)+h \sum_{i=1}^{n-1} f\left(x_{i}\right)  \tag{120}\\
& +\frac{h}{2} f(b) \\
& +\frac{h^{2}}{12}\left[f^{\prime}(a)-f^{\prime}(b)\right] .
\end{align*}
$$

Consider the nonlinear Fredholm integral equation of second kind defined in (109) as follows:

$$
\begin{array}{r}
u(x)=f(x)+\int_{a}^{b} K(x, t) F(u(t)) d t  \tag{121}\\
a \leq x \leq b .
\end{array}
$$

For solving (121), we approximate the right-hand integral of (121) with Simpson's rule and modified trapezoid rule; then we get the following.

### 5.2.3. Simpson's Rule. One has

$$
\begin{align*}
& u(x)=f(x) \\
&+\frac{h}{3}\left[K\left(x, t_{0}\right) F\left(u_{0}\right)\right. \\
&+4 \sum_{j=1}^{n / 2} K\left(x, t_{2 j-1}\right) F\left(u_{2 j-1}\right)  \tag{122}\\
&+2 \sum_{j=1}^{(n / 2)-1} K\left(x, t_{2 j}\right) F\left(u_{2 j}\right) \\
&\left.+K\left(x, t_{n}\right) F\left(u_{n}\right)\right] .
\end{align*}
$$

Hence, for $x=x_{0}, x_{1}, \ldots, x_{n}$ and $t=t_{0}, t_{1}, \ldots, t_{n}$ in (122), we have

$$
\begin{aligned}
& u\left(x_{i}\right)=f\left(x_{i}\right) \\
&+\frac{h}{3} {\left[K\left(x_{i}, t_{0}\right) F\left(u_{0}\right)\right.} \\
&+4 \sum_{j=1}^{n / 2} K\left(x_{i}, t_{2 j-1}\right) F\left(u_{2 j-1}\right)
\end{aligned}
$$

$$
\begin{align*}
& +2 \sum_{j=1}^{(n / 2)-1} K\left(x_{i}, t_{2 j}\right) F\left(u_{2 j}\right) \\
& \left.+K\left(x_{i}, t_{n}\right) F\left(u_{n}\right)\right] . \tag{123}
\end{align*}
$$

Equation (123) is a nonlinear system of equations and, by solving (123), we obtain the unknowns $u\left(x_{i}\right)$ for $i=0,1, \ldots, n$.

### 5.2.4. Modified Trapezoid Rule. One has

$$
\begin{align*}
u(x)= & f(x) \\
& +\frac{h}{2} K\left(x, t_{0}\right) F\left(u_{0}\right) \\
+ & h \sum_{j=1}^{n-1} K\left(x, t_{j}\right) F\left(u_{j}\right) \\
+ & \frac{h}{2} K\left(x, t_{n}\right) F\left(u_{n}\right)  \tag{124}\\
+ & \frac{h^{2}}{12}\left[J\left(x, t_{0}\right) F\left(u_{0}\right)\right. \\
& +K\left(x, t_{0}\right) u_{0}^{\prime} F^{\prime}\left(u_{0}\right) \\
& \quad-J\left(x, t_{n}\right) F\left(u_{n}\right) \\
& \left.\quad-K\left(x, t_{n}\right) u_{n}^{\prime} F^{\prime}\left(u_{n}\right)\right],
\end{align*}
$$

where $J(x, t)=\partial K(x, t) / \partial t$.
For $x=x_{0}, x_{1}, \ldots, x_{n}$ and $t=t_{0}, t_{1}, \ldots, t_{n}$ in (124), we have

$$
\begin{align*}
u\left(x_{i}\right)= & f\left(x_{i}\right) \\
& +\frac{h}{2} K\left(x_{i}, t_{0}\right) F\left(u_{0}\right) \\
+ & h \sum_{j=1}^{n-1} K\left(x_{i}, t_{j}\right) F\left(u_{j}\right) \\
+ & \frac{h}{2} K\left(x_{i}, t_{n}\right) F\left(u_{n}\right)  \tag{125}\\
+ & \frac{h^{2}}{12}\left[J\left(x_{i}, t_{0}\right) F\left(u_{0}\right)\right. \\
& +K\left(x_{i}, t_{0}\right) u_{0}^{\prime} F^{\prime}\left(u_{0}\right) \\
& \quad-J\left(x_{i}, t_{n}\right) F\left(u_{n}\right) \\
& \left.\quad-K\left(x_{i}, t_{n}\right) u_{n}^{\prime} F^{\prime}\left(u_{n}\right)\right]
\end{align*}
$$

for $i=0,1, \ldots, n$.

This is a system of $(n+1)$ equations and $(n+3)$ unknowns. By taking derivative from (121) and setting $H(x, t)=$ $\partial K(x, t) / \partial x$, we obtain

$$
\begin{array}{r}
u^{\prime}(x)=f^{\prime}(x)+\int_{a}^{b} H(x, t) F(u(t)) d t  \tag{126}\\
a \leq x \leq b
\end{array}
$$

If $u$ is a solution of (121), then it is also solution of (126). By using trapezoid rule for (126) and replacing $x=x_{i}$, we get

$$
\begin{align*}
u^{\prime}\left(x_{i}\right)= & f^{\prime}\left(x_{i}\right) \\
& +\frac{h}{2} H\left(x_{i}, t_{0}\right) F\left(u_{0}\right) \\
& +h \sum_{j=1}^{n-1} H\left(x_{i}, t_{j}\right) F\left(u_{j}\right)  \tag{127}\\
& +\frac{h}{2} H\left(x_{i}, t_{n}\right) F\left(u_{n}\right),
\end{align*}
$$

for $i=0,1, \ldots, n$. In case of $i=0, n$ from system (127), we obtain two equations.

Now (127) combined with (125) generates the nonlinear system of equations as follows:

$$
\begin{aligned}
u\left(x_{i}\right)= & \left(\frac{h}{2} K\left(x_{i}, t_{0}\right)+\frac{h^{2}}{12} J\left(x_{i}, t_{0}\right)\right) F\left(u_{0}\right) \\
& +h \sum_{j=1}^{n-1} K\left(x_{i}, t_{j}\right) F\left(u_{j}\right) \\
& +\left(\frac{h}{2} K\left(x_{i}, t_{n}\right)-\frac{h^{2}}{12} J\left(x_{i}, t_{n}\right)\right) F\left(u_{n}\right) \\
& +\frac{h^{2}}{12}\left(K\left(x_{i}, t_{0}\right) u_{0}^{\prime} F^{\prime}\left(u_{0}\right)\right. \\
u^{\prime}\left(x_{0}\right)= & f^{\prime}\left(x_{0}\right) \\
& +\frac{h}{2} H\left(x_{0}, t_{0}\right) F\left(u_{0}\right) \\
& \left.\left.+h \sum_{j=1}^{n-1} H\left(x_{n}\right) u_{n}^{\prime} F^{\prime}\left(u_{n}\right)\right) \text {, }\right) F\left(u_{j}\right) \\
& +\frac{h}{2} H\left(x_{0}, t_{n}\right) F\left(u_{n}\right)
\end{aligned}
$$

$$
\begin{align*}
u^{\prime}\left(x_{n}\right)= & f^{\prime}\left(x_{n}\right) \\
& +\frac{h}{2} H\left(x_{n}, t_{0}\right) F\left(u_{0}\right) \\
& +h \sum_{j=1}^{n-1} H\left(x_{n}, t_{j}\right) F\left(u_{j}\right)  \tag{128}\\
& +\frac{h}{2} H\left(x_{n}, t_{n}\right) F\left(u_{n}\right) .
\end{align*}
$$

By solving this system with $(n+3)$ nonlinear equations and $(n+3)$ unknowns, we can obtain the solution of (109).
5.3. Wavelet Galerkin Method. In this section, the continuous Legendre wavelets [12], constructed on the interval [ 0,1 ], are applied to solve the nonlinear Fredholm integral equation of the second kind. The nonlinear part of the integral equation is approximated by Legendre wavelets, and the nonlinear integral equation is reduced to a system of nonlinear equations.

We have the following family of continuous wavelets with dilation parameter $a$ and the translation parameter $b$

$$
\begin{array}{r}
\psi_{a, b}(t)=|a|^{-1 / 2} \psi\left(\frac{t-b}{a}\right),  \tag{129}\\
a, b \in R, \quad a \neq 0
\end{array}
$$

Legendre wavelets $\psi_{m, n}(t)=\psi(k, \widehat{n}, m, t)$ have four arguments; $k=2,3, \ldots, \widehat{n}=2 n-1, n=1,2, \ldots, 2^{k-1}, m$ is the order for Legendre polynomials and $t$ is the normalized time.

Legendre wavelets are defined on $[0,1)$ by

$$
\begin{align*}
& \psi_{m, n}(t) \\
& \qquad= \begin{cases}\left(m+\frac{1}{2}\right)^{1 / 2} 2^{k / 2} L_{m}\left(2^{k} t-\widehat{n}\right), & \frac{\widehat{n}-1}{2^{k}} \leq t<\frac{\widehat{n}+1}{2^{k}}, \\
0, & \text { otherwise, }\end{cases} \tag{130}
\end{align*}
$$

where $L_{m}(t)$ are the well-known Legendre polynomials of order $m$, which are orthogonal with respect to the weight function $w(t)=1$ and satisfy the following recursive formula:

$$
\begin{gather*}
L_{0}(t)=1, \\
L_{1}(t)=t, \\
L_{m+1}(t)=\frac{2 m+1}{m+1} t L_{m}(t)  \tag{131}\\
-\frac{m}{m+1} L_{m-1}(t), \quad m=1,2,3, \ldots
\end{gather*}
$$

The set of Legendre wavelets are an orthonormal set.
5.3.1. Function Approximation. A function $f(x) \in L^{2}[0,1]$ can be expanded as

$$
\begin{equation*}
f(x)=\sum_{n=1}^{\infty} \sum_{m=0}^{\infty} c_{n, m} \psi_{n, m}(x) \tag{132}
\end{equation*}
$$

where

$$
\begin{equation*}
c_{n, m}=\left\langle f(x), \psi_{n, m}(x)\right\rangle \tag{133}
\end{equation*}
$$

If the infinite series in (132) is truncated, then (132) can be written as

$$
\begin{equation*}
f(x) \approx \sum_{n=1}^{2^{k-1}} \sum_{m=0}^{M-1} c_{n, m} \psi_{n, m}(x)=C^{T} \Psi(x) \tag{134}
\end{equation*}
$$

where $C$ and $\Psi(x)$ are $2^{k-1} M \times 1$ matrices given by

$$
\begin{gather*}
C=\left[c_{1,0}, c_{1,1}, \ldots, c_{1, M-1}, c_{2,0}, \ldots,\right. \\
\left.c_{2, M-1}, \ldots, c_{2^{k-1}, 0}, \ldots, c_{2^{k-1}, M-1}\right]^{T},  \tag{135}\\
\Psi(x)=\left[\psi_{1,0}(x), \ldots, \psi_{1, M-1}(x),\right. \\
\psi_{2,0}(x), \ldots, \psi_{2, M-1}(x), \ldots,  \tag{136}\\
\left.\psi_{2^{k-1}, 0}(x), \ldots, \psi_{2^{k-1}, M-1}(x)\right]^{T} .
\end{gather*}
$$

Similarly, a function $k(x, t) \in L^{2}([0,1] \times[0,1])$ can be approximated as

$$
\begin{equation*}
k(x, t) \approx \Psi^{T}(t) K \Psi(x) \tag{137}
\end{equation*}
$$

where $K$ is $\left(2^{k-1} M \times 2^{k-1} M\right)$ matrix, with

$$
\begin{equation*}
K_{i, j}=\left\langle\psi_{i}(t),\left\langle k(x, t), \psi_{j}(x)\right\rangle\right\rangle . \tag{138}
\end{equation*}
$$

Also, the integer power of a function can be approximated as

$$
\begin{equation*}
[y(x)]^{p}=\left[Y^{T} \Psi(x)\right]^{p}=Y_{p}^{* T} \Psi(x), \tag{139}
\end{equation*}
$$

where $Y_{p}^{*}$ is a column vector, whose elements are nonlinear combinations of the elements of the vector $Y . Y_{p}^{*}$ is called the operational vector of the $p$ th power of the function $y(x)$.
5.3.2. The Operational Matrices. The integration of the vector $\Psi(x)$ defined in (136) can be obtained as

$$
\begin{equation*}
\int_{0}^{t} \Psi\left(t^{\prime}\right) d t^{\prime}=P \Psi(t) \tag{140}
\end{equation*}
$$

where $P$ is the $\left(2^{k-1} M \times 2^{k-1} M\right)$ operational matrix for integration and is given in [23] as

$$
P=\left[\begin{array}{ccccc}
L & H & \cdots & H & H  \tag{141}\\
0 & L & \cdots & H & H \\
\vdots & \vdots & \ddots & \vdots & \vdots \\
0 & 0 & \cdots & L & H \\
0 & 0 & \cdots & 0 & L
\end{array}\right]
$$

In (141), $H$ and $L$ are $(M \times M)$ matrices given in [23] as

$$
\begin{align*}
& H=\frac{1}{2^{k}}\left[\begin{array}{cccc}
2 & 0 & \cdots & 0 \\
0 & 0 & \cdots & 0 \\
\vdots & \vdots & \ddots & \vdots \\
0 & 0 & \cdots & 0
\end{array}\right], \\
& L=\frac{1}{2^{k}}\left[\begin{array}{ccccccc}
1 & \frac{1}{\sqrt{3}} & 0 & 0 & \cdots & 0 & 0 \\
-\frac{\sqrt{3}}{3} & 0 & \frac{\sqrt{3}}{3 \sqrt{5}} & 0 & \cdots & 0 & 0 \\
0 & -\frac{\sqrt{5}}{5 \sqrt{3}} & 0 & \frac{\sqrt{5}}{5 \sqrt{7}} & \cdots & 0 & 0 \\
0 & 0 & -\frac{\sqrt{7}}{7 \sqrt{5}} & 0 & \cdots & 0 & 0 \\
\vdots & \vdots & \vdots & \vdots & \ddots & & \vdots \\
0 & 0 & 0 & 0 & \cdots & 0 & \frac{\sqrt{2 M-3}}{(2 M-3) \sqrt{2 M-1}} \\
0 & 0 & 0 & 0 & \cdots & \frac{-\sqrt{2 M-1}}{(2 M-1) \sqrt{2 M-3}} & 0
\end{array}\right] . \tag{142}
\end{align*}
$$

The integration of the product of two Legendre wavelets vector functions is obtained as

$$
\begin{equation*}
\int_{0}^{1} \Psi(t) \Psi^{T}(t) d t=I \tag{143}
\end{equation*}
$$

where $I$ is an identity matrix.
The product of two Legendre wavelet vector functions is defined as

$$
\begin{equation*}
\Psi(t) \Psi^{T}(t) C=\widetilde{C}^{T} \Psi(t) \tag{144}
\end{equation*}
$$

where $C$ is a vector given in (135) and $\widetilde{C}$ is $\left(2^{k-1} M \times 2^{k-1} M\right)$ matrix, which is called the product operation of Legendre wavelet vector functions [23, 24].
5.3.3. Solution of Fredholm Integral Equation of Second Kind. Consider the nonlinear Fredholm-Hammerstein integral equation of second kind of the form

$$
\begin{equation*}
y(x)=f(x)+\int_{0}^{1} k(x, t)[y(t)]^{p} d t \tag{145}
\end{equation*}
$$

where $f \in L^{2}[0,1], k \in L^{2}([0,1] \times[0,1]), y$ is an unknown function, and $p$ is a positive integer.

We can approximate the following functions as

$$
\begin{gather*}
f(x) \approx F^{T} \Psi(x), \\
y(x) \approx Y^{T} \Psi(x),  \tag{146}\\
k(x, t) \approx \Psi^{T}(t) K \Psi(x), \\
{[y(x)]^{p} \approx Y^{* T} \Psi(x)} \tag{149}
\end{gather*}
$$

For solving (148) by Homotopy perturbation method (HPM) [14-16], we consider (148) as

$$
L(u)=u(x)-f(x)-\int_{0}^{1} K(x, t) F(u(t)) d t=0
$$

As a possible remedy, we can define $H(u, p)$ by

$$
\begin{align*}
& H(u, 0)=N(u),  \tag{150}\\
& H(u, 1)=L(u),
\end{align*}
$$

where $N(u)$ is an integral operator with known solution $u_{0}$.
We may choose a convex homotopy by

$$
\begin{equation*}
H(u, p)=(1-p) N(u)+p L(u)=0 \tag{151}
\end{equation*}
$$

and continuously trace an implicitly defined curve from a starting point $H\left(u_{0}, 0\right)$ to a solution function $H(U, 1)$. The embedding parameter $p$ monotonically increases from zero to unit as the trivial problem $L(u)=0$. The embedding parameter $p \in(0,1]$ can be considered as an expanding parameter. The HPM uses the homotopy parameter $p$ as an expanding parameter; that is,

$$
\begin{equation*}
u=u_{0}+p u_{1}+p^{2} u_{2}+\cdots \tag{152a}
\end{equation*}
$$

When $p \rightarrow 1$, (152a) corresponding to (151) become the approximate solution of (149) as follows:

$$
\begin{equation*}
U=\lim _{p \rightarrow 1} u=u_{0}+u_{1}+u_{2}+\cdots \tag{152b}
\end{equation*}
$$

The series in (152b) converges in most cases, and the rate of convergence depends on $L(u)$ [14].

Consider

$$
\begin{equation*}
N(u)=u(x)-f(x) . \tag{153}
\end{equation*}
$$

The nonlinear term $F(u)$ can be expressed in He polynomials [25] as

$$
\begin{align*}
F(u)= & \sum_{m=0}^{\infty} p^{m} H_{m}\left(u_{0}, u_{1}, \ldots, u_{m}\right) \\
= & H_{0}\left(u_{0}\right)+p H_{1}\left(u_{0}, u_{1}\right)  \tag{154}\\
& +\cdots+p^{m} H_{m}\left(u_{0}, u_{1}, \ldots u_{m}\right)+\cdots,
\end{align*}
$$

where

$$
\begin{align*}
H_{m} & \left(u_{0}, u_{1}, \ldots, u_{m}\right) \\
& =\left.\frac{1}{m!} \frac{\partial^{m}}{\partial p^{m}}\left(F\left(\sum_{k=0}^{m} p^{k} u_{k}\right)\right)\right|_{p=0}, \quad m \geq 0 . \tag{155}
\end{align*}
$$

Substituting (152a), (153), and (154) into (151), we have

$$
\begin{align*}
& (1-p)\left(\left(u_{0}+p u_{1}+\cdots\right)-f(x)\right) \\
& +p\left(\left(u_{0}+p u_{1}+\cdots\right)-f(x)\right. \\
& \left.\quad-\int_{0}^{1} K(x, t) \sum_{m=0}^{\infty} p^{m} H_{m}\left(u_{0}, u_{1}, \ldots, u_{m}\right) d t\right)=0 \\
& \Longrightarrow\left(u_{0}+p u_{1}+\cdots\right)-f(x) \\
& \quad-p \int_{0}^{1} K(x, t) \sum_{m=0}^{\infty} p^{m} H_{m}\left(u_{0}, u_{1}, \ldots, u_{m}\right) d t=0 . \tag{156}
\end{align*}
$$

Equating the terms with identical power of $p$ in (156), we have

$$
\begin{align*}
& p^{0}: u_{0}(x)-f(x)=0 \Longrightarrow u_{0}(x)=f(x) \\
& p^{1}: u_{1}(x)-\int_{0}^{1} K(x, t) H_{0} d t=0 \Longrightarrow u_{1}(x) \\
& =\int_{0}^{1} K(x, t) H_{0} d t \\
& p^{2}: u_{2}(x)-\int_{0}^{1} K(x, t) H_{1} d t=0 \Longrightarrow u_{2}(x)  \tag{157}\\
& \quad=\int_{0}^{1} K(x, t) H_{1} d t \\
& \vdots
\end{align*}
$$

and in general form we have

$$
\begin{gathered}
u_{0}(x)=f(x) \\
u_{n+1}(x)=\int_{0}^{1} K(x, t) H_{n} d t, \quad n=0,1,2, \ldots
\end{gathered}
$$

Hence, we can obtain the approximate solution of aforesaid equation (148) from (152b).
5.5. Adomian Decomposition Method. Adomian decomposition method (ADM) [16-18] has been applied to a wide class of functional equations. This method gives the solution as an infinite series usually converging to an accurate solution. Let us consider the nonlinear Fredholm integral equation of second kind as follows:

$$
\begin{array}{r}
u(x)=f(x)+\int_{a}^{b} K(x, t)(L u(t)+N u(t)) d t  \tag{159}\\
a \leq x \leq b,
\end{array}
$$

where $L(u(t))$ and $N(u(t))$ are the linear and nonlinear terms, respectively.

The Adomian decomposition method (ADM) consists of representing $u(x)$ as a series

$$
\begin{equation*}
u(x)=\sum_{m=0}^{\infty} u_{m}(x) \tag{160}
\end{equation*}
$$

In the view of ADM , the nonlinear term $N u$ can be represented as

$$
\begin{gather*}
N u=\sum_{n=0}^{\infty} A_{n}  \tag{161}\\
\text { where } A_{n}=\left.\frac{1}{n!}\left(\frac{\partial^{n}}{\partial \lambda^{n}} N\left(\sum_{k=0}^{\infty} \lambda^{k} u_{k}\right)\right)\right|_{\lambda=0} \tag{162}
\end{gather*}
$$

Now substituting (160) and (161) into (159), we have

$$
\begin{align*}
\sum_{m=0}^{\infty} u_{m}(x)= & f(x) \\
& +\int_{a}^{b} K(x, t)\left(L\left(\sum_{m=0}^{\infty} u_{m}(t)\right)+\sum_{m=0}^{\infty} A_{m}\right) d t \tag{163}
\end{align*}
$$

and, then, ADM uses the recursive relations

$$
\begin{gather*}
u_{0}(x)=f(x) \\
u_{m}(x)=\int_{a}^{b} K(x, t)\left(L\left(u_{m-1}(t)\right)+A_{m-1}(t)\right) d t  \tag{164}\\
m \geq 1
\end{gather*}
$$

where $A_{m}$ is so-called Adomian polynomial.
Therefore, we obtain the $n$-terms approximate solution as

$$
\begin{equation*}
\varphi_{n}=u_{0}+u_{1}+\cdots+u_{n} \tag{165}
\end{equation*}
$$

with

$$
\begin{equation*}
u(x)=\lim _{n \rightarrow \infty} \varphi_{n} \tag{166}
\end{equation*}
$$

## 6. Conclusion and Discussion

In this work, we have examined many numerical methods to solve Fredholm integral equations. Using these methods except variational iteration method, the Fredholm integral equations have been reduced to a system of algebraic equations and this system can be easily solved by any usual methods. In this work, we have applied compactly supported semiorthogonal $B$-spline wavelets along with their dual wavelets for solving both linear and nonlinear Fredholm integral equations of second kind. The problem has been reduced to solve a system of algebraic equations. In order to increase the accuracy of the approximate solution, it is necessary to apply higher-order $B$-spline wavelet method. The method of moments based on compactly supported semiorthogonal $B$ spline wavelets via Galerkin method has been used to solve Fredholm integral equation of second kind. This method determines a strong reduction in the computation time and memory requirement in inverting the matrix. Variational iteration method has been successfully applied to find the approximate solution of Fredholm integral equation of both linear and nonlinear types. Taylor series expansion method reduces the system of integral equations to a linear system of ordinary differential equation. After including the required boundary conditions, this system reduces to a system of algebraic equations that can be solved easily. Block-Pulse functions and Haar wavelet method can be applied to the system of Fredholm integral equations by reducing into a system of algebraic equations. These methods give more accuracy if we increase their order. Quadrature method can be applied to solve the nonlinear Fredholm-Hammerstein integral equation of second kind by reducing it to a system of algebraic equations. Homotopy perturbation method (HPM)
and Adomian decomposition method (ADM) can be also applied to approximate the solution of nonlinear Fredholm integral equation of second kind. The solutions obtained by HPM and ADM are applicable for not only weakly nonlinear equations, but also strong ones. The approximate solutions by these aforesaid methods highly agree with exact solutions.

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## Research Article

# Classification of Exact Solutions for Generalized Form of $K(m, n)$ Equation 

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Received 24 May 2013; Revised 1 August 2013; Accepted 18 August 2013
Academic Editor: Santanu Saha Ray
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#### Abstract

The classification of exact solutions, including solitons and elliptic solutions, to the generalized $K(m, n)$ equation by the complete discrimination system for polynomial method has been obtained. From here, we find some interesting results for nonlinear partial differential equations with generalized evolution.


## 1. Introduction

In science and engineering applications, it is often very difficult to obtain analytical solutions of partial differential equations. Recently, many exact solutions of partial differential equations have been examined by the use of trial equation method. Also there are a lot of important methods that have been defined such as Hirota method, tanh-coth method, sine-cosine method, the trial equation method, and the extended trial equation method [1-15] to find exact solutions to nonlinear partial differential equations. There are a lot of nonlinear evolution equations that are solved by the use of various mathematical methods. Soliton solutions, singular solitons, and other solutions have been found by using these approaches. These obtained solutions have been encountered in various areas of applied mathematics and are very important.

In Section 2, we introduce an extended trial equation method for nonlinear evolution equations with higher order nonlinearity. In Section 3, as applications, we procure some exact solutions to nonlinear partial differential equations such as the generalized form of $K(m, n)$ equation [16-18]:

$$
\begin{equation*}
\left(q^{l}\right)_{t}+a q^{m} q_{x}+b\left(q^{n}\right)_{x x x}=0 \tag{1}
\end{equation*}
$$

where $a, b \in R$ are constants since $l, m$, and $n \in Z^{+}$. Here, the first term is the generalized evolution term, while the second term represents the nonlinear term and the third term is the dispersion term. This equation is the generalized form
of the KdV equation, where, in particular, the case $l=m=$ $n=1$ leads to the KdV equation. The Korteweg de Vries equation is one of the most important equations in applied mathematics and physics. There have been several kinds of solutions, such as compactons, that are studied in the context of $K(m, n)$ equation, for various situations. We now offer a more general trial equation method for discussion as follows.

## 2. The Extended Trial Equation Method

Step 1. For a given nonlinear partial differential equation

$$
\begin{equation*}
P\left(u, u_{t}, u_{x}, u_{x x}, \ldots\right)=0 \tag{2}
\end{equation*}
$$

take the general wave transformation

$$
\begin{equation*}
u\left(x_{1}, x_{2}, \ldots, x_{N}, t\right)=u(\eta), \quad \eta=\lambda\left(\sum_{j=1}^{N} x_{j}-c t\right) \tag{3}
\end{equation*}
$$

where $\lambda \neq 0$ and $c \neq 0$. Substituting (3) into (2) yields a nonlinear ordinary differential equation:

$$
\begin{equation*}
N\left(u, u^{\prime}, u^{\prime \prime}, \ldots\right)=0 \tag{4}
\end{equation*}
$$

Step 2. Take the finite series and trial equation as follows:

$$
\begin{equation*}
u=\sum_{i=0}^{\delta} \tau_{i} \Gamma^{i} \tag{5}
\end{equation*}
$$

where

$$
\begin{equation*}
\left(\Gamma^{\prime}\right)^{2}=\Lambda(\Gamma)=\frac{\Phi(\Gamma)}{\Psi(\Gamma)}=\frac{\xi_{\theta} \Gamma^{\theta}+\cdots+\xi_{1} \Gamma+\xi_{0}}{\zeta_{\epsilon} \Gamma^{\epsilon}+\cdots+\zeta_{1} \Gamma+\zeta_{0}} \tag{6}
\end{equation*}
$$

Using (5) and (6), we can write

$$
\begin{gather*}
\left(u^{\prime}\right)^{2}=\frac{\Phi(\Gamma)}{\Psi(\Gamma)}\left(\sum_{i=0}^{\delta} i \tau_{i} \Gamma^{i-1}\right)^{2} \\
u^{\prime \prime}=\frac{\Phi^{\prime}(\Gamma) \Psi(\Gamma)-\Phi(\Gamma) \Psi^{\prime}(\Gamma)}{2 \Psi^{2}(\Gamma)}\left(\sum_{i=0}^{\delta} i \tau_{i} \Gamma^{i-1}\right)  \tag{7}\\
+\frac{\Phi(\Gamma)}{\Psi(\Gamma)}\left(\sum_{i=0}^{\delta} i(i-1) \tau_{i} \Gamma^{i-2}\right),
\end{gather*}
$$

where $\Phi(\Gamma)$ and $\Psi(\Gamma)$ are polynomials. Substituting these relations into (4) yields an equation of polynomial $\Omega(\Gamma)$ of $\Gamma$ :

$$
\begin{equation*}
\Omega(\Gamma)=\varrho_{s} \Gamma^{s}+\cdots+\varrho_{1} \Gamma+\varrho_{0}=0 . \tag{8}
\end{equation*}
$$

According to the balance principle, we can find a relation of $\theta, \epsilon$, and $\delta$. We can calculate some values of $\theta, \epsilon$, and $\delta$.

Step 3. Letting the coefficients of $\Omega(\Gamma)$ all be zero will yield an algebraic equations system:

$$
\begin{equation*}
\varrho_{i}=0, \quad i=0, \ldots, s \tag{9}
\end{equation*}
$$

Solving this system, we will determine the values of $\xi_{0}, \ldots, \xi_{\theta} ; \zeta_{0}, \ldots, \zeta_{\epsilon}$; and $\tau_{0}, \ldots, \tau_{\delta}$.

Step 4. Reduce (6) to the elementary integral form

$$
\begin{equation*}
\pm\left(\eta-\eta_{0}\right)=\int \frac{d \Gamma}{\sqrt{\Lambda(\Gamma)}}=\int \sqrt{\frac{\Psi(\Gamma)}{\Phi(\Gamma)}} d \Gamma \tag{10}
\end{equation*}
$$

Using a complete discrimination system for polynomial to classify the roots of $\Phi(\Gamma)$, we solve (10) and obtain the exact solutions to (4). Furthermore, we can write the exact traveling wave solutions to (2), respectively.

## 3. Application to the Generalized Form of $K(m, n)$ Equation

In order to look for travelling wave solutions of (1), we make the transformation $q(x, t)=u(\eta), \eta=x-c t$, where $c$ is the wave speed. Therefore it can be converted to the ODE

$$
\begin{equation*}
-c\left(u^{l}(\eta)\right)^{\prime}+\frac{a}{m+1}\left(u^{m+1}(\eta)\right)^{\prime}+b\left(u^{n}(\eta)\right)^{\prime \prime \prime}=0 \tag{11}
\end{equation*}
$$

where prime denotes the derivative with respect to $\eta$. Then, integrating this equation with respect to $\eta$ one time and setting the integration constant to zero, we obtain

$$
\begin{equation*}
-c u^{l}(\eta)+\frac{a}{m+1} u^{m+1}(\eta)+b\left(u^{n}(\eta)\right)^{\prime \prime}=0 \tag{12}
\end{equation*}
$$

Let $l=n$, applying balance and using the following transformation:

$$
\begin{equation*}
u=v^{1 /(m-n+1)} . \tag{13}
\end{equation*}
$$

Equation (12) turns into the following equation:

$$
\begin{align*}
& -c(m+1)(m+1-n)^{2} v^{2}+a(m+1-n)^{2} v^{3} \\
& \quad+b n(m+1)(2 n-m-1)\left(v^{\prime}\right)^{2}  \tag{14}\\
& \quad+b n(m+1)(m+1-n) v v^{\prime \prime}=0 .
\end{align*}
$$

Substituting (7) into (14) and using balance principle yield

$$
\begin{equation*}
\theta=\epsilon+\delta+2 . \tag{15}
\end{equation*}
$$

After this solution procedure, we obtain the results as follows.
Case 1. If we take $\epsilon=0, \delta=1$, and $\theta=3$, then

$$
\begin{gather*}
\left(v^{\prime}\right)^{2}=\frac{\left(\tau_{1}\right)^{2}\left(\xi_{3} \Gamma^{3}+\xi_{2} \Gamma^{2}+\xi_{1} \Gamma+\xi_{0}\right)}{\zeta_{0}},  \tag{16}\\
v^{\prime \prime}=\frac{\tau_{1}\left(3 \xi_{3} \Gamma^{2}+2 \xi_{2} \Gamma+\xi_{1}\right)}{2 \zeta_{0}},
\end{gather*}
$$

where $\xi_{3} \neq 0$ and $\zeta_{0} \neq 0$. Respectively, solving the algebraic equation system (9) yields

$$
\begin{align*}
& \xi_{0}=-\frac{\xi_{1}^{2}(3+3 m-5 n)(1+m+n)}{16 \xi_{2}(1+m-2 n)^{2}}, \\
& \xi_{1}=\xi_{1}, \quad \xi_{2}=\xi_{2}, \\
& \xi_{3}=-\frac{8 \xi_{2}^{2}(1+m-2 n)(1+m-n)}{\xi_{1}(1+m+n)^{2}}, \\
& \tau_{0}=\tau_{0}, \quad \tau_{1}=-\frac{4(1+m-2 n) \xi_{2} \tau_{0}}{(1+m+n) \xi_{1}},  \tag{17}\\
& \zeta_{0}=-\frac{b n \xi_{2}(1+m)}{a(1+m-n) \tau_{0}}, \\
& c=\frac{a n(5+5 m-7 n) \tau_{0}}{(1+m)(1+m-n)(1+m+n)} .
\end{align*}
$$

Substituting these results into (6) and (10), we have

$$
\begin{equation*}
\pm\left(\eta-\eta_{0}\right)=\frac{A}{2} \int \frac{d \Gamma}{\sqrt{\Gamma^{3}-\frac{\xi_{1}(1+m+n)^{2}}{8 \xi_{2}(1+m-2 n)(1+m-n)} \Gamma^{2}-\frac{\xi_{1}^{2}(1+m+n)^{2}}{8 \xi_{2}^{2}(1+m-2 n)(1+m-n)} \Gamma+\frac{\xi_{1}^{3}(3+3 m-5 n)(1+m+n)^{3}}{128 \xi_{2}^{2}(1+m-2 n)^{3}(1+m-n)}}} \tag{18}
\end{equation*}
$$

where

$$
\begin{equation*}
A=\sqrt{\frac{b n \xi_{1}(1+m)(1+m+n)^{2}}{2 a \xi_{2} \tau_{0}(1+m-n)^{2}(1+m-2 n)}} \tag{19}
\end{equation*}
$$

Integrating (18), we obtain the solutions to (1) as follows:

$$
\begin{gather*}
\pm\left(\eta-\eta_{0}\right)=-\frac{A}{\sqrt{\Gamma-\alpha_{1}}} \\
\pm\left(\eta-\eta_{0}\right)=\frac{A}{\sqrt{\alpha_{2}-\alpha_{1}}} \arctan \sqrt{\frac{\Gamma-\alpha_{2}}{\alpha_{2}-\alpha_{1}}}, \quad \alpha_{2}>\alpha_{1} \\
\pm\left(\eta-\eta_{0}\right)=\frac{A}{\sqrt{\alpha_{1}-\alpha_{2}}} \ln \left|\frac{\sqrt{\Gamma-\alpha_{2}}-\sqrt{\alpha_{1}-\alpha_{2}}}{\sqrt{\Gamma-\alpha_{2}}+\sqrt{\alpha_{1}-\alpha_{2}}}\right|, \quad \alpha_{1}>\alpha_{2} \\
\pm\left(\eta-\eta_{0}\right)=-\frac{A}{\sqrt{\alpha_{1}-\alpha_{3}}} F(\varphi, l), \quad \alpha_{1}>\alpha_{2}>\alpha_{3} \tag{20}
\end{gather*}
$$

where

$$
\begin{gather*}
F(\varphi, l)=\int_{0}^{\varphi} \frac{d \psi}{\sqrt{1-l^{2} \sin ^{2} \psi}}, \quad \varphi=\arcsin \sqrt{\frac{\Gamma-\alpha_{3}}{\alpha_{2}-\alpha_{3}}} \\
l^{2}=\frac{\alpha_{2}-\alpha_{3}}{\alpha_{1}-\alpha_{3}} \tag{21}
\end{gather*}
$$

Also $\alpha_{1}, \alpha_{2}$, and $\alpha_{3}$ are the roots of the polynomial equation

$$
\begin{equation*}
\Gamma^{3}+\frac{\xi_{2}}{\xi_{3}} \Gamma^{2}+\frac{\xi_{1}}{\xi_{3}} \Gamma+\frac{\xi_{0}}{\xi_{3}}=0 \tag{22}
\end{equation*}
$$

Substituting solutions (20) into (5) and (13), we have

$$
\begin{aligned}
& u(x, t) \\
& \qquad \begin{aligned}
&=\left[\tau_{0}+\tau_{1} \alpha_{1}\right. \\
&+\left(A ^ { 2 } \tau _ { 1 } \left(\left(x-\left(\operatorname{an}(5+5 m-7 n) \tau_{0}\right)\right.\right.\right. \\
& \times((1+m)(1+m-n)(1+m+n))^{-1} \\
&\left.\left.\left.\left.\times t-\eta_{0}\right)^{2}\right)^{-1}\right)\right]^{1 /(m-n+1)}
\end{aligned}
\end{aligned}
$$

$$
\begin{aligned}
& u(x, t)=\left[\tau_{0}+\tau_{1} \alpha_{1}+\tau_{1}\left(\alpha_{2}-\alpha_{1}\right)\right. \\
& \times \operatorname{sech}^{2}\left(\frac{\sqrt{\alpha_{2}-\alpha_{1}}}{A}( \right. x-\left(\operatorname{an}(5+5 m-7 n) \tau_{0}\right) \\
& \times((1+m)(1+m-n) \\
&\times(1+m+n))^{-1} \\
&\left.\left.\left.\times t-\eta_{0}\right)\right)\right]^{1 /(m-n+1)}
\end{aligned}
$$

$$
u(x, t)
$$

$$
=\left[\tau_{0}+\tau_{1} \alpha_{1}+\tau_{1}\left(\alpha_{1}-\alpha_{2}\right)\right.
$$

$$
\times \operatorname{cosech}^{2}\left(\frac { \sqrt { \alpha _ { 1 } - \alpha _ { 2 } } } { 2 A } \left(x-\left(\operatorname{an}(5+5 m-7 n) \tau_{0}\right)\right.\right.
$$

$$
\times((1+m)(1+m-n)
$$

$$
\times(1+m+n))^{-1}
$$

$$
\left.\left.\left.\times t-\eta_{0}\right)\right)\right]^{1 /(m-n+1)}
$$

$$
u(x, t)
$$

$$
=\left[\tau_{0}+\tau_{1} \alpha_{1}+\left(\tau_{1}\left(\alpha_{2}-\alpha_{1}\right)\right)\right.
$$

$$
\times\left(\operatorname { s n } ^ { 2 } \left( \pm \frac{\sqrt{\alpha_{2}-\alpha_{1}}}{A}\right.\right.
$$

$$
\times\left(x-\frac{a n(5+5 m-7 n) \tau_{0}}{(1+m)(1+m-n)(1+m+n)}\right.
$$

$$
\left.\times t-\eta_{0}\right)
$$

$$
\begin{equation*}
\left.\left.\left.\frac{\alpha_{1}-\alpha_{3}}{\alpha_{1}-\alpha_{2}}\right)\right)^{-1}\right]^{1 /(m-n+1)} \tag{23}
\end{equation*}
$$

If we take $\tau_{0}=-\tau_{1} \alpha_{1}$ and $\eta_{0}=0$, then solutions (23) can reduce to rational function solution

$$
\begin{equation*}
u(x, t)=\left(\frac{\tilde{A}}{x-c t}\right)^{2 /(m-n+1)} \tag{24}
\end{equation*}
$$

1-soliton wave solution

$$
\begin{equation*}
u(x, t)=\frac{\widetilde{B}}{\cosh ^{2 /(m-n+1)}(B(x-c t))}, \tag{25}
\end{equation*}
$$

singular soliton solution

$$
\begin{equation*}
u(x, t)=\frac{\widetilde{C}}{\sinh ^{2 /(m-n+1)}(C(x-c t))}, \tag{26}
\end{equation*}
$$

and elliptic soliton solution

$$
\begin{equation*}
u(x, t)=\frac{\widetilde{B}}{s n^{2 /(m-n+1)}(\varphi, l)} \tag{27}
\end{equation*}
$$

where $\widetilde{A}=A \sqrt{\tau_{1}}, \widetilde{B}=\left(\tau_{1}\left(\alpha_{2}-\alpha_{1}\right)\right)^{1 /(m-n+1)}, B=$ $\sqrt{\alpha_{2}-\alpha_{1}} / A, \widetilde{C}=\left(\tau_{1}\left(\alpha_{1}-\alpha_{2}\right)\right)^{1 /(m-n+1)}, C=\sqrt{\alpha_{1}-\alpha_{2}} / 2 A$, $\varphi= \pm\left(\sqrt{\alpha_{2}-\alpha_{1}} / A\right)(x-c t), l^{2}=\left(\alpha_{1}-\alpha_{3}\right) /\left(\alpha_{1}-\alpha_{2}\right)$, and $c=a n(5+5 m-7 n) \tau_{1} \alpha_{1} /(1+m)(1+m-n)(1+m+n)$. Here, $\widetilde{B}$ and $\widetilde{C}$ are the amplitudes of the solitons, while $B$ and $C$ are the inverse widths of the solitons and $c$ is the velocity. Thus, we can say that the solitons exist for $\tau_{1}>0$.

Remark 1. If we choose the corresponding values for some parameters, solution (25) is in full agreement with solution (21) mentioned in [17].

Case 2. If we take $\epsilon=0, \delta=2$, and $\theta=4$, then

$$
\begin{equation*}
\left(v^{\prime}\right)^{2}=\frac{\left(\tau_{1}+2 \tau_{2} \Gamma\right)^{2}\left(\xi_{4} \Gamma^{4}+\xi_{3} \Gamma^{3}+\xi_{2} \Gamma^{2}+\xi_{1} \Gamma+\xi_{0}\right)}{\zeta_{0}} \tag{28}
\end{equation*}
$$

where $\zeta_{4} \neq 0$ and $\zeta_{0} \neq 0$. Respectively, solving the algebraic equation system (9) yields

$$
\begin{gathered}
\xi_{0}=\xi_{0}, \quad \xi_{1}=\xi_{1}, \quad \xi_{2}=\frac{\xi_{1}^{2}}{3 \xi_{0}}, \quad \xi_{3}=\frac{\xi_{1}^{3}}{24 \xi_{0}^{2}} \\
\xi_{4}=\frac{\xi_{1}^{4}}{576 \xi_{0}^{3}}, \quad \zeta_{0}=-\frac{b n(m+1)(m+n+1) \xi_{1}^{3}}{24 a(m-n+1)^{2} \xi_{0}^{2} \tau_{1}} \\
\tau_{0}=\frac{2 \xi_{0} \tau_{1}}{\xi_{1}}, \quad \tau_{1}=\tau_{1}, \quad \tau_{2}=\frac{\xi_{1} \tau_{1}}{12 \xi_{0}} \\
c=-\frac{2 a n \xi_{0} \tau_{1}}{(m+1)(m+n+1) \xi_{1}}
\end{gathered}
$$

Substituting these resultss into (6) and (10), we get

$$
\begin{align*}
& \pm\left(\eta-\eta_{0}\right) \\
& \qquad \begin{aligned}
=2 A_{1} \int((d \Gamma) & \\
& \times\left(\Gamma^{4}+\left(\frac{24 \xi_{0}}{\xi_{1}}\right) \Gamma^{3}+\left(\frac{192 \xi_{0}^{2}}{\xi_{1}^{2}}\right) \Gamma^{2}\right. \\
& \left.\left.+\left(\frac{576 \xi_{0}^{3}}{\xi_{1}^{3}}\right) \Gamma+\left(\frac{576 \xi_{0}^{4}}{\xi_{1}^{4}}\right)\right)^{-1 / 2}\right)
\end{aligned}
\end{align*}
$$

where $A_{1}=\sqrt{-6 b n \xi_{0}(1+m)(1+m+n) / a \xi_{1} \tau_{1}(1+m-n)^{2}}$. Integrating (30), we obtain the solutions to (1) as follows:

$$
\begin{gather*}
\pm\left(\eta-\eta_{0}\right)=-\frac{2 A_{1}}{\Gamma-\alpha_{1}}, \\
\pm\left(\eta-\eta_{0}\right)=\frac{4 A_{1}}{\alpha_{1}-\alpha_{2}} \sqrt{\frac{\Gamma-\alpha_{2}}{\Gamma-\alpha_{1}}}, \quad \alpha_{1}>\alpha_{2}, \\
\pm\left(\eta-\eta_{0}\right)=\frac{2 A_{1}}{\alpha_{1}-\alpha_{2}} \ln \left|\frac{\Gamma-\alpha_{1}}{\Gamma-\alpha_{2}}\right|, \\
=\frac{4 A_{1}}{\sqrt{\left(\alpha_{1}-\alpha_{2}\right)\left(\alpha_{1}-\alpha_{3}\right)}} \\
\times \ln \left|\frac{\sqrt{\left(\Gamma-\alpha_{2}\right)\left(\alpha_{1}-\alpha_{3}\right)}-\sqrt{\left(\Gamma-\alpha_{3}\right)\left(\alpha_{1}-\alpha_{2}\right)}}{\sqrt{\left(\Gamma-\alpha_{2}\right)\left(\alpha_{1}-\alpha_{3}\right)}+\sqrt{\left(\Gamma-\alpha_{3}\right)\left(\alpha_{1}-\alpha_{2}\right)}}\right|, \\
\pm\left(\eta-\eta_{0}\right)=\frac{4 A_{1}}{\sqrt{\left(\alpha_{1}-\alpha_{4}\right)\left(\alpha_{2}-\alpha_{3}\right)}} F(\varphi, l), \\
\alpha_{1}>\alpha_{2}>\alpha_{3}>\alpha_{4},
\end{gather*}
$$

where

$$
\begin{align*}
& \varphi_{1}=\arcsin \sqrt{\frac{\left(\Gamma-\alpha_{2}\right)\left(\alpha_{1}-\alpha_{4}\right)}{\left(\Gamma-\alpha_{1}\right)\left(\alpha_{2}-\alpha_{4}\right)}}  \tag{32}\\
& l_{1}^{2}=\frac{\left(\alpha_{1}-\alpha_{3}\right)\left(\alpha_{2}-\alpha_{4}\right)}{\left(\alpha_{2}-\alpha_{3}\right)\left(\alpha_{1}-\alpha_{4}\right)}
\end{align*}
$$

Also $\alpha_{1}, \alpha_{2}, \alpha_{3}$, and $\alpha_{4}$ are the roots of the polynomial equation

$$
\begin{equation*}
\Gamma^{4}+\frac{\xi_{3}}{\xi_{4}} \Gamma^{3}+\frac{\xi_{2}}{\xi_{4}} \Gamma^{2}+\frac{\xi_{1}}{\xi_{4}} \Gamma+\frac{\xi_{0}}{\xi_{4}}=0 \tag{33}
\end{equation*}
$$

Substituting solutions (31) into (5) and (13), we have

$$
\begin{aligned}
& u(x, t) \\
& =\left[\tau_{0}+\tau_{1} \alpha_{1} \pm\left(2 \tau_{1} A_{1}\right)\right. \\
& \quad \times\left(x+\frac{2 a n \xi_{0} \tau_{1}}{(m+1)(m+n+1) \xi_{1}} t-\eta_{0}\right)^{-1}
\end{aligned}
$$

$$
\begin{aligned}
&+\tau_{2}\left(\alpha_{1} \pm\left(2 A_{1}\right)\right. \\
& \times\left(x+\frac{2 a n \xi_{0} \tau_{1}}{(m+1)(m+n+1) \xi_{1}}\right. \\
&\left.\left.\left.\times t-\eta_{0}\right)^{-1}\right)^{2}\right]^{1 /(m-n+1)}
\end{aligned}
$$

$$
\begin{aligned}
& u(x, t) \\
& =\left[\tau_{0}+\tau_{1} \alpha_{1}\right. \\
& \quad+\left(16 A_{1}^{2}\left(\alpha_{2}-\alpha_{1}\right) \tau_{1}\right) \\
& \quad \times\left(16 A_{1}^{2}-\left[\left(\alpha_{1}-\alpha_{2}\right)\right.\right. \\
& \left.\left.\quad \times\left(x+\frac{2 a n \xi_{0} \tau_{1}}{(m+1)(m+n+1) \xi_{1}} t-\eta_{0}\right)\right]^{2}\right)^{-1} \\
& \quad+\tau_{2}\left(\alpha_{1}+\left(16 A_{1}^{2}\left(\alpha_{2}-\alpha_{1}\right)\right)\right. \\
& \quad \times\left(16 A_{1}^{2}-\left[\left(\alpha_{1}-\alpha_{2}\right)\right.\right. \\
&
\end{aligned} \quad \times\left(x+\frac{2 a n \xi_{0} \tau_{1}}{(m+1)(m+n+1) \xi_{1}}\right)
$$

$$
u(x, t)
$$

$$
=\left[\tau_{0}+\tau_{1} \alpha_{2}+\left(\left(\alpha_{2}-\alpha_{1}\right) \tau_{1}\right)\right.
$$

$$
\times\left(\operatorname { e x p } \left[\frac{\alpha_{1}-\alpha_{2}}{2 A_{1}}\right.\right.
$$

$$
\left.\left.\times\left(x+\frac{2 a n \xi_{0} \tau_{1}}{(m+1)(m+n+1) \xi_{1}} t-\eta_{0}\right)\right]-1\right)^{-1}
$$

$$
+\tau_{2}\left(\alpha_{2}+\left(\alpha_{2}-\alpha_{1}\right)\right.
$$

$$
\times\left(\operatorname { e x p } \left[\frac{\alpha_{1}-\alpha_{2}}{2 A_{1}}\right.\right.
$$

$$
\times\left(x+\frac{2 a n \xi_{0} \tau_{1}}{(m+1)(m+n+1) \xi_{1}}\right.
$$

$$
\left.\left.\left.\left.\left.\times t-\eta_{0}\right)\right]-1\right)^{-1}\right)^{2}\right]^{1 /(m-n+1)}
$$

$$
\begin{aligned}
& u(x, t) \\
& =\left[\tau_{0}+\tau_{1} \alpha_{1}+\left(\left(\alpha_{1}-\alpha_{2}\right) \tau_{1}\right)\right. \\
& \times\left(\operatorname { e x p } \left[\frac{\alpha_{1}-\alpha_{2}}{2 A_{1}}\right.\right. \\
& \times\left(x+\left(2 a n \xi_{0} \tau_{1}\right)\right. \\
& \times\left((m+1)(m+n+1) \xi_{1}\right)^{-1} \\
& \left.\left.\left.\times t-\eta_{0}\right)\right]-1\right)^{-1} \\
& +\tau_{2}\left(\alpha_{1}+\left(\alpha_{1}-\alpha_{2}\right)\right. \\
& \times\left(\operatorname { e x p } \left[\frac{\alpha_{1}-\alpha_{2}}{2 A_{1}}\right.\right. \\
& \times\left(x+\left(2 a n \xi_{0} \tau_{1}\right)\right. \\
& \times\left((m+1)(m+n+1) \xi_{1}\right)^{-1} \\
& \left.\left.\left.\left.\left.\times t-\eta_{0}\right)\right]-1\right)^{-1}\right)^{2}\right]^{1 /(m-n+1)}, \\
& u(x, t) \\
& =\left[\tau_{0}+\tau_{1} \alpha_{1}-\left(2\left(\alpha_{1}-\alpha_{2}\right)\left(\alpha_{1}-\alpha_{3}\right) \tau_{1}\right)\right. \\
& \times\left(2 \alpha_{1}-\alpha_{2}-\alpha_{3}+\left(\alpha_{3}-\alpha_{2}\right)\right. \\
& \times \cosh \left[\frac{\sqrt{\left(\alpha_{1}-\alpha_{2}\right)\left(\alpha_{1}-\alpha_{3}\right)}}{2 A_{1}}\right. \\
& \times\left(x+\frac{2 a n \xi_{0} \tau_{1}}{(m+1)(m+n+1) \xi_{1}}\right. \\
& \left.\left.\left.\times t-\eta_{0}\right)\right]\right)^{-1} \\
& +\tau_{2}\left(\alpha_{1}-\left(2\left(\alpha_{1}-\alpha_{2}\right)\left(\alpha_{1}-\alpha_{3}\right)\right)\right. \\
& \times\left(2 \alpha_{1}-\alpha_{2}-\alpha_{3}+\left(\alpha_{3}-\alpha_{2}\right)\right.
\end{aligned}
$$

$$
\begin{aligned}
\times \cosh & {\left[\frac{\sqrt{\left(\alpha_{1}-\alpha_{2}\right)\left(\alpha_{1}-\alpha_{3}\right)}}{2 A_{1}}\right.} \\
& \times\left(x+\frac{2 a n \xi_{0} \tau_{1}}{(m+1)(m+n+1) \xi_{1}}\right. \\
& \left.\left.\left.\left.\times t-\eta_{0}\right)\right]\right]^{-1}\right]^{2 /(m-n+1)}
\end{aligned}
$$

$u(x, t)$

$$
=\left[\tau_{0}+\tau_{1} \alpha_{2}+\left(\left(\alpha_{1}-\alpha_{2}\right)\left(\alpha_{4}-\alpha_{2}\right) \tau_{1}\right)\right.
$$

$$
\begin{aligned}
& \times\binom{\left(\alpha_{1}-\alpha_{4}\right)}{\quad \times \operatorname{sn}^{2}\left(\frac{\sqrt{\left(\alpha_{1}-\alpha_{3}\right)\left(\alpha_{2}-\alpha_{4}\right)}}{4 A_{1}}\right.}
\end{aligned}
$$

$$
\times\left(x+\frac{2 a n \xi_{0} \tau_{1}}{(m+1)(m+n+1) \xi_{1}}\right.
$$

$$
\left.\times t-\eta_{0}\right)
$$

$$
\left.\frac{\left(\alpha_{2}-\alpha_{3}\right)\left(\alpha_{1}-\alpha_{4}\right)}{\left(\alpha_{1}-\alpha_{3}\right)\left(\alpha_{2}-\alpha_{4}\right)}\right)
$$

$$
\begin{aligned}
& \left.+\alpha_{4}-\alpha_{2}\right)^{-1} \\
& +\tau_{2}\left(\alpha_{2}+\left(\left(\alpha_{1}-\alpha_{2}\right)\left(\alpha_{4}-\alpha_{2}\right) \tau_{1}\right)\right.
\end{aligned}
$$

$$
\times\left(\left(\alpha_{1}-\alpha_{4}\right)\right.
$$

$$
\times \operatorname{sn}^{2}\left(\frac{\sqrt{\left(\alpha_{1}-\alpha_{3}\right)\left(\alpha_{2}-\alpha_{4}\right)}}{4 A_{1}}\right.
$$

$$
\times\left(x+\frac{2 a n \xi_{0} \tau_{1}}{(m+1)(m+n+1) \xi_{1}}\right.
$$

$$
\left.\times t-\eta_{0}\right)
$$

$$
\begin{array}{r}
\left.\frac{\left(\alpha_{2}-\alpha_{3}\right)\left(\alpha_{1}-\alpha_{4}\right)}{\left(\alpha_{1}-\alpha_{3}\right)\left(\alpha_{2}-\alpha_{4}\right)}\right) \\
\left.\left.\left.+\alpha_{4}-\alpha_{2}\right)^{-1}\right)^{2}\right]^{1 /(m-n+1)} \tag{34}
\end{array}
$$

For simplicity, if we take $\eta_{0}=0$, then we can write solutions (34) as follows:

$$
\begin{align*}
& u(x, t)=\left[\sum_{i=0}^{2} \tau_{i}\left(\alpha_{1} \pm \frac{2 A_{1}}{x-c t}\right)^{i}\right]^{1 /(m-n+1)}, \\
& u(x, t) \\
& =\left[\sum _ { i = 0 } ^ { 2 } \tau _ { i } \left(\alpha_{1}+\left(16 A_{1}^{2}\left(\alpha_{1}-\alpha_{2}\right)\right)\right.\right. \\
& \left.\left.\times\left(16 A_{1}^{2}-\left[\left(\alpha_{1}-\alpha_{2}\right)(x-c t)\right]^{2}\right)^{-1}\right)^{i}\right]^{1 /(m-n+1)}, \\
& u(x, t)=\left[\sum_{i=0}^{2} \tau_{i}\left(\alpha_{2}+\frac{\alpha_{2}-\alpha_{1}}{\exp \left[B_{1}(x-c t)\right]-1}\right)^{i}\right]^{1 /(m-n+1)}, \\
& u(x, t)=\left[\sum_{i=0}^{2} \tau_{i}\left(\alpha_{1}+\frac{\alpha_{1}-\alpha_{2}}{\exp \left[B_{1}(x-c t)\right]-1}\right)^{i}\right]^{1 /(m-n+1)}, \\
& u(x, t)=\left[\sum _ { i = 0 } ^ { 2 } \tau _ { i } \left(\alpha_{1}-\left(2\left(\alpha_{1}-\alpha_{2}\right)\left(\alpha_{1}-\alpha_{3}\right)\right)\right.\right. \\
& \times\left(2 \alpha_{1}-\alpha_{2}-\alpha_{3}+\left(\alpha_{3}-\alpha_{2}\right)\right. \\
& \left.\left.\left.\times \cosh \left[C_{1}(x-c t)\right]\right)^{-1}\right)^{i}\right]^{1 /(m-n+1)}, \\
& u(x, t)=\left[\sum _ { i = 0 } ^ { 2 } \tau _ { i } \left(\alpha_{2}+\left(\left(\alpha_{1}-\alpha_{2}\right)\left(\alpha_{4}-\alpha_{2}\right)\right)\right.\right. \\
& \times\left(\left(\alpha_{1}-\alpha_{4}\right) s n^{2}(\varphi, l)\right. \\
& \left.\left.\left.+\alpha_{4}-\alpha_{2}\right)^{-1}\right)^{i}\right]^{1 /(m-n+1)}, \tag{35}
\end{align*}
$$

where $B_{1}=\left(\alpha_{1}-\alpha_{2}\right) / 2 A_{1}, C_{1}=\sqrt{\left(\alpha_{1}-\alpha_{2}\right)\left(\alpha_{1}-\alpha_{3}\right)} / 2 A_{1}$, $\varphi_{1}=\left(\sqrt{\left(\alpha_{1}-\alpha_{3}\right)\left(\alpha_{2}-\alpha_{4}\right)} / 4 A_{1}\right)(x-c t), l_{1}^{2}=\left(\alpha_{2}-\alpha_{3}\right)\left(\alpha_{1}-\right.$ $\left.\alpha_{4}\right) /\left(\alpha_{1}-\alpha_{3}\right)\left(\alpha_{2}-\alpha_{4}\right)$, and $c=-2 a n \xi_{0} \tau_{1} /(m+1)(m+n+1) \xi_{1}$. Here, $A_{1}$ is the amplitude of the soliton, while $c$ is the velocity
and $B_{1}$ and $C_{1}$ are the inverse widths of the solitons. Thus, we can say that the solitons exist for $\tau_{1}>0$.

Case 3. If we take $\epsilon=0, \delta=3$, and $\theta=5$, then

$$
\begin{align*}
\left(v^{\prime}\right)^{2}= & \left(\tau_{1}+2 \tau_{2} \Gamma+3 \tau_{3} \Gamma^{2}\right)^{2} \\
& \times\left(\xi_{5} \Gamma^{5}+\xi_{4} \Gamma^{4}+\xi_{3} \Gamma^{3}+\xi_{2} \Gamma^{2}+\xi_{1} \Gamma+\xi_{0}\right)  \tag{36}\\
& \times\left(\zeta_{0}\right)^{-1}
\end{align*}
$$

where $\xi_{5} \neq 0$ and $\zeta_{0} \neq 0$. Respectively, solving the algebraic equation system (9) yields

$$
\begin{gather*}
\xi_{0}=\frac{\xi_{5}\left(\tau_{2}^{2}-4 \tau_{1} \tau_{3}\right)\left(2 \tau_{2}^{3}-9 \tau_{1} \tau_{2} \tau_{3}+2 \sqrt{\left(\tau_{2}^{2}-3 \tau_{1} \tau_{3}\right)^{3}}\right)}{81 \tau_{3}^{5}}, \\
\xi_{1}=-\frac{\xi_{5}\left(4 \tau_{2}^{4}+9 \tau_{1} \tau_{2}^{2} \tau_{3}-108 \tau_{1}^{2} \tau_{3}^{2}+4 \tau_{2} \sqrt{\left(\tau_{2}^{2}-3 \tau_{1} \tau_{3}\right)^{3}}\right)}{81 \tau_{3}^{4}}, \\
\xi_{2}=\frac{\xi_{5}\left(-11 \tau_{2}^{3}+63 \tau_{1} \tau_{2} \tau_{3}-2 \sqrt{\left(\tau_{2}^{2}-3 \tau_{1} \tau_{3}\right)^{3}}\right)}{27 \tau_{3}^{3}}, \\
\xi_{3}=\frac{\xi_{5}\left(\tau_{2}^{2}+7 \tau_{1} \tau_{3}\right)}{3 \tau_{3}^{2}}, \quad \xi_{4}=\frac{5 \xi_{5} \tau_{2}}{3 \tau_{3}}, \\
\zeta_{0}=-\frac{9 b n(m+1)(m+n+1) \xi_{5}}{2 a \tau_{3}(m-n+1)^{2}}, \\
\tau_{0}=-\frac{2 \tau_{2}^{3}-9 \tau_{1} \tau_{2} \tau_{3}+2 \sqrt{\left(\tau_{2}^{2}-3 \tau_{1} \tau_{3}\right)^{3}}}{27 \tau_{3}^{2}} \\
c=-\frac{8 a n \sqrt{\left(\tau_{2}^{2}-3 \tau_{1} \tau_{3}\right)^{3}}}{27(m+1)(m+n+1) \tau_{3}^{2}}
\end{gather*}
$$

Substituting these results into (6) and (10), we get

$$
\begin{aligned}
& \pm\left(\eta-\eta_{0}\right) \\
& \qquad \begin{aligned}
=3 A_{2} \int( & (d \Gamma) \\
& \times\left(\Gamma^{5}+\frac{\xi_{4}}{\xi_{5}} \Gamma^{4}+\frac{\xi_{3}}{\xi_{5}} \Gamma^{3}\right. \\
& \left.\left.+\frac{\xi_{2}}{\xi_{5}} \Gamma^{2}+\frac{\xi_{1}}{\xi_{5}} \Gamma+\frac{\xi_{0}}{\xi_{5}}\right)^{-1 / 2}\right)
\end{aligned}
\end{aligned}
$$

where $A_{2}=\sqrt{-b n(1+m)(1+m+n) / 2 a \tau_{3}(1+m-n)^{2}}$. Integrating (38), we obtain the solutions to (1) as follows:

$$
\begin{align*}
& \pm\left(\eta-\eta_{0}\right)=-\frac{2 A_{2}}{\sqrt{\left(\Gamma-\alpha_{1}\right)^{3}}}, \\
& \pm\left(\eta-\eta_{0}\right)= \frac{3 A_{2} \operatorname{arctanh}\left[\sqrt{\left(\Gamma-\alpha_{2}\right) /\left(\alpha_{1}-\alpha_{2}\right)}\right]}{\left(\alpha_{1}-\alpha_{2}\right)^{3 / 2}} \\
&-\frac{3 A_{2} \sqrt{\Gamma-\alpha_{2}}}{\left(\alpha_{1}-\alpha_{2}\right)\left(\Gamma-\alpha_{1}\right)}, \alpha_{1}>\alpha_{2} \\
& \pm\left(\eta-\eta_{0}\right)=-\frac{6 A_{2} \arctan \left[\sqrt{\left(\Gamma-\alpha_{1}\right) /\left(\alpha_{1}-\alpha_{2}\right)}\right]}{\left(\alpha_{1}-\alpha_{2}\right)^{3 / 2}} \\
&-\frac{6 A_{2}}{\sqrt{\Gamma-\alpha_{1}}\left(\alpha_{1}-\alpha_{2}\right)}, \\
& \pm\left(\eta-\eta_{0}\right)= \frac{6 A_{2} \operatorname{arctanh}\left[\sqrt{\left(\Gamma-\alpha_{3}\right) /\left(\alpha_{2}-\alpha_{3}\right)}\right]}{\alpha_{1}-\alpha_{2}} \\
& \pm\left(\eta-\eta_{0}\right) \\
& \times \frac{1}{\sqrt{\Gamma-\alpha_{1}}\left(\alpha_{1}-\alpha_{2}\right)\left(\alpha_{1}-\alpha_{3}\right)} \\
& \quad \times\left[\sqrt{\left(\Gamma-\alpha_{2}\right)\left(\Gamma-\alpha_{3}\right)}+i(E(\varphi, l)-F(\varphi, l))\right]
\end{align*}
$$

where

$$
E(\varphi, l)=\int_{0}^{\varphi} \sqrt{1-l^{2} \sin ^{2} \psi} d \psi
$$

$$
\varphi_{2}=-\arcsin \sqrt{\frac{\Gamma-\alpha_{1}}{\alpha_{2}-\alpha_{1}}}
$$

$$
\begin{equation*}
l_{2}^{2}=\frac{\alpha_{1}-\alpha_{2}}{\alpha_{1}-\alpha_{3}} \tag{40}
\end{equation*}
$$

$$
\begin{aligned}
& \pm\left(\eta-\eta_{0}\right) \\
& \qquad \begin{array}{l}
=\frac{-6 i A_{2}}{\sqrt{\alpha_{2}-\alpha_{3}}\left(\alpha_{1}-\alpha_{2}\right)}(F(\varphi, l)-\pi(\varphi, n, l)) \\
\\
\alpha_{1}>\alpha_{2}>\alpha_{3}>\alpha_{4}
\end{array}
\end{aligned}
$$

where

$$
\begin{gather*}
\varphi_{3}=-\arcsin \sqrt{\frac{\alpha_{3}-\alpha_{2}}{\Gamma-\alpha_{2}}}, \quad l_{3}^{2}=\frac{\alpha_{2}-\alpha_{4}}{\alpha_{2}-\alpha_{3}},  \tag{41}\\
n=\frac{\alpha_{2}-\alpha_{1}}{\alpha_{2}-\alpha_{3}}
\end{gather*}
$$

Also $\alpha_{1}, \alpha_{2}, \alpha_{3}, \alpha_{4}$, and $\alpha_{5}$ are the roots of the polynomial equation

$$
\begin{equation*}
\Gamma^{5}+\frac{\xi_{4}}{\xi_{5}} \Gamma^{4}+\frac{\xi_{3}}{\xi_{5}} \Gamma^{3}+\frac{\xi_{2}}{\xi_{5}} \Gamma^{2}+\frac{\xi_{1}}{\xi_{5}} \Gamma+\frac{\xi_{0}}{\xi_{5}}=0 . \tag{42}
\end{equation*}
$$

Case 4. If we take $\epsilon=1, \delta=1$, and $\theta=4$, then

$$
\begin{equation*}
\left(v^{\prime}\right)^{2}=\frac{\tau_{1}^{2}\left(\xi_{4} \Gamma^{4}+\xi_{3} \Gamma^{3}+\xi_{2} \Gamma^{2}+\xi_{1} \Gamma+\xi_{0}\right)}{\zeta_{0}+\zeta_{1} \Gamma} \tag{43}
\end{equation*}
$$

where $\xi_{4} \neq 0$ and $\zeta_{1} \neq 0$. Respectively, solving the algebraic equation system (9) yields

$$
\begin{align*}
& \xi_{0}= \frac{\zeta_{0} \tau_{0}^{2}\left(M+2 a(1+m-n)^{2}\left(2 \zeta_{1} \tau_{0}+\zeta_{0} \tau_{1}\right)\right)}{b n(1+m)(1+m+n) \zeta_{1} \tau_{1}^{2}}, \\
& \xi_{3}= \xi_{3}, \\
& \xi_{4}=-\frac{2 a(1+m-n)^{2} \zeta_{1} \tau_{1}}{b n(1+m)(1+m+n)}, \\
& \xi_{1}=\left(\tau _ { 0 } \left(4 a(1+m-n)^{2} \zeta_{1}^{2} \tau_{0}^{2}\right.\right. \\
&+2 \zeta_{0} \tau_{1}\left(M+2 a(1+m-n)^{2} \zeta_{0} \tau_{1}\right) \\
&\left.\left.+\zeta_{1} \tau_{0}\left(M+8 a(1+m-n)^{2} \zeta_{0} \tau_{1}\right)\right)\right) \\
& \times\left(b n(1+m)(1+m+n) \zeta_{1} \tau_{1}^{2}\right)^{-1},  \tag{44}\\
& \xi_{2}=\left(6 a(1+m-n)^{2} \zeta_{1}^{2} \tau_{0}^{2}\right. \\
&+2 \zeta_{1} \tau_{0}\left(M+2 a(1+m-n)^{2} \zeta_{0} \tau_{1}\right) \\
&\left.+\zeta_{0} \tau_{1}\left(M+2 a(1+m-n)^{2} \zeta_{0} \tau_{1}\right)\right) \\
& \times\left(b n(1+m)(1+m+n) \zeta_{1} \tau_{1}^{2}\right)^{-1} \\
& \zeta_{0}= \zeta_{0}, \\
& \tau_{0}= \tau_{0}, \\
& \zeta_{1}=\tau_{1}=\tau_{1}, \\
& c=\frac{n\left(M+2 a(1+m-n)^{2}\left(3 \zeta_{1} \tau_{0}+\zeta_{0} \tau_{1}\right)\right)}{(1+m)(1+m+n)(1+m-n)^{2} \zeta_{1}}
\end{align*}
$$

where $M=b n(1+m)(1+m+n) \xi_{3}$. Substituting these results into (6) and (10), we get

$$
\begin{aligned}
& \pm\left(\eta-\eta_{0}\right) \\
& \qquad \begin{array}{l}
=A_{3} \int\left(\left(\Gamma+\frac{\zeta_{0}}{\zeta_{1}}\right)\right. \\
\\
\times\left(\Gamma^{4}+\left(\frac{\xi_{3}}{\xi_{4}}\right) \Gamma^{3}+\left(\frac{\xi_{2}}{\xi_{4}}\right) \Gamma^{2}\right. \\
\\
\left.\left.\quad+\left(\frac{\xi_{1}}{\xi_{4}}\right) \Gamma+\left(\frac{\xi_{0}}{\xi_{4}}\right)\right)^{-1}\right)^{1 / 2} d \Gamma
\end{array}
\end{aligned}
$$

$$
\begin{align*}
& \times\left(\zeta_{0}(F(\varphi, l)-\pi(\varphi, n, l))\right. \\
& \left.\quad+\zeta_{1}\left(\alpha_{2} F(\varphi, l)-\alpha_{2} \pi(\varphi, n, l)\right)\right) \tag{48}
\end{align*}
$$

where

$$
\begin{align*}
\varphi_{6} & =-\arcsin \sqrt{\frac{\zeta_{0}+\zeta_{1} \alpha_{2}}{\zeta_{1}\left(\alpha_{2}-\Gamma\right)}}  \tag{49}\\
l_{6}^{2} & =\frac{\zeta_{1}\left(\alpha_{2}-\alpha_{3}\right)}{\zeta_{0}+\zeta_{1} \alpha_{2}}, \quad n_{1}=\frac{\zeta_{1}\left(\alpha_{2}-\alpha_{1}\right)}{\zeta_{0}+\zeta_{1} \alpha_{2}}
\end{align*}
$$

Case 5. If we take $\epsilon=1, \delta=2$, and $\theta=5$, then

$$
\begin{equation*}
\left(v^{\prime}\right)^{2}=\frac{\left(\tau_{1}+2 \tau_{2} \Gamma\right)^{2}\left(\xi_{5} \Gamma^{5}+\xi_{4} \Gamma^{4}+\xi_{3} \Gamma^{3}+\xi_{2} \Gamma^{2}+\xi_{1} \Gamma+\xi_{0}\right)}{\zeta_{0}+\zeta_{1} \Gamma} \tag{50}
\end{equation*}
$$

where $\zeta_{5} \neq 0$ and $\zeta_{1} \neq 0$. Respectively, solving the algebraic equation system (9) yields

$$
\begin{gather*}
\xi_{0}=\frac{\tau_{0}^{2}\left(-2 \xi_{5} \tau_{1}+\xi_{4} \tau_{2}\right)}{\tau_{2}^{3}}, \\
\xi_{1}=\frac{\tau_{0}\left(2 \xi_{4} \tau_{1} \tau_{2}+\xi_{5}\left(-4 \tau_{1}^{2}+\tau_{0} \tau_{2}\right)\right)}{\tau_{2}^{3}}, \\
\xi_{2}=\frac{\xi_{4} \tau_{2}\left(\tau_{1}^{2}+2 \tau_{0} \tau_{2}\right)-2 \xi_{5}\left(\tau_{1}^{3}+\tau_{0} \tau_{1} \tau_{2}\right)}{\tau_{2}^{3}}, \\
\zeta_{0}=\frac{-2 b n(1+m)(1+m+n)\left(\xi_{4} \tau_{2}-2 \xi_{5} \tau_{1}\right)}{a(1+m-n)^{2} \tau_{2}^{2}} \\
\xi_{3}=\frac{-3 \xi_{5} \tau_{1}^{2}+2 \tau_{2}\left(\xi_{5} \tau_{0}+\xi_{4} \tau_{1}\right)}{\tau_{2}^{2}}  \tag{51}\\
\zeta_{1}=\frac{-2 b n(1+m)(1+m+n) \xi_{5}}{a(1+m-n)^{2} \tau_{2}^{2}} \\
\xi_{5}=\xi_{5} \\
\tau_{0}=\tau_{0}, \\
\tau_{4}=\xi_{4}, \\
c=-\frac{\tau_{1},}{2(1+m)(1+m+n) \tau_{2}}
\end{gather*}
$$

Substituting these results into (6) and (10), we get

$$
\begin{align*}
& \pm\left(\eta-\eta_{0}\right)=A_{4} \\
& \qquad \begin{aligned}
& \times \int\left(\left(\Gamma+\frac{\zeta_{0}}{\zeta_{1}}\right)\right. \\
& \times\left(\Gamma^{5}+\frac{\xi_{4}}{\xi_{5}} \Gamma^{4}+\frac{\xi_{3}}{\xi_{5}} \Gamma^{3}\right. \\
& \left.\left.+\frac{\xi_{2}}{\xi_{5}} \Gamma^{2}+\frac{\xi_{1}}{\xi_{5}} \Gamma+\frac{\xi_{0}}{\xi_{5}}\right)^{-1}\right)^{1 / 2} d \Gamma
\end{aligned}
\end{align*}
$$

where $A_{4}=\sqrt{-2 b n(1+m)(1+m+n) / a \tau_{2}^{2}(1+m-n)^{2}}$.
Integrating (52), we obtain the solution to (1) as follows:

$$
\begin{aligned}
& \pm\left(\eta-\eta_{0}\right)=\frac{-2 A_{4}}{3 \sqrt{\zeta_{1}}\left(\zeta_{0}+\zeta_{1} \alpha_{1}\right)}\left(\frac{\zeta_{0}+\zeta_{1} \Gamma}{\Gamma-\alpha_{1}}\right)^{3 / 2}, \\
& \pm\left(\eta-\eta_{0}\right) \\
& =\frac{-A_{4}\left(\zeta_{0}+\zeta_{1} \alpha_{2}\right)}{2\left(\alpha_{1}-\alpha_{2}\right)^{3 / 2} \sqrt{\zeta_{1}\left(\zeta_{0}+\zeta_{1} \alpha_{1}\right)}} \\
& \times \ln \mid\left(\Gamma-\alpha_{1}\right) \\
& \times\left(\zeta_{0}\left(\Gamma+\alpha_{1}-2 \alpha_{2}\right)\right. \\
& +2 \sqrt{\left(\zeta_{0}+\zeta_{1} \Gamma\right)\left(\zeta_{0}+\zeta_{1} \alpha_{1}\right)\left(\Gamma-\alpha_{2}\right)\left(\alpha_{1}-\alpha_{2}\right)} \\
& \left.+\zeta_{1}\left(2 \Gamma \alpha_{1}-\alpha_{2}\left(\Gamma+\alpha_{1}\right)\right)\right)^{-1} \mid \\
& -\frac{A_{4}}{\left(\alpha_{1}-\alpha_{2}\right)\left(\Gamma-\alpha_{1}\right)} \sqrt{\frac{\left(\zeta_{0}+\zeta_{1} \Gamma\right)\left(\Gamma-\alpha_{2}\right)}{\zeta_{1}}}, \quad \alpha_{1}>\alpha_{2}, \\
& \pm\left(\eta-\eta_{0}\right) \\
& =\frac{-2 A_{4}}{\left(\alpha_{1}-\alpha_{2}\right)} \sqrt{\frac{\zeta_{0}+\zeta_{1} \Gamma}{\zeta_{1}\left(\Gamma-\alpha_{1}\right)}} \\
& -\frac{2 A_{4}}{\left(\alpha_{1}-\alpha_{2}\right)^{3 / 2}} \sqrt{\frac{\zeta_{0}+\zeta_{1} \alpha_{2}}{\zeta_{1}}} \\
& \times \arctan \left[\sqrt{\frac{\left(\Gamma-\alpha_{1}\right)\left(\zeta_{0}+\zeta_{1} \alpha_{2}\right)}{\left(\alpha_{1}-\alpha_{2}\right)\left(\zeta_{0}+\zeta_{1} \Gamma\right)}}\right], \\
& \pm\left(\eta-\eta_{0}\right) \\
& =\frac{-A_{4}}{\alpha_{1}-\alpha_{3}} \sqrt{\frac{\zeta_{0}+\zeta_{1} \alpha_{2}}{\zeta_{1}\left(\alpha_{2}-\alpha_{3}\right)}} \\
& \times \ln \mid\left(\alpha_{2}-\Gamma\right) \\
& \times\left(\zeta_{0}\left(\Gamma+\alpha_{2}-2 \alpha_{3}\right)\right. \\
& +2 \sqrt{\left(\zeta_{0}+\zeta_{1} \Gamma\right)\left(\zeta_{0}+\zeta_{1} \alpha_{2}\right)\left(\Gamma-\alpha_{3}\right)\left(\alpha_{2}-\alpha_{3}\right)} \\
& \left.+\zeta_{1}\left(2 \Gamma \alpha_{2}-\alpha_{3}\left(\Gamma+\alpha_{2}\right)\right)\right)^{-1} \mid \\
& -\frac{A_{4}}{\alpha_{1}-\alpha_{3}} \\
& \times \sqrt{\frac{\zeta_{0}+\zeta_{1} \alpha_{1}}{\zeta_{1}\left(\alpha_{1}-\alpha_{3}\right)}} \\
& \times \ln \mid\left(\zeta_{0}\left(\Gamma+\alpha_{1}-2 \alpha_{3}\right)\right. \\
& +2 \sqrt{\left(\zeta_{0}+\zeta_{1} \Gamma\right)\left(\zeta_{0}+\zeta_{1} \alpha_{1}\right)\left(\Gamma-\alpha_{3}\right)\left(\alpha_{1}-\alpha_{3}\right)}
\end{aligned}
$$

$$
\begin{gather*}
\left.+\zeta_{1}\left(2 \Gamma \alpha_{1}-\alpha_{3}\left(\Gamma+\alpha_{1}\right)\right)\right) \\
\times\left(\Gamma-\alpha_{2}\right)^{-1} \mid, \quad \alpha_{1}>\alpha_{2}>\alpha_{3} \\
\pm\left(\eta-\eta_{0}\right)= \\
\frac{-2 A_{4}}{\alpha_{1}-\alpha_{3}} \\
 \tag{53}\\
\times \sqrt{\frac{\zeta_{0}+\zeta_{1} \alpha_{3}}{\zeta_{1}\left(\alpha_{1}-\alpha_{2}\right)}} E(\varphi, l), \\
\alpha_{1}>\alpha_{2}>\alpha_{3}
\end{gather*}
$$

where

$$
\begin{align*}
& \varphi_{7}= \arcsin \sqrt{\frac{\left(\Gamma-\alpha_{3}\right)\left(\alpha_{2}-\alpha_{1}\right)}{\left(\Gamma-\alpha_{1}\right)\left(\alpha_{2}-\alpha_{3}\right)}}, \\
& l_{7}^{2}= \frac{\left(\alpha_{3}-\alpha_{2}\right)\left(\zeta_{0}+\zeta_{1} \alpha_{1}\right)}{\left(\alpha_{1}-\alpha_{2}\right)\left(\zeta_{0}+\zeta_{1} \alpha_{3}\right)}, \\
& \pm\left(\eta-\eta_{0}\right) \\
&= \frac{2 A_{4}\left(\alpha_{2}-\alpha_{4}\right)}{\left(\alpha_{1}-\alpha_{2}\right)\left(\alpha_{3}-\alpha_{4}\right) \sqrt{\zeta_{1}\left(\alpha_{2}-\alpha_{3}\right)\left(\zeta_{0}+\zeta_{1} \alpha_{4}\right)}}  \tag{54}\\
& \times\left(\frac{\left(\zeta_{0}+\zeta_{1} \Gamma\right)\left(\alpha_{3}-\alpha_{4}\right)}{\alpha_{1}-\alpha_{2}} \pi(\varphi, n, l)\right. \\
&\left.\quad+\frac{\left(\zeta_{0}+\zeta_{1} \alpha_{2}\right)\left(\alpha_{4}-\alpha_{3}\right)}{\alpha_{2}-\alpha_{4}} F(\varphi, l)\right)
\end{align*}
$$

where

$$
\begin{gather*}
\varphi_{8}=\arcsin \sqrt{\frac{\left(\Gamma-\alpha_{3}\right)\left(\alpha_{2}-\alpha_{1}\right)}{\left(\Gamma-\alpha_{1}\right)\left(\alpha_{2}-\alpha_{3}\right)}} \\
l_{8}^{2}=\frac{\left(\alpha_{3}-\alpha_{2}\right)\left(\zeta_{0}+\zeta_{1} \alpha_{1}\right)}{\left(\alpha_{1}-\alpha_{2}\right)\left(\zeta_{0}+\zeta_{1} \alpha_{3}\right)}  \tag{55}\\
n_{2}=-\frac{\left(\alpha_{1}-\alpha_{2}\right)\left(\alpha_{3}-\alpha_{4}\right)}{\left(\alpha_{2}-\alpha_{3}\right)\left(\alpha_{1}-\alpha_{4}\right)} \\
\alpha_{1}>\alpha_{2}>\alpha_{3}>\alpha_{4}
\end{gather*}
$$

## 4. Discussion

Thus we introduce a more general extended trial equation method for nonlinear partial differential equations as follows.

Step 1. Extended trial equation (6) can be reduced to the following more general form:

$$
\begin{equation*}
u=\frac{A(\Gamma)}{B(\Gamma)}=\frac{\sum_{i=0}^{\delta} \tau_{i} \Gamma^{i}}{\sum_{j=0}^{\mu} \omega_{j} \Gamma^{j}}, \tag{56}
\end{equation*}
$$

where

$$
\begin{equation*}
\left(\Gamma^{\prime}\right)^{2}=\Lambda(\Gamma)=\frac{\Phi(\Gamma)}{\Psi(\Gamma)}=\frac{\xi_{\theta} \Gamma^{\theta}+\cdots+\xi_{1} \Gamma+\xi_{0}}{\zeta_{\epsilon} \Gamma^{\epsilon}+\cdots+\zeta_{1} \Gamma+\zeta_{0}} \tag{57}
\end{equation*}
$$

Here, $\tau_{i}(i=0, \ldots, \delta), \omega_{j}(j=0, \ldots, \mu), \xi_{\varsigma}(\varsigma=0, \ldots, \theta)$, and $\zeta_{\sigma}(\sigma=0, \ldots, \epsilon)$ are the constants to be specified.

Step 2. Taking trial equations (56) and (57), we derive the following equations:

$$
\begin{align*}
& \left(u^{\prime}\right)^{2}=\frac{\Phi(\Gamma)}{\Psi(\Gamma)} \frac{\left(A^{\prime}(\Gamma) B(\Gamma)-A(\Gamma) B^{\prime}(\Gamma)\right)^{2}}{B^{4}(\Gamma)},  \tag{58}\\
u^{\prime \prime}= & \left(A^{\prime}(\Gamma) B(\Gamma)-A(\Gamma) B^{\prime}(\Gamma)\right) \\
& \times\left\{\left(\Phi^{\prime}(\Gamma) \Psi(\Gamma)-\Phi(\Gamma) \Psi^{\prime}(\Gamma)\right) B(\Gamma)\right. \\
& \left.-4 \Phi(\Gamma) \Psi(\Gamma) B^{\prime}(\Gamma)\right\} \\
& +2 \Phi(\Gamma) \Psi(\Gamma) B(\Gamma)\left(A^{\prime \prime}(\Gamma) B(\Gamma)-A(\Gamma) B^{\prime \prime}(\Gamma)\right) \\
& \times\left(2 B^{3}(\Gamma) \Psi^{2}(\Gamma)\right)^{-1} \tag{59}
\end{align*}
$$

and other derivation terms such as $u^{\prime \prime \prime}$.
Step 3. Substituting $u^{\prime}, u^{\prime \prime}$, and other derivation terms into (5) yields the following equation:

$$
\begin{equation*}
\Omega(\Gamma)=\varrho_{s} \Gamma^{s}+\cdots+\varrho_{1} \Gamma+\varrho_{0}=0 \tag{60}
\end{equation*}
$$

According to the balance principle we can determine a relation of $\theta, \epsilon, \delta$, and $\mu$.

Step 4. Letting the coefficients of $\Omega(\Gamma)$ all be zero will yield an algebraic equations system $\varrho_{i}=0 \quad(i=0, \ldots, s)$. Solving this equations system, we will determine the values $\tau_{0}, \ldots \tau_{\delta}$; $\omega_{0}, \ldots, \omega_{\mu} ; \xi_{0}, \ldots, \xi_{\theta}$; and $\zeta_{0}, \ldots, \zeta_{\epsilon}$.

Step 5. Substituting the results obtained in Step 4 into (57) and integrating (57), we can find the exact solutions of (3).

## 5. Conclusions and Remarks

In this study, we proposed an extended trial equation method and used it to obtain some soliton and elliptic function solutions to the generalized $K(m, n)$ equation. Otherwise, we discussed a more general trial equation method. The proposed method can also be applied to other nonlinear differential equations with nonlinear evolution.

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# Numerical Solution of the Fractional Partial Differential Equations by the Two-Dimensional Fractional-Order Legendre Functions 

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Received 13 May 2013; Revised 8 September 2013; Accepted 8 September 2013
Academic Editor: Santanu Saha Ray
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#### Abstract

A numerical method is presented to obtain the approximate solutions of the fractional partial differential equations (FPDEs). The basic idea of this method is to achieve the approximate solutions in a generalized expansion form of two-dimensional fractionalorder Legendre functions (2D-FLFs). The operational matrices of integration and derivative for 2D-FLFs are first derived. Then, by these matrices, a system of algebraic equations is obtained from FPDEs. Hence, by solving this system, the unknown 2D-FLFs coefficients can be computed. Three examples are discussed to demonstrate the validity and applicability of the proposed method.


## 1. Introduction

Fractional partial differential equations play a significant role in modeling physical and engineering processes. Therefore, there is an urgent need to develop efficient and fast convergent methods for FPDEs. Recently, several different techniques, including Adomian's decomposition method (ADM) [1, 2], homotopy perturbation method (HPM) [3-5], variational iteration method (VIM) [6-8], spectral methods [913], orthogonal polynomials method [14-17], and wavelets method [18-21] have been presented and applied to solve FPDEs.

The method based on the orthogonal functions is a wonderful and powerful tool for solving the FDEs and has enjoyed many successes in this realm. The operational matrix of fractional integration has been determined for some types of orthogonal polynomials, such as Chebyshev polynomials [16], Legendre polynomials [22], Laguerre polynomials [2325], and Jacobi polynomials [26]. Moreover, the operational matrix of fractional derivative for Chebyshev polynomials [9] and Legendre polynomials [9,14] also has been derived. However, since these polynomials using integer power series to approximate fractional ones, it cannot accurately represent
properties of fractional calculus. Recently, Rida and Yousef [27] presented a fractional extension of the classical Legendre polynomials by replacing the integer order derivative in Rodrigues formula with fractional order derivatives. The defect is that the complexity of these functions made them unsuitable for solving FDEs. Subsequently, Kazem et al. [28] presented the orthogonal fractional order Legendre functions based on shifted Legendre polynomials to find the numerical solution of FDEs and drew a conclusion that their method is accurate, effective, and easy to implement.

Benefiting from their "exponential-convergence" property when smooth solutions are involved, spectral methods have been widely and effectively used for the numerical solution of partial differential equations. The basic idea of spectral methods is to expand a function into sets of smooth global functions, called the trial functions. Because of their special properties, the orthogonal polynomials are usually chosen to be trial functions. Spectral methods can obtain very accurate approximations for a smooth solution while only need a few degrees of freedom. Recently, Chebyshev spectral method [9], Legendre spectral method [10], and adaptive pseudospectral method [11] were proposed for solving fractional boundary value problems. Moreover, generalized Laguerre spectral
algorithms and Legendre spectral Galerkin method were developed by Baleanu et al. [12] and Bhrawy and Alghamdi [13] for fractional initial value problems, respectively.

Motivated and inspired by the ongoing research in orthogonal polynomials methods and spectral methods, we construct two-dimensional fractional-order Legendre functions and derive the operational matrices of integration and derivative for the solution of FPDEs. To the best of the authors' knowledge, such approach has not been employed for solving FPDEs.

The rest of the paper is organized as follows. In Section 2, we introduce some mathematical preliminaries of the fractional calculus theory and fractional-order Legendre functions. In Section 3, a basis of 2D-FLFs is defined and some properties are given. Section 4 is devoted to the operational matrices of fractional derivative and integration for 2D-FLFs. Some numerical examples are presented in Section 5. Finally, we conclude the paper with some remarks.

## 2. Preliminaries and Notations

2.1. Fractional Calculus Theory. Some necessary definitions and Lemma of the fractional calculus theory $[29,30]$ are listed here for our subsequent development.

Definition 1. A real function $h(t), t>0$, is said to be in the space $C_{\mu}, \mu \in R$, if there exists a real number $p>\mu$, such that $h(t)=t^{p} h_{1}(t)$, where $h_{1}(t) \in C(0, \infty)$, and it is said to be in the space $C_{\mu}^{n}$ if and only if $h^{(n)} \in C_{\mu}, n \in N$.

Definition 2. Riemann-Liouville fractional integral operator $\left(J^{\alpha}\right)$ of order $\alpha \geq 0$, of a function $f \in C_{\mu}, \mu \geq-1$ is defined as

$$
\begin{gather*}
J^{\alpha} f(t)=\frac{1}{\Gamma(\alpha)} \int_{0}^{t}(t-\tau)^{\alpha-1} f(\tau) d \tau, \quad t>0  \tag{1}\\
J^{0} f(t)=f(t)
\end{gather*}
$$

where $\Gamma(\alpha)$ is the well-known Gamma function. Some properties of the operator $J^{\alpha}$ can be found, for example, in $[29,30]$.

Definition 3. The fractional derivative of $f(x)$ in the Caputo sense is defined as

$$
\begin{align*}
& \left(D^{\alpha} f\right)(x) \\
& \quad= \begin{cases}\frac{1}{\Gamma(m-\alpha)} \\
\times \int_{0}^{x} \frac{f^{(m)}(\xi)}{(x-\xi)^{\alpha-m+1}} d \xi, & (\alpha>0, m-1<\alpha<m) \\
\frac{d^{m} f(x)}{d x^{m}}, & \alpha=m,\end{cases} \tag{2}
\end{align*}
$$

where $f: R \rightarrow R, x \rightarrow f(x)$ denotes a continuous (but not necessarily differentiable) function.

Lemma 4. Let $n-1<\alpha \leq n, n \in N, t>0, h \in C_{\mu}^{n}, \mu \geq-1$. Then

$$
\begin{equation*}
\left(J^{\alpha} D^{\alpha}\right) h(t)=h(t)-\sum_{k=0}^{n-1} h^{(k)}\left(0^{+}\right) \frac{t^{k}}{k!} \tag{3}
\end{equation*}
$$

2.2. Fractional-Order Legendre Functions. In this section, we introduce the fractional-order Legendre functions which were first proposed by Kazem et al. [28]. The normalized eigenfunctions problem for FLFs is

$$
\begin{array}{r}
\left(\left(x-x^{1+\alpha}\right) L_{i}^{\prime \alpha}(x)\right)^{\prime}+\alpha^{2} i(i+1) x^{\alpha-1} L_{i}^{\alpha}(x)=0  \tag{4}\\
x \in(0,1)
\end{array}
$$

which is a singular Sturm-Liouville problem. The fractionalorder Legendre polynomials, denoted by $\mathrm{FL}_{i}^{\alpha}(x)$, are defined on the interval $[0,1]$ and can be determined with the aid of following recurrence formulae:

$$
\begin{align*}
\mathrm{FL}_{0}^{\alpha}(x) & =1, \quad \mathrm{FL}_{1}^{\alpha}(x)=2 x^{\alpha}-1 \\
\mathrm{FL}_{i+1}^{\alpha}(x)= & \frac{(2 i+1)\left(2 x^{\alpha}-1\right)}{i+1} \mathrm{FL}_{i}^{\alpha}(x)  \tag{5}\\
& -\frac{i}{i+1} \mathrm{FL}_{i-1}^{\alpha}(x), \quad i=1,2, \ldots
\end{align*}
$$

and the analytic form of $\mathrm{FL}_{i}^{\alpha}(x)$ of degree $i$ is given by

$$
\begin{equation*}
\mathrm{FL}_{i}^{\alpha}(x)=\sum_{s=0}^{i} b_{s, i} i^{s \alpha}, \quad b_{s, i}=\frac{(-1)^{i+s}(i+s)!}{(i-s)!(s!)^{2}} \tag{6}
\end{equation*}
$$

where $\mathrm{FL}_{i}^{\alpha}(0)=(-1)^{i}$ and $\mathrm{FL}_{i}^{\alpha}(1)=1$. The orthogonality condition is

$$
\begin{equation*}
\int_{0}^{1} \mathrm{FL}_{n}^{\alpha}(x) \mathrm{FL}_{m}^{\alpha}(x) \omega(x) d x=\frac{1}{(2 n+1) \alpha} \delta_{n m} \tag{7}
\end{equation*}
$$

where $\omega(x)=x^{\alpha-1}$ is the weight function and $\delta$ is the Kronecker delta. For more details, please see [28].

## 3. 2D-FLFs

In this section, the definitions and theorems of 2D-FLFs are given by Liu's method described in [31].

### 3.1. Definitions and Properties of the 2D-FLFs

Definition 5. Let $\left\{\mathrm{FL}_{n}^{\alpha}(x)\right\}_{n=0}^{\infty}$ be the fractional Legendre polynomials on $[0,1]$; we call $\left\{\mathrm{FL}_{i}^{\alpha}(x) \mathrm{FL}_{j}^{\beta}(y)\right\}_{i, j=0}^{\infty}$ the two-dimensional fractional Legendre polynomials on $[0,1] \times[0,1]$.

Theorem 6. The basis $\left\{F L_{i}^{\alpha}(x) F L_{j}^{\beta}(y)\right\}_{i, j=0}^{\infty}$ is orthogonal on $[0,1] \times[0,1]$ with the weight function $\omega(x, y)=\omega(x) \omega(y)=$ $x^{\alpha-1} y^{\beta-1}$.

Proof. Let $i \neq m$ or $j \neq n$

$$
\begin{align*}
& \int_{0}^{1} \int_{0}^{1} \omega(x, y) \mathrm{FL}_{i}^{\alpha}(x) \mathrm{FL}_{j}^{\beta}(y) \mathrm{FL}_{m}^{\alpha}(x) \mathrm{FL}_{n}^{\beta}(y) d x d y \\
& \quad=\int_{0}^{1} \omega(x) \mathrm{FL}_{i}^{\alpha}(x) \mathrm{FL}_{m}^{\alpha}(x) d x  \tag{8}\\
& \quad \times \int_{0}^{1} \omega(y) \mathrm{FL}_{j}^{\beta}(y) \mathrm{FL}_{n}^{\beta}(y) d y=0 .
\end{align*}
$$

## Theorem 7. Consider

$$
\begin{aligned}
& \int_{0}^{1} \int_{0}^{1} \omega(x, y)\left[F L_{i}^{\alpha}(x) F L_{j}^{\beta}(y)\right]^{2} d x d y \\
& \quad=\frac{1}{(2 i+1) \alpha} \frac{1}{(2 j+1) \beta}, \\
& \int_{0}^{1} \int_{0}^{1} \omega(x, y)\left[F L_{i}^{\alpha}(x) F L_{j}^{\beta}(y)\right]^{2} d x d y \\
& \quad=\int_{0}^{1} \omega(x)\left[F L_{i}^{\alpha}(x)\right]^{2} d x \int_{0}^{1} \omega(y)\left[F L_{j}^{\beta}(y)\right]^{2} d y \\
& \quad=\frac{1}{(2 i+1) \alpha} \frac{1}{(2 j+1) \beta} .
\end{aligned}
$$

### 3.2. 2D-FLFs Expansion

Definition 8. A function of two independent variables $f(x, y)$ which is integrable in square $[0,1] \times[0,1]$ can be expanded as

$$
\begin{equation*}
f(x, y)=\sum_{i=0}^{\infty} \sum_{j=0}^{\infty} a_{i j} \mathrm{FL}_{i}^{\alpha}(x) \mathrm{FL}_{j}^{\beta}(y), \tag{10}
\end{equation*}
$$

where

$$
\begin{align*}
a_{i j}= & (2 i+1)(2 j+1) \alpha \beta \\
& \times \int_{0}^{1} \int_{0}^{1} f(x, y) \omega(x, y) \mathrm{FL}_{i}^{\alpha}(x) \mathrm{FL}_{j}^{\beta}(y) d x d y \tag{11}
\end{align*}
$$

Theorem 9. If the series $\sum_{i=0}^{\infty} \sum_{j=0}^{\infty} a_{i j} F L_{i}^{\alpha}(x) F L_{j}^{\beta}(y)$ converges uniformly to $f(x, y)$ on the square $[0,1] \times[0,1]$, then we have

$$
\begin{align*}
a_{i j}= & (2 i+1)(2 j+1) \alpha \beta \\
& \times \int_{0}^{1} \int_{0}^{1} f(x, y) \omega(x, y) F L_{i}^{\alpha}(x) F L_{j}^{\beta}(y) d x d y . \tag{12}
\end{align*}
$$

Proof. By multiplying $\omega(x, y) \mathrm{FL}_{n}^{\alpha}(x) \mathrm{FL}_{m}^{\beta}(y)$ on both sides of (10), where $n$ and $m$ are fixed and integrating termwise with regard to $x$ and $y$ on $[0,1] \times[0,1]$, then

$$
\begin{align*}
& \int_{0}^{1} \int_{0}^{1} f(x, y) \omega(x, y) \mathrm{FL}_{n}^{\alpha}(x) \mathrm{FL}_{m}^{\beta}(y) d x d y \\
& =\sum_{i=0}^{\infty} \sum_{j=0}^{\infty} a_{i j} \int_{0}^{1} \int_{0}^{1} \omega(x, y) \mathrm{FL}_{i}^{\alpha}(x) \mathrm{FL}_{j}^{\beta}(y) \\
& \quad \times \mathrm{FL}_{n}^{\alpha}(x) \mathrm{FL}_{m}^{\beta}(y) d x d y \\
& =\sum_{i=0}^{\infty} \sum_{j=0}^{\infty} a_{i j} \int_{0}^{1} \omega(x) \mathrm{FL}_{i}^{\alpha}(x) \mathrm{FL}_{n}^{\alpha}(x) d x \\
& \quad \times \int_{0}^{1} \omega(y) \mathrm{FL}_{j}^{\beta}(y) \mathrm{FL}_{m}^{\beta}(y) d y \\
& = \\
& =a_{n m} \int_{0}^{1} \omega(x)\left[\mathrm{FL}_{n}^{\alpha}(x)\right]^{2} d x \int_{0}^{1} \omega(y)\left[\mathrm{FL}_{m}^{\beta}(y)\right]^{2} d y  \tag{13}\\
& = \\
& a_{n m} \frac{1}{(2 n+1) \alpha} \frac{1}{(2 m+1) \beta} .
\end{align*}
$$

Finally one can get (11).
If the infinite series in (10) is truncated, then it can be written as

$$
\begin{equation*}
f(x, y) \approx \sum_{i=0}^{m} \sum_{j=0}^{m^{\prime}} a_{i j} \mathrm{FL}_{i}^{\alpha}(x) \mathrm{FL}_{j}^{\beta}(y)=C^{T} \Psi\left(x^{\alpha}, y^{\alpha}\right) \tag{14}
\end{equation*}
$$

where $C$ and $\Psi\left(x^{\alpha}, y^{\beta}\right)$ are given by

$$
\begin{align*}
C= & {\left[c_{0,0}, c_{0,1}, \ldots, c_{0, m^{\prime}-1}, c_{1,0}, c_{1,1}, \ldots,\right.} \\
& \left.c_{1, m^{\prime}-1}, \ldots, c_{m-1,0}, c_{m-1,1}, \ldots, c_{m-1, m^{\prime}-1}\right]^{T} \tag{15}
\end{align*}
$$

$$
\begin{align*}
\Psi\left(x^{\alpha}, y^{\beta}\right)=[ & \psi_{0,0}, \psi_{0,1}, \ldots, \psi_{0, m^{\prime}-1}, \psi_{1,0}, \psi_{1,1}, \ldots \\
& \left.\psi_{1, m^{\prime}-1}, \ldots, \psi_{m-1,0}, \psi_{m-1,1}, \ldots, \psi_{m-1, m^{\prime}-1}\right]^{T} \tag{16}
\end{align*}
$$

where $\psi_{i j}=\operatorname{FL}_{i}^{\alpha}(x) \operatorname{FL}_{j}^{\beta}(y), i=0,1, \ldots, m$, and $j=$ $0,1, \ldots, m^{\prime}$.

According to the definition of FLFs, one can find that fractional Legendre polynomials are identical to Legendre polynomials shifted to $[0,1]$ when using the transform $x^{\alpha} \rightarrow$ $x, y^{\beta} \rightarrow y$. Therefore, in a similar method described in [31], we can easily get the convergence and stability theorems of proposed method.

Lemma 10. If the function $f(x, y)$ is a continuous function on $[0,1] \times[0,1]$ and the series $\sum_{i=0}^{\infty} \sum_{j=0}^{\infty} a_{i j} F L_{i}^{\alpha}(x) F L_{j}^{\beta}(y)$ converges uniformly to $f(x, y)$, then $\sum_{i=0}^{\infty} \sum_{j=0}^{\infty} a_{i j} F L_{i}^{\alpha}(x) F L_{j}^{\beta}(y)$ is the 2D-FLFs expansion of $f(x, y)$.

Proof (by contradiction). Let

$$
\begin{align*}
& f(x, y)=\sum_{i=0}^{\infty} \sum_{j=0}^{\infty} b_{i j} \mathrm{FL}_{i}^{\alpha}(x) \mathrm{FL}_{j}^{\beta}(y), \\
& f(x, y) \sim \sum_{i=0}^{\infty} \sum_{j=0}^{\infty} a_{i j} \mathrm{FL}_{i}^{\alpha}(x) \mathrm{FL}_{j}^{\beta}(y) . \tag{17}
\end{align*}
$$

Then there is at least one coefficient such that $a_{n m} \neq b_{n m}$. However,

$$
\begin{align*}
b_{n m}= & (2 n+1)(2 m+1) \alpha \beta \\
& \times \int_{0}^{1} \int_{0}^{1} f(x, y) \omega(x, y) \mathrm{FL}_{n}^{\alpha}(x) \mathrm{FL}_{m}^{\beta}(y) d x d y  \tag{18}\\
= & a_{n m} .
\end{align*}
$$

Lemma 11. If two continuous functions defined on $[0,1] \times$ $[0,1]$ have the identical 2D-FLFs expansions, then these two functions are identical.

Proof. Suppose that $f(x, y)$ and $g(x, y)$ can be expanded by 2D-FLFs as follows:

$$
\begin{align*}
& f(x, y) \sim \sum_{i=0}^{\infty} \sum_{j=0}^{\infty} a_{i j} \mathrm{FL}_{i}^{\alpha}(x) \mathrm{FL}_{j}^{\beta}(y),  \tag{19}\\
& g(x, y) \sim \sum_{i=0}^{\infty} \sum_{j=0}^{\infty} a_{i j} \mathrm{FL}_{i}^{\alpha}(x) \mathrm{FL}_{j}^{\beta}(y) .
\end{align*}
$$

By subtracting the above two equations with each other, one has

$$
\begin{align*}
f(x, y)-g(x, y) & \sim \sum_{i=0}^{\infty} \sum_{j=0}^{\infty}\left(a_{i j}-a_{i j}\right) \mathrm{FL}_{i}^{\alpha}(x) \mathrm{FL}_{j}^{\beta}(y)  \tag{20}\\
& =0=\sum_{i=0}^{\infty} \sum_{j=0}^{\infty} 0 \times \mathrm{FL}_{i}^{\alpha}(x) \mathrm{FL}_{j}^{\beta}(y) .
\end{align*}
$$

Then Lemma 11 can be proved.
Theorem 12. If the 2D-FLFs expansion of a continuous function $f(x, y)$ converges uniformly, then the 2D-FLFs expansion converges to the function $f(x, y)$.

Proof. Theorem 12 can be proved by Theorems 7 and 9.
Theorem 13. If the sum of the absolute values of the 2D-FLFs coefficients of a continuous function $f(x, y)$ forms a convergent series, then the 2D-FLFs expansion is absolutely uniformly convergent, and converges to the function $f(x, y)$.

Proof. Consider

$$
\begin{align*}
\left|\sum_{i=0}^{\infty} \sum_{j=0}^{\infty} a_{i j} \mathrm{FL}_{i}^{\alpha}(x) \mathrm{FL}_{j}^{\beta}(y)\right| & \leq \sum_{i=0}^{\infty} \sum_{j=0}^{\infty}\left|a_{i j}\right|\left|\mathrm{FL}_{i}^{\alpha}(x)\right|\left|\mathrm{FL}_{j}^{\beta}(y)\right| \\
& \leq \sum_{i=0}^{\infty} \sum_{j=0}^{\infty}\left|a_{i j}\right| . \tag{21}
\end{align*}
$$

Then $\sum_{i=0}^{\infty} \sum_{j=0}^{\infty} a_{i j} \mathrm{FL}_{i}^{\alpha}(x) \mathrm{FL}_{j}^{\beta}(y)$ converges uniformly to the function $f(x, y)$.

Theorem 14. If a continuous function $f(x, y)$, defined on $[0$, $1] \times[0,1]$, has bounded mixed partial derivative $D_{x}^{2 \alpha} D_{y}^{2 \beta} f(x$, $y$ ), then the 2D-FLFs expansion of the function converges uniformly to the function.

Proof. Let $f(x, y)$ be a function defined on $[0,1] \times[0,1]$ such that

$$
\begin{equation*}
\left|D_{x}^{2 \alpha} D_{y}^{2 \beta} f(x, y)\right| \leq M \tag{22}
\end{equation*}
$$

where $M$ is a positive constant and

$$
\begin{align*}
a_{i j}= & (2 i+1)(2 j+1) \alpha \beta \\
& \times \int_{0}^{1} \int_{0}^{1} f(x, y) \omega(x, y) \mathrm{FL}_{i}^{\alpha}(x) \mathrm{FL}_{j}^{\beta}(y) d x d y \tag{23}
\end{align*}
$$

By employing the transform $X=2 x^{\alpha}-1$ and $Y=2 y^{\beta}-1$, one can obtain

$$
\begin{equation*}
a_{i j}=\frac{2 i+1}{2} \frac{2 j+1}{2} \int_{-1}^{1} \int_{-1}^{1} f(X, Y) p_{i}(X) p_{j}(Y) d X d Y \tag{24}
\end{equation*}
$$

Consequently, in a similar method described in [31], Theorem 14 can be proved.

## 4. Operational Matrices of 2D-FLFs

### 4.1. Integration Operational Matrices of 2D-FLFs

Lemma 15. The Riemann-Liouville fractional integration of order $\gamma>0$ of the 2D-FLFs $\psi_{i j}$ can be obtained in the form of

$$
\begin{equation*}
J_{x}^{\gamma}\left\{\psi_{i j}\left(x^{\alpha}, y^{\beta}\right)\right\}=F L_{j}^{\beta}(y) \sum_{s=0}^{i} b_{s i} \frac{\Gamma(1+s \alpha)}{\Gamma(1+s \alpha+\gamma)} x^{s \alpha+\gamma} \tag{25}
\end{equation*}
$$

Proof. Consider

$$
\begin{align*}
J_{x}^{\gamma}\left\{\psi_{i j}\left(x^{\alpha}, y^{\beta}\right)\right\} & =J_{x}^{\gamma}\left\{\mathrm{FL}_{i}^{\alpha}(x) \mathrm{FL}_{j}^{\beta}(y)\right\} \\
& =J_{x}^{\gamma}\left\{\mathrm{FL}_{i}^{\alpha}(x)\right\} \mathrm{FL}_{j}^{\beta}(y) \\
& =J_{x}^{\gamma}\left\{\sum_{s=0}^{i} b_{s i} x^{s \alpha}\right\} \mathrm{FL}_{j}^{\beta}(y)  \tag{26}\\
& =\mathrm{FL}_{j}^{\beta}(y) \sum_{s=0}^{i} b_{s i} \frac{\Gamma(1+s \alpha)}{\Gamma(1+s \alpha+\gamma)} x^{s \alpha+\gamma} .
\end{align*}
$$

Lemma 16. Let $\gamma>0$; then one has

$$
\begin{align*}
& \int_{0}^{1} \int_{0}^{1} J_{x}^{\gamma}\left\{\psi_{i j}\right\} \psi_{i^{\prime} j^{\prime}} \omega(x, y) d x d y \\
& \quad= \begin{cases}\sum_{s=0}^{i} \sum_{s^{\prime}=0}^{i^{\prime}} \frac{b_{s i} b_{s^{\prime} i^{\prime}}}{\left(s+s^{\prime}+1\right) \alpha+\gamma} \\
& \times \frac{\Gamma(1+s \alpha)}{\Gamma(1+s \alpha+\gamma)} \frac{1}{(2 j+1) \beta}, \\
0, & j=j^{\prime} \\
0, & j \neq j^{\prime}\end{cases} \tag{27}
\end{align*}
$$

Proof. Using previous Lemma 15 and (6), one can have

$$
\begin{align*}
& \int_{0}^{1} \int_{0}^{1} J_{x}^{\gamma}\left\{\psi_{i j}\left(x^{\alpha}, y^{\beta}\right)\right\} \psi_{i^{\prime} j^{\prime}}\left(x^{\alpha}, y^{\beta}\right) \omega(x, y) d x d y \\
& =\int_{0}^{1} \int_{0}^{1} \omega(x, y) \mathrm{FL}_{i^{\prime}}^{\alpha}\left(x^{\alpha}\right) \mathrm{FL}_{j^{\prime}}^{\beta}\left(y^{\beta}\right) \mathrm{FL}_{j}^{\beta}\left(y^{\beta}\right) \\
& \times \sum_{s=0}^{i} b_{s i} \frac{\Gamma(1+s \alpha)}{\Gamma(1+s \alpha+\gamma)} x^{s \alpha+\gamma} d x d y \\
& =\int_{0}^{1} \int_{0}^{1} \omega(y) \mathrm{FL}_{j}^{\beta}\left(y^{\beta}\right) \mathrm{FL}_{j^{\prime}}^{\beta}\left(y^{\beta}\right) \\
& \times \sum_{s=0}^{i} \sum_{s^{\prime}=0}^{i^{\prime}} b_{s i} b_{s^{\prime} i^{\prime}} \frac{\Gamma(1+s \alpha)}{\Gamma(1+s \alpha+\gamma)} \\
& \times x^{\left(s+s^{\prime}+1\right) \alpha+\gamma-1} d x d y \\
& =\int_{0}^{1} \omega(y) \mathrm{FL}_{j}^{\beta}\left(y^{\beta}\right) \mathrm{FL}_{j^{\prime}}^{\beta}\left(y^{\beta}\right) \\
& \times\left(\int_{0}^{1} \sum_{s=0}^{i} \sum_{s^{\prime}=0}^{i^{\prime}} b_{s i} b_{s^{\prime} i^{\prime}} \frac{\Gamma(1+s \alpha)}{\Gamma(1+s \alpha+\gamma)}\right.  \tag{28}\\
& \left.\times x^{\left(s+s^{\prime}+1\right) \alpha+\gamma-1} d x\right) d y \\
& =\sum_{s=0}^{i} \sum_{s^{\prime}=0}^{i^{\prime}} \frac{b_{s i} b_{s^{\prime} i^{\prime}}}{\left(s+s^{\prime}+1\right) \alpha+\gamma} \frac{\Gamma(1+s \alpha)}{\Gamma(1+s \alpha+\gamma)} \\
& \times \int_{0}^{1} \omega(y) \mathrm{FL}_{j}^{\beta}\left(y^{\beta}\right) \mathrm{FL}_{j^{\prime}}^{\beta}\left(y^{\beta}\right) d y \\
& = \begin{cases}\sum_{s=0}^{i} \sum_{s^{\prime}=0}^{i^{\prime}} \frac{b_{s i} b_{s^{\prime} i^{\prime}}}{\left(s+s^{\prime}+1\right) \alpha+\gamma} & \\
\quad \times \frac{\Gamma(1+s \alpha)}{\Gamma(1+s \alpha+\gamma)} \frac{1}{(2 j+1) \beta}, & j=j^{\prime} \\
0, & j \neq j^{\prime} .\end{cases}
\end{align*}
$$

Theorem 17. Let $\Psi\left(x^{\alpha}, y^{\beta}\right)$ be the 2D-FLFs vector defined in (16); then one has

$$
\begin{equation*}
J_{x}^{\gamma} \Psi\left(x^{\alpha}, y^{\beta}\right) \simeq \mathbf{P}_{x}^{\gamma} \Psi\left(x^{\alpha}, y^{\beta}\right) \tag{29}
\end{equation*}
$$

where $\mathbf{P}_{x}^{\gamma}$ is the $\mathrm{mm}^{\prime} \times \mathrm{mm}^{\prime}$ operational matrix of RiemannLiouville fractional integration of order $\gamma>0$, and has the form as follows:

$$
\mathbf{P}_{x}^{\gamma}=\left[\begin{array}{cccc}
E_{0,0} & E_{0,1} & \cdots & E_{0, m-1}  \tag{30}\\
E_{1,0} & E_{1,1} & \cdots & E_{1, m-1} \\
\vdots & \vdots & \ddots & \vdots \\
E_{m-1,0} & E_{m-1,1} & \cdots & E_{m-1, m-1}
\end{array}\right]
$$

in which $E_{i, i^{\prime}}$ is $m^{\prime} \times m^{\prime}$ matrix and the elements are defined as follows:

$$
\begin{array}{r}
E_{i, i^{\prime}}=I \sum_{s=0}^{i} \sum_{s^{\prime}=0}^{i^{\prime}} \frac{b_{s i} b_{s^{\prime} i^{\prime}}\left(2 i^{\prime}+1\right) \alpha}{\left(s+s^{\prime}+1\right) \alpha+\gamma} \frac{\Gamma(1+s \alpha)}{\Gamma(1+s \alpha+\gamma)}  \tag{31}\\
i, i^{\prime}=0,1, \ldots, m-1
\end{array}
$$

and $I$ is $m^{\prime} \times m^{\prime}$ identity matrix.
Proof. Using (29) and orthogonality property of FLFs, one can get

$$
\begin{equation*}
\mathbf{P}_{x}^{\gamma}=\left\langle J_{x}^{\gamma} \Psi\left(x^{\alpha}, y^{\beta}\right), \Psi^{T}\left(x^{\alpha}, y^{\beta}\right)\right\rangle H^{-1} \tag{32}
\end{equation*}
$$

where $\left\langle J_{x}^{\gamma} \Psi\left(x^{\alpha}, y^{\beta}\right), \Psi^{T}\left(x^{\alpha}, y^{\beta}\right)\right\rangle$ and $H^{-1}$ are two $m^{\prime} \times m^{\prime}$ matrices defined as

$$
\begin{align*}
& \left\langle J_{x}^{\gamma} \Psi\left(x^{\alpha}, y^{\beta}\right), \Psi^{T}\left(x^{\alpha}, y^{\beta}\right)\right\rangle \\
& =\left\{\int_{0}^{1} \int_{0}^{1} J_{x}^{\gamma}\left\{\Psi_{k}\left(x^{\alpha}, y^{\beta}\right)\right\}\right. \\
& \left.\quad \times \Psi_{k^{\prime}}\left(x^{\alpha}, y^{\beta}\right) \omega(x, y) d x d y\right\}_{k, k^{\prime}}^{m m^{\prime}} \\
& =\left\{\sum_{s=0}^{i} \sum_{s^{\prime}=0}^{i^{\prime}} b_{s i} b_{s^{\prime} i^{\prime}} \frac{\Gamma(1+s \alpha)}{\Gamma(1+s \alpha+\gamma)}\right. \\
& \left.\quad \times \frac{1}{\left(s+s^{\prime}+1\right) \alpha+\gamma} \frac{1}{(2 j+1) \beta}\right\}_{i, i^{\prime} ; j=j^{\prime}}^{m ; m^{\prime}} \\
& H^{-1}=\left\{\left(2 i^{\prime}+1\right)(2 j+1) \alpha \beta\right\}_{i, i^{\prime} ; j=j^{\prime}}^{m ; m^{\prime}} \tag{33}
\end{align*}
$$

Now by substituting above equations in (32), Theorem 12 can be proved.

In a similar way as previous, one can obtain the operational matrix of Riemann-Liouville fractional integration with respect to variable $y$.

Theorem 18. Let $\Psi\left(x^{\alpha}, y^{\beta}\right)$ be the $2 D$-FLFs vector defined in (16); one has

$$
\begin{equation*}
J_{y}^{\gamma} \Psi\left(x^{\alpha}, y^{\beta}\right) \simeq \mathbf{P}_{y}^{\gamma} \Psi\left(x^{\alpha}, y^{\beta}\right) \tag{34}
\end{equation*}
$$

where $\mathbf{P}_{y}^{\gamma}$ is the $\mathrm{mm}^{\prime} \times \mathrm{mm}^{\prime}$ operational matrix of RiemannLiouville fractional integration of order $\gamma>0$, and has the form as follows:

$$
\mathbf{P}_{y}^{\gamma}=\left[\begin{array}{cccc}
E & O & \cdots & O  \tag{35}\\
O & E & \cdots & O \\
\vdots & \vdots & \ddots & \vdots \\
O & O & \cdots & E
\end{array}\right]
$$

in which $E$ is $m^{\prime} \times m^{\prime}$ matrix and the elements are defined as follows:

$$
\begin{array}{r}
E_{j, j^{\prime}}=\sum_{r=0}^{j} \sum_{r^{\prime}=0}^{j^{\prime}} \frac{b_{r r} b_{r}^{\prime} j^{\prime}}{\left(r+r^{\prime}+1\right) \beta+\gamma}+\frac{\left.j^{\prime}+1\right) \beta}{(r(1+r \beta)},  \tag{36}\\
j, j^{\prime}=0,1, \ldots, m^{\prime}-1 .
\end{array}
$$

4.2. Derivative Operational Matrices of 2D-FLFs

Lemma 19. The FLFs Caputo fractional derivative of $\gamma>0$ can be obtained in the form of

$$
\begin{equation*}
D_{x}^{\gamma}\left\{\psi_{i j}\left(x^{\alpha}, y^{\beta}\right)\right\}=F L_{j}^{\beta}\left(y^{\beta}\right) \sum_{s=0}^{i} b_{s i}^{\prime} \frac{\Gamma(1+s \alpha)}{\Gamma(1+s \alpha-\gamma)} x^{s \alpha-\gamma} \tag{37}
\end{equation*}
$$

where $b_{s, i}^{\prime}=0$ when $s \alpha \in N_{0}$ and $s \alpha<\gamma$ in other case $b_{s, i}^{\prime}=b_{s, i}$.
Proof. Consider

$$
\begin{align*}
D_{x}^{\gamma}\left\{\psi_{i j}\left(x^{\alpha}, y^{\beta}\right)\right\} & =D_{x}^{\gamma}\left\{\operatorname{FL}_{i}^{\alpha}\left(x^{\alpha}\right) \mathrm{FL}_{j}^{\beta}\left(y^{\beta}\right)\right\} \\
& =\mathrm{FL}_{j}^{\beta}\left(y^{\beta}\right) D_{x}^{\gamma}\left\{\mathrm{FL}_{i}^{\alpha}\left(x^{\alpha}\right)\right\} \\
& =D_{x}^{\gamma}\left\{\sum_{s=0}^{i} b_{s i} x^{s \alpha}\right\} \mathrm{FL}_{j}^{\beta}\left(y^{\beta}\right) \\
& =\mathrm{FL}_{j}^{\beta}\left(y^{\beta}\right) \sum_{s=0}^{i} b_{s i}^{\prime} \frac{\Gamma(1+s \alpha)}{\Gamma(1+s \alpha-\gamma)} x^{s \alpha-\gamma} \tag{38}
\end{align*}
$$

Lemma 20. Let $\gamma>0, \alpha \notin N$; then one has

$$
\begin{aligned}
& \int_{0}^{1} \int_{0}^{1} D_{x}^{\gamma}\left\{\psi_{i j}\right\} \psi_{i^{\prime} j^{\prime}} \omega(x, y) d x d y \\
& \quad= \begin{cases}\sum_{s=0_{s^{\prime}=0}^{i}}^{\sum^{i^{\prime}} \frac{b_{s i} b_{s^{\prime} i^{\prime}}}{\left(s+s^{\prime}+1\right) \alpha-\gamma}} \\
\quad \times \frac{\Gamma(1+s \alpha)}{\Gamma(1+s \alpha-\gamma)} \frac{1}{(2 j+1) \beta}, & j=j^{\prime} \\
0, & j \neq j^{\prime}\end{cases}
\end{aligned}
$$

Proof. Using previous Lemma 19 and (6), one can have

$$
\begin{align*}
& \int_{0}^{1} \int_{0}^{1} D_{x}^{\gamma}\left\{\psi_{i j}\left(x^{\alpha}, y^{\beta}\right)\right\} \psi_{i^{\prime} j^{\prime}}\left(x^{\alpha}, y^{\beta}\right) \omega(x, y) d x d y \\
& =\int_{0}^{1} \int_{0}^{1} \omega(x, y) \mathrm{FL}_{i^{\prime}}^{\alpha}\left(x^{\alpha}\right) \mathrm{FL}_{j^{\prime}}^{\beta}\left(y^{\beta}\right) \mathrm{FL}_{j}^{\beta}\left(y^{\beta}\right) \\
& \times \sum_{s=0}^{i} b_{s i}^{\prime} \frac{\Gamma(1+s \alpha)}{\Gamma(1+s \alpha-\gamma)} x^{s \alpha-\gamma} d x d y \\
& =\int_{0}^{1} \int_{0}^{1} \omega(y) \mathrm{FL}_{j}^{\beta}\left(y^{\beta}\right) \mathrm{FL}_{j^{\prime}}^{\beta}\left(y^{\beta}\right) \\
& \times \sum_{s=0}^{i} \sum_{s^{\prime}=0}^{i^{\prime}} b_{s i}^{\prime} b_{s^{\prime} i^{\prime}} \frac{\Gamma(1+s \alpha)}{\Gamma(1+s \alpha-\gamma)} \\
& \times x^{\left(s+s^{\prime}+1\right) \alpha-\gamma-1} d x d y \\
& =\int_{0}^{1} \omega(y) \mathrm{FL}_{j}^{\beta}\left(y^{\beta}\right) \mathrm{FL}_{j^{\prime}}^{\beta}\left(y^{\beta}\right) \\
& \times\left(\int_{0}^{1} \sum_{s=0}^{i} \sum_{s^{\prime}=0}^{i^{\prime}} b_{s i}^{\prime} b_{s^{\prime} i^{\prime}} \frac{\Gamma(1+s \alpha)}{\Gamma(1+s \alpha-\gamma)}\right.  \tag{40}\\
& \left.\times x^{\left(s+s^{\prime}+1\right) \alpha-\gamma-1} d x\right) d y \\
& =\sum_{s=0}^{i} \sum_{s^{\prime}=0}^{i^{\prime}} \frac{b_{s i}^{\prime} b_{s^{\prime} i^{\prime}}}{\left(s+s^{\prime}+1\right) \alpha-\gamma} \frac{\Gamma(1+s \alpha)}{\Gamma(1+s \alpha-\gamma)} \\
& \times \int_{0}^{1} \omega(y) \mathrm{FL}_{j}^{\beta}\left(y^{\beta}\right) \mathrm{FL}_{j^{\prime}}^{\beta}\left(y^{\beta}\right) d y \\
& = \begin{cases}\sum_{s=0}^{i} \sum_{s^{\prime}=0}^{i^{\prime}} \frac{b_{s i}^{\prime} b_{s^{\prime} i^{\prime}}}{\left(s+s^{\prime}+1\right) \alpha-\gamma} & \\
\times \frac{\Gamma(1+s \alpha)}{\Gamma(1+s \alpha-\gamma)} \frac{1}{(2 j+1) \beta}, & j=j^{\prime} \\
0, & j \neq j^{\prime} .\end{cases}
\end{align*}
$$

Theorem 21. Let $\Psi\left(x^{\alpha}, y^{\beta}\right)$ be the 2D-FLFs vector defined in (16); one has

$$
\begin{equation*}
D_{x}^{\gamma} \Psi\left(x^{\alpha}, y^{\beta}\right) \simeq \mathbf{D}_{x}^{\gamma} \Psi\left(x^{\alpha}, y^{\beta}\right) \tag{41}
\end{equation*}
$$

where $\mathbf{D}_{x}^{\gamma}$ is the $\mathrm{mm}^{\prime} \times \mathrm{mm}^{\prime}$ operational matrix of Caputo fractional derivative of order $\gamma>0$, and has the form as follows:

$$
\mathbf{D}_{x}^{\gamma}=\left[\begin{array}{cccc}
O & O & \cdots & O  \tag{42}\\
F_{1,0} & O & \cdots & O \\
\vdots & \vdots & \ddots & \vdots \\
F_{m-1,0} & F_{m-1,1} & \cdots & O
\end{array}\right]
$$

in which $F_{i, i^{\prime}}$ is $m^{\prime} \times m^{\prime}$ matrix and the elements are defined as follows:

$$
\begin{array}{r}
F_{i, i^{\prime}}=I \sum_{s=0}^{i} \sum_{s^{\prime}=0}^{i^{\prime}} \frac{b_{s i}^{\prime} b_{s^{\prime} i^{\prime}}\left(2 i^{\prime}+1\right) \alpha}{\left(s+s^{\prime}+1\right) \alpha-\gamma} \frac{\Gamma(1+s \alpha)}{\Gamma(1+s \alpha-\gamma)},  \tag{43}\\
i, i^{\prime}=0,1, \ldots, m-1
\end{array}
$$

and $I$ is a $m^{\prime} \times m^{\prime}$ identity matrix.
Proof. Using (41) and the orthogonality property of FLFs, one can have

$$
\begin{equation*}
\mathbf{D}_{x}^{\gamma}=\left\langle D_{x}^{\gamma} \Psi\left(x^{\alpha}, y^{\beta}\right), \Psi^{T}\left(x^{\alpha}, y^{\beta}\right)\right\rangle H^{-1} \tag{44}
\end{equation*}
$$

where $\left\langle D_{x}^{\gamma} \Psi\left(x^{\alpha}, y^{\beta}\right), \Psi^{T}\left(x^{\alpha}, y^{\beta}\right)\right\rangle$ and $H^{-1}$ are two $\mathrm{mm}^{\prime} \times$ $\mathrm{mm}^{\prime}$ matrices defined as

$$
\begin{align*}
& \left\langle D_{x}^{\gamma} \Psi\left(x^{\alpha}, y^{\beta}\right), \Psi^{T}\left(x^{\alpha}, y^{\beta}\right)\right\rangle \\
& =\left\{\int_{0}^{1} \int_{0}^{1} D_{x}^{\gamma}\left\{\Psi_{k}\left(x^{\alpha}, y^{\beta}\right)\right\}\right. \\
& \left.\quad \times \Psi_{k^{\prime}}\left(x^{\alpha}, y^{\beta}\right) \omega(x, y) d x d y\right\}_{k, k^{\prime}}^{m m^{\prime}} \\
& =\left\{\sum_{s=0}^{i} \sum_{s^{\prime}=0}^{i^{\prime}} \frac{b_{s i}^{\prime} b_{s^{\prime} i^{\prime}}}{\left(s+s^{\prime}+1\right) \alpha-\gamma}\right.  \tag{45}\\
& \left.\quad \times \frac{\Gamma(1+s \alpha)}{\Gamma(1+s \alpha-\gamma)} \frac{1}{(2 j+1) \beta}\right\}_{i, i^{\prime} ; j=j^{\prime}}^{m ; m^{\prime}} \\
& \\
& H^{-1}=\left\{\left(2 i^{\prime}+1\right)(2 j+1) \alpha \beta\right\}_{i, i^{\prime} ; j=j^{\prime}}^{m ; m^{\prime}}
\end{align*}
$$

Now by substituting above equations in (44), Theorem 21 can be proved.

In a similar way as above, one can get Caputo fractional derivative of order $\gamma>0$ with respect to variable $y$.

Theorem 22. Let $\Psi\left(x^{\alpha}, y^{\beta}\right)$ be the 2D-FLFs vector defined in (16); one can have

$$
\begin{equation*}
D_{y}^{\gamma} \Psi\left(x^{\alpha}, y^{\beta}\right) \simeq \mathbf{D}_{y}^{\gamma} \Psi\left(x^{\alpha}, y^{\beta}\right) \tag{46}
\end{equation*}
$$

where $\mathbf{D}_{y}^{\gamma}$ is the $\mathrm{mm}^{\prime} \times \mathrm{mm}^{\prime}$ operational matrix of Caputo fractional derivative of order $\gamma>0$, and has the form as follows:

$$
\mathbf{D}_{y}^{y}=\left[\begin{array}{cccc}
F & O & \cdots & O  \tag{47}\\
O & F & \cdots & O \\
\vdots & \vdots & \ddots & \vdots \\
O & O & \cdots & F
\end{array}\right]
$$

in which $F$ is $m^{\prime} \times m^{\prime}$ matrix and the elements are defined as follows:

$$
\begin{array}{r}
F_{j, j^{\prime}}=\sum_{r=0}^{j} \sum_{r^{\prime}=0}^{j^{\prime}} \frac{b_{r j}^{\prime} b_{r r^{\prime} j^{\prime}}\left(2 j^{\prime}+1\right) \beta}{\left(r+r^{\prime}+1\right) \beta+\gamma} \frac{\Gamma(1+r \beta)}{\Gamma(1+r \beta+\gamma)}  \tag{48}\\
j, j^{\prime}=0,1, \ldots, m^{\prime}-1
\end{array}
$$

## 5. Applications and Results

Consider the following FPDEs:

$$
\begin{align*}
& D_{x}^{\alpha} u(x, t)+D_{t}^{\beta} u(x, t) \\
& \quad+N[u(x, t)]+L[u(x, t)]=g(x, t), \quad \alpha, \beta \in(0,1] \tag{49}
\end{align*}
$$

where $L$ and $N$ are linear operator and nonlinear operator; respectively. $D^{\alpha}$ and $D^{\beta}$ are the Caputo fractional derivatives of order $\alpha$ and $\beta$, respectively; $g$ is a known analytic function.

By employing operator $J_{t}^{\beta}$ on both sides of (49) and then using the Lemma 4, one can have

$$
\begin{gather*}
u(x, t)+J_{t}^{\beta}\left\{D_{x}^{\alpha} u(x, t)+N u(x, t)+L u(x, t)\right\} \\
-\sum_{k=0}^{m-1} u^{(k)}(x, 0) \frac{x^{k}}{k!}-J_{t}^{\beta} g(x, t)=0 . \tag{50}
\end{gather*}
$$

We first express unknown function $u(x, t)$ and derivative term $D_{x}^{\alpha} u(x, t)$ as

$$
\begin{equation*}
u(x, t)=C^{T} \Psi\left(x^{\alpha}, t^{\beta}\right), \quad D_{x}^{\alpha} u(x, t)=C^{T} \mathbf{D}_{x}^{\alpha} \Psi\left(x^{\alpha}, t^{\beta}\right) \tag{51}
\end{equation*}
$$

Now for the nonlinear part, by employing the nonlinear term approximation method described in [32] and then by using transform $x \rightarrow x^{\alpha}, t \rightarrow t^{\beta}$, one can get the 2D-FLFs expansion of nonlinear term as

$$
\begin{equation*}
N u(x, t)=N^{T} \Psi\left(x^{\alpha}, t^{\beta}\right) \tag{52}
\end{equation*}
$$

where $N^{T}$ is coefficient matrix of nonlinear term which must be computed and its order is $\mathrm{mm}^{\prime} \times \mathrm{mm}^{\prime}$.

For the linear part, we have

$$
\begin{equation*}
L u(x, t)=L^{T} \Psi\left(x^{\alpha}, t^{\beta}\right) \tag{53}
\end{equation*}
$$

where $L$ is a matrix of order $\mathrm{mm}^{\prime} \times \mathrm{mm}^{\prime}$.
After substituting (51)-(53) into (50), one can obtain

$$
\begin{equation*}
C^{T}+\left(C^{T} \mathbf{D}_{x}^{\alpha}+N^{T}+L^{T}\right) \mathbf{P}_{y}^{\beta}-C_{\text {guess }}^{T}=0 \tag{54}
\end{equation*}
$$

According to the Wu's [33] technology for determining the initial iteration value, the initial iteration value is chosen as $u_{\text {guess }}=\sum_{k=0}^{m-1} u^{(k)}(x, 0)\left(x^{k} / k!\right)+J_{t}^{\beta}\{g(x, t)\}=C_{\text {guess }}^{T} \Psi\left(x^{\alpha}, t^{\beta}\right)$. The coefficient matrix $C^{T}$ can be computed by using the MATLAB function fsolve( ) or the method described in [34].


Figure 1: Numerical results for Example 23 when $\beta=0.25,0.50$.

Now, the present method is applied to solve the linear and nonlinear FPDEs, and their results are compared with the solution of other methods. The accuracy of our approach is estimated by the following error functions:

$$
\begin{gather*}
e_{j}=\left(u_{\text {exact }}\right)_{j}-\left(u_{\text {approx }}\right)_{j}, \quad e=u_{\text {exact }}-u_{\text {approx }} \\
\|e\|_{L_{\infty}}=\max _{1 \leq j \leq N}\left|e_{j}\right|, \quad\|e\|_{L_{2}}=\sqrt{\sum_{j=1}^{N}\left|\left(e_{j}\right)^{2}\right|}  \tag{55}\\
\|e\|_{\mathrm{RMS}}=\sqrt{\frac{1}{N} \sum_{j=1}^{N}\left|\left(e_{j}\right)^{2}\right|}
\end{gather*}
$$

Example 23. Consider the one-dimensional linear inhomogeneous fractional Burger's equation [35]:

$$
\begin{align*}
& \frac{\partial^{\beta} u(x, t)}{\partial t^{\beta}}+\frac{\partial u(x, t)}{\partial x}-\frac{\partial^{2} u(x, t)}{\partial x^{2}} \\
& \quad=\frac{2 t^{2-\beta}}{\Gamma(3-\beta)}+2 x-2, \quad 0<\beta \leq 1 \tag{56}
\end{align*}
$$

with the initial condition $u(x, 0)=x^{2}$ and the exact solution being $u(x, t)=x^{2}+t^{2}$.

By employing 2D-FLFs method, one can get

$$
\begin{equation*}
C^{T}\left[I+\left(\mathbf{D}_{x}^{\alpha}-\left(\mathbf{D}_{x}^{\alpha}\right)^{2}\right) \mathbf{P}_{t}^{\beta}\right]=C_{\text {guess }}^{T}, \tag{57}
\end{equation*}
$$

where $\alpha=1$. Then we can get $C^{T}=C_{\text {guess }}^{T} \operatorname{inv}\left(I+\left(\mathbf{D}_{x}^{\alpha}-\right.\right.$ $\left.\left(\mathbf{D}_{x}^{\alpha}\right)^{2}\right) \mathbf{P}_{t}^{\beta}$.

Figures 1(a) and 1(b) show the numerical results for $\beta=$ 0.25 with $m=3, m^{\prime}=9$ and $\beta=0.5$ with $m=3, m^{\prime}=5$, respectively. It should be found that the accuracy of 2D-FLFs method is very high while only a small number of 2D-FLFs are needed.

Example 24. Consider nonlinear fractional Klein-Gordon equation [36, 37]:

$$
\begin{array}{r}
D_{t}^{\beta} u(x, t)-D_{x}^{\alpha} u(x)+u^{3}(x)=g(x, t)  \tag{58}\\
x \geq 0, \quad t>0, \quad \alpha, \beta \in(1,2]
\end{array}
$$

subject to the initial conditions

$$
\begin{equation*}
u(x, 0)=0, \quad u_{t}(x, 0)=0 \tag{59}
\end{equation*}
$$

and $g(x, t)=\Gamma(\beta+1) x^{\alpha}-\Gamma(\alpha+1) t^{\beta}+x^{3 \alpha} t^{3 \beta}$. The exact solution of (58) is $u(x, t)=x^{\alpha} t^{\beta}$.


Figure 2: Numerical results of Example 24 for different values of $\alpha$ and $\beta$.

By employing 2D-FLFs method with $m=3$ and $m^{\prime}=3$, one can have

$$
\begin{equation*}
C^{T}+\left(-C^{T} \mathbf{D}_{x}^{\alpha}+N^{T}\right) \mathbf{P}_{t}^{\beta}-C_{\text {guess }}^{T}=0 \tag{60}
\end{equation*}
$$

The numerical results of Example 24 for different values of $\alpha$ and $\beta$ are shown in Figure 2. In addition, $L_{2}$ and $L_{\infty}$ errors are presented in Table 1. From Table 1, one can conclude that the solutions of 2D-FLFs method are in good agreement with the exact results. Compared with homotopy analysis method (HAM) [36] and homotopy perturbation method (HPM) [37], 2D-FLFs method can get high accuracy solution while only need a few terms of 2D-FLFs.

Example 25. Consider the nonlinear time-fractional advection partial differential equation [37-39]

$$
\begin{gather*}
D_{t}^{\beta} u(x, t)+u(x, t) u_{x}(x, t)=x+x t^{2}  \tag{61}\\
t>0, \quad x \in R, \quad 0<\beta \leq 1
\end{gather*}
$$

subject to the initial condition

$$
\begin{equation*}
u(x, 0)=0 \tag{62}
\end{equation*}
$$

Figure 3 gives the approximation solutions of (61) for $\beta=0.50$ with $m=4, m^{\prime}=5$ and $\beta=0.75$ with $m=4$,
$m^{\prime}=9$. Moreover, Table 2 shows the approximate solutions for (61) obtained for different values of $\beta$ using the fractional variational iteration method (FVIM) [39] and 2D-FLFs method. The values of $\beta=1$ are the only case for which we know the exact solution $u(x, t)=x t$. It should be noted that only the fourth-order term of the FVIM was used in evaluating the approximate solutions for Table 2. From Table 2, it clearly appears that 2D-FLFs method is more accurate than FVIM and the obtained results are in good agreement with exact solution.

Example 26. We finally consider the linear time-fractional wave equation:

$$
\begin{equation*}
\frac{\partial^{2 \beta} u}{\partial t^{2 \beta}}=\frac{1}{2} x^{2} \frac{\partial^{2} u}{\partial x^{2}}, \quad t>0, x \in R, 0.5<\beta \leq 1 \tag{63}
\end{equation*}
$$

subject to the initial conditions

$$
\begin{equation*}
u(x, 0)=x, \quad \frac{\partial u(x, 0)}{\partial t}=x^{2} \tag{64}
\end{equation*}
$$

Table 3 gives a comparison of the approximate solutions at different values of $\beta$ using the FVIM [39] and 2D-FLFs method. Figure 4 shows the numerical solutions of 2D-FLFs method for (63) at different values of $\beta$ with $m=3, m^{\prime}=9$. The values of $\beta=1$ are the only case for which we know


Figure 3: Numerical results of Example 25 for different value of $\beta$.

(a)

(c)

(b)

$$
\beta=1.000
$$


(d)

Figure 4: Numerical results of Example 26 for different value of $\beta$.

Table 1: Errors of Example 24 for different values of $\alpha$ and $\beta$ with $M=M^{\prime}=4$.

| Error | $\alpha=\beta=1.25$ | $\alpha=\beta=1.50$ | $\alpha=\beta=1.75$ | $\alpha=\beta=2.00$ |
| :--- | :--- | :--- | :--- | :--- |
| $L_{2}$ | $5.6437 e-015$ | $1.2075 e-015$ | $3.4584 e-015$ | $8.9917 e-016$ |
| $L_{\infty}$ | $4.4409 e-016$ | $1.1102 e-016$ | $3.3307 e-016$ | $1.1102 e-016$ |

Table 2: Numerical values when $\beta=0.50,0.75$, and 1.0 for (61).

| $t$ | $x$ | $\beta=0.50$ |  | $\beta=0.75$ |  | $\beta=1.00$ |  |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| $t$ |  | FVIM | 2D-FLFs | FVIM | 2D-FLFs | FVIM | 2D-FLFs | Exact |
| 0.25 | 0.25 | 0.12422501 | 0.12225461 | 0.09230374 | 0.09224583 | 0.06250058 | 0.062500 | 0.062500 |
|  | 0.50 | 0.24845002 | 0.24450922 | 0.18460748 | 0.18449165 | 0.12500117 | 0.125000 | 0.125000 |
|  | 0.75 | 0.37267504 | 0.36676383 | 0.27691122 | 0.27673748 | 0.18750175 | 0.187500 | 0.187500 |
|  | 1.00 | 0.49690005 | 0.48901844 | 0.36921496 | 0.36898331 | 0.25000234 | 0.250000 | 0.250000 |
| 0.50 | 0.25 | 0.18377520 | 0.16584130 | 0.15148283 | 0.14985508 | 0.12507592 | 0.125000 | 0.125000 |
|  | 0.50 | 0.36755040 | 0.33168259 | 0.30296566 | 0.29971016 | 0.25015184 | 0.250000 | 0.250000 |
|  | 0.75 | 0.55132559 | 0.49752389 | 0.45444848 | 0.44956524 | 0.37522776 | 0.375000 | 0.375000 |
|  | 1.00 | 0.73510079 | 0.66336518 | 0.60593131 | 0.59942032 | 0.50030368 | 0.500000 | 0.500000 |
| 0.75 | 0.25 | 0.27227270 | 0.20678964 | 0.21407798 | 0.20119503 | 0.18881843 | 0.187500 | 0.187500 |
|  | 0.50 | 0.54454540 | 0.41357929 | 0.42815596 | 0.40239005 | 0.37763687 | 0.375000 | 0.375000 |
|  | 0.75 | 0.81681810 | 0.62036893 | 0.64223394 | 0.60358508 | 0.56645530 | 0.562500 | 0.562500 |
|  | 1.00 | 1.08909080 | 0.82715857 | 0.85631192 | 0.80478011 | 0.75527373 | 0.750000 | 0.750000 |

Table 3: Numerical values when $\beta=0.750,0.875$, and 1.000 for (63).

| $t$ | $x$ | $\beta=0.750$ |  | $\beta=0.875$ |  | $\beta=1.000$ |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  |  | FVIM | 2D-FLFs | FVIM | 2D-FLFs | FVIM | Exact |
| 0.25 | 0.25 | 0.26622298 | 0.26622021 | 0.26593959 | 0.26594005 | 0.26578827 | 0.26578827 |
|  | 0.50 | 0.56489190 | 0.56488083 | 0.56375836 | 0.56376020 | 0.56315308 | 0.56315308 |
|  | 0.75 | 0.89600678 | 0.89598187 | 0.89345630 | 0.89346046 | 0.89209443 | 0.89209443 |
|  | 1.00 | 1.25956762 | 1.25952332 | 1.25503343 | 1.25504082 | 1.25261232 | 1.25261232 |
| 0.50 | 0.25 | 0.28474208 | 0.28474415 | 0.28340402 | 0.28340659 | 0.28256846 | 0.28256846 |
|  | 0.50 | 0.63896831 | 0.63897662 | 0.63361610 | 0.63362636 | 0.63027383 | 0.63027383 |
|  | 0.75 | 1.06267869 | 1.06269739 | 1.05063622 | 1.05065931 | 1.04311611 | 1.04311611 |
|  | 1.00 | 1.55587323 | 1.55590647 | 1.53446439 | 1.53450544 | 1.52109530 | 1.52109531 |
| 0.75 | 0.25 | 0.30690489 | 0.30690747 | 0.30361709 | 0.30361656 | 0.30139478 | 0.30139480 |
|  | 0.50 | 0.72761955 | 0.72762986 | 0.71446834 | 0.71446625 | 0.70557913 | 0.70557918 |
|  | 0.75 | 1.26214400 | 1.26216719 | 1.23255378 | 1.23254905 | 1.21255304 | 1.21255316 |
|  | 1.00 | 1.91047821 | 1.91051944 | 1.85787338 | 1.85786498 | 1.82231652 | 1.82231673 |

the exact solution $u(x, t)=x+x^{2} \sinh (t)$. As previous, only the fourth-order term of the FVIM was used in evaluating the numerical solutions for Table 3. In the case of $\beta=1$, it can be found that absolute error of 2D-FLFs is not bigger than $1.0 e-10$ which is very small compared with that obtained by FVIM.

## 6. Conclusion

We define a basis of 2D-FLFs and derived its operational matrices of fractional derivative and integration, which are used to approximate the numerical solution of FPDEs. Compared with other numerical methods, 2D-FLFs method can accurately represent properties of fractional calculus. Moreover, only a small number of 2D-FLFs are needed to obtain a satisfactory result. The obtained results demonstrate the validity and applicability of proposed method for solving the FPFEs.

## Acknowledgments

This work is supported by the National Natural Science Foundation of China (Grant no. 11272352). The authors are grateful to the anonymous referees for their comments which substantially improved the quality of this paper.

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## Research Article

# Persistence Property and Estimate on Momentum Support for the Integrable Degasperis-Procesi Equation 

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Received 24 April 2013; Accepted 4 October 2013
Academic Editor: T. Raja Sekhar
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It is shown that a strong solution of the Degasperis-Procesi equation possesses persistence property in the sense that the solution with algebraically decaying initial data and its spatial derivative must retain this property. Moreover, we give estimates of measure for the momentum support.

## 1. Introduction

Recently, Degasperis and Procesi [1] consider the following family of third order dispersive conservation laws:

$$
\begin{equation*}
u_{t}+c_{0} u_{x}+\gamma u_{x x x}-\alpha^{2} u_{x x t}=\left(c_{1} u^{2}+c_{2} u_{x}^{2}+c_{3} u u_{x x}\right)_{x} \tag{1}
\end{equation*}
$$

where $\alpha, \gamma, c_{0}, c_{1}, c_{2}$, and $c_{3}$ are real constants. Within this family, only three equations that satisfy asymptotic integrability condition up to third order are singled out, namely, the KdV equation

$$
\begin{equation*}
u_{t}+u_{x}+u u_{x}+u_{x x x}=0 \tag{2}
\end{equation*}
$$

the Camassa-Holm equation

$$
\begin{equation*}
u_{t}-u_{x x t}+3 u u_{x}=2 u_{x} u_{x x}+u u_{x x x} \tag{3}
\end{equation*}
$$

and a new equation (the Degasperis-Procesi equation, the DP equation, for simplicity) which can be written as (after rescaling) the dispersionless form [1]

$$
\begin{equation*}
u_{t}-u_{x x t}+4 u u_{x}=3 u_{x} u_{x x}+u u_{x x x} \tag{4}
\end{equation*}
$$

It is worth noting that in [2] both the Camassa-Holm and DP equations are derived as members of a one-parameter family of asymptotic shallow water approximations to the Euler equations: this is important because it shows that (after
the addition of linear dispersion terms) both the CamassaHolm and DP equations are physically relevant; otherwise the DP equation would be of purely theoretical interest.

When $c_{1}=-3 c_{3} / 2 \alpha^{2}$ and $c_{2}=c_{3} / 2$ in (1), we recover the Camassa-Holm equation derived physically by Camassa and Holm in [3] by approximating directly the Hamiltonian for Euler's equations in the shallow water regime, where $u(x, t)$ represents the free surface above a flat bottom. There is also a geometric approach which is used to prove the least action principle holding for the Camassa-Holm equation, compared with [4]. It is worth pointing out that a fundamental aspect of the Camassa-Holm equation, the fact that it is a completely integrable system, was shown in $[5,6]$. Some satisfactory results have been obtained for this shallow water equation recently, we refer the readers to see [7-19].

Although, the DP equation (4) has a similar form to the Camassa-Holm equation and admits exact peakon solutions analogous to the Camassa-Holm peakons [20], these two equations are pretty different. The isospectral problem for equation (4) is

$$
\begin{equation*}
\Psi_{x}-\Psi_{x x x}-\lambda y \Psi=0 \tag{5}
\end{equation*}
$$

while for Camassa-Holm equation it is

$$
\begin{equation*}
\Psi_{x x}-\frac{1}{4} \Psi-\lambda y \Psi=0 \tag{6}
\end{equation*}
$$

where $y=u-u_{x x}$ for both cases. This implies that the inside structures of the DP equation (4) and the Camassa-Holm equation are truly different. However, we not only have some similar results [21-23], but also have considerable differences in the scattering/inverse scattering approach, compared with the discussion in $[5,6]$ and in the paper [24].

Analogous to the Camassa-Holm equation, (4) can be written in Hamiltonian form and has infinitely many conservation laws. Here we list some of the simplest conserved quantities [20]:

$$
\begin{align*}
H_{-1} & =\int_{\mathbb{R}} u^{3} d x, & H_{0}=\int_{\mathbb{R}} y d x, \quad H_{1}=\int_{\mathbb{R}} y v d x \\
H_{5} & =\int_{\mathbb{R}} y^{1 / 3} d x, & H_{7}=\int_{\mathbb{R}}\left(y_{x}^{2} y^{-7 / 3}+9 y^{-1 / 3}\right) d x \tag{7}
\end{align*}
$$

where $v=\left(4-\partial_{x}^{2}\right)^{-1} u$. So they are different from the invariants of the Camassa-Holm equation

$$
\begin{equation*}
E(u)=\int_{\mathbb{R}}\left(u^{2}+u_{x}^{2}\right) d x, \quad F(u)=\int_{\mathbb{R}}\left(u^{3}+u u_{x}^{2}\right) d x . \tag{8}
\end{equation*}
$$

Set $Q=\left(1-\partial_{x}^{2}\right)$; then the operator $Q^{-1}$ in $\mathbb{R}$ can be expressed by

$$
\begin{equation*}
Q^{-1} f=G * f=\frac{1}{2} \int_{\mathbb{R}} e^{-|x-y|} f(y) d y . \tag{9}
\end{equation*}
$$

Equation (4) can be written as

$$
\begin{equation*}
u_{t}+u u_{x}+\partial_{x} G *\left(\frac{3}{2} u^{2}\right)=0 \tag{10}
\end{equation*}
$$

while the Camassa-Holm equation can be written as

$$
\begin{equation*}
u_{t}+u u_{x}+\partial_{x} G *\left(u^{2}+\frac{1}{2} u_{x}^{2}\right)=0 . \tag{11}
\end{equation*}
$$

On the other hand, the DP equation can also be expressed in the following momentum form:

$$
\begin{gather*}
y_{t}+y_{x} u=-3 y u_{x} \\
y=\left(1-\partial_{x}^{2}\right) u . \tag{12}
\end{gather*}
$$

This formulation is important to motivate us to consider the measure of momentum support which is the second object of this paper, since we found that (12) is similar to the vorticity equation of the three-dimensional Euler equation for incompressible perfect fluids ( $U$ is the speed, and $\omega$ is its vorticity)

$$
\begin{align*}
\omega_{t}+(U \cdot \nabla) \omega & =(\omega \cdot \nabla) U, \\
\operatorname{div} U & =0  \tag{13}\\
\operatorname{curl} U & =\omega
\end{align*}
$$

The stretching term $(\omega \cdot \nabla) U$ in (13) is similar to the term $-3 y u_{x}$ in (12).

One can follow the argument for the Camassa-Holm equation [8] to establish the following well posedness theorem for the Degasperis-Procesi equation.

Theorem 1 (see [23]). Given $u(x, t=0)=u_{0} \in H^{s}(\mathbb{R}), s>$ $3 / 2$, then there exist a $T$ and a unique solution $u$ to (4) (also (10)) such that

$$
\begin{equation*}
u(x, t) \in C\left([0, T) ; H^{s}(\mathbb{R})\right) \cap C^{1}\left([0, T) ; H^{s-1}(\mathbb{R})\right) \tag{14}
\end{equation*}
$$

It should be mentioned that due to the form of (10) (no derivative appears in the convolution term), Coclite and Karlsen [25] established global existence and uniqueness result for entropy weak solutions belonging to the class $L^{1}(\mathbb{R}) \cap B V(\mathbb{R})$.

## 2. Unique Continuation

The purpose of this section is to show that the solution to (10) and its first-order spatial derivative retain algebraic decay at infinity as their initial values do. Precisely, we prove.

Theorem 2. Assume that for some $T>0$ and $s>3 / 2, u \in$ $C\left([0, T] ; H^{s}(\mathbb{R})\right)$ is a strong solution of the initial value problem associated with (10), and that $u_{0}(x)=u(x, 0)$ satisfies that for some $\theta>1$

$$
\begin{equation*}
\left|u_{0}(x)\right|, \quad\left|\partial_{x} u_{0}(x)\right|=O\left(x^{-\theta}\right) \quad \text { as } x \uparrow \infty \tag{15}
\end{equation*}
$$

Then

$$
\begin{equation*}
|u(x, t)|, \quad\left|\partial_{x} u(x, t)\right|=O\left(x^{-\theta}\right) \quad \text { as } x \uparrow \infty, \tag{16}
\end{equation*}
$$

uniformly in the time interval $[0, T]$.
Notation. We will say that

$$
\begin{equation*}
|f(x)|=O\left(x^{-\theta}\right) \quad \text { as } x \uparrow \infty \quad \text { if } \lim _{x \rightarrow \infty} \frac{|f(x)|}{x^{-\theta}}=L, \tag{17}
\end{equation*}
$$

where $L$ is a nonnegative constant.
Proof. We introduce the following notations:

$$
\begin{gather*}
F(u)=\frac{3}{2} u^{2},  \tag{18}\\
M=\sup _{t \in[0, T]}\|u(t)\|_{H^{s}} . \tag{19}
\end{gather*}
$$

Multiplying (10) by $u^{2 p-1}$ with $p \in Z^{+}$and integrating the result in the $x$-variable, one gets

$$
\begin{equation*}
\int_{-\infty}^{\infty} u^{2 p-1}\left(u_{t}+u u_{x}+\partial_{x} G * F(u)\right) d x=0 \tag{20}
\end{equation*}
$$

The first term in (20) is

$$
\begin{align*}
\int_{-\infty}^{\infty} u^{2 p-1} u_{t} d x & =\int_{-\infty}^{\infty} \frac{1}{2 p} \frac{d u^{2 p}}{d t} d x \\
& =\frac{1}{2 p} \frac{d}{d t} \int_{-\infty}^{\infty} u^{2 p} d x=\|u(t)\|_{2 p}^{2 p-1} \frac{d}{d t}\|u(t)\|_{2 p}, \tag{21}
\end{align*}
$$

and for the rest, we have

$$
\begin{align*}
\left|\int_{-\infty}^{\infty} u^{2 p-1} u u_{x} d x\right| & =\left|\int_{-\infty}^{\infty} u^{2 p} u_{x} d x\right| \\
& \leq\left\|u_{x}(t)\right\|_{\infty}\|u(t)\|_{2 p}^{2 p}, \\
\left|\int_{-\infty}^{\infty} u^{2 p-1} \partial_{x} G * F(u) d x\right| & \leq\|u(t)\|_{2 p}^{2 p-1}\left\|\partial_{x} G * F(u)(t)\right\|_{2 p} . \tag{22}
\end{align*}
$$

From the above inequalities, we get

$$
\begin{equation*}
\frac{d}{d t}\|u(t)\|_{2 p} \leq\left\|u_{x}(t)\right\|_{\infty}\|u(t)\|_{2 p}+\left\|\partial_{x} G * F(u)\right\|_{2 p} \tag{23}
\end{equation*}
$$

and therefore, by Sobolev embedding theorem and Gronwall's inequality, there exists a constant $M$ such that

$$
\begin{equation*}
\|u(t)\|_{2 p} \leq\left(\|u(0)\|_{2 p}+\int_{0}^{t}\left\|\partial_{x} G * F(u)\right\|_{2 p} d \tau\right) e^{M t} \tag{24}
\end{equation*}
$$

Since $f \in L^{1}(\mathbb{R}) \cap L^{\infty}(\mathbb{R})$ implies

$$
\begin{equation*}
\lim _{q \rightarrow \infty}\|f\|_{q}=\|f\|_{\infty}, \tag{25}
\end{equation*}
$$

taking the limits in (24) (note that $\partial_{x} G \in L^{1}$ and $F(u) \in L^{1} \cap$ $L^{\infty}$ ) from (25) we get

$$
\begin{equation*}
\|u(t)\|_{\infty} \leq\left(\|u(0)\|_{\infty}+\int_{0}^{t}\left\|\partial_{x} G * F(u)\right\|_{\infty} d \tau\right) e^{M t} \tag{26}
\end{equation*}
$$

We will now repeat the above arguments using the barrier function

$$
\varphi_{N}(x)= \begin{cases}1, & x \leq 1  \tag{27}\\ x^{\theta}, & x \in(1, N) \\ N^{\theta}, & x \geq N\end{cases}
$$

where $N \in \mathbb{Z}^{+}$. Observe that for all $N$ we have

$$
\begin{equation*}
0 \leq \varphi_{N}^{\prime}(x) \leq \theta \varphi_{N}(x) \quad \text { a.e. } x \in \mathbb{R} \tag{28}
\end{equation*}
$$

Using notation in (18), from (10) we obtain

$$
\begin{equation*}
\left(u \varphi_{N}\right)_{t}+\left(u \varphi_{N}\right) u_{x}+\varphi_{N} \partial_{x} G * F(u)=0 . \tag{29}
\end{equation*}
$$

Hence, as in the weightless case (26), we get

$$
\begin{align*}
\left\|u(t) \varphi_{N}\right\|_{\infty} \leq & e^{M t}\left\|u(0) \varphi_{N}\right\|_{\infty} \\
& +e^{M t} \int_{0}^{t}\left\|\varphi_{N} \partial_{x} G * F(u)\right\|_{\infty} d \tau \tag{30}
\end{align*}
$$

A simple calculation shows that there exists $C_{0}>0$ depending only on $\theta$ such that, for any $N \in \mathbb{Z}^{+}$,

$$
\begin{equation*}
\frac{1}{2} \varphi_{N}(x) \int_{-\infty}^{\infty} e^{-|x-y|} \frac{1}{\varphi_{N}(y)} d y \leq C_{0} \tag{31}
\end{equation*}
$$

Thus, for any appropriate function $f$ one finds that

$$
\begin{align*}
& \left|\varphi_{N} \partial_{x} G * f^{2}(x)\right| \\
& \quad=\left|\frac{1}{2} \varphi_{N}(x) \int_{-\infty}^{\infty} \operatorname{sgn}(x-y) e^{-|x-y|} f^{2}(y) d y\right| \\
& \quad \leq \frac{\varphi_{N}(x)}{2} \int_{-\infty}^{\infty} e^{-|x-y|} \frac{1}{\varphi_{N}(y)} \varphi_{N}(y) f(y) f(y) d y \\
& \quad \leq\left(\frac{\varphi_{N}(x)}{2} \int_{-\infty}^{\infty} \frac{e^{-|x-y|}}{\varphi_{N}(y)} d y\right)\left\|\varphi_{N} f\right\|_{\infty}\|f\|_{\infty} \\
& \quad \leq C_{0}\left\|\varphi_{N} f\right\|_{\infty}\|f\|_{\infty} . \tag{32}
\end{align*}
$$

Combining with (30), we get

$$
\begin{equation*}
\left\|u(t) \varphi_{N}\right\|_{\infty} \leq C_{1}\left(\left\|u_{0} \varphi_{N}\right\|_{\infty}+\int_{0}^{t}\left\|\varphi_{N} u\right\|_{\infty} d \tau\right) \tag{33}
\end{equation*}
$$

where $C_{1}=C_{1}(M ; T)>$,0 . By Gronwall's inequality, there exists a constant $\widetilde{C}$ for any $t \in[0, T]$ such that

$$
\begin{equation*}
\left\|\varphi_{N} u\right\|_{\infty} \leq \widetilde{C}\left\|u_{0} \varphi_{N}\right\|_{\infty} \leq \widetilde{C}\left\|u_{0} \cdot \max \left(1, x^{\theta}\right)\right\|_{\infty} \tag{34}
\end{equation*}
$$

Finally, taking the limit as $N$ goes to infinity in (34) we find that for any $t \in[0, T]$

$$
\begin{equation*}
\left|u(x, t) x^{\theta}\right| \leq \widetilde{C}\left\|u_{0} \cdot \max \left(1, x^{\theta}\right)\right\|_{\infty} . \tag{35}
\end{equation*}
$$

From (15), we get $|u(x, t)|=O\left(x^{-\theta}\right)$ as $x \uparrow \infty$.
Next, differentiating (10) in the $x$-variable produces the equation

$$
\begin{equation*}
u_{x t}+u u_{x x}+u_{x}^{2}+\partial_{x}^{2} G *\left(\frac{3}{2} u^{2}\right)=0 \tag{36}
\end{equation*}
$$

Again, multiplying (36) by $u_{x}^{2 p-1},\left(p \in \mathbb{Z}^{+}\right)$, integrating the result in the $x$-variable, and using integration by parts

$$
\begin{align*}
\int_{-\infty}^{\infty} u u_{x x}\left(u_{x}\right)^{2 p-1} d x & =\int_{-\infty}^{\infty} u \frac{\left(u_{x}\right)^{2 p}}{2 p} d x  \tag{37}\\
& =-\frac{1}{2 p} \int_{-\infty}^{\infty} u_{x}\left(u_{x}\right)^{2 p} d x
\end{align*}
$$

one gets the inequality

$$
\begin{equation*}
\frac{d}{d t}\left\|u_{x}(t)\right\|_{2 p} \leq 2\left\|u_{x}(t)\right\|_{\infty}\left\|u_{x}(t)\right\|_{2 p}+\left\|\partial_{x}^{2} G * F(u)\right\|_{2 p} \tag{38}
\end{equation*}
$$

and therefore as before

$$
\begin{equation*}
\left\|u_{x}(t)\right\|_{2 p} \leq\left(\left\|u_{x}(0)\right\|_{2 p}+\int_{0}^{t}\left\|\partial_{x}^{2} G * F(u)\right\|_{2 p} d \tau\right) e^{2 M t} \tag{39}
\end{equation*}
$$

Since $\partial_{x}^{2} G=G-\delta$, we can use (25) and pass to the limit in (39) to obtain

$$
\begin{equation*}
\left\|u_{x}(t)\right\|_{\infty} \leq\left(\left\|u_{x}(0)\right\|_{\infty}+\int_{0}^{t}\left\|\partial_{x}^{2} G * F(u)\right\|_{\infty} d \tau\right) e^{2 M t} \tag{40}
\end{equation*}
$$

from (36) we get

$$
\begin{equation*}
\partial_{t}\left(u_{x} \varphi_{N}\right)+u u_{x x} \varphi_{N}+\left(u_{x} \varphi_{N}\right) u_{x}+\varphi_{N} \partial_{x}^{2} G * F(u)=0 \tag{41}
\end{equation*}
$$

We need to eliminate the second derivatives in the second term in (41). Thus, combining integration by parts and (28), we find

$$
\begin{align*}
& \left|\int_{-\infty}^{\infty} u u_{x x} \varphi_{N}\left(u_{x} \varphi_{N}\right)^{2 p-1} d x\right| \\
& \quad=\left|\int_{-\infty}^{\infty} u\left(u_{x} \varphi_{N}\right)^{2 p-1}\left(\partial_{x}\left(u_{x} \varphi_{N}\right)-u_{x} \varphi_{N}^{\prime}\right) d x\right| \\
& \quad=\left|\int_{-\infty}^{\infty} u\left(\partial_{x}\left(\frac{\left(u_{x} \varphi_{N}\right)^{2 p}}{2 p}\right)-u_{x} \varphi_{N}^{\prime}\left(u_{x} \varphi_{N}\right)^{2 p-1}\right) d x\right| \\
& \quad \leq \kappa \cdot\left(\|u(t)\|_{\infty}+\left\|\partial_{x} u(t)\right\|_{\infty}\right)\left\|\partial_{x} u \varphi_{N}\right\|_{2 p}^{2 p} . \tag{42}
\end{align*}
$$

Since $\partial_{x}^{2} G=G-\delta$, the argument in (32) also shows that

$$
\begin{equation*}
\left|\varphi_{N} \partial_{x}^{2} G * f^{2}(x)\right| \leq C_{0}\left\|\varphi_{N} f\right\|_{\infty}\|f\|_{\infty} . \tag{43}
\end{equation*}
$$

Similarly, we get

$$
\begin{align*}
& \left\|u_{x}(t) \varphi_{N}\right\|_{\infty} \\
& \quad \leq C_{2}\left(\left\|u_{x}(0) \varphi_{N}\right\|_{\infty}+\int_{0}^{t}\left\|u(\tau) \varphi_{N}\right\|_{\infty} d \tau\right), \tag{44}
\end{align*}
$$

where $C_{2}=C_{2}(M ; T)$.
Then, taking the limit as $N$ goes to infinity, we find that for any $t \in[0, T]$

$$
\begin{equation*}
\left|u_{x}(t) x^{\theta}\right| \leq C_{2}\left(\left\|u_{x}(0) x^{\theta}\right\|_{\infty}+\int_{0}^{t}\left\|u(\tau) x^{\theta}\right\|_{\infty} d \tau\right) \tag{45}
\end{equation*}
$$

Since $|u(x, t)|=O\left(x^{-\theta}\right)$ as $x \uparrow \infty$ and (15), we get

$$
\begin{equation*}
\left|\partial_{x} u(x, t)\right|=O\left(x^{-\theta}\right), \quad \text { as } x \uparrow \infty . \tag{46}
\end{equation*}
$$

This completes the proof.

## 3. Measure of Momentum Support

It is known that, for the Degasperis-Procesi equation, the momentum density $y(x, t)$ with compactly supported initial data $y_{0}(x)$ will retain this property; that is, $y(x, t)$ is also compactly supported [21]. However, the same argument for $u(x, t)$ is false [21]. Note that a detailed description of solution $u(x, t)$ outside of the support of $y(x, t)$ is given in [26,27].

Moreover, the exponential behavior of $u$ in $x$ outside this support is obvious. The comparison of the DP equation and the incompressible Euler equation above implies that the momentum $y(x, t)$ in (12) plays a similar role as the vorticity does in (13). This motivates us to estimate the size of supp $y(t, \cdot)$ for strong solutions. The approach is inspired by the work of Kim [28] and the recent work [29].

We first introduce the particle trajectory method. Let $u \in$ $C\left([0, T], H^{3}(\mathbb{R})\right) \cap C^{1}\left([0, T], H^{2}(\mathbb{R})\right)$ be a strong solution of (4) guaranteed by the well posedness Theorem 1 . Let $s \in$ $[0, T], q(t ; \alpha, s)$ be the solution of the following initial value problem:

$$
\begin{gather*}
\frac{d q(t ; \alpha, s)}{d t}=u(s+t, q(t ; \alpha, s)), s, s+t \in[0, T], \quad \alpha \in \mathbb{R}, \\
q(0 ; \alpha, s)=\alpha, \quad \alpha \in \mathbb{R} . \tag{47}
\end{gather*}
$$

Then, $q(t ; \cdot, s): \mathbb{R} \rightarrow \mathbb{R}$ is an increasing diffeomorphism. It is shown $[21,23]$ that

$$
\begin{equation*}
y(q(t ; x, 0), t) q_{x}^{3}(t ; x, 0)=y(x, 0) ; \tag{48}
\end{equation*}
$$

this implies that the support of $y$ propagates along the flow. Set $D(t)$ to be the support of $y(\cdot, t)$. Let $\psi \in L^{2}(D(s))$, and let $\psi^{t} \in L^{2}(D(s+t))$ be given by the following:

$$
\begin{equation*}
\psi^{t}(q(t ; \alpha, s))=\psi(\alpha) . \tag{49}
\end{equation*}
$$

Moreover, we also want to mention the standard argument on the first Dirichlet eigenvalue problem. Let $\Omega$ be an open interval in $\mathbb{R}$, and, $\lambda_{1}(\Omega)$ be the first Dirichlet eigenvalue of the Laplacian on $\Omega$. Then we have

$$
\begin{equation*}
\lambda_{1}(\Omega)=\inf \left\{\left\|\phi^{\prime}\right\|_{L^{2}(\Omega)}^{2} \mid \phi \in H_{0}^{1}(\Omega) \text { with }\|\phi\|_{L^{2}(\Omega)}=1\right\} . \tag{50}
\end{equation*}
$$

It is just $(\pi /|\Omega|)^{2}$ and the normalized eigenfunctions are the suitable translations of

$$
\begin{equation*}
\pm\left(\frac{2}{|\Omega|}\right)^{1 / 2} \sin \left(\frac{\pi x}{|\Omega|}\right) \tag{51}
\end{equation*}
$$

Theorem 3. Let $y \in C\left([0, T] ; H^{1}(\mathbb{R})\right) \cap C^{1}\left([0, T] ; L^{2}(\mathbb{R})\right)$ be a strong solution of (12). Let $D(t)$ be the support of $y(\cdot, t)$ for $t \in[0, T]$ with its initial $D(0)$ being connected.
(I) Suppose there exists a positive constant $K$ such that $u_{x}(x, k)>-K$ for $(x, t) \in \mathbb{R} \times[0, T]$. Then

$$
\begin{align*}
& |D(0)| e^{-\left(\exp (5 K T / 2)\left\|y_{0}\right\|_{L^{2}(\mathbb{R})}\right) t} \\
& \quad \leq|D(t)| \leq|D(0)| e^{\left(\exp (5 K T / 2)\left\|y_{0}\right\|_{L^{2}(\mathbb{R})}\right) t} . \tag{52}
\end{align*}
$$

(II) $y_{0}$ does not change sign or

$$
\begin{align*}
& y_{0}(x) \leq 0, \quad x \in\left(-\infty, x_{0}\right) \\
& y_{0}(x) \geq 0, \quad x \in\left(x_{0}, \infty\right) \tag{53}
\end{align*}
$$

and $y_{0} \in H^{1}(\mathbb{R}) \cap L^{1}(\mathbb{R})$; then, for all $t \geq 0$

$$
\begin{align*}
|D(0)| e^{-\left\|y_{0}\right\|_{L^{1}(\mathbb{R})^{t}}} & \leq|D(t)|  \tag{54}\\
& \leq|D(0)| e^{\left.\left\|y_{0}\right\|_{L^{1}(\mathbb{R})}\right)^{t}} .
\end{align*}
$$

Proof. (I) The relation of momenta $y$ and $u$ gives

$$
\begin{gather*}
u(x, t)=\frac{1}{2} \int_{\mathbb{R}} e^{-|x-\xi|} y(\xi, t) d \xi  \tag{55}\\
u_{x}(x, t)=\frac{1}{2} \int_{\mathbb{R}} \operatorname{sgn}(\xi-x) e^{-|x-\xi|} y(\xi, t) d \xi \tag{56}
\end{gather*}
$$

Then, we have by (12) and the lower bound of $u_{x}$

$$
\begin{align*}
& \frac{d}{d t} \int_{\mathbb{R}} y^{2}(x, t) d x \\
& \quad=-5 \int_{\mathbb{R}} u_{x}(x, t) y^{2}(x, t) d x \leq 5 K \int_{\mathbb{R}} y^{2}(x, t) d x \tag{57}
\end{align*}
$$

Thus

$$
\begin{equation*}
\frac{d}{d t}\|y(x, t)\|_{L^{2}}^{2} \leq 5 K\|y(x, t)\|_{L^{2}}^{2} \tag{58}
\end{equation*}
$$

Therefore, (56), (58), and Gronwall inequality imply that

$$
\begin{equation*}
\left|u_{x}(x, t)\right| \leq \frac{1}{2}\|y(x, t)\|_{L^{2}} \leq \frac{1}{2} e^{5 K T / 2}\left\|y_{0}\right\|_{L^{2}} \tag{59}
\end{equation*}
$$

On the other hand, due to Propositions A. 2 and A.3, $\lambda_{1}(D(s))$ is Lipschitz and differentiable almost everywhere. Moreover, we have

$$
\begin{equation*}
-4 M_{1} \lambda_{1}(D(s)) \leq \frac{d}{d s} \lambda_{1}(D(s)) \leq 4 M_{1} \lambda_{1}(D(s)) \tag{60}
\end{equation*}
$$

Then, it follows that

$$
\begin{equation*}
e^{-4 M_{1} s} \lambda_{1}(D(0)) \leq \lambda_{1}(D(s)) \leq e^{4 M_{1} s} \lambda_{1}(D(0)) \tag{61}
\end{equation*}
$$

with $\lambda_{1}(D(s))=\pi^{2} /|D(s)|^{2}$. So (52) follows from (61) and (59).
(II) If $y_{0} \in H^{1}(\mathbb{R}) \cap L^{1}(\mathbb{R})$ does not change sign, we conclude that solutions of (10) exist globally in time. Equality (56) and the conservation of $\int_{\mathbb{R}} y(x, t) d x$ yield

$$
\begin{equation*}
\left|u_{x}(x, t)\right| \leq \frac{1}{2}\|y(x, t)\|_{L^{1}(\mathbb{R})}=\frac{1}{2}\left\|y_{0}(x)\right\|_{L^{1}(\mathbb{R})} \tag{62}
\end{equation*}
$$

By similar arguments of (I), constant $M_{1}$ in (61) can be replaced by $\left\|y_{0}(x)\right\|_{L^{1}(\mathbb{R})} / 2$; then (54) follows. If (53) is satisfied, we know that the solution of (10) exists globally in time [21, 30]. From (53) and (48), it is easy to get

$$
\begin{align*}
& y(x, t) \leq 0,  \tag{63}\\
& y \in\left(-\infty, q\left(x_{0}, t\right)\right) \\
& y(x, t) \geq 0, \\
& x \in\left(q\left(x_{0}, t\right), \infty\right)
\end{align*}
$$

where we denote $q(t ; x, s)$ with $s=0$ by $q(x, t)$. By direct computation, we have

$$
\begin{equation*}
\int_{\mathbb{R}}|y(x, t)| d x=\int_{q\left(x_{0}, t\right)}^{\infty} y(x, t) d x-\int_{-\infty}^{q\left(x_{0}, t\right)} y(x, t) d x . \tag{64}
\end{equation*}
$$

Next, we prove that $\|y(x, t)\|_{L^{1}(\mathbb{R})}$ is decreasing with respect to time. To this end, one gets, by differentiating (64) with respect to $t$ and integrating by parts,

$$
\begin{align*}
\frac{d}{d t} \int_{\mathbb{R}}|y(x, t)| d x= & \int_{q\left(x_{0}, t\right)}^{\infty} y_{t}(x, t) d x \\
& -\int_{-\infty}^{q\left(x_{0}, t\right)} y_{t}(x, t) d x \\
& -2(y u)\left(q\left(x_{0}, t\right), t\right) \\
= & -\int_{q\left(x_{0}, t\right)}^{\infty}\left(y_{x} u+3 y u_{x}\right) d x \\
& +\int_{-\infty}^{q\left(x_{0}, t\right)}\left(y_{x} u+3 y u_{x}\right) d x \\
& -2(y u)\left(q\left(x_{0}, t\right), t\right) \\
= & -2 \int_{q\left(x_{0}, t\right)}^{\infty} y u_{x} d x+2 \int_{-\infty}^{q\left(x_{0}, t\right)} y u_{x} d x \\
= & u^{2}\left(q\left(x_{0}, t\right), t\right)-u_{x}^{2}\left(q\left(x_{0}, t\right), t\right) \\
= & \int_{q\left(x_{0}, t\right)}^{\infty} e^{-\xi} y(\xi, t) d x \int_{-\infty}^{q\left(x_{0}, t\right)} e^{\xi} y(\xi, t) d x \\
\leq & 0 . \tag{65}
\end{align*}
$$

This implies that

$$
\begin{equation*}
\left|u_{x}(x, t)\right| \leq \frac{1}{2}\|y(x, t)\|_{L^{1}(\mathbb{R})} \leq \frac{1}{2}\left\|y_{0}(x)\right\|_{L^{1}(\mathbb{R})} \tag{66}
\end{equation*}
$$

Therefore, (54) follows by replacing $M_{1}$ with $\left\|y_{0}(x)\right\|_{L^{1}(\mathbb{R})} / 2$ in (61).

## Appendix

The following propositions with standard proofs are known in [29]; we list them here only for convenience of readers.

Proposition A.1. Let $s, s+t \in[0, T], \alpha \in D(s)$, and $\psi \in$ $H_{0}^{1}(D(s)) ; u_{x}$ can be bounded by a constant $M_{1}$; then
(a)

$$
\begin{equation*}
e^{-M_{1}|t|} \leq q_{\alpha}(t ; \alpha, s) \leq e^{M_{1}|t|} \tag{A.1}
\end{equation*}
$$

(b)

$$
\begin{align*}
\left|\psi^{\prime}(\alpha)\right| e^{-M_{1}|t|} & \leq\left|\left(\psi^{t}\right)^{\prime}(q(t ; \alpha, s))\right|  \tag{A.2}\\
& \leq\left|\psi^{\prime}(\alpha)\right| e^{M_{1}|t|}
\end{align*}
$$

(c)

$$
\begin{align*}
\|\psi\|_{L^{2}(D(s))} e^{-M_{1}|t| / 2} & \leq\left\|\psi^{t}\right\|_{L^{2}(D(s+t))} \\
& \leq\|\psi\|_{L^{2}(D(s))^{e_{1}|t| / 2}} . \tag{A.3}
\end{align*}
$$

Proof. (a) Differentiating (47) with respect to $\alpha$, we obtain

$$
\begin{equation*}
\frac{d q_{t}}{d \alpha}=u_{q} q_{\alpha} \tag{A.4}
\end{equation*}
$$

Since $q(t ; \cdot, s): \mathbb{R} \rightarrow \mathbb{R}$ is an increasing diffeomorphism, then $q_{\alpha}>0$. Combining the bound of $u_{x}$, there holds

$$
\begin{equation*}
-M_{1} q_{\alpha} \leq q_{\alpha t} \leq M_{1} q_{\alpha} \tag{A.5}
\end{equation*}
$$

This can be solved as (a).
(b) Differentiating (49) with respect to $\alpha$ to get

$$
\begin{equation*}
\psi_{q}^{t} q_{\alpha}=\psi^{\prime}(\alpha), \tag{A.6}
\end{equation*}
$$

then (A.2) is a direct consequence of (A.1).
(c) Equation (49) and the definition of Sobolev norm give that

$$
\begin{equation*}
\left\|\psi^{t}\right\|_{L^{2}(D(s+t))}^{2}=\int_{D(s+t)} \psi^{t}(x)^{2} d x=\int_{D(s)} \psi^{2}(\alpha) q_{\alpha} d \alpha \tag{A.7}
\end{equation*}
$$

where we have used the change of variable $x=q(t ; \alpha, s)$. So (A.3) follows from (A.1).

Proposition A.2. Under the hypothesis of Theorem 3, for $s, s+$ $t \in[0, T]$,

$$
\begin{align*}
& \lim _{t \rightarrow 0^{+}} \sup \frac{\lambda_{1}(D(s+t))-\lambda_{1}(D(s))}{t} \leq 4 M_{1} \lambda_{1}(D(s)), \\
& \lim _{t \rightarrow 0^{-}} \inf \frac{\lambda_{1}(D(s+t))-\lambda_{1}(D(s))}{t} \geq-4 M_{1} \lambda_{1}(D(s)) . \tag{A.8}
\end{align*}
$$

Proof. Let $t>0, \phi_{1} \in H_{0}^{1}(D(s))$ with $\left\|\phi_{1}\right\|_{L^{2}(D(s))}=1$ be a first normalized eigenfunction on $D(s)$. Then, for $\varphi \in H_{0}^{1}(D(s+t))$ with $\|\varphi\|_{L^{2}(D(s+t))}=1$, we have

$$
\begin{align*}
\lambda_{1}(D(s+t))-\lambda_{1}(D(s))= & \inf \left\|\varphi^{\prime}\right\|_{L^{2}(D(s+t))}^{2}-\left\|\phi_{1}^{\prime}\right\|_{L^{2}(D(s))}^{2} \\
\leq & \left\|\phi_{1}^{t}\right\|_{L^{2}(D(s+t))}^{-2}\left\|\left(\phi_{1}^{t}\right)^{\prime}\right\|_{L^{2}(D(s+t))}^{2} \\
& -\left\|\phi_{1}^{\prime}\right\|_{L^{2}(D(s))}^{2} . \tag{A.9}
\end{align*}
$$

Furthermore

$$
\begin{aligned}
& \left\|\phi_{1}^{t}\right\|_{L^{2}(D(s+t))}^{-2}\left\|\left(\phi_{1}^{t}\right)^{\prime}\right\|_{L^{2}(D(s+t))}^{2} \\
& \quad=\left\|\phi_{1}^{t}\right\|_{L^{2}(D(s+t))}^{-2} \int_{D(s)}\left[\left(\phi_{1}^{t}\right)^{\prime}\right]^{2} q_{\alpha} d \alpha \\
& \quad \leq\left\|\phi_{1}^{t}\right\|_{L^{2}(D(s+t))}^{-2} e^{3 M_{1} t}\left\|\phi_{1}^{\prime}\right\|_{L^{2}(D(s))}^{2} \\
& \quad \leq e^{4 M_{1} t}\left\|\phi_{1}^{\prime}\right\|_{L^{2}(D(s))}^{2} .
\end{aligned}
$$

Combing (A.9) and (A.10) together yields

$$
\begin{align*}
\lim _{t \rightarrow 0^{+}} & \sup \frac{\lambda_{1}(D(s+t))-\lambda_{1}(D(s))}{t} \\
& \leq \lim _{t \rightarrow 0^{+}} \sup \frac{e^{4 M_{1} t}\left\|\phi_{1}^{\prime}\right\|_{L^{2}(D(s))}^{2}-\left\|\phi_{1}^{\prime}\right\|_{L^{2}(D(s))}^{2}}{t}  \tag{A.11}\\
& =4 M_{1} \lambda_{1}(D(s)) .
\end{align*}
$$

The second one follows by similar arguments for $t<0$.
Proposition A.3. Under the hypothesis of Theorem 3, for $s, s+$ $t \in[0, T]$,

$$
\begin{align*}
& \lim _{t \rightarrow 0^{-}} \sup \frac{\lambda_{1}(D(s+t))-\lambda_{1}(D(s))}{t} \leq 4 M_{1} \lambda_{1}(D(s)), \\
& \lim _{t \rightarrow 0^{+}} \inf \frac{\lambda_{1}(D(s+t))-\lambda_{1}(D(s))}{t} \geq-4 M_{1} \lambda_{1}(D(s)) . \tag{A.12}
\end{align*}
$$

Proof. Let $\phi_{1} \in H_{0}^{1}(D(s))$ with $\left\|\phi_{1}\right\|_{L^{2}(D(s))}=1$ be a first normalized eigenfunction on $D(s)$, and let $\phi_{2} \in L^{2}(D(s))$ be such that its $t$-transport is a normalized first eigenfunction on $D(s+t)$. For $t>0$, using the left halves of (A.1) and (A.2) and then the right half of (A.3) we get

$$
\begin{align*}
\left\|\left(\phi_{2}^{t}\right)^{\prime}\right\|_{L^{2}(D(s+t))}^{2} & =\int_{D(s+t)}\left[\left(\phi_{2}^{t}(x)\right)^{\prime}\right]^{2} d x \\
& =\int_{D(s)}\left[\left(\phi_{2}^{t}\right)^{\prime}\right]^{2} q_{\alpha} d \alpha \\
& \geq e^{-3 M_{1} t} \int_{D(s)}\left[\phi_{2}^{\prime}(\alpha)\right]^{2} d \alpha \\
& =e^{-3 M_{1} t}\left\|\phi_{2}\right\|_{L^{2}(D(s))}^{2}\left\|\left(\frac{\phi_{2}}{\left\|\phi_{2}\right\|_{L^{2}(D(s))}^{2}}\right)^{\prime}\right\|_{L^{2}(D(s))}^{2} \\
& \geq e^{-4 M_{1} t}\left\|\phi_{2}^{t}\right\|_{L^{2}(D(s+t))}^{2} \lambda_{1}(D(s)) \\
& =e^{-4 M_{1} t} \lambda_{1}(D(s)) . \tag{A.13}
\end{align*}
$$

Hence

$$
\begin{align*}
& \lim _{t \rightarrow 0^{+}} \inf \frac{\lambda_{1}(D(s+t))-\lambda_{1}(D(s))}{t} \\
& \quad \geq \lim _{t \rightarrow 0^{+}} \inf \frac{e^{-4 M_{1} t}-1}{t} \lambda_{1}(D(s))  \tag{A.14}\\
& \quad=-4 M_{1} \lambda_{1}(D(s))
\end{align*}
$$

The other part is similar.

## Acknowledgments

This work was partially supported by ZJNSF, under Grant nos. LQ12A01009 and LQ13A010008, and NSFC, under Grant nos. 11301394,11226176, and 11226172.

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## Research Article

# Existence and Decay Estimate of Global Solutions to Systems of Nonlinear Wave Equations with Damping and Source Terms 

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Received 30 April 2013; Revised 1 September 2013; Accepted 2 September 2013
Academic Editor: T. Raja Sekhar
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The initial-boundary value problem for a class of nonlinear wave equations system in bounded domain is studied. The existence of global solutions for this problem is proved by constructing a stable set and obtain the asymptotic stability of global solutions through the use of a difference inequality.

## 1. Introduction

In this paper, we are concerned with the global solvability and decay stabilization for the following nonlinear wave equations system:

$$
\begin{align*}
u_{t t} & -\operatorname{div}\left(|\nabla u|^{p-2} \nabla u\right)+\left|u_{t}\right|^{q-2} u_{t}-\Delta u_{t}  \tag{1}\\
& =|v|^{r+2}|u|^{r} u, \quad(x, t) \in \Omega \times R^{+}, \\
v_{t t} & -\operatorname{div}\left(|\nabla v|^{p-2} \nabla v\right)+\left|v_{t}\right|^{q-2} v_{t}-\Delta v_{t}  \tag{2}\\
& =|u|^{r+2}|v|^{r} v, \quad(x, t) \in \Omega \times R^{+}
\end{align*}
$$

with the initial-boundary value conditions

$$
u(x, 0)=u_{0}(x) \in W_{0}^{1, p}(\Omega), \quad u_{t}(x, 0)=u_{1}(x) \in L^{2}(\Omega)
$$

$$
\begin{array}{r}
v(x, 0)=v_{0}(x) \in W_{0}^{1, p}(\Omega), \quad v_{t}(x, 0)=v_{1}(x) \in L^{2}(\Omega)  \tag{3}\\
x \in \Omega
\end{array}
$$

$$
\begin{equation*}
u(x, t)=0, \quad v(x, t)=0, \quad(x, t) \in \partial \Omega \times R^{+} \tag{5}
\end{equation*}
$$

where $\Omega$ is a bounded open domain in $R^{n}$ with a smooth boundary $\partial \Omega, p, q \geq 2, r>0$ and $p<2(r+2) \leq n p /(n-p)$ for $n \geq p$ and $p<2(r+2)<+\infty$ for $n<p$.

When $p=2$, Medeiros and Miranda [1] proved the existence and uniqueness of global weak solutions. Cavalcanti et al. in $[2-4]$ considered the asymptotic behavior for wave equation and an analogous hyperbolic-parabolic system with boundary damping and boundary source term. In paper [5, 6], the authors dealt with the existence, uniform decay rates, and blowup for solutions of systems of nonlinear wave equations with damping and source terms.

Rammaha and Wilstein [7] and Yang [8] are concerned with the initial boundary value problem for a class of quasilinear evolution equations with nonlinear damping and source terms. Under appropriate conditions, by a Galerkin approximation scheme combined with the potential well method, they proved the existence and asymptotic behavior of global weak solutions when $m<p$, where $m \geq 0$ and $p$ are, respectively, the growth orders of the nonlinear strain terms and the source term.

Ono [9] considers the following initial-boundary value problem for nonlinear wave equations with nonlinear dissipative terms:

$$
\begin{gather*}
u_{t t}-\Delta u+\delta_{1} u_{t}+\delta_{2}\left|u_{t}\right|^{\beta} u_{t}-\delta_{3} \Delta u_{t}=|u|^{\alpha} u \\
\quad(x, t) \in \Omega \times R^{+}  \tag{6}\\
u(x, 0)=u_{0}(x), \quad u_{t}(x, 0)=u_{1}(x), \quad x \in \Omega \\
u(x, t)=0, \quad x \in \partial \Omega, t \geq 0
\end{gather*}
$$

where $\delta_{i} \geq 0, i=1,2,3$, and $\alpha, \beta>0$ are constants. The author mainly investigates on the blowup phenomenon to problem (6). On the other hand, in the case of $\delta_{1}+\delta_{2}+\delta_{3}>0$, he shows that the problem (6) admits a unique global solution, and its energy has some decay properties under some assumptions on $u_{0}$ and initial energy $E(0) \equiv E\left(u_{0}, u_{1}\right)$. In particular, when $\delta_{2}>0$ and $\delta_{1}+\delta_{3}>0$ in (6), the energy $E(t) \equiv$ $E\left(u(t), u_{t}(t)\right)$ has some polynomial and exponential decay rates, respectively.

For the following strongly damped nonlinear wave equation

$$
\begin{equation*}
u_{t t}-\Delta u_{t}-\Delta u+f\left(u_{t}\right)+g(u)=h \tag{7}
\end{equation*}
$$

Dell'Oro and Pata [10] obtain the long-time behavior of the related solution semigroup, which is shown to possess the global attractor in the natural weak energy space. In addition, the existence of global and local solutions, decay estimates, and blowup for solutions of nonlinear wave equation with source and damping terms and exponential nonlinearities are studied in [11-14].

In this paper, we prove the global existence for the problem (1)-(5) by applying the potential well theory introduced by Sattinger [15] and Payne and Sattinger [16]. Meanwhile, we obtain the asymptotic stabilization of global solutions by using a difference inequality [17].

For simplicity of notations, hereafter we denote by $\|\cdot\|_{p}$ the norm of $L^{p}(\Omega) ;\|\cdot\|$ denotes $L^{2}(\Omega)$ norm, and we write equivalent norm $\|\cdot \nabla\|_{p}$ instead of $W_{0}^{1, p}(\Omega)$ norm $\|\cdot\|_{W_{0}^{1, p}(\Omega)}$. Moreover, $C$ denotes various positive constants depending on the known constants and may be different at each appearance.

## 2. Local Existence

In this section, we investigate the local existence and uniqueness of the solutions of the problem (1)-(5). For this purpose, we list up two useful lemmas which will be used later and give the definition of weak solutions.

Lemma 1. Let $u \in W_{0}^{1, p}(\Omega)$, then $u \in L^{s}(\Omega)$; and the inequality $\|u\|_{s} \leq C\|u\|_{W_{0}^{1, p}(\Omega)}$ holds with a constant $C>0$ depending on $\Omega$, $p$, and $s$, provided that $2 \leq s<+\infty, 2 \leq n \leq p$ and $2 \leq s \leq n p /(n-p), 2<p<n$.

Lemma 2 (Young inequality). Let $a, b \geq 0$ and $1 / p+1 / q=1$ for $1<p, q<+\infty$; then one has the inequality

$$
\begin{equation*}
a b \leq \delta a^{p}+C(\delta) b^{q}, \tag{8}
\end{equation*}
$$

where $\delta>0$ is an arbitrary constant, and $C(\delta)$ is a positive constant depending on $\delta$.

Definition 3. A pair of functions $(u, v)$ is said to be a weak solution of (1)-(5) on $[0, T]$ if $u, v \in C\left([0, T], W_{0}^{1, p}(\Omega)\right)$,
$u_{t}, v_{t} \in C\left([0, T], L^{2}(\Omega)\right),[u(0), v(0)]=\left[u_{0}, v_{0}\right] \in W_{0}^{1, p}(\Omega) \times$ $W_{0}^{1, p}(\Omega),\left[u_{t}(0), v_{t}(0)\right]=\left[u_{1}, v_{1}\right] \in L^{2}(\Omega) \times L^{2}(\Omega)$, and $[u, v]$ satisfies

$$
\begin{align*}
& \left\langle u_{t}(t), \phi\right\rangle_{L^{2}(\Omega)}-\left\langle u_{1}, \phi\right\rangle_{L^{2}(\Omega)} \\
& \quad+\int_{0}^{t}\left\langle\left(|\nabla u|^{p-2} \nabla u\right), \nabla \phi\right\rangle_{L^{2}(\Omega)} d \tau \\
& \left.\quad+\left.\int_{0}^{t}\langle | u_{t}\right|^{q-2} u_{t}, \phi\right\rangle_{L^{2}(\Omega)} d \tau+\int_{0}^{t}\left\langle\nabla u_{t}, \nabla \phi\right\rangle_{L^{2}(\Omega)} \\
& \left.=\left.\int_{0}^{t}\langle | v\right|^{r+2}|u|^{r} u, \phi\right\rangle_{L^{2}(\Omega)} d \tau \\
& \left\langle v_{t}(t), \psi\right\rangle_{L^{2}(\Omega)}-\left\langle v_{1}, \psi\right\rangle_{L^{2}(\Omega)}  \tag{9}\\
& \quad+\int_{0}^{t}\left\langle\left(|\nabla v|^{p-2} \nabla v\right), \nabla \psi\right\rangle_{L^{2}(\Omega)} d \tau \\
& \left.\quad+\left.\int_{0}^{t}\langle | v_{t}\right|^{q-2} v_{t}, \psi\right\rangle_{L^{2}(\Omega)} d \tau+\int_{0}^{t}\left\langle\nabla v_{t}, \nabla \psi\right\rangle_{L^{2}(\Omega)} \\
& \left.=\left.\int_{0}^{t}\langle | u\right|^{r+2}|v|^{r} v, \psi\right\rangle_{L^{2}(\Omega)} d \tau,
\end{align*}
$$

for all test functions $\phi, \psi \in W_{0}^{1, p}(\Omega)$ and for almost all $t \in$ $[0, T]$.

The local existence and uniqueness of solutions for problem (1)-(5) can be proved through the use of Galerkin method. The result reads as follows.

Theorem 4 (local solution). Supposed that $\left[u_{0}, v_{0}\right] \in$ $W_{0}^{1, p}(\Omega) \times W_{0}^{1, p}(\Omega),\left[u_{1}, v_{1}\right] \in L^{2}(\Omega) \times L^{2}(\Omega)$, and $p<2(r+$ 2) $\leq n p /(n-p)$ if $n \geq p$ and $p<2(r+2)<+\infty$ for $n<p$, then there exists $T>0$ such that the problem (1)-(5) has a unique local solution $[u(t), v(t)]$ satisfying

$$
\begin{gather*}
{[u, v] \in L^{\infty}\left([0, T) ; W_{0}^{1, p}(\Omega) \times W_{0}^{1, p}(\Omega)\right) ;}  \tag{10}\\
{\left[u_{t}, v_{t}\right] \in L^{\infty}\left([0, T) ; L^{2}(\Omega) \times L^{2}(\Omega)\right)} \\
E(t)+\int_{0}^{t}\left(\left\|\nabla u_{\tau}(\tau)\right\|^{2}+\left\|\nabla v_{\tau}(\tau)\right\|^{2}\right.  \tag{11}\\
\left.+\|u(\tau)\|_{q}^{q}+\|v(\tau)\|_{q}^{q}\right) d \tau=E(0)
\end{gather*}
$$

where

$$
\begin{align*}
E(t)= & \frac{1}{2}\left(\left\|u_{t}\right\|^{2}+\left\|v_{t}\right\|^{2}\right)+\frac{1}{p}\left(\|\nabla u\|_{p}^{p}+\|\nabla v\|_{p}^{p}\right)  \tag{12}\\
& -\frac{1}{r+2}\|u v\|_{r+2}^{r+2} .
\end{align*}
$$

Proof. Let $\left\{\omega_{i}\right\}_{i=1}^{\infty}$ be a basis for $W_{0}^{1, p}(\Omega)$. Supposed that $V_{k}$ is the subspace of $W_{0}^{1, p}(\Omega)$ generated by $\left\{\omega_{1}, \omega_{2}, \ldots, \omega_{k}\right\}, k \in N$. We are going to look for the approximate solution

$$
\begin{equation*}
u_{k}(t)=\sum_{i=1}^{k} g_{i k}(t) \omega_{i}, \quad v_{k}(t)=\sum_{i=1}^{k} h_{i k}(t) \omega_{i} \tag{13}
\end{equation*}
$$

which satisfies the following Cauchy problem:

$$
\begin{align*}
& \int_{\Omega}\left(u_{k}^{\prime \prime}-\operatorname{div}\left(\left|\nabla u_{k}\right|^{p-2} \nabla u_{k}\right)+\left|u_{k}^{\prime}\right|^{q-2} u_{k}^{\prime}-\Delta u_{k}^{\prime}\right) \omega_{i} d x  \tag{14}\\
& =\int_{\Omega}\left|v_{k}\right|^{r+2}\left|u_{k}\right|^{r} u_{k} \omega_{i} d x, \\
& \int_{\Omega}\left(v_{k}^{\prime \prime}-\operatorname{div}\left(\left|\nabla v_{k}\right|^{p-2} \nabla v_{k}\right)+\left|v_{k}^{\prime}\right|^{q-2} v_{k}^{\prime}-\Delta v_{k}^{\prime}\right) \omega_{i} d x \\
& =\int_{\Omega}\left|u_{k}\right|^{r+2}\left|v_{k}\right|^{r} v_{k} \omega_{i} d x,  \tag{15}\\
& u_{k}(0)=u_{0 k}=\sum_{i=1}^{k}\left(u_{0}, \omega_{i}\right) \omega_{i} \longrightarrow u_{0}, \quad \text { in } W_{0}^{1, p}(\Omega),  \tag{16}\\
& k \longrightarrow \infty \text {, } \\
& v_{k}(0)=v_{0 k}=\sum_{i=1}^{k}\left(v_{0}, \omega_{i}\right) \omega_{i} \longrightarrow v_{0} \quad \text { in } W_{0}^{1, p}(\Omega),  \tag{17}\\
& k \longrightarrow \infty \text {, } \\
& u_{k}^{\prime}(0)=u_{1 k}=\sum_{i=1}^{k}\left(u_{1}, \omega_{i}\right) \omega_{i} \longrightarrow u_{1} \quad \text { in } L^{2}(\Omega),  \tag{18}\\
& k \longrightarrow \infty \text {, } \\
& v_{k}^{\prime}(0)=v_{1 k}=\sum_{i=1}^{k}\left(v_{1}, \omega_{i}\right) \omega_{i} \longrightarrow v_{1} \quad \text { in } L^{2}(\Omega),  \tag{19}\\
& k \longrightarrow \infty .
\end{align*}
$$

Note that, we can solve the problem (14)-(19) by a Picard's iteration method in ordinary differential equations. Hence, there exists a solution in $\left[0, T_{k}\right.$ ) for some $T_{k}>0$, and we can extend this solution to the whole interval $[0, T]$ for any given $T>0$ by making use of the a priori estimates below.

Multiplying (14) by $g_{i k}^{\prime}(t)$ and (15) by $h_{i k}^{\prime}(t)$ and summing over $i$ from 1 to $k$, we obtain

$$
\begin{align*}
\frac{1}{2} & \frac{d}{d t}\left(\left\|u_{k}^{\prime}(t)\right\|^{2}+\left\|\nabla u_{k}\right\|_{p}^{p}\right)+\left\|u_{k}^{\prime}(t)\right\|_{q}^{q}+\left\|\nabla u_{k}^{\prime}(t)\right\|^{2} \\
\quad & =\int_{\Omega}\left|v_{k}\right|^{r+2}\left|u_{k}\right|^{r} u_{k} u_{k}^{\prime} d x,  \tag{20}\\
\frac{1}{2} & \frac{d}{d t}\left(\left\|v_{k}^{\prime}(t)\right\|^{2}+\left\|\nabla v_{k}\right\|_{p}^{p}\right)+\left\|v_{k}^{\prime}(t)\right\|_{q}^{q}+\left\|\nabla v_{k}^{\prime}(t)\right\|^{2}  \tag{21}\\
\quad & =\int_{\Omega}\left|u_{k}\right|^{r+2}\left|v_{k}\right|^{r} v_{k} v_{k}^{\prime} d x .
\end{align*}
$$

By summing (20) and (21) and integrating the resulting identity over $[0, t]$, we have

$$
\begin{align*}
& \frac{1}{2}\left(\left\|u_{k}^{\prime}(t)\right\|^{2}+\left\|v_{k}^{\prime}(t)\right\|^{2}+\left\|\nabla u_{k}\right\|_{p}^{p}+\left\|\nabla v_{k}\right\|_{p}^{p}\right) \\
& \quad+\int_{0}^{t}\left(\left\|\nabla u_{k}^{\prime}(t)\right\|^{2}+\left\|\nabla v_{k}^{\prime}(t)\right\|^{2}\right. \\
& \left.\quad+\left\|u_{k}^{\prime}(\tau)\right\|_{q}^{q}+\left\|v_{k}^{\prime}(\tau)\right\|_{q}^{q}\right) d \tau  \tag{22}\\
& \leq C_{0}+\int_{0}^{t} \int_{\Omega}\left(\left|v_{k}\right|^{r+2}\left|u_{k}\right|^{r} u_{k} u_{k}^{\prime}\right. \\
& \left.\quad+\left|u_{k}\right|^{r+2}\left|v_{k}\right|^{r} v_{k} v_{k}^{\prime}\right) d x d \tau
\end{align*}
$$

We estimate the right-hand terms of (22) as follows: we get from Hölder inequality and Lemmas 1 and 2 that

$$
\begin{align*}
& \left|\int_{0}^{t} \int_{\Omega}\left(\left|v_{k}\right|^{r+2}\left|u_{k}\right|^{r} u_{k} u_{k}^{\prime}+\left|u_{k}\right|^{r+2}\left|v_{k}\right|^{r} v_{k} v_{k}^{\prime}\right) d x d \tau\right| \\
& \quad \leq \int_{0}^{t}\left(\left\|u_{k}^{\prime}(\tau)\right\|^{2}+\left\|v_{k}^{\prime}(\tau)\right\|^{2}\right) d \tau \\
& \quad+\int_{0}^{t} \int_{\Omega}\left|u_{k} v_{k}\right|^{2(r+1)}\left(\left|u_{k}\right|^{2}+\left|v_{k}\right|^{2}\right) d x d \tau \\
& \leq \int_{0}^{t}\left(\left\|u_{k}^{\prime}(\tau)\right\|^{2}+\left\|v_{k}^{\prime}(\tau)\right\|^{2}\right) d \tau \\
& \quad+C \int_{0}^{t}\left(\left\|u_{k}\right\|_{2(r+2)}^{2(r+2)}+\left\|v_{k}\right\|_{2(r+2)}^{2(r+2)}\right) d \tau  \tag{23}\\
& \leq C \int_{0}^{t}\left(\left\|u_{k}^{\prime}(\tau)\right\|^{2}+\left\|v_{k}^{\prime}(\tau)\right\|^{2}\right. \\
& \left.\quad+\left\|\nabla u_{k}\right\|_{p}^{2(r+2)}+\left\|\nabla v_{k}\right\|_{p}^{2(r+2)}\right) d \tau \\
& \leq C \int_{0}^{t}\left(\left\|u_{k}^{\prime}(\tau)\right\|^{2}+\left\|v_{k}^{\prime}(\tau)\right\|^{2}\right. \\
& \left.\quad+\left\|\nabla u_{k}\right\|_{p}^{p}+\left\|\nabla v_{k}\right\|_{p}^{p}\right)^{2(r+2) / p} d \tau .
\end{align*}
$$

It follows from (22) and (23) that

$$
\begin{align*}
& \left\|u_{k}^{\prime}(t)\right\|^{2}+\left\|v_{k}^{\prime}(t)\right\|^{2}+\left\|\nabla u_{k}\right\|_{p}^{p}+\left\|\nabla v_{k}\right\|_{p}^{p} \\
& +2 \int_{0}^{t}\left(\left\|u_{k}^{\prime}(\tau)\right\|_{q}^{q}\|+\| v_{k}^{\prime}(\tau) \|_{q}^{q}\right. \\
& \left.+\left\|\nabla u_{k}^{\prime}(t)\right\|^{2}+\left\|\nabla v_{k}^{\prime}(t)\right\|^{2}\right) d \tau  \tag{24}\\
& \leq 2 C_{0}+C \int_{0}^{t}\left(\left\|u_{k}^{\prime}(\tau)\right\|^{2}+\left\|v_{k}^{\prime}(\tau)\right\|^{2}\right. \\
& \left.\quad+\left\|\nabla u_{k}\right\|_{p}^{p}+\left\|\nabla v_{k}\right\|_{p}^{p}\right)^{2(r+2) / p} d \tau
\end{align*}
$$

which implies that

$$
\begin{align*}
&\left\|u_{k}^{\prime}(t)\right\|^{2}+\left\|v_{k}^{\prime}(t)\right\|^{2}+\left\|\nabla u_{k}\right\|_{p}^{p}+\left\|\nabla v_{k}\right\|_{p}^{p} \\
& \leq 2 C_{0}+C \int_{0}^{t}\left(\left\|u_{k}^{\prime}(\tau)\right\|^{2}+\left\|v_{k}^{\prime}(\tau)\right\|^{2}\right.  \tag{25}\\
&\left.+\left\|\nabla u_{k}\right\|_{p}^{p}+\left\|\nabla v_{k}\right\|_{p}^{p}\right)^{2(r+2) / p} d \tau .
\end{align*}
$$

We get from (25) and Gronwall type inequality that

$$
\begin{align*}
& \left\|u_{k}^{\prime}(t)\right\|^{2}+\left\|v_{k}^{\prime}(t)\right\|^{2}+\left\|\nabla u_{k}\right\|_{p}^{p}+\left\|\nabla v_{k}\right\|_{p}^{p} \\
& \quad \leq\left[2 C_{0}-\frac{2(r+2)-p}{p} C t\right]^{-p /(2(r+2)-p)} \tag{26}
\end{align*}
$$

Thus, we deduce from (26) that there exists a time $T>0$ such that

$$
\begin{equation*}
\left\|u_{k}^{\prime}(t)\right\|^{2}+\left\|v_{k}^{\prime}(t)\right\|^{2}+\left\|\nabla u_{k}\right\|_{p}^{p}+\left\|\nabla v_{k}\right\|_{p}^{p} \leq C_{1}, \quad \forall t \in[0, T] \tag{27}
\end{equation*}
$$

where $C_{1}$ is a positive constant independent of $k$.
We have from (24) and (26) that

$$
\begin{align*}
2 \int_{0}^{t} & \left(\left\|u_{k}^{\prime}(\tau)\right\|_{q}^{q}+\left\|v_{k}^{\prime}(\tau)\right\|_{q}^{q}\right. \\
& \left.\quad+\left\|\nabla u_{k}^{\prime}(\tau)\right\|^{2}+\left\|\nabla v_{k}^{\prime}(\tau)\right\|^{2}\right) d \tau \leq C_{2}, \quad \forall t \in[0, T] \tag{28}
\end{align*}
$$

It follows from (27) and (28) that

$$
\begin{gather*}
\left\|u_{k}^{\prime}(t)\right\|^{2} \leq C_{1}, \quad\left\|v_{k}^{\prime}(t)\right\|^{2} \leq C_{1} \\
\left\|\nabla u_{k}\right\|_{p}^{p} \leq C_{1}, \quad\left\|\nabla v_{k}\right\|_{p}^{p} \leq C_{1} . \tag{29}
\end{gather*}
$$

$u_{k}^{\prime}(t)$ and $v_{k}^{\prime}(t)$ are bounded in $L^{2}\left([0, T] ; L^{q}(\Omega)\right)$

$$
\text { and } L^{2}\left([0, T] ; H_{0}^{1}(\Omega)\right)
$$

Using the same process as the proof of Theorem 2.1 in paper [18], we derive that $[u(t), v(t)]$ is a local solution of the problem (1)-(5). By (20) and (21), we conclude that (11) is valid.

## 3. Global Existence

In order to state our main results, we first introduce the following functionals:

$$
\begin{gather*}
J([u, v])=\frac{1}{p}\left(\|\nabla u\|_{p}^{p}+\|\nabla v\|_{p}^{p}\right)-\frac{1}{r+2}\|u v\|_{r+2}^{r+2}  \tag{30}\\
K([u, v])=\left(\|\nabla u\|_{p}^{p}+\|\nabla v\|_{p}^{p}\right)-2\|u v\|_{r+2}^{r+2} \tag{31}
\end{gather*}
$$

for $[u, v] \in W_{0}^{1, p}(\Omega) \times W_{0}^{1, p}(\Omega)$.
We put that

$$
\begin{gather*}
d=\inf \left\{\sup _{\lambda \geq 0} J(\lambda[u, v]):[u, v] \in W_{0}^{1, p}(\Omega)\right.  \tag{32}\\
\left.\times W_{0}^{1, p}(\Omega) /\{[0,0]\}\right\} .
\end{gather*}
$$

Then, we are able to define the stable set as follows for problem (1)-(5):

$$
\begin{gather*}
W=\left\{[u, v] \in W_{0}^{1, p}(\Omega) \times W_{0}^{1, p}(\Omega) \mid K([u, v])>0,\right. \\
J([u, v])<d\} \cup\{[0,0]\} . \tag{33}
\end{gather*}
$$

We denote the total energy related to (1) and (2) by (12), and

$$
\begin{align*}
E(0)= & \frac{1}{2}\left(\left\|u_{1}\right\|^{2}+\left\|v_{1}\right\|^{2}\right)+\frac{1}{p}\left(\left\|\nabla u_{0}\right\|_{p}^{p}+\left\|\nabla v_{0}\right\|_{p}^{p}\right)  \tag{34}\\
& -\frac{1}{r+2}\left\|u_{0} v_{0}\right\|_{r+2}^{r+2}
\end{align*}
$$

is the total energy of the initial data.

Lemma 5. Let $[u, v]$ be a solution to problem (1)-(5); then, $E(t)$ is a nonincreasing function for $t>0$ and

$$
\begin{equation*}
\frac{d}{d t} E(t)=-\left(\left\|u_{t}\right\|_{q}^{q}+\left\|v_{t}\right\|_{q}^{q}+\left\|\nabla u_{t}\right\|_{2}^{2}+\left\|\nabla v_{t}\right\|_{2}^{2}\right) \tag{35}
\end{equation*}
$$

We have from (11) that $E(t)$ is the primitive of an integrable function. Therefore, $E(t)$ is absolutely continuous, and equality (35) is satisfied.

Lemma 6. Supposed that $[u, v] \in W_{0}^{1, p}(\Omega) \times W_{0}^{1, p}(\Omega)$, and $p<2(r+2) \leq n p /(n-p)$ if $n \geq p ; p<2(r+2)<+\infty$ if $n<p$, then $d>0$.

Proof. Since

$$
\begin{equation*}
J(\lambda[u, v])=\frac{\lambda^{p}}{p}\left(\|\nabla u\|_{p}^{p}+\|\nabla v\|_{p}^{p}\right)-\frac{\lambda^{2(r+2)}}{r+2}\|u v\|_{r+2}^{r+2} \tag{36}
\end{equation*}
$$

so we get

$$
\begin{equation*}
\frac{d}{d \lambda} J(\lambda[u, v])=\lambda^{p-1}\left(\|\nabla u\|_{p}^{p}+\|\nabla v\|_{p}^{p}\right)-2 \lambda^{2 r+3}\|u v\|_{r+2}^{r+2} \tag{37}
\end{equation*}
$$

In case $u v \neq 0$, let $(d / d \lambda) J(\lambda[u, v])=0$, which implies that

$$
\begin{equation*}
\lambda_{1}=\left(\frac{\|\nabla u\|_{p}^{p}+\|\nabla v\|_{p}^{p}}{2\|u v\|_{r+2}^{r+2}}\right)^{1 /(2 r-p+4)} \tag{38}
\end{equation*}
$$

As $\lambda=\lambda_{1}$, an elementary calculation shows that $\left.\left(d^{2} / d \lambda^{2}\right) J(\lambda[u, v])\right|_{\lambda=\lambda_{1}}<0$. Therefore, we have that

$$
\begin{align*}
& \sup _{\lambda \geq 0} J(\lambda[u, v]) \\
& \quad=J\left(\lambda_{1}[u, v]\right)  \tag{39}\\
& \quad=\frac{2(r+2)-p}{2 p(r+2)}\left(\frac{\|\nabla u\|_{p}^{p}+\|\nabla v\|_{p}^{p}}{2^{p /(2 r+4)}\|u v\|_{r+2}^{p / 2}}\right)^{(2 r+4) /(2 r-p+4)} .
\end{align*}
$$

It follows from Hölder inequality and Lemma 1 that

$$
\begin{align*}
\|u v\|_{r+2}^{p / 2} & \leq\|u\|_{2(r+2)}^{p / 2}\|v\|_{2(r+2)}^{p / 2} \\
& \leq \frac{1}{2}\left(\|u\|_{2(r+2)}^{p}+\|v\|_{2(r+2)}^{p}\right)  \tag{40}\\
& \leq C\left(\|\nabla u\|_{p}^{p}+\|\nabla v\|_{p}^{p}\right) .
\end{align*}
$$

We get from (39) and (40) that
$\sup _{\lambda \geq 0} J(\lambda[u, v]) \geq \frac{2(r+2)-p}{2 p(r+2)}\left(2^{p /(2 r+4)} C\right)^{-(2 r+4) /(2 r-p+4)}>0$.

In case $u v=0$ and $u=0$ or $v=0$, then

$$
\begin{equation*}
J(\lambda[u, v])=\frac{\lambda^{p}}{p}\left(\|\nabla u\|_{p}^{p}+\|\nabla v\|_{p}^{p}\right) \tag{42}
\end{equation*}
$$

Therefore, we have

$$
\begin{equation*}
J(\lambda[u, v])=+\infty . \tag{43}
\end{equation*}
$$

We conclude from (41) and (43) that

$$
\begin{equation*}
d \geq \frac{2(r+2)-p}{2 p(r+2)}\left(2^{p /(2 r+4)} C\right)^{-(2 r+4) /(2 r-p+4)}>0 \tag{44}
\end{equation*}
$$

Thus, we complete the proof of Lemma 6.
Lemma 7. Supposed that $p<2(r+2) \leq n p /(n-p)$ for $n \geq p$ and $p<2(r+2)<+\infty$ for $n<p$, if $\left[u_{0}, v_{0}\right] \in W,\left[u_{1}, v_{1}\right] \in$ $L^{2}(\Omega) \times L^{2}(\Omega)$ and $E(0)<d$, then $[u, v] \in W$ for $\forall t \in[0, T)$.

Proof. Assume that there exists a number $t^{*} \in(0, T)$ such that $[u(t), v(t)] \in W$ on $\left[0, t^{*}\right)$ and $u\left(t^{*}\right) \notin W$. Then, in virtue of the continuity of $u(t)$, we see $u\left(t^{*}\right) \in \partial W$, where $\partial W$ denotes the boundary of domain $W$. From the definition of $W$ and the continuity of $J([u(t), v(t)])$ and $K([u(t), v(t)])$ in $t$, we have either

$$
\begin{equation*}
J\left(\left[u\left(t^{*}\right), v\left(t^{*}\right)\right]\right)=d \tag{45}
\end{equation*}
$$

or

$$
\begin{equation*}
K\left(\left[u\left(t^{*}\right), v\left(t^{*}\right)\right]\right)=0 . \tag{46}
\end{equation*}
$$

It follows from (12) and (30) that

$$
\begin{equation*}
J\left(\left[u\left(t^{*}\right), v\left(t^{*}\right)\right]\right) \leq E\left(t^{*}\right) \leq E(0)<d . \tag{47}
\end{equation*}
$$

So, case (45) is impossible.
Assume that (46) holds; then, we get that

$$
\begin{align*}
& \frac{d}{d \lambda} J\left(\lambda\left[u\left(t^{*}\right), v\left(t^{*}\right)\right]\right)  \tag{48}\\
& \quad=\lambda^{p-1}\left(1-\lambda^{2 r-p+4}\right)\left(\|\nabla u\|_{p}^{p}+\|\nabla v\|_{p}^{p}\right) .
\end{align*}
$$

We obtain from $(d / d \lambda) J\left(\lambda\left[u\left(t^{*}\right), v\left(t^{*}\right)\right]\right)=0$ that $\lambda=1$. Since

$$
\begin{align*}
& \left.\frac{d^{2}}{d \lambda^{2}} J\left(\lambda\left[u\left(t^{*}\right), v\left(t^{*}\right)\right]\right)\right|_{\lambda=1}  \tag{49}\\
& \quad=-[(2(r+2)-p)+(2 r+3)]<0
\end{align*}
$$

Consequently, we get from (47) that

$$
\begin{equation*}
\sup _{\lambda \geq 0} J\left(\lambda\left[u\left(t^{*}\right), v\left(t^{*}\right)\right]\right)=J\left(\left[u\left(t^{*}\right), v\left(t^{*}\right)\right]\right)<d \tag{50}
\end{equation*}
$$

which contradicts the definition of $d$. Hence, case (46) is impossible as well. Thus we conclude that $[u(t), v(t)] \in W$ on $[0, T)$.

Theorem 8 (global solution). Supposed that $p<2(r+2) \leq$ $n p /(n-p)$ as $n \geq p$ and $p<2(r+2)<+\infty$ as $n<p$, and $[u(t), v(t)]$ is a local solution of problem (1)-(5) on $[0, T)$. If $\left[u_{0}, v_{0}\right] \in W,\left[u_{1}, v_{1}\right] \in L^{2}(\Omega) \times L^{2}(\Omega)$ and $E(0)<d$, then $[u(t), v(t)]$ is a global solution of problem (1)-(5).

Proof. It suffices to show that $\left\|u_{t}\right\|^{2}+\left\|v_{t}\right\|^{2}+\|\nabla u\|_{p}^{p}+\|\nabla v\|_{p}^{p}$ is bounded uniformly with respect to $t$. Under the hypotheses in Theorem 8, we get from Lemma 7 that $[u, v] \in W$ on $[0, T)$. So the following formula holds on $[0, T)$ :

$$
\begin{align*}
J([u, v]) & =\frac{1}{p}\left(\|\nabla u\|_{p}^{p}+\|\nabla v\|_{p}^{p}\right)-\frac{1}{r+2}\|u v\|_{r+2}^{r+2} \\
& \geq \frac{2(r+2)-p}{2 p(r+2)}\left(\|\nabla u\|_{p}^{p}+\|\nabla v\|_{p}^{p}\right) . \tag{51}
\end{align*}
$$

We have from (51) that

$$
\begin{align*}
& \frac{1}{2}\left(\left\|u_{t}\right\|^{2}+\left\|v_{t}\right\|^{2}\right)+\frac{2(r+2)-p}{2 p(r+2)}\left(\|\nabla u\|_{p}^{p}+\|\nabla v\|_{p}^{p}\right) \\
& \quad \leq \frac{1}{2}\left(\left\|u_{t}\right\|^{2}+\left\|v_{t}\right\|^{2}\right)+J([u(t), v(t)])  \tag{52}\\
& \quad=E(t) \leq E(0)<d
\end{align*}
$$

Hence, we get

$$
\begin{align*}
& \left(\left\|u_{t}\right\|^{2}+\left\|v_{t}\right\|^{2}\right)+\left(\|\nabla u\|_{p}^{p}+\|\nabla v\|_{p}^{p}\right) \\
& \quad \leq \max \left(2, \frac{2 p(r+2)}{2(r+2)-p}\right) d<+\infty \tag{53}
\end{align*}
$$

The above inequality and the continuation principle lead to the global existence of the solution $[u, v]$ for problem (1)-(5).

## 4. Asymptotic Behavior of Global Solutions

The following lemma plays an important role in studying the decay estimate of global solutions for the problem (1)-(5).

Lemma 9 (see [9]). Suppose that $\varphi(t)$ is a nonincreasing nonnegative function on $[0,+\infty)$ and satisfies

$$
\begin{equation*}
\varphi(t)^{r+1} \leq k(\varphi(t)-\varphi(t+1)), \quad \forall t \geq 0 \tag{54}
\end{equation*}
$$

Then, $\varphi(t)$ has the decay property

$$
\begin{equation*}
\varphi(t) \leq\left[\frac{r}{k}(t-1)+M^{-r}\right]^{-1 / r}, \quad \forall t \geq 1 \tag{55}
\end{equation*}
$$

where $k, r>0$ are constants and $M=\max _{t \in[0,1]} \varphi(t)$.
Lemma 10. Under the assumptions of Theorem 8, if initial value $\left[u_{0}, v_{0}\right] \in W$ and $\left[u_{1}, v_{1}\right] \in L^{2}(\Omega) \times L^{2}(\Omega)$ are sufficiently small such that

$$
\begin{equation*}
C^{2(r+2)}\left(\frac{2 p(r+2)}{2 p(r+2)-p} E(0)\right)^{(2(r+2)-p) / p}<1, \tag{56}
\end{equation*}
$$

then

$$
\begin{equation*}
\left(\|\nabla u\|_{p}^{p}+\|\nabla v\|_{p}^{p}\right) \leq \frac{1}{\theta} K([u, v]), \tag{57}
\end{equation*}
$$

where $\theta=1-C^{2(r+2)}((2 p(r+2) /(2 p(r+2)-$ $p)) E(0))^{(2(r+2)-p) / p}>0$ is a positive constant and $C$ is the optimal Sobolev's constant from $W_{0}^{1, p}(\Omega)$ to $L^{2(r+2)}(\Omega)$.

Proof. We have from Lemma 1 and (52) that

$$
\begin{align*}
2\|u v\|_{r+2}^{r+2} \leq & 2\|u\|_{2(r+2)}^{r+2}\|v\|_{2(r+2)}^{r+2} \\
\leq & \|u\|_{2(r+2)}^{2(r+2)}+\|v\|_{2(r+2)}^{2(r+2)} \\
\leq & C^{2(r+2)}\left(\|\nabla u\|_{p}^{2(r+2)}+\|\nabla v\|_{p}^{2(r+2)}\right) \\
\leq & C^{2(r+2)}\left(\|\nabla u\|_{p}^{2(r+2)-p}\|\nabla u\|_{p}^{p}\right.  \tag{58}\\
& \left.\quad+\|\nabla v\|_{p}^{2(r+2)-p}\|\nabla v\|_{p}^{p}\right) \\
\leq & C^{2(r+2)}\left(\frac{2 p(r+2)}{2 p(r+2)-p} E(0)\right)^{(2(r+2)-p) / p} \\
& \times\left(\|\nabla u\|_{p}^{p}+\|\nabla v\|_{p}^{p}\right)
\end{align*}
$$

Therefore, we get from (58) and (31) that

$$
\begin{align*}
& {\left[1-C^{2(r+2)}\left(\frac{2 p(r+2)}{2 p(r+2)-p} E(0)\right)^{(2(r+2)-p) / p}\right]}  \tag{59}\\
& \quad \times\left(\|\nabla u\|_{p}^{p}+\|\nabla v\|_{p}^{p}\right) \leq K([u, v])
\end{align*}
$$

Let

$$
\begin{equation*}
\theta=1-C^{2(r+2)}\left(\frac{2 p(r+2)}{2 p(r+2)-p} E(0)\right)^{(2(r+2)-p) / p}>0 \tag{60}
\end{equation*}
$$

then, we have from (59) that

$$
\begin{equation*}
\|\nabla u\|_{p}^{p}+\|\nabla v\|_{p}^{p} \leq \frac{1}{\theta} K([u, v]) \tag{61}
\end{equation*}
$$

Theorem 11. Under the assumptions of Theorem 8, if $p<q<$ $r+2$ and (56) hold, then the global solution $[u, v]$ in $W$ of the problem (1)-(5) has the following decay property:

$$
\begin{equation*}
E(t) \leq\left[\frac{p-2}{p C}(t-1)+M^{(p+q-p q) / p}\right]^{p /(p+q-p q)}, \quad \forall t>1, \tag{62}
\end{equation*}
$$

where $M=\max _{t \in[0,1]} E(t)>0$ is some constant depending only on $\left[u_{0}, v_{0}\right]$ and $\left[u_{1}, v_{1}\right]$.

Proof. Multiplying (1) by $u_{t}$ and (2) by $v_{t}$ and integrating over $\Omega \times[t, t+1]$, and summing up together, we get

$$
\begin{align*}
& \int_{t}^{t+1}\left(\left\|u_{t}(s)\right\|_{q}^{q}+\left\|v_{t}(s)\right\|_{q}^{q}+\left\|\nabla u_{t}(s)\right\|_{2}^{2}\right.  \tag{63}\\
&\left.\quad+\left\|\nabla v_{t}(s)\right\|_{2}^{2}\right) d s=E(t)-E(t+1) .
\end{align*}
$$

Thus, there exists $t_{1} \in[t, t+1 / 4], t_{2} \in[t+3 / 4, t+1]$ such that

$$
\begin{align*}
& 4\left(\left\|u_{t}\left(t_{i}\right)\right\|_{q}^{q}+\left\|v_{t}\left(t_{i}\right)\right\|_{q}^{q}+\left\|\nabla u_{t}\left(t_{i}\right)\right\|_{2}^{2}+\left\|\nabla v_{t}\left(t_{i}\right)\right\|_{2}^{2}\right)  \tag{64}\\
& \quad=E(t)-E(t+1), \quad i=1,2
\end{align*}
$$

On the other hand, we multiply (1) by $u$ and (2) by $v$ and integrate over $\Omega \times\left[t_{1}, t_{2}\right]$. Adding them together, we obtain

$$
\begin{align*}
\int_{t_{1}}^{t_{2}} K([u, v]) d s= & \int_{t_{1}}^{t_{2}}\left\|u_{t}\right\|^{2} d s+\int_{t_{1}}^{t_{2}}\left\|v_{t}\right\|^{2} d s \\
& +\left(u_{t}\left(t_{1}\right), u\left(t_{1}\right)\right)-\left(u_{t}\left(t_{2}\right), u\left(t_{2}\right)\right) \\
& +\left(v_{t}\left(t_{1}\right), v\left(t_{2}\right)\right)-\left(v_{t}\left(t_{2}\right) v\left(t_{2}\right)\right) \\
& -\left(\int_{t_{1}}^{t_{2}} \int_{\Omega}\left|u_{t}\right|^{q-2} u_{t} u d x d s\right. \\
& \left.+\int_{t_{1}}^{t_{2}} \int_{\Omega}\left|v_{t}\right|^{q-2} v_{t} v d x d s\right) \\
& -\int_{t_{1}}^{t_{2}} \int_{\Omega} \nabla u_{t} \nabla u d x d s-\int_{t_{1}}^{t_{2}} \int_{\Omega} \nabla v_{t} \nabla v d x d s . \tag{65}
\end{align*}
$$

From (63), Sobolev inequality, and Hölder inequality, we have

$$
\begin{align*}
& \int_{t_{1}}^{t_{2}}\left\|u_{t}\right\|^{2} d s \leq C \int_{t_{1}}^{t_{2}}\left\|\nabla u_{t}\right\|^{2} d s \leq C(E(t)-E(t+1)) \\
& \int_{t_{1}}^{t_{2}}\left\|v_{t}\right\|^{2} d s \leq C \int_{t_{1}}^{t_{2}}\left\|\nabla v_{t}\right\|^{2} d s \leq C(E(t)-E(t+1)) \tag{66}
\end{align*}
$$

We get from (52), (64), and Lemmas 1 and 2 that

$$
\begin{align*}
\left|u_{t}\left(t_{i}\right), u\left(t_{i}\right)\right| \leq & \left\|u_{t}\left(t_{i}\right)\right\| \cdot\left\|u\left(t_{i}\right)\right\| \leq C\left\|\nabla u_{t}\left(t_{i}\right)\right\| \cdot\left\|\nabla u\left(t_{i}\right)\right\|_{p} \\
\leq & C(E(t)-E(t+1))^{1 / 2} \sup _{t \leq s \leq t+1} E(s)^{1 / p} \\
\leq & C(\varepsilon)(E(t)-E(t+1))^{p / 2(p-1)} \\
& +\varepsilon \sup _{t \leq s \leq t+1} E(s), \quad i=1,2, \\
\left|\left(v_{t}\left(t_{i}\right), v\left(t_{i}\right)\right)\right| \leq & \left\|v_{t}\left(t_{i}\right)\right\| \cdot\left\|v\left(t_{i}\right)\right\| \leq C\left\|\nabla v_{t}\left(t_{i}\right)\right\| \cdot\left\|\nabla v\left(t_{i}\right)\right\|_{p} \\
\leq & C(E(t)-E(t+1))^{1 / 2} \sup _{t \leq s \leq t+1} E(s)^{1 / p} \\
\leq & C(\varepsilon)(E(t)-E(t+1))^{p / 2(p-1)} \\
& +\varepsilon \sup _{t \leq s \leq t+1} E(s), \quad i=1,2 . \tag{67}
\end{align*}
$$

From Hölder inequality and Lemma 2, we get

$$
\begin{align*}
\left.\left|\int_{t_{1}}^{t_{2}} \int_{\Omega}\right| u_{t}\right|^{q-2} u_{t} u d x d s \mid & \leq \int_{t_{1}}^{t_{2}}\left\|u_{t}\right\|_{q}^{q-1}\|u\|_{q} d s \\
& \leq\left(\int_{t_{1}}^{t_{2}}\left\|u_{t}\right\|_{q}^{q} d s\right)^{(q-1) / q}\left(\int_{t_{1}}^{t_{2}}\|u\|_{q}^{q} d s\right)^{1 / q} \\
& \leq C(\varepsilon) \int_{t_{1}}^{t_{2}}\left\|u_{t}\right\|_{q}^{q} d s+\varepsilon \int_{t_{1}}^{t_{2}}\|u\|_{q}^{q} d s, \tag{68}
\end{align*}
$$

$$
\begin{align*}
\left.\left|\int_{t_{1}}^{t_{2}} \int_{\Omega}\right| v_{t}\right|^{q-2} v_{t} v d x d s \mid & \leq \int_{t_{1}}^{t_{2}}\left\|v_{t}\right\|_{q}^{q-1}\|v\|_{q} d s \\
& \leq\left(\int_{t_{1}}^{t_{2}}\left\|v_{t}\right\|_{q}^{q} d s\right)^{(q-1) / q}\left(\int_{t_{1}}^{t_{2}}\|v\|_{q}^{q} d s\right)^{1 / q} \\
& \leq C(\varepsilon) \int_{t_{1}}^{t_{2}}\left\|v_{t}\right\|_{q}^{q} d s+\varepsilon \int_{t_{1}}^{t_{2}}\|v\|_{q}^{q} d s . \tag{69}
\end{align*}
$$

Since $p<q<r+2$ and the property of the function $f(x)=\alpha^{x} / x, \alpha \geq 0, x>0$, we obtain

$$
\begin{equation*}
\frac{\|u\|_{q}^{q}}{q} \leq C \frac{\|u\|_{p}^{p}}{p}+C \frac{\|u\|_{r+2}^{r+2}}{r+2}, \quad \frac{\|v\|_{q}^{q}}{q} \leq C \frac{\|v\|_{p}^{p}}{p}+C \frac{\|v\|_{r+2}^{r+2}}{r+2} . \tag{70}
\end{equation*}
$$

We conclude from (69), (70), $[u, v] \in W$, and Lemma 1 that

$$
\begin{align*}
\|u\|_{q}^{q}+\|v\|_{q}^{q} & \leq C\left(\|u\|_{p}^{p}+\|u\|_{r+2}^{r+2}+\|v\|_{p}^{p}+\|v\|_{r+2}^{r+2}\right) \\
& \leq C\left(\|u\|_{p}^{p}+\|\nabla u\|_{p}^{p}+\|v\|_{p}^{p}+\|\nabla v\|_{p}^{p}\right)  \tag{71}\\
& \leq C\left(\|\nabla u\|_{p}^{p}+\|\nabla v\|_{p}^{p}\right) \leq C E(t) .
\end{align*}
$$

It follows from (63), (68), (69), and (71) that

$$
\begin{align*}
& \left|-\left(\int_{t_{1}}^{t_{2}} \int_{\Omega}\left|u_{t}\right|^{q-2} u_{t} u d x d s+\int_{t_{1}}^{t_{2}} \int_{\Omega}\left|v_{t}\right|^{q-2} v_{t} v d x d s\right)\right|  \tag{72}\\
& \quad \leq C(\varepsilon)(E(t)-E(t+1))+\varepsilon C \int_{t_{1}}^{t_{2}} E(s) d s
\end{align*}
$$

and we obtain from (63), Sobolev inequality, Hölder inequality, and Lemma 2 that

$$
\begin{align*}
\left|-\int_{t_{1}}^{t_{2}} \int_{\Omega} \nabla u_{t} \nabla u d s\right| \leq & \int_{t_{1}}^{t_{2}}\left\|\nabla u_{t}\right\| \cdot\|\nabla u\| d s \\
\leq & \left(\int_{t_{1}}^{t_{2}}\left\|\nabla u_{t}\right\|^{2} d s\right)^{1 / 2}\left(\int_{t_{1}}^{t_{2}}\|\nabla u\|^{2} d s\right)^{1 / 2} \\
\leq & C(E(t)-E(t+1))^{1 / 2}\left(\int_{t_{1}}^{t_{2}}\|\nabla u\|_{p}^{2} d s\right)^{1 / 2} \\
\leq & C(E(t)-E(t+1))^{1 / 2}\left(\int_{t_{1}}^{t_{2}}\|\nabla u\|_{p}^{p} d s\right)^{1 / p} \\
\leq & C(E(t)-E(t+1))^{p / 2(p-1)} \\
& +\varepsilon \int_{t_{1}}^{t_{2}}\|\nabla u\|_{p}^{p} d s . \tag{73}
\end{align*}
$$

Similarly, we have the following formula:

$$
\begin{aligned}
\left|-\int_{t_{1}}^{t_{2}} \int_{\Omega} \nabla v_{t} \nabla v d s\right| & \leq \int_{t_{1}}^{t_{2}}\left\|\nabla v_{t}\right\| \cdot\|\nabla v\| d s \\
& \leq\left(\int_{t_{1}}^{t_{2}}\left\|\nabla v_{t}\right\|^{2} d s\right)^{1 / 2}\left(\int_{t_{1}}^{t_{2}}\|\nabla v\|^{2} d s\right)^{1 / 2}
\end{aligned}
$$

$$
\begin{align*}
\leq & C(E(t)-E(t+1))^{1 / 2}\left(\int_{t_{1}}^{t_{2}}\|\nabla v\|_{p}^{2} d s\right)^{1 / 2} \\
\leq & C(E(t)-E(t+1))^{1 / 2}\left(\int_{t_{1}}^{t_{2}}\|\nabla v\|_{p}^{p} d s\right)^{1 / p} \\
\leq & C(E(t)-E(t+1))^{p / 2(p-1)} \\
& +\varepsilon \int_{t_{1}}^{t_{2}}\|\nabla v\|_{p}^{p} d s . \tag{74}
\end{align*}
$$

We get from (57), (73), and (74) that

$$
\begin{align*}
& \left|\int_{t_{1}}^{t_{2}} \int_{\Omega}\left(\nabla u_{t} \nabla u+\nabla v_{t} \nabla v\right) d s\right| \\
& \leq C(E(t)-E(t+1))^{p / 2(p-1)}+\varepsilon \int_{t_{1}}^{t_{2}}\left(\|\nabla u\|_{p}^{p}+\|\nabla v\|_{p}^{p}\right) d s \\
& \leq C(E(t)-E(t+1))^{p / 2(p-1)}+\frac{\varepsilon}{\theta} \int_{t_{1}}^{t_{2}} K([u, v]) d s \tag{75}
\end{align*}
$$

Choosing small enough $\varepsilon$, we have from (65), (66), (67), (72), and (75) that

$$
\begin{align*}
& \int_{t_{1}}^{t_{2}} K([u, v]) d s \leq C[(E(t)-E(t+1)) \\
&\left.+(E(t)-E(t+1))^{p / 2(p-1)}\right]  \tag{76}\\
&+\varepsilon \sup _{t \leq s \leq t+1} E(s)+\varepsilon \int_{t_{1}}^{t_{2}} E(s) d s
\end{align*}
$$

It follows from (30) and (31) that

$$
\begin{align*}
J([u, v])= & \frac{2(r+2)-p}{2 p(r+2)}\left(\|\nabla u\|_{p}^{p}+\|\nabla v\|_{p}^{p}\right)  \tag{77}\\
& +\frac{1}{2(r+2)} K([u, v]) .
\end{align*}
$$

On the other hand, from (12) and using (57) and (77), we deduce that

$$
\begin{aligned}
E(t)= & \frac{1}{2}\left(\left\|u_{t}\right\|^{2}+\left\|v_{t}\right\|^{2}\right)+J([u, v]) \\
= & \frac{1}{2}\left(\left\|u_{t}\right\|^{2}+\left\|v_{t}\right\|^{2}\right)+\frac{2(r+2)-p}{2 p(r+2)} \\
& \times\left(\|\nabla u\|_{p}^{p}+\|\nabla v\|_{p}^{p}\right)+\frac{1}{2(r+2)} K([u, v]) \\
\leq & \frac{1}{2}\left(\left\|u_{t}\right\|^{2}+\left\|v_{t}\right\|^{2}\right)+\left(\frac{2(r+2)-p}{2 \theta p(r+2)}+\frac{1}{2(r+2)}\right) \\
& \times K([u, v]) .
\end{aligned}
$$

By integrating (78) over [ $t_{1}, t_{2}$ ], we obtain

$$
\begin{align*}
\int_{t_{1}}^{t_{2}} E(s) d s \leq & \frac{1}{2} \int_{t_{1}}^{t_{2}}\left(\left\|u_{t}\right\|^{2}+\left\|v_{t}\right\|^{2}\right) d s \\
& +\left(\frac{2(r+2)-p}{2 \theta p(r+2)}+\frac{1}{2(r+2)}\right) \int_{t_{1}}^{t_{2}} K([u, v]) d s . \tag{79}
\end{align*}
$$

For small enough $\varepsilon$, we have from (76) and (79) that

$$
\begin{array}{rl}
\int_{t_{1}}^{t_{2}} & E(s) d s \\
\leq & C\left[(E(t)-E(t+1))+(E(t)-E(t+1))^{p /(2(p-1))}\right] \\
& +\varepsilon \sup _{t \leq s \leq t+1} E(s) \tag{80}
\end{array}
$$

Thus, there exists $t^{*} \in\left[t_{1}, t_{2}\right]$, such that

$$
\begin{align*}
E\left(t^{*}\right) \leq & C\left[(E(t)-E(t+1))+(E(t)-E(t+1))^{p / 2(p-1)}\right] \\
& +\varepsilon \sup _{t \leq s \leq t+1} E(s) . \tag{81}
\end{align*}
$$

Multiplying (1) by $u_{t}$ and (2) by $v_{t}$ and integrating over $\Omega \times$ [ $\left.t^{*}, t_{2}\right]$, and summing up, we get

$$
\begin{equation*}
E\left(t_{2}\right)=E\left(t^{*}\right)-\int_{t^{*}}^{t_{2}}\left(\left\|u_{t}\right\|_{q}^{q}+\left\|v_{t}\right\|_{q}^{q}+\left\|\nabla u_{t}\right\|^{2}+\left\|\nabla v_{t}\right\|^{2}\right) d s \tag{82}
\end{equation*}
$$

Therefore, we obtain from (63), (81), and (82) that

$$
\begin{align*}
\sup _{t \leq s \leq t+1} E(s) \leq C[ & (E(t)-E(t+1)) \\
& \left.+(E(t)-E(t+1))^{p / 2(p-1)}\right]+\varepsilon \sup _{t \leq s \leq t+1} E(s) . \tag{83}
\end{align*}
$$

Choosing small enough $\varepsilon$, we have from (83) that

$$
\begin{align*}
\sup _{t \leq s \leq t+1} E(s) \leq C[ & (E(t)-E(t+1))  \tag{84}\\
& \left.+(E(t)-E(t+1))^{p / 2(p-1)}\right] .
\end{align*}
$$

Since $p>2$ and $E(t)<E(0)$, we get

$$
\begin{equation*}
\sup _{t \leq s \leq t+1} E(s) \leq C(E(t)-E(t+1))^{p / 2(p-1)} \tag{85}
\end{equation*}
$$

Consequently,

$$
\begin{equation*}
\sup _{t \leq s \leq t+1} E(s)^{(2(p-1)) / p} \leq C(E(t)-E(t+1)) \tag{86}
\end{equation*}
$$

Thus, applying Lemma 9 to (86), we get

$$
\begin{equation*}
E(t) \leq\left[\frac{p-2}{p C}(t-1)+M^{(p-2) / p}\right]^{p /(2-p)}, \quad \forall t>1 \tag{87}
\end{equation*}
$$

where $M=\max _{t \in[0,1]} E(t)>0$ is some constant depending only on $\left[u_{0}, v_{0}\right.$ ] and $\left[u_{1}, v_{1}\right]$.

## Acknowledgments

This research was supported by the National Natural Science Foundation of China (no. 61273016), The Natural Science Foundation of Zhejiang Province (no. Y6100016), The Middle-aged and Young Leader in Zhejiang University of Science and Technology (2008-2012), and the Interdisciplinary Pre-research Project of Zhejiang University of Science and Technology (2010-2012).

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# A Class of Spectral Element Methods and Its A Priori/A Posteriori Error Estimates for 2nd-Order Elliptic Eigenvalue Problems 

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Received 24 May 2013; Accepted 1 September 2013
Academic Editor: Rasajit Bera
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#### Abstract

This paper discusses spectral and spectral element methods with Legendre-Gauss-Lobatto nodal basis for general 2nd-order elliptic eigenvalue problems. The special work of this paper is as follows. (1) We prove a priori and a posteriori error estimates for spectral and spectral element methods. (2) We compare between spectral methods, spectral element methods, finite element methods and their derived $p$-version, $h$-version, and $h p$-version methods from accuracy, degree of freedom, and stability and verify that spectral methods and spectral element methods are highly efficient computational methods.


## 1. Introduction

As we know, finite element methods are local numerical methods for partial differential equations and particularly well suitable for problems in complex geometries, whereas spectral methods can provide a superior accuracy, at the expense of domain flexibility. Spectral element methods combine the advantages of the above methods (see [1]). So far, spectral and spectral element methods are widely applied to boundary value problems (see $[1,2]$ ), as well as applied to symmetric eigenvalue problems (see [3]). However, it is still a new subject to apply them to nonsymmetric elliptic eigenvalue problems.

A posteriorii error estimates and highly efficient computational methods for finite elements of eigenvalue problems are the subjects focused on by the academia these years; see [3-16], and among them, for nonsymmetric 2nd-order elliptic eigenvalue problems, $[5,15$ ] provide a posteriori error estimates and adaptive algorithms, [9] the function value recovery techniques and $[8,10]$ two-level discretization schemes.

Based on the work mentioned above, this paper shall further apply spectral and spectral element methods to
nonsymmetric elliptic eigenvalue problems. This paper will mainly perform the following work.
(1) We prove a priori and a posteriori error estimates of spectral and spectral element methods with Legendre-Gauss-Lobatto nodal basis, respectively, for the general 2nd-order elliptic eigenvalue problems.
(2) We compare between spectral methods, spectral element methods with Legendre-Gauss-Lobatto nodal basis, finite element methods, and their derived $p$ version, $h$-version, and $h p$-version methods from accuracy, degree of freedom, and stability and verify that spectral methods and spectral element methods are highly efficient computational methods for nonsymmetric 2 nd-order elliptic eigenvalue problems.

This paper is organized as follows. Section 2 introduces basic knowledge of second elliptic eigenvalue problems. Sections 3 and 4 are devoted to a priori and a posteriori error estimates of spectral and spectral element methods, respectively. In Section 5, some numerical experiments are performed by the methods mentioned above.

## 2. Preliminaries

Consider the 2nd-order elliptic boundary value problem

$$
\begin{gather*}
L w=-\nabla \cdot(D \nabla w)+\mathbf{b} \cdot \nabla w+c w=f, \quad \text { in } \Omega \\
w=0, \quad \text { on } \partial \Omega \tag{1}
\end{gather*}
$$

where $\Omega \subset R^{d}(d=2,3)$ is a bounded domain, $\mathbf{b}$ and $c$ are a real-valued vector function and a real-valued function, respectively, and $D$ is a positive scalar function with $D(x) \geq$ $D_{0}>0(\forall x \in \Omega)$.

We denote the complex Sobolev spaces with norm $\|\cdot\|_{m}$ by $H^{m}(\Omega), H_{0}^{1}(\Omega)=\left\{v \in H^{1}(\Omega),\left.v\right|_{\partial \Omega}=0\right\}$. Let $(\cdot, \cdot)$ and $\|\cdot\|_{0, \Omega}$ be a inner product and a norm in the complex space $L^{2}(\Omega)$, respectively.

In this paper, $C$ denotes a generic positive constant independent of the polynomial degrees and mesh scales, which may not be the same at different occurrences.

Define the bilinear form $a(\cdot, \cdot)$ as follows:

$$
\begin{align*}
a(w, v)=\int_{\Omega} & D \nabla w \nabla \bar{v}+(\mathbf{b} \cdot \nabla w) \bar{v}  \tag{2}\\
& +c w \bar{v}, \quad \forall w, v \in H_{0}^{1}(\Omega) .
\end{align*}
$$

We assume that $f \in L^{2}(\Omega), D, \mathbf{b}$, and $c$ are bounded functions on $\Omega$, namely $D, c \in L^{\infty}(\Omega), \mathbf{b} \in\left(L^{\infty}(\Omega)\right)^{d}$. Further more, we assume that $\nabla \cdot \mathbf{b}$ exists and satisfies

$$
\begin{equation*}
-\frac{1}{2} \nabla \cdot \mathbf{b}+c \geq 0, \quad \text { in } \Omega \tag{3}
\end{equation*}
$$

Under these assumptions, the bilinear form $a(\cdot, \cdot)$ is continuous in $H_{0}^{1}(\Omega)$ and $H_{0}^{1}(\Omega)$-elliptic; that is, there exist two constants $A, B>0$ independent of $w, v$ such that

$$
\begin{gather*}
|a(w, v)| \leq A\|w\|_{1, \Omega}\|v\|_{1, \Omega}, \quad \forall w, v \in H_{0}^{1}(\Omega) \\
\operatorname{Re} a(v, v) \geq B\|v\|_{1, \Omega}^{2}, \quad \forall v \in H_{0}^{1}(\Omega) \tag{4}
\end{gather*}
$$

The corresponding variational formulation of (1) is given as follows: find $w \in H_{0}^{1}(\Omega)$, such that

$$
\begin{equation*}
a(w, v)=(f, v), \quad \forall v \in H_{0}^{1}(\Omega) \tag{5}
\end{equation*}
$$

The adjoint problem of (5) is as follows: find $w^{*} \in H_{0}^{1}(\Omega)$, such that

$$
\begin{equation*}
a\left(v, w^{*}\right)=(v, f), \quad \forall v \in H_{0}^{1}(\Omega) \tag{6}
\end{equation*}
$$

As the general 2nd-order elliptic boundary value problems, we assume that the regularity estimates for problem (5) and its adjoint problem (6) hold, respectively. Namely

$$
\begin{array}{cc}
\|w\|_{r_{1}+1, \Omega} \leq C\|f\|_{0, \Omega}, & 0<r_{1} \leq 1 \\
\left\|w^{*}\right\|_{r_{2}+1, \Omega} \leq C\|f\|_{0, \Omega}, & 0<r_{2} \leq 1 \tag{8}
\end{array}
$$

We assume that $K_{h}=\{\kappa\}$ is a regular rectangle (resp. cuboid) or simplex partition of the domain $\Omega$ and satisfies
$\bar{\Omega}=\bigcup \bar{\kappa}$. We associate with the partition a polynomial degree vector $\mathbf{N}=\left\{N_{\kappa}\right\}$, where $N_{\kappa}$ is the polynomial degree in $\kappa$. Let $h_{\kappa}$ be the diameter of the element $\kappa$, and let $h=\max _{\kappa \in K_{h}} h_{\kappa}$.

We define spectral and spectral element spaces as follows:

$$
\begin{gather*}
S_{N}(\Omega)=\left\{v \in P_{N}(\Omega),\left.v\right|_{\partial \Omega}=0\right\}, \\
S_{N, h}(\Omega)=\left\{v \in C(\bar{\Omega}):\left.v\right|_{\kappa} \in P_{N_{\kappa}}(\kappa), \forall \kappa \in K_{h},\left.v\right|_{\partial \Omega}=0\right\}, \tag{9}
\end{gather*}
$$

where $P_{N}(\Omega)$ and $P_{N_{\kappa}}(\kappa)$ are polynomial spaces of degree $N$ (resp. degree $N$ in every direction) in $\Omega$ and degree $N_{\kappa}$ (resp. degree $N_{\kappa}$ in every direction) in the element $\kappa$, respectively.

The spectral approximation of (5) is as follows: find $w_{N} \in$ $S_{N}(\Omega)$, such that

$$
\begin{equation*}
a\left(w_{N}, v\right)=(f, v), \quad \forall v \in S_{N}(\Omega) \tag{10}
\end{equation*}
$$

The spectral element approximation of (5) is as follows: find $w_{N, h} \in S_{N, h}(\Omega)$, such that

$$
\begin{equation*}
a\left(w_{N, h}, v\right)=(f, v), \quad \forall v \in S_{N, h}(\Omega) \tag{11}
\end{equation*}
$$

We assume that $f \in L^{2}(\Omega)$ and derive from Lax-Milgram theorem that the variational formations (5), (6), (10), and (11) have a unique solution, respectively.

Define the interpolation operators

$$
\begin{gather*}
I_{N_{\kappa}, h_{\kappa}}: C(\kappa) \longrightarrow P_{N_{\kappa}}(\kappa),  \tag{12}\\
I_{N}: C(\Omega) \longrightarrow S_{N}(\Omega)
\end{gather*}
$$

as the interpolations in the element $\kappa$ and the domain $\Omega$, respectively, with the tensorial Legendre-Gauss-Lobatto (LGL) interpolation nodes.

Define the interpolation operator

$$
\begin{equation*}
I_{N, h}: C(\Omega) \longrightarrow S_{N, h}(\Omega), \quad \text { satisfying }\left.I_{N, h}\right|_{\kappa}=I_{N_{\kappa}, h_{\kappa}} \tag{13}
\end{equation*}
$$

We quote from [2] (see (5.8.27) therein) the interpolation estimates for spectral and spectral element methods with LGL Nodal-basis as follows.

For all $v \in H^{s_{\kappa}}(\kappa), s_{\kappa} \geq(d+1) / 2$,

$$
\begin{gather*}
\left\|v-I_{N_{\kappa}, h_{\kappa}} v\right\|_{1, \kappa} \leq C\left(s_{\kappa}\right) h_{\kappa}^{\min \left(N_{\kappa}+1, s_{\kappa}\right)-1} N_{\kappa}^{-s_{\kappa}+1}\|v\|_{s_{\kappa}, \kappa},  \tag{14}\\
\left\|v-I_{N_{\kappa}, h_{\kappa}} v\right\|_{0, \kappa} \leq C\left(s_{\kappa}\right) h_{\kappa}^{\min \left(N_{\kappa}+1, s_{\kappa}\right)} N_{\kappa}^{-s_{\kappa}}\|v\|_{s_{\kappa}, \kappa} . \tag{15}
\end{gather*}
$$

For all $v \in H^{s}(\Omega), s \geq(d+1) / 2$,

$$
\begin{gather*}
\left\|v-I_{N} v\right\|_{1, \Omega} \leq C(s) N^{-s+1}\|v\|_{s, \Omega}  \tag{16}\\
\left\|v-I_{N} v\right\|_{0, \Omega} \leq C(s) N^{-s}\|v\|_{s, \Omega} \tag{17}
\end{gather*}
$$

We assume that the solution of boundary value problem (5) $w \in H_{0}^{1}(\Omega) \cap H^{m}(\Omega)(m>1)$, that $w_{N}$ and $w_{N, h}$ are the
solutions of (10) and (11), respectively; then we derive from Céa lemma and the interpolation estimates that

$$
\begin{align*}
&\left\|w_{N}-w\right\|_{1, \Omega} \leq C(m) N^{-m+1}\|w\|_{m, \Omega}  \tag{18}\\
&\left\|w_{N, h}-w\right\|_{1, \Omega} \leq\left\{\sum_{\kappa} C\left(s_{\kappa}\right) h_{\kappa}^{2\left\{\min \left(N_{\kappa}+1, s_{\kappa}\right)-1\right\}}\right. \\
&\left.\times N_{\kappa}^{2\left(-s_{\kappa}+1\right)}\|w\|_{s_{\kappa}, \kappa}^{2}\right\}^{1 / 2} \tag{19}
\end{align*}
$$

where $s_{\kappa} \geq(d+1) / 2, \forall \kappa \in K_{h}$.
Particularly, if $N_{\kappa}=N, \forall \kappa \in K_{h}$, then we have

$$
\begin{align*}
\left\|w_{N, h}-w\right\|_{1, \Omega} \leq & C(m) h^{\min (N+1, m)-1} \\
& \times N^{1-m}\|w\|_{m, \Omega} \tag{20}
\end{align*}
$$

Note that (18) is also suited to spectral methods with modal basis (see [1, 2]).

Using Aubin-Nitsche technique, we deduce from the regularity estimate (8) and the estimates (18)-(20) the priori estimates of boundary value problem (5) for spectral and spectral element methods; that is,

$$
\begin{gather*}
\left\|w_{N}-w\right\|_{0, \Omega} \leq C N^{-m-r_{2}+1}\|w\|_{m, \Omega}  \tag{21}\\
\left\|w_{N, h}-w\right\|_{0, \Omega} \leq C N^{-m-r_{2}+1} h^{r_{2}+\min (N+1, m)-1}\|w\|_{m, \Omega} \tag{22}
\end{gather*}
$$

## 3. Spectral and Spectral-Element Approximations and Error Estimates for Eigenvalue Problems

3.1. Spectral and Spectral-Element Approximations for Eigenvalue Problems. Consider the following eigenvalue problem corresponding to the boundary value problem (1):

$$
\begin{gather*}
L u=\lambda u, \quad \text { in } \Omega, \\
u=0, \quad \text { on } \partial \Omega . \tag{23}
\end{gather*}
$$

The variational formation of (23) is given by the following: find $\lambda \in \mathbb{C}, 0 \neq u \in H_{0}^{1}(\Omega)$, such that

$$
\begin{equation*}
a(u, v)=\lambda(u, v), \quad \forall v \in H_{0}^{1}(\Omega) \tag{24}
\end{equation*}
$$

The spectral approximation scheme of (24) is given by the following: find $\lambda_{N} \in \mathbb{C}, 0 \neq u_{N} \in S_{N}(\Omega)$, such that

$$
\begin{equation*}
a\left(u_{N}, v_{N}\right)=\lambda_{N}\left(u_{N}, v_{N}\right), \quad \forall v_{N} \in S_{N}(\Omega) \tag{25}
\end{equation*}
$$

The spectral element approximation scheme of (24) is given by the following: find $\lambda_{N, h} \in \mathbb{C}, 0 \neq u_{N, h} \in S_{N, h}(\Omega)$, such that

$$
\begin{equation*}
a\left(u_{N, h}, v_{N, h}\right)=\lambda_{N, h}\left(u_{N, h}, v_{N, h}\right), \quad \forall v_{N, h} \in S_{N, h}(\Omega) \tag{26}
\end{equation*}
$$

Define the solution operators $T: L^{2}(\Omega) \rightarrow H_{0}^{1}(\Omega), T_{N}$ : $L^{2}(\Omega) \rightarrow S_{N}(\Omega)$, and $T_{N, h}: L^{2}(\Omega) \rightarrow S_{N, h}(\Omega)$ by

$$
\begin{array}{r}
a(T f, v)=(f, v), \quad \forall f \in L^{2}(\Omega), \forall v \in H_{0}^{1}(\Omega), \\
a\left(T_{N} f, v_{N}\right)=\left(f, v_{N}\right), \quad \forall f \in L^{2}(\Omega), \\
\forall v_{N} \in S_{N}(\Omega), \tag{27}
\end{array}
$$

$$
\begin{array}{r}
a\left(T_{N, h} f, v_{N, h}\right)=\left(f, v_{N, h}\right), \quad \forall f \in L^{2}(\Omega), \\
\forall v_{N, h} \in S_{N, h}(\Omega)
\end{array}
$$

Obviously (see [17]), the equivalent operator forms for (24) and (26) are the following.

Find $\lambda \in \mathbb{C}, 0 \neq u \in H_{0}^{1}(\Omega)$, such that

$$
\begin{equation*}
T u=\lambda^{-1} u \tag{28}
\end{equation*}
$$

Find $\lambda_{N, h} \in \mathbb{C}, 0 \neq u_{N, h} \in S_{N, h}(\Omega)$, such that

$$
\begin{equation*}
T_{N, h} u_{N, h}=\lambda_{N, h}^{-1} u_{N, h} \tag{29}
\end{equation*}
$$

The adjoint problem of the eigenvalue problem (23) is

$$
\begin{align*}
L^{*} u^{*} & =\lambda^{*} u^{*}, \quad \text { in } \Omega \\
u^{*} & =0, \quad \text { on } \partial \Omega \tag{30}
\end{align*}
$$

where $L^{*} u^{*}=-\nabla \cdot\left(D \nabla u^{*}\right)-\mathbf{b} \cdot \nabla u^{*}+(c-\nabla \cdot \mathbf{b}) u^{*}$.
The variational formation of (30) is given by the following: find $\lambda^{*} \in \mathbb{C}, 0 \neq u^{*} \in H_{0}^{1}(\Omega)$, such that

$$
\begin{equation*}
a^{*}\left(u^{*}, v\right):=\overline{a\left(v, u^{*}\right)}=\lambda^{*}\left(u^{*}, v\right), \quad \forall v \in H_{0}^{1}(\Omega) \tag{31}
\end{equation*}
$$

The spectral element approximation scheme of (31) is given by the following: find $\lambda_{N, h}^{*} \in \mathbb{C}, 0 \neq u_{N, h}^{*} \in S_{N, h}(\Omega)$, such that

$$
\begin{equation*}
a^{*}\left(u_{N, h}^{*}, v_{N, h}\right)=\lambda_{N, h}^{*}\left(u_{N, h}^{*}, v_{N, h}\right), \quad \forall v_{N, h} \in S_{N, h}(\Omega) \tag{32}
\end{equation*}
$$

We can likewise define the equivalent operator forms for the eigenvalue problems (31) and (32) as

$$
\begin{gather*}
T^{*} u^{*}=\lambda^{*-1} u^{*}, \\
T_{N, h}^{*} u_{N, h}^{*}=\lambda_{N, h}^{*-1} u_{N, h}^{*} . \tag{33}
\end{gather*}
$$

Let $\lambda$ be an eigenvalue of (23). There exists a smallest integer $\mu$, called the ascent of $\lambda$, such that $\operatorname{ker}\left(\left(\lambda^{-1}-T\right)^{\mu}\right)=$ $\operatorname{ker}\left(\left(\lambda^{-1}-T\right)^{\mu+1}\right) \cdot q=\operatorname{dim} \operatorname{ker}\left(\left(\lambda^{-1}-T\right)^{\mu}\right)$ is called the algebraic multiplicity of $\lambda$. The functions in $\operatorname{ker}\left(\left(\lambda^{-1}-T\right)^{\mu}\right)$ are called generalized eigenfunctions of $T$ corresponding to $\lambda$. Likewise the ascent, algebraic multiplicity and generalized eigenfunctions of $\lambda_{N, h}, \lambda^{*}$ and $\lambda_{N, h}^{*}$ can be defined.

Let $\lambda$ be an eigenvalue of (23) with the algebraic multiplicity $q$ and the ascent $\mu$. Assume $\left\|T_{N, h}-T\right\|_{1, \Omega} \rightarrow 0(N \rightarrow \infty$, $h \rightarrow 0)$; then there are eigenvalues $\lambda_{j, N, h}(j=1,2, \ldots, q)$ of (26) which converge to $\lambda$. Let $M(\lambda)$ be the space spanned by all generalized eigenfunctions corresponding to $\lambda$ of $T$,
and let $M_{N, h}(\lambda)$ be the space spanned by all generalized eigenfunctions corresponding to all eigenvalues of $T_{N, h}$ that converge to $\lambda$.

In view of adjoint problems (31) and (32), the definitions of $M^{*}\left(\lambda^{*}\right)$ and $M_{N, h}^{*}\left(\lambda^{*}\right)$ are analogous to $M(\lambda)$ and $M_{N, h}(\lambda)$. Let $\widehat{M}(\lambda)=\left\{v \in M(\lambda):\|v\|_{1, \Omega}=1\right\}$, and let $\widehat{M^{*}}\left(\lambda^{*}\right)=\{v \in$ $\left.M^{*}\left(\lambda^{*}\right):\|v\|_{1, \Omega}=1\right\}$.

Note that when $\mathbf{b}=0$, both (24) and (26) are symmetric. Thus, the ascent $\mu=1$ of $\lambda$, and the ascent $l=1$ of $\lambda_{N, h}$.
3.2. A Priori Error Estimates. We will analyze a prior error estimates for spectral element methods which are suitable for spectral methods with mesh fineness $h$ not considered.

Assume that $R$ and $U$ are two closed subspace in $H_{0}^{1}(\Omega)$. Denote

$$
\begin{gather*}
\delta(R, U)=\sup _{\substack{v \in R \\
\|v\|_{1, \Omega}=1}} \operatorname{dist}(v, U)  \tag{34}\\
\theta(R, U)=\max (\delta(R, U), \delta(U, R))
\end{gather*}
$$

We say that $\theta(R, U)$ is the gap between $R$ and $U$.
Denote

$$
\begin{gather*}
\varepsilon_{N, h}=\varepsilon_{N, h}(\lambda)=\sup _{u \in \widehat{M}(\lambda)} \inf _{v \in S_{N, h}(\Omega)}\|u-v\|_{1, \Omega},  \tag{35}\\
\varepsilon_{N, h}^{*}=\varepsilon_{N, h}^{*}\left(\lambda^{*}\right)=\sup _{u \in \widehat{M^{*}\left(\lambda^{*}\right)}} \inf _{v \in S_{N, h}(\Omega)}\|u-v\|_{1, \Omega} .
\end{gather*}
$$

We give the following four lemmas from Theorem 8.1-8.4 in [17], which are applications to spectral element methods.

Lemma 1. Assume $\left\|T_{N, h}-T\right\|_{1, \Omega} \rightarrow 0(N \rightarrow \infty, h \rightarrow 0)$. For small enough $h$ and big enough $N$, there holds

$$
\begin{equation*}
\theta\left(M(\lambda), M_{N, h}(\lambda)\right) \leq C \varepsilon_{N, h} . \tag{36}
\end{equation*}
$$

Lemma 2. Assume $\left\|T_{N, h}-T\right\|_{1, \Omega} \rightarrow 0(N \rightarrow \infty, h \rightarrow 0)$; then

$$
\begin{equation*}
\left|\lambda^{-1}-\frac{1}{q} \sum_{j=1}^{q} \lambda_{j, N, h}^{-1}\right| \leq C \varepsilon_{N, k} \varepsilon_{N, h}^{*} \tag{37}
\end{equation*}
$$

Lemma 3. Assume that $\left\|T_{N, h}-T\right\|_{1, \Omega} \rightarrow 0(N \rightarrow \infty, h \rightarrow$ 0 ); then there holds

$$
\begin{equation*}
\left|\lambda-\lambda_{j, N, h}\right| \leq C\left(\varepsilon_{N, h} \varepsilon_{N, h}^{*}\right)^{1 / \mu} \quad(j=1,2, \ldots, q) . \tag{38}
\end{equation*}
$$

Since $\operatorname{ker}\left(\left(\lambda^{-1}-T\right)^{l}\right)(l \geq 1)$ is a finite dimensional space, there exists a direct-sum decomposition $H_{0}^{1}(\Omega)=\operatorname{ker}\left(\left(\lambda^{-1}-T\right)^{l}\right) \oplus$ $M_{l}$. We define the operator $E_{l}$ as a projection along $M_{l}$ from $H_{0}^{1}(\Omega)$ to $\operatorname{ker}\left(\left(\lambda^{-1}-T\right)^{l}\right)$.

Lemma 4. Assume $\left\|T_{N, h}-T\right\|_{1, \Omega} \rightarrow 0(N \rightarrow \infty, h \rightarrow 0)$. Let $\lambda_{N, h}$ be an eigenvalue of $T_{N, h}$ and $\lim _{N \rightarrow \infty, h \rightarrow 0} \lambda_{N, h}=\lambda$. $u_{N, h}$ satisfies $\left(\lambda_{N, h}^{-1}-T_{N, h}\right)^{k} u_{N, h}=0$ and $\left\|u_{N, h}\right\|_{1, \Omega}=1$, where $k \leq \mu$ is a positive integer. Then, for every integer $l \in[k, \mu]$, there holds

$$
\begin{equation*}
\left\|u_{N, h}-E_{l} u_{N, h}\right\|_{1, \Omega} \leq C \varepsilon_{N, h}^{(l-k+1) / \mu} \tag{39}
\end{equation*}
$$

We assume that in this section, for the sake of simplicity, $N_{\kappa}=N, \forall \kappa \in K_{h}$.

Theorem 5. If $M(\lambda) \subset H^{t_{1}}(\Omega)$ and $M^{*}\left(\lambda^{*}\right) \subset H^{t_{2}}(\Omega)$, then there holds the following error estimates:

$$
\begin{align*}
& \left|\frac{1}{q} \sum_{j=1}^{q} \lambda_{j, N, h}-\lambda\right|  \tag{40}\\
& \leq C\left(\frac{h^{\tau_{1}+\tau_{2}-2}}{N^{t_{1}+t_{2}-2}}\right) \sup _{u \in \widehat{M}(\lambda)}\|u\|_{t_{1}, \Omega} \sup _{v \in \widehat{M^{*}}\left(\lambda^{*}\right)}\|v\|_{t_{2}, \Omega}, \\
& \left|\lambda_{j, N, h}-\lambda\right| \\
& \leq C\left(\left(\frac{h^{\tau_{1}+\tau_{2}-2}}{N^{t_{1}+t_{2}-2}}\right) \sup _{u \in \widehat{M}(\lambda)}\|u\|_{t_{1}, \Omega} \sup _{v \in \widehat{M^{*}}\left(\lambda^{*}\right)}\|v\|_{t_{2}, \Omega}\right)^{1 / \mu}  \tag{41}\\
& (j=1,2, \ldots, q), \\
& \theta\left(M(\lambda), M_{N, h}(\lambda)\right) \leq C \frac{h^{\tau_{1}-1}}{N^{t_{1}-1}} \sup _{u \in \widetilde{M}(\lambda)}\|u\|_{t_{1}, \Omega} . \tag{42}
\end{align*}
$$

Let $\left\|u_{N, h}\right\|_{1, \Omega}=1$, and let $\left(\lambda_{N, h}^{-1}-T_{N, h}\right)^{l_{1}} u_{N, h}=0$, for some $l_{1} \leq \mu$. Then, for every integer $l_{2}\left(l_{1} \leq l_{2} \leq \mu\right)$, there exists a function $u^{\prime}$, such that $\left(\lambda^{-1}-T\right)^{l_{2}} u^{\prime}=0$ and

$$
\begin{equation*}
\left\|u_{N, h}-u^{\prime}\right\|_{1, \Omega} \leq C\left(\left(\frac{h^{\tau_{1}-1}}{N^{t_{1}-1}}\right) \sup _{u \in \widehat{M}(\lambda)}\|u\|_{t_{1}, \Omega}\right)^{\left(l_{2}-l_{1}+1\right) / \mu}, \tag{43}
\end{equation*}
$$

where $\tau_{1}=\min \left(N+1, t_{1}\right), \tau_{2}=\min \left(N+1, t_{2}\right)$.
Proof. We derive from the error estimate (20) that

$$
\begin{align*}
& \left\|T_{N, h}-T\right\|_{1, \Omega} \\
& \quad=\sup _{f \in H_{0}^{1}(\Omega)} \frac{\left\|\left(T-T_{N, h}\right) f\right\|_{1, \Omega}}{\|f\|_{1, \Omega}}  \tag{44}\\
& \quad \leq C\left(1+r_{1}\right) h^{r_{1}} N^{-r_{1}} \longrightarrow 0 \quad(N \longrightarrow \infty, h \longrightarrow 0) .
\end{align*}
$$

By (14),

$$
\begin{align*}
\varepsilon_{N, h}=\varepsilon_{N, h}(\lambda) & =\sup _{u \in \widehat{M}(\lambda)} \inf _{v \in S_{N, h}(\Omega)}\|u-v\|_{1, \Omega} \\
& \leq C\left(\frac{h^{\tau_{1}-1}}{N^{t_{1}-1}}\right) \sup _{u \in \widehat{M}(\lambda)}\|u\|_{t_{1}, \Omega} . \tag{45}
\end{align*}
$$

Analogically,

$$
\begin{equation*}
\varepsilon_{N, h}^{*} \leq C\left(\frac{h^{\tau_{2}-1}}{N^{t_{2}-1}}\right) \sup _{u \in \overline{M^{*}}\left(\lambda^{*}\right)}\|u\|_{t_{2}, \Omega} \tag{46}
\end{equation*}
$$

Plugging the two inequalities above into (36), (38), and (39) yields (42), (41), and (43), respectively. We find from (37) that

$$
\begin{align*}
\left|\frac{1}{q} \sum_{j=1}^{q} \lambda_{j, N, h}-\lambda\right| & =\left|\frac{1}{q} \sum_{j=1}^{q} \frac{\lambda_{j, N, h}^{-1}-\lambda^{-1}}{\lambda^{-1} \lambda_{j, N, h}^{-1}}\right|  \tag{47}\\
& \leq C\left|\frac{1}{q} \sum_{j=1}^{q} \lambda_{j, N, h}^{-1}-\lambda^{-1}\right| \leq C \varepsilon_{N, h} \varepsilon_{N, h}^{*}
\end{align*}
$$

combining with (45) and (46) yields (40).
Supposing that $\left\|T_{N, h}-T\right\|_{0, \Omega} \rightarrow 0(N \rightarrow \infty, h \rightarrow 0)$, $\rho(T)$ is a regular set of $T$, and $\Gamma \subset \rho(T)$ is a closed Jordan curve enclosing $\lambda^{-1}$.

Denote

$$
\begin{gather*}
R(z)=(T-z)^{-1} \\
R\left(T_{N, h}, z\right)=\left(T_{N, h}-z\right)^{-1} . \tag{48}
\end{gather*}
$$

Define the spectral projection operators

$$
\begin{gather*}
E=\frac{-1}{2 i \pi} \int_{\Gamma} R(T, z) \mathrm{dz}: H_{0}^{1}(\Omega) \longrightarrow M(\lambda),  \tag{49}\\
E_{N, h}=\frac{-1}{2 i \pi} \int_{\Gamma} R\left(T_{N, h}, z\right) \mathrm{dz}: H_{0}^{1}(\Omega) \longrightarrow M_{N, h}(\lambda) .
\end{gather*}
$$

We give the following lemma by referring to $[18,19]$ (see proposition 5.3 in [18] and theorem 1.3.2 in [19]).

Lemma 6. If $\left\|T_{N, h}-T\right\|_{0, \Omega} \rightarrow 0(N \rightarrow \infty, h \rightarrow 0)$, then there holds that $E_{N, h} \rightarrow E(p), R\left(T_{N, h}, z\right)$ is uniformly bounded with $N$ and $h$, and

$$
\begin{array}{r}
\left\|\left(E_{N, h}-E\right) v\right\|_{0, \Omega} \leq \max _{z \in \Gamma}\left\|\left(T-T_{N, h}\right) R(z) v\right\|_{0, \Omega}, \\
\forall v \in H_{0}^{1}(\Omega), \\
\left\|\left(E_{N, h}-E\right) v\right\|_{0, \Omega} \leq C \max _{z \in \Gamma}\left\|\left(T-T_{N, h}\right) R\left(T_{N, h}, z\right) v\right\|_{0, \Omega}, \\
\forall v \in H_{0}^{1}(\Omega) . \tag{50}
\end{array}
$$

Theorem 7. Under the assumptions of Theorem 5, further assume that the ascent of $\lambda$ is $\mu=1$. Let $\left(\lambda_{N, h}, u_{N, h}\right)$ be an eigenpair of (26) with $\left\|u_{N, h}\right\|_{0, \Omega}=1$; then there exists an eigenpair $(\lambda, u)$ of $(24)$, such that

$$
\begin{gather*}
\left\|u_{N, h}-u\right\|_{1, \Omega} \leq \frac{C h^{\tau_{1}-1}\|u\|_{t_{1}, \Omega}}{N^{t_{1}-1}},  \tag{51}\\
\left\|u_{N, h}-u\right\|_{0, \Omega} \leq \frac{C h^{r_{2}+\tau_{1}-1}\|u\|_{t_{1}, \Omega}}{N^{r_{2}+t_{1}-1}},  \tag{52}\\
\left|\lambda_{N, h}-\lambda\right| \\
\leq C\left(\left(\frac{h^{\tau_{1}+\tau_{2}-2}}{N^{t_{1}+t_{2}-2}}\right) \sup _{u \in \widetilde{M}(\lambda)}\|u\|_{t_{1}, \Omega} \sup _{v \in \widehat{M}^{*}\left(\lambda^{*}\right)}\|v\|_{t_{2}, \Omega}\right), \tag{53}
\end{gather*}
$$

where $\tau_{1}=\min \left(N+1, t_{1}\right)$ and $\tau_{2}=\min \left(N+1, t_{2}\right)$.

Let $(\lambda, u)$ be an eigenpair of (24). If $\lambda_{N, h}$ is an eigenvalue of (26) convergence to $\lambda$, then there exists $u_{N, h} \in \operatorname{ker}\left(\lambda_{N, h}^{-1}-T_{N, h}\right)$, such that (51)-(53) hold.

Proof. We deduce (53) immediately from (41). We derive from (22) and (7) that

$$
\begin{equation*}
\left\|T f-T_{N, h} f\right\|_{0, \Omega} \leq C N^{-r_{1}-r_{2}} h^{r_{1}+r_{2}}\|f\|_{0, \Omega} \tag{54}
\end{equation*}
$$

thus, $\left\|T-T_{N, h}\right\|_{0, \Omega} \rightarrow 0(N \rightarrow \infty, h \rightarrow 0)$. Taking $u=$ $E u_{N, h}$ and by virtue of $R\left(T_{N, h}, z\right) u_{N, h}=\left(\lambda_{N, h}^{-1}-z\right)^{-1} u_{N, h}$, Lemma 6 and (22), we have

$$
\begin{align*}
\left\|u-u_{N, h}\right\|_{0, \Omega}= & \left\|E u_{N, h}-E_{N, h} u_{N, h}\right\|_{0, \Omega} \\
\leq & C\left\|\left(T-T_{N, h}\right) u_{N, h}\right\|_{0, \Omega} \\
\leq & C\left(\left\|\left(T-T_{N, h}\right) u\right\|_{0, \Omega}\right.  \tag{55}\\
& \left.+\left\|\left(T-T_{N, h}\right)\left(u_{N, h}-u\right)\right\|_{0, \Omega}\right),
\end{align*}
$$

from which follows

$$
\begin{align*}
\left\|u-u_{N, h}\right\|_{0, \Omega} & \leq C\left\|\left(T-T_{N, h}\right) u\right\|_{0, \Omega} \\
& \leq \frac{C h^{r_{2}+\tau_{1}-1}\|u\|_{t_{1}, \Omega}}{N^{r_{2}+t_{1}-1}} \tag{56}
\end{align*}
$$

which is (52). By direct calculation, we have

$$
\begin{align*}
\| u & -u_{N, h} \|_{1, \Omega} \\
& =\left\|\lambda T u-\lambda_{h} T_{N, h} u_{N, h}\right\|_{1, \Omega} \\
& \leq\left\|\lambda T u-\lambda T_{N, h} u\right\|_{1, \Omega}+\left\|\lambda T_{N, h} u-\lambda_{h} T_{N, h} u_{N, h}\right\|_{1, \Omega}  \tag{57}\\
& \leq\left\|\left(T-T_{N, h}\right)(\lambda u)\right\|_{1, \Omega}+C\left\|\lambda u-\lambda_{N, h} u_{N, h}\right\|_{0, \Omega} .
\end{align*}
$$

Plugging (20), (52), and (53) into (57) yields (51).
If $(\lambda, u)$ is an eigenpair of (24), let $u_{N, h}=E_{N, h} u$; by the same argument we can prove (51) and (52).

## 4. A Posteriori Error Estimates

Based on [20], we will discuss a posteriori error estimates. We further assume that $\Omega \subset R^{2}$, the partition $K_{h}$ is $\gamma$-shape regular, and the polynomial degree of neighboring elements are comparable; that is, there exists $\gamma>0$, such that for all $\kappa, \kappa^{\prime} \in K_{h}, \bar{\kappa} \cap \overline{\kappa^{\prime}} \neq \emptyset$,

$$
\begin{gather*}
\gamma^{-1} h_{\kappa} \leq h_{\kappa^{\prime}} \leq \gamma h_{\kappa}  \tag{58}\\
\gamma^{-1}\left(N_{\kappa}+1\right) \leq N_{\kappa^{\prime}}+1 \leq \gamma\left(N_{\kappa}+1\right)
\end{gather*}
$$

We refer to the $h p$-clément interpolation estimates given by [20,21] (see theorems 2.2 and 2.3, respectively), which generalize the well-known clément type interpolation operators studied in [22] and [23] to the hp context.

Lemma 8. Assume that the partition $K_{h}$ is $\gamma$-shape regular and the polynomial distribution $\mathbf{N}$ is comparable. Then there
exists a positive constant $C=C(\gamma)$ and the clément operator $I: H_{0}^{1}(\Omega) \rightarrow S_{N, h}(\Omega)$, such that

$$
\begin{align*}
& \|v-I v\|_{0, \kappa} \leq C \frac{h_{\kappa}}{N_{\kappa}}\|\nabla v\|_{0, \omega_{\kappa}},  \tag{59}\\
& \|v-I v\|_{0, e} \leq C \sqrt{\frac{h_{e}}{N_{e}}}\|\nabla v\|_{0, \omega_{e}}, \tag{60}
\end{align*}
$$

where $h_{e}$ is the length of the edge $e$ and $N_{e}=\max \left(N_{\kappa_{1}}, N_{\kappa_{2}}\right)$, where $\kappa_{1}, \kappa_{2}$ are elements sharing the edge e and $\omega_{\kappa}, \omega_{e}$ are patches covering $\kappa$ and $e$ with a few layers, respectively.

Define interval $\widehat{I}=(0,1)$ and weight function $\Phi_{\widehat{I}}(x):=$ $x(1-x)$. Denote the reference square and triangle element by $\widehat{\kappa}=(0,1)^{2}$ and $\widehat{\kappa}=\{(x, y) \mid 0<x<1,0<y<\sqrt{3}(1 / 2-\mid 1 / 2-$ $x \mid)\}$, respectively. Define weight function $\Phi_{\widehat{\kappa}}(x):=\operatorname{dist}(x, \partial \widehat{\kappa})$.

The following three lemmas are given by [20]. Lemmas $9-10$ provide the polynomial inverse estimates in standard interval and element, while Lemma 11 provides a result for the extension from an edge to the element.

Lemma 9. Let $-1<\alpha<\beta, \sigma \in[0,1]$. Then there exists $C=$ $C(\alpha, \beta)$, such that for all $N \in \mathbb{N}$ and all univariate polynomials $\pi_{N}$ of degree $N$,

$$
\begin{align*}
& \int_{\hat{I}} \Phi_{\hat{I}}^{\alpha}(x)\left|\pi_{N}(x)\right|^{2} d x \\
& \quad \leq C N^{2(\beta-\alpha)} \int_{\hat{I}} \Phi_{\hat{I}}^{\beta}\left|\pi_{N}(x)\right|^{2} d x . \tag{61}
\end{align*}
$$

Lemma 10. Let $-1<\alpha<\beta, \sigma \in[0,1]$. Then there exist $C_{1}=C(\alpha, \beta), C_{2}=C_{\sigma}>0$, such that for all $N \in \mathbb{N}$ and all polynomials $\pi_{N}$ of degree bi- $N$,

$$
\begin{align*}
& \int_{\widehat{\kappa}} \Phi_{\widehat{\kappa}}^{\alpha}\left|\pi_{N}\right|^{2} d x d y  \tag{62}\\
& \quad \leq C_{1} N^{2(\beta-\alpha)} \int_{\widehat{\kappa}} \Phi_{\widehat{\kappa}}^{\beta}\left|\pi_{N}\right|^{2} d x d y \\
& \int_{\widehat{\kappa}} \Phi_{\widehat{\kappa}}^{2 \sigma}\left|\nabla \pi_{N}\right|^{2} d x d y  \tag{63}\\
& \quad \leq C_{2} N^{2(2-\sigma)} \int_{\widehat{\kappa}} \Phi_{\widehat{\kappa}}^{\sigma}\left|\pi_{N}\right|^{2} d x d y
\end{align*}
$$

Lemma 11. Let $\alpha \in(1 / 2,1] . \widehat{e}:=(0,1) \times\{0\}, \Phi_{\widehat{e}}:=x(1-x)$; then there exists $C_{\alpha}>0$ such that for all $N \in \mathbb{N}, \varepsilon \in(0,1]$, and all univariate polynomials $\pi$ of degree $N$, there exists an extension $v_{\widehat{e}} \in H^{1}(\widehat{\kappa})$ and holds

$$
\begin{gather*}
\left.v_{\widehat{e}}\right|_{\widehat{e}}=\pi \cdot \Phi_{\widehat{e}}^{\alpha},\left.\quad v_{\widehat{e}}\right|_{\partial \widehat{\kappa} \mid \widehat{e}}=0,  \tag{64}\\
\left\|v_{\widehat{e}}\right\|_{0, \widehat{\kappa}}^{2} \leq C_{\alpha} \varepsilon\left\|\pi \Phi_{\widehat{e}}^{\alpha / 2}\right\|_{0, \vec{e}}^{2}  \tag{65}\\
\left\|\nabla v_{\hat{e}}\right\|_{0, \widehat{\kappa}}^{2} \leq C_{\alpha}\left(\varepsilon N^{2(2-\alpha)}+\varepsilon^{-1}\right)\left\|\pi \Phi_{\widehat{e}}^{\alpha / 2}\right\|_{0, \vec{e}}^{2} \tag{66}
\end{gather*}
$$

It is easy to know that the three lemmas above hold for complex-valued polynomials.

Let $D_{\kappa}, \mathbf{b}_{\kappa}$, and $c_{\kappa}$ be the interpolations of $D, \mathbf{b}$, and $c$ in $\kappa$ with the polynomial degree $N_{\kappa}$ (resp. degree $N_{\kappa}$ in every direction), respectively, or the $L^{2}(\kappa)$-projection on the space of polynomials with degree $N_{\kappa}$. For convenient argument, here and hereafter we assume that $(\lambda, u)$ and $\left(\lambda^{*}=\bar{\lambda}, u^{*}\right)$ are the eigenpairs of the eigenvalue problem (24) and its adjoint problem (31), respectively. $\left(\lambda_{N, h}, u_{N, h}\right)$ and ( $\lambda_{N, h}^{*}=$ $\left.\overline{\lambda_{N, h}}, u_{N, h}^{*}\right)$ are the solutions of the corresponding spectral element approximations (26) and (32), respectively.

Denote

$$
\begin{align*}
L_{\kappa} u_{N, h}:= & -\nabla \cdot\left(D_{\kappa} \nabla u_{N, h}\right) \\
& +\mathbf{b}_{\kappa} \cdot \nabla u_{N, h}+c_{\kappa} u_{N, h}, \\
L_{\kappa}^{*} u_{N, h}^{*}:= & -\nabla \cdot\left(D_{\kappa} \nabla u_{N, h}^{*}\right)  \tag{67}\\
& -\mathbf{b}_{\kappa} \cdot \nabla u_{N, h}^{*}+\left(c_{\kappa}-\nabla \cdot \mathbf{b}_{\kappa}\right) u_{N, h}^{*} .
\end{align*}
$$

Define the local error indicators

$$
\begin{align*}
& \eta_{\alpha ; k}^{2}:=\eta_{\alpha ; B_{k}}^{2}+\eta_{\alpha ; E_{k}}^{2},  \tag{68}\\
& \eta_{\alpha ; k}^{* 2}:=\eta_{\alpha ; B_{k}}^{* 2}+\eta_{\alpha ; E_{k}}^{* 2} .
\end{align*}
$$

Their first terms $\eta_{\alpha ; B_{\kappa} \kappa}^{2}, \eta_{\alpha ; B_{\kappa}}^{* 2}$ are the weighted element internal residuals given by

$$
\begin{align*}
& \eta_{\alpha ; B_{\kappa}}^{2}:=\frac{h_{\kappa}^{2}}{N_{\kappa}^{2}}\left\|\left(-L_{\kappa} u_{N, h}+\lambda_{N, h} u_{N, h}\right) \Phi_{\kappa}^{\alpha / 2}\right\|_{0, \kappa^{\prime}}^{2} \\
& \eta_{\alpha ; B_{\kappa}}^{* 2}:=\frac{h_{\kappa}^{2}}{N_{\kappa}^{2}}\left\|\left(-L_{\kappa}^{*} u_{N, h}^{*}+\lambda_{N, h}^{*} u_{N, h}^{*}\right) \Phi_{\kappa}^{\alpha / 2}\right\|_{0, \kappa}^{2} . \tag{69}
\end{align*}
$$

Their second terms $\eta_{\alpha ; E_{\kappa}}^{2}, \eta_{\alpha ; E_{\kappa}}^{* 2}$ are the weighted element boundary residuals given by

$$
\begin{align*}
& \eta_{\alpha ; E_{\kappa}}^{2}:=\sum_{e \subset \partial \kappa \cap \Omega} \frac{h_{e}}{2 N_{e}}\left\|D_{\kappa}\left[\frac{\partial u_{N, h}}{\partial n}\right] \Phi_{e}^{\alpha / 2}\right\|_{0, e}^{2}, \\
& \eta_{\alpha ; E_{\kappa}}^{* 2}:=\sum_{e \subset \partial \kappa \cap \Omega} \frac{h_{e}}{2 N_{e}}\left\|D_{\kappa}\left[\frac{\partial u_{N, h}^{*}}{\partial n}\right] \Phi_{e}^{\alpha / 2}\right\|_{0, e}^{2}, \tag{70}
\end{align*}
$$

where we denote the jump of the normal derivatives of $u_{N, h}$ and $u_{N, h}^{*}$ across the edges by $\left[\partial u_{N, h} / \partial n\right]$ and $\left[\partial u_{N, h}^{*} / \partial n\right]$, respectively. $h_{e}$ is the length of edge $e$. The weight functions $\Phi_{\kappa}$ and $\Phi_{e}$ are scaled transformations of the weight functions $\Phi_{\widehat{\kappa}}$ and $\Phi_{\widehat{e}}$; that is, if $F_{\kappa}$ is the element map for element $\kappa$ and $e$ is the image of the edge $\hat{e}$ under $F_{\kappa}$, then

$$
\begin{equation*}
\Phi_{\kappa}=C_{\kappa} \Phi_{\widehat{\kappa}} \circ F_{\kappa}^{-1}, \quad \Phi_{e}=C_{e} \Phi_{\widehat{e}} \circ F_{\kappa}^{-1} \tag{71}
\end{equation*}
$$

where we choose $C_{\kappa}, C_{e}>0$, such that

$$
\begin{equation*}
\int_{\kappa} \Phi_{\kappa} d x d y=\int_{\kappa} d x d y, \quad \int_{e} \Phi_{e} d s=\int_{e} d s \tag{72}
\end{equation*}
$$

We define the global error indicators as follows:

$$
\begin{align*}
& \eta_{\alpha}^{2}:=\sum_{\kappa \in K_{h}} \eta_{\alpha ; \kappa}^{2} \\
& \eta_{\alpha}^{* 2}:=\sum_{\kappa \in K_{h}} \eta_{\alpha ; \kappa}^{* 2} \tag{73}
\end{align*}
$$

Theorem 12. Let $\alpha \in[0,1]$. Then there exists a constant $C>0$ independent of $h, \mathbf{N}$, and $\kappa$, such that

$$
\begin{align*}
&\left\|u-u_{N, h}\right\|_{1, \Omega}^{2} \leq C \sum_{\kappa \in K_{h}} N_{\kappa}^{2 \alpha} \eta_{\alpha ; \kappa}^{2} \\
&+C \sum_{\kappa \in K_{h}}\left\{\frac{h_{\kappa}^{2}}{N_{\kappa}^{2}}\left\|L_{\kappa} u_{N, h}-L u_{N, h}\right\|_{0, \kappa}^{2}\right. \\
&+\sum_{e \subset \partial \kappa \cap \Omega} \frac{h_{e}}{N_{e}}\left\|D-D_{\kappa}\right\|_{0, e}^{2}  \tag{74}\\
&\left.\quad \times\left\|\frac{\partial u_{N, h}}{\partial n}\right\|_{0, \infty, e}^{2}\right\} \\
&+C\left\|\lambda u-\lambda_{N, h} u_{N, h}\right\|_{0, \Omega}^{2}
\end{align*}
$$

Proof. We denote $w:=u-u_{N, h}-I\left(u-u_{N, h}\right)$, where $I$ is $h p$ clément operator given by Lemma 8 . We derive from $H_{0}^{1}(\Omega)$ elliptic of $a(\cdot, \cdot)$ that

$$
\begin{aligned}
C\left\|u-u_{N, h}\right\|_{1, \Omega}^{2} \leq & a\left(u-u_{N, h}, w\right) \\
& +a\left(u-u_{N, h}, I\left(u-u_{N, h}\right)\right) \\
= & \lambda \int_{\Omega} u \bar{w}-a\left(u_{N, h}, w\right) \\
& +\int_{\Omega}\left(\lambda u-\lambda_{N, h} u_{N, h}\right) \\
& \times \overline{I\left(u-u_{N, h}\right)} \\
= & \int_{\Omega}\left(\lambda_{N, h} u_{N, h}\right) \bar{w} \\
& +\int_{\Omega}\left(\lambda u-\lambda_{N, h} u_{N, h}\right) \overline{u-u_{N, h}} \\
& -a\left(u_{N, h}, w\right), \\
a\left(u_{N, h}, w\right)= & \sum_{\kappa \in \kappa_{h}} \int_{\kappa} L u_{N, h} \bar{w} \\
& -\sum_{\kappa \in \kappa_{h}} \int_{\partial \kappa} D \frac{\partial u_{N, h}}{\partial n} \bar{w} \\
= & \sum_{\kappa \in \kappa_{h}} \int_{\kappa} L u_{N, h} \bar{w} \\
& -\frac{1}{2} \sum_{\kappa \in \kappa_{h}} \sum_{e c \partial \kappa \cap \Omega} \int_{e} D\left[\frac{\partial u_{N, h}}{\partial n}\right] \bar{w} .
\end{aligned}
$$

Therefore,

$$
\begin{aligned}
& C\left\|u-u_{N, h}\right\|_{1, \Omega}^{2} \\
& \quad \leq \sum_{\kappa \in K_{h}} \int_{\kappa}\left(-L u_{N, h}+\lambda_{N, h} u_{N, h}\right) \bar{w}
\end{aligned}
$$

$$
\begin{align*}
& \quad+\frac{1}{2} \sum_{\kappa \in K_{h}} \sum_{e \subset \partial \kappa \cap \Omega} \int_{e} D\left[\frac{\partial u_{N, h}}{\partial n}\right] \bar{w} \\
& +\int_{\Omega}\left(\lambda u-\lambda_{N, h} u_{N, h}\right) \overline{\left(u-u_{N, h}\right)}, \tag{76}
\end{align*}
$$

which together with

$$
\begin{align*}
\int_{e} D\left[\frac{\partial u_{N, h}}{\partial n}\right] \bar{w}= & \int_{e}\left(D-D_{\kappa}\right)\left[\frac{\partial u_{N, h}}{\partial n}\right] \bar{w} \\
& +\int_{e} D_{\kappa}\left[\frac{\partial u_{N, h}}{\partial n}\right] \bar{w} \\
\int_{\kappa}\left(-L u_{N, h}+\lambda_{N, h} u_{N, h}\right) \bar{w}= & \int_{\kappa}\left(-L_{\kappa} u_{N, h}+\lambda_{N, h} u_{N, h}\right) \bar{w} \\
& +\int_{\kappa}\left(L_{\kappa} u_{N, h}-L u_{N, h}\right) \bar{w} \tag{77}
\end{align*}
$$

and using Cauchy-Schwartz inequality, the $h p$-clément interpolation estimates in Lemma 8 then yield

$$
\begin{align*}
& \left\|u-u_{N, h}\right\|_{1, \Omega}^{2} \\
& \leq \quad C\left\{\sum _ { \kappa \in K _ { h } } \left[\eta_{0 ; B_{\kappa}}^{2}+\eta_{0 ; E_{\kappa}}^{2}+\frac{h_{\kappa}^{2}}{N_{\kappa}^{2}}\left\|L_{\kappa} u_{N, h}-L u_{N, h}\right\|_{0, \kappa}^{2}\right.\right. \\
& \\
& \left.\left.\quad+\sum_{e \subset \partial \kappa \cap \Omega} \frac{h_{e}}{N_{e}}\left\|\left(D-D_{\kappa}\right)\left[\frac{\partial u_{N, h}}{\partial n}\right]\right\|_{0, e}^{2}\right]\right\}^{1 / 2}  \tag{78}\\
& \quad \times\left\|u-u_{N, h}\right\|_{1, \Omega}+C\left\|\lambda u-\lambda_{N, h} u_{N, h}\right\|_{0, \Omega}\left\|u-u_{N, h}\right\|_{0, \Omega} .
\end{align*}
$$

Using scaled transformation and setting $\alpha=0, \beta=\alpha$ in (61) and (62), we get $\eta_{0 ; E_{\kappa}} \leq C N_{\kappa}^{\alpha} \eta_{\alpha ; E_{\kappa}}$ and $\eta_{0 ; B_{\kappa}} \leq C N_{\kappa}^{\alpha} \eta_{\alpha ; B_{\kappa}}$; then this proof concludes.

For the adjoint eigenvalue problem, we still have the following.

Theorem 13. Let $\alpha \in[0,1]$. Then there exists a constant $C>0$ independent of $h, \mathbf{N}$, and $\kappa$, such that

$$
\begin{align*}
& \left\|u^{*}-u_{N, h}^{*}\right\|_{1, \Omega}^{2} \\
& \quad \leq \begin{array}{l}
C \sum_{\kappa \in K_{h}} N_{\kappa}^{2 \alpha} \eta_{\alpha ; \kappa}^{* 2} \\
\\
\quad+C \sum_{\kappa \in K_{h}}\left\{\frac{h_{\kappa}^{2}}{N_{\kappa}^{2}}\left\|L_{\kappa}^{*} u_{N, h}^{*}-L^{*} u_{N, h}^{*}\right\|^{2}\right. \\
\\
\left.\quad+\sum_{e \subset \partial \kappa \cap \Omega} \frac{h_{e}}{N_{e}}\left\|D-D_{\kappa}\right\|_{0, e}^{2}\left\|\frac{\partial u_{N, h}^{*}}{\partial n}\right\|_{0, \infty, e}^{2}\right\} \\
\quad+C\left\|\lambda^{*} u^{*}-\lambda_{N, h}^{*} u_{N, h}^{*}\right\|_{0, \Omega}^{2}
\end{array}
\end{align*}
$$

Lemma 14. Let $\alpha \in[0,1], \varepsilon>0$. Then there exists a constant $C(\varepsilon)>0$ independent of $h, \mathbf{N}$, and $\kappa$, such that

$$
\begin{align*}
\eta_{\alpha ; B_{\kappa}}^{2} \leq C(\varepsilon)\{ & N_{\kappa}^{2(1-\alpha)}\left\|u-u_{N, h}\right\|_{1, \kappa}^{2} \\
& +N_{\kappa}^{\max \{1+2 \varepsilon-2 \alpha, 0\}} \frac{h_{\kappa}^{2}}{N_{\kappa}^{2}} \\
& \left.\times\left(\left\|\lambda_{N, h} u_{N, h}-\lambda u+L u_{N, h}-L_{\kappa} u_{N, h}\right\|_{0, k}^{2}\right)\right\} . \tag{80}
\end{align*}
$$

Proof. We denote $v_{\kappa}:=\left(-L_{\kappa} u_{N, h}+\lambda_{N, h} u_{N, h}\right) \Phi_{\kappa}^{\alpha} \in H_{0}^{1}(\kappa)$ with $\alpha \in(0,1]$ and extend $v_{\kappa}$ to $\Omega$ by $v_{\kappa}=0$ on $\Omega \backslash \kappa$; then

$$
\begin{align*}
\left\|v_{\kappa} \Phi_{\kappa}^{-\alpha / 2}\right\|_{0, \kappa}^{2}= & \int_{\kappa}\left(-L_{\kappa} u_{N, h}+\lambda_{N, h} u_{N, h}\right) \overline{v_{\kappa}} \\
= & -\int_{\kappa}\left(L_{\kappa} u_{N, h}\right) \overline{v_{\kappa}}+a\left(u, v_{\kappa}\right) \\
& +\int_{\kappa}\left(\lambda_{N, h} u_{N, h}-\lambda u\right) \overline{v_{\kappa}} \\
= & a\left(u-u_{N, h}, v_{\kappa}\right) \\
& +\int_{\kappa}\left(\lambda_{N, h} u_{N, h}-\lambda u+L u_{N, h}-L_{\kappa} u_{N, h}\right) \overline{v_{\kappa}} \\
\leq & C\left\|u-u_{N, h}\right\|_{1, \kappa}\left|v_{\kappa}\right|_{1, \kappa} \\
& +\|\left(\lambda_{N, h} u_{N, h}-\lambda u\right. \\
& \left.+L u_{N, h}-L_{\kappa} u_{N, h}\right) \Phi_{\kappa}^{\alpha / 2} \|_{0, \kappa} \\
& \times\left\|v_{\kappa} \Phi_{\kappa}^{-\alpha / 2}\right\|_{0, \kappa} . \tag{81}
\end{align*}
$$

We consider the $H^{1}$ semi norm for $v_{\kappa}$. Using the polynomial inverse estimates (62)-(63) in Lemma 10, by transformation between the reference element $\widehat{\kappa}$ and $\kappa$, we find for $\alpha>1 / 2$ that

$$
\begin{align*}
\left|v_{\kappa}\right|_{1, \kappa}^{2} \leq & 2 \int_{\kappa} \Phi_{\kappa}^{2 \alpha}\left|\nabla\left(\lambda_{N, h} u_{N, h}-L_{\kappa} u_{N, h}\right)\right|^{2} \\
& +2 \int_{\kappa}\left|\nabla \Phi_{\kappa}^{\alpha}\right|^{2}\left|\lambda_{N, h} u_{N, h}-L_{\kappa} u_{N, h}\right|^{2} \\
\leq & C \frac{N_{\kappa}^{2(2-\alpha)}}{h_{\kappa}^{2}} \int_{\kappa} \Phi_{\kappa}^{\alpha}\left|\lambda_{N, h} u_{N, h}-L_{\kappa} u_{N, h}\right|^{2} \\
& +C \frac{1}{h_{\kappa}^{2}} \int_{\kappa} \Phi_{\kappa}^{2(\alpha-1)}\left|\lambda_{N, h} u_{N, h}-L_{\kappa} u_{N, h}\right|^{2}  \tag{82}\\
\leq & C \frac{N_{\kappa}^{2(2-\alpha)}}{h_{\kappa}^{2}} \int_{\kappa} \Phi_{\kappa}^{\alpha}\left|\lambda_{N, h} u_{N, h}-L_{\kappa} u_{N, h}\right|^{2} \\
= & C N_{\kappa}^{2(1-\alpha)} \frac{N_{\kappa}^{2}}{h_{\kappa}^{2}}\left\|v_{\kappa} \Phi_{\kappa}^{-\alpha / 2}\right\|_{0, \kappa}^{2} .
\end{align*}
$$

Note that (62) is applicable since $\alpha>1 / 2$ implies $2(\alpha-1)>$ -1 ; thus, we set $\beta=\alpha, \alpha=2(\alpha-1)$ in (62); then the third inequality above holds.

Since $\eta_{\alpha ; B_{\kappa}}=h_{\kappa} / N_{\kappa}\left\|v_{\kappa} \Phi_{\kappa}^{-\alpha / 2}\right\|_{0, \kappa}$, we obtain

$$
\begin{align*}
\eta_{\alpha ; B_{\kappa}} \leq C\left(N_{\kappa}^{1-\alpha} \| u\right. & -u_{N, h} \|_{1, \kappa} \\
& +\frac{h_{\kappa}}{N_{\kappa}} \| L u_{N, h}-L_{\kappa} u_{N, h}  \tag{83}\\
& \left.+\lambda_{N, h} u_{N, h}-\lambda u \|_{0, \kappa}\right)
\end{align*}
$$

To obtain an upper bound in the case of $0 \leq \alpha \leq 1 / 2$, we use the polynomial inverse estimate (62) in Lemma 10; for $\beta>$ $1 / 2$, we derive from (62) that

$$
\begin{align*}
\frac{N_{\kappa}}{h_{\kappa}} \eta_{\alpha ; B_{\kappa}} & =\Phi_{\kappa}^{\alpha / 2}\left\|\left(\lambda_{N, h} u_{N, h}-L_{\kappa} u_{N, h}\right)\right\|_{0, \kappa} \\
\leq & C N_{\kappa}^{\beta-\alpha}\left\|\left(\lambda_{N, h} u_{N, h}-L_{\kappa} u_{N, h}\right) \Phi_{\kappa}^{\beta / 2}\right\|_{0, \kappa} \\
= & C N_{\kappa}^{\beta-\alpha} \frac{N_{\kappa}}{h_{\kappa}} \eta_{\beta ; B_{\kappa}} \\
\leq & C N_{\kappa}^{\beta-\alpha}\left(N_{\kappa}^{1-\beta} \frac{N_{\kappa}}{h_{\kappa}}\left\|u-u_{N, h}\right\|_{1, \kappa}\right. \\
& \left.\quad+\left\|\lambda_{N, h} u_{N, h}-\lambda u+L u_{N, h}-L_{\kappa} u_{N, h}\right\|_{0, \kappa}\right) \tag{84}
\end{align*}
$$

Setting $\beta=1 / 2+\varepsilon, \varepsilon>0$,

$$
\begin{align*}
\eta_{\alpha ; B_{\kappa}} \leq C(\varepsilon)\{ & N_{\kappa}^{1-\alpha}\left\|u-u_{N, h}\right\|_{1, \kappa} \\
& +N_{\kappa}^{1 / 2+\varepsilon-\alpha} \frac{h_{\kappa}}{N_{\kappa}} \\
& \left.\times\left\|\lambda_{N, h} u_{N, h}-\lambda u+L u_{N, h}-L_{\kappa} u_{N, h}\right\|_{0, \kappa}\right\} \tag{85}
\end{align*}
$$

We obtain the desired result immediately from (83) and (85).

In order to obtain a local upper bound for the error indicator $\eta_{\alpha ; \kappa}$, we consider the edge residual term $\eta_{\alpha ; E_{\kappa}}$. we introduce the set

$$
\begin{equation*}
\omega_{\kappa}:=\cup\left\{\kappa^{\prime} \mid \kappa^{\prime} \text { and } \kappa \text { share at least one edge }\right\} . \tag{86}
\end{equation*}
$$

Lemma 15. Let $\alpha \in[0,1], \varepsilon>0$. Then there exists a constant $C(\varepsilon)>0$ independent of $h, \mathbf{N}$, and $\kappa$, such that

$$
\begin{align*}
\eta_{\alpha ; E_{\kappa}}^{2} \leq & C(\varepsilon) N_{\kappa}^{\max (1-2 \alpha+2 \varepsilon, 0)} \\
& \times\left\{N_{\kappa}\left\|u-u_{N, h}\right\|_{1, \omega_{\kappa}}^{2}+N_{\kappa}^{2 \varepsilon} \frac{h_{\kappa}^{2}}{N_{\kappa}^{2}}\right. \\
& \left.\times\left\|\lambda_{N, h} u_{N, h}-\lambda u+L u_{N, h}-L_{\kappa} u_{N, h}\right\|_{0, \omega_{\kappa}}^{2}\right\} \tag{87}
\end{align*}
$$

Proof. We will use weight functions on edge and a suitable extension operator. For a given element $\kappa$ with edge $e$, we choose the element $\kappa_{1}$ so that $\partial \kappa_{1} \cap \partial \kappa=e$. Denote $\overline{\kappa_{e}}$ := $\overline{\kappa_{1}} \cup \bar{\kappa}$; we construct a function $w_{e} \in H_{0}^{1}\left(\kappa_{e}\right)$ with $\left.w_{e}\right|_{e}=$ $D_{\kappa}\left[\partial u_{N, h} / \partial n\right] \Phi_{e}^{\alpha}$ as follows.

Let $v_{\widehat{e}}=C_{e} D_{\kappa}\left[\partial u_{N, h} / \partial n\right] \Phi_{\widehat{e}}^{\alpha}\left(C_{e}\right.$ is defined by (71)). Using Lemma 11, we extend $v_{\hat{e}}$ to $\widehat{\kappa}$, where the polynomial $\pi$ corresponds to $C_{e} D_{\kappa}\left[\partial u_{N, h} / \partial n\right]$. Define $\left.w_{e}\right|_{\kappa}$ and $\left.w_{e}\right|_{\kappa_{1}}$ as the affine transformation of $v_{\widehat{e}}$ in $\widehat{\kappa}$; Thus, $w_{e}$ is a piecewise $H^{1}$ function. From (64), we know $w_{e}$ vanishes on $\partial \kappa_{e}$; Therefore, $w_{e} \in H_{0}^{1}\left(\kappa_{e}\right)$. It is trivial to extend $w_{e}$ to $\Omega$, such that $w_{e}=0$ in $\Omega \backslash \kappa_{e}$. We find

$$
\begin{align*}
\| D_{\kappa} & {\left[\frac{\partial u_{N, h}}{\partial n}\right] \Phi_{e}^{\alpha / 2} \|_{0, e}^{2} } \\
= & \int_{e} D\left[\frac{\partial u_{N, h}}{\partial n}\right] \overline{w_{e}}+\int_{e}\left(D_{\kappa}-D\right)\left[\frac{\partial u_{N, h}}{\partial n}\right] \overline{w_{e}} \\
= & \int_{\mathcal{K}_{e}} L u_{N, h} \overline{w_{e}}-a\left(u_{N, h}, w_{e}\right) \\
& +\int_{e}\left(D_{\kappa}-D\right)\left[\frac{\partial u_{N, h}}{\partial n}\right] \overline{w_{e}} \\
= & \int_{\mathcal{K}_{e}}\left(L u_{N, h}-\lambda u\right) \overline{w_{e}}+a\left(u-u_{N, h}, w_{e}\right)  \tag{88}\\
& +\int_{e}\left(D_{\kappa}-D\right)\left[\frac{\partial u_{N, h}}{\partial n}\right] \overline{w_{e}} \\
\leq & \left\|L u_{N, h}-\lambda u\right\|_{0, \kappa_{e}}\left\|w_{e}\right\|_{0, \kappa_{e}} \\
& +C\left\|u-u_{N, h}\right\|_{1, \kappa_{e}}\left|w_{e}\right|_{1, \kappa_{e}} \\
& +\left\|\frac{\left(D_{\kappa}-D\right)}{D_{\kappa}}\right\|_{0, \infty, e}\left\|D_{\kappa}\left[\frac{\partial u_{N, h}}{\partial n}\right] \Phi_{e}^{\alpha / 2}\right\|_{0, e}^{2}
\end{align*}
$$

Therefore,

$$
\begin{align*}
& \left\|D_{\kappa}\left[\frac{\partial u_{N, h}}{\partial n}\right] \Phi_{e}^{\alpha / 2}\right\|_{0, e}^{2} \\
& \leq  \tag{89}\\
& \quad C\left\|L u_{N, h}-\lambda u\right\|_{0, \kappa_{e}}\left\|w_{e}\right\|_{0, \kappa_{e}} \\
& \quad+C\left\|u-u_{N, h}\right\|_{1, k_{e}}\left|w_{e}\right|_{1, \kappa_{e}}
\end{align*}
$$

We consider the case of $\alpha \in(1 / 2,1]$ first. Using the affine equivalence and (65)-(66) in Lemma 11, we obtain the upper bounds for $\left\|w_{e}\right\|_{0, \kappa_{e}}$ and $\left|w_{e}\right|_{1, \kappa_{e}}$ as follows:

$$
\begin{aligned}
\left|w_{e}\right|_{1, \kappa_{e}}^{2} \leq & C \frac{1}{h_{\kappa}}\left(\varepsilon N_{\kappa}^{2(2-\alpha)}+\varepsilon^{-1}\right) \\
& \times\left\|D_{\kappa}\left[\frac{\partial u_{N, h}}{\partial n}\right] \Phi_{e}^{\alpha / 2}\right\|_{0, e}^{2} \\
\left\|w_{e}\right\|_{0, \kappa_{e}} \leq & C h_{\kappa} \varepsilon\left\|D_{\kappa}\left[\frac{\partial u_{N, h}}{\partial n}\right] \Phi_{e}^{\alpha / 2}\right\|_{0, e}^{2} .
\end{aligned}
$$

It follows from (89)-(90) that

$$
\begin{align*}
&\left\|D_{\kappa}\left[\frac{\partial u_{N, h}}{\partial n}\right] \Phi_{e}^{\alpha / 2}\right\|_{0, e} \\
& \leq C \\
&\left\{\left(\frac{1}{h_{\kappa}}\left(\varepsilon N_{\kappa}^{2(2-\alpha)}+\varepsilon^{-1}\right)\right)^{1 / 2}\right.  \tag{91}\\
&\left.\times\left\|u-u_{N, h}\right\|_{1, \kappa_{e}}+\left(h_{\kappa} \varepsilon\right)^{1 / 2}\| \| L u_{N, h}-\lambda u \|_{0, \kappa_{e}}\right\}
\end{align*}
$$

By the definition of $\eta_{\alpha ; E_{\kappa}}^{2}$ and setting $\alpha=0$ in Lemma 14, we get

$$
\left.\begin{array}{rl}
\| L_{\kappa} u_{N, h}- & \lambda_{N, h} u_{N, h} \|_{0, k}^{2} \\
\leq C(\varepsilon) & \left\{\frac{N_{\kappa}^{4}}{h_{\kappa}^{2}}\left\|u-u_{N, h}\right\|_{1, k}^{2}+N_{\kappa}^{1+2 \varepsilon}\right. \tag{92}
\end{array}\right\}
$$

by the triangle inequality

$$
\begin{align*}
& \left\|L u_{N, h}-\lambda u\right\|_{0, \kappa_{e}} \\
& \quad \leq  \tag{93}\\
& \quad\left\|L_{\kappa} u_{N, h}-\lambda_{N, h} u_{N, h}\right\|_{0, \kappa_{e}} \\
& \quad+\left\|\lambda_{N, h} u_{N, h}-\lambda u+L u_{N, h}-L_{\kappa} u_{N, h}\right\|_{0, \kappa_{e}}
\end{align*}
$$

Combining the three inequalities above and summing, we have

$$
\begin{align*}
\eta_{\alpha ; E_{\kappa}}^{2} \leq & C\left\{\frac{1}{N_{\kappa}}\left(\varepsilon N_{\kappa}^{2(2-\alpha)}+\varepsilon^{-1}\right)+N_{\kappa}^{3} \varepsilon\right\} \\
& \times\left\|u-u_{N, h}\right\|_{1, \omega_{\kappa}}^{2}+C \varepsilon N_{\kappa}^{2(1+\varepsilon)} \frac{h_{\kappa}^{2}}{N_{\kappa}^{2}}  \tag{94}\\
& \times\left\|\lambda_{N, h} u_{N, h}-\lambda u+L u_{N, h}-L_{\kappa} u_{N, h}\right\|_{0, \omega_{\kappa}}^{2}
\end{align*}
$$

Setting $\varepsilon=1 / N_{\kappa}^{2}$ in the above inequality yields the assertion for $\alpha>1 / 2$. For the case of $\alpha \in[0,1 / 2]$, we set $\beta=1 / 2+\varepsilon$, use (62) in Lemma 10 to get $\eta_{\alpha ; E_{\kappa}} \leq N_{\kappa}^{\beta-\alpha} \eta_{\beta ; E_{\kappa}}$, and find the desired result.

Combining Lemmas 14 and 15, we obtain the following theorem.

Theorem 16. Let $\alpha \in[0,1], \varepsilon>0$. Then there exists a constant $C>0$ independent of $h, \mathbf{N}$, and $\kappa$, such that

$$
\begin{align*}
\eta_{\alpha ; \kappa}^{2} \leq & C(\varepsilon) N_{\kappa}^{\max (1-2 \alpha+2 \varepsilon, 0)} \\
& \times\left\{N_{\kappa}\left\|u-u_{N, h}\right\|_{1, \omega_{\kappa}}^{2}+N_{\kappa}^{2 \varepsilon} \frac{h_{\kappa}^{2}}{N_{\kappa}^{2}}\right.  \tag{95}\\
& \left.\times\left\|\lambda_{N, h} u_{N, h}-\lambda u+L u_{N, h}-L_{\kappa} u_{N, h}\right\|_{0, \omega_{\kappa}}^{2}\right\} .
\end{align*}
$$

Similarly, we have Theorem 17.

Theorem 17. Let $\alpha \in[0,1], \varepsilon>0$. Then there exists a constant $C>0$ independent of $h, \mathbf{N}$, and $\kappa$, such that

$$
\begin{align*}
\eta_{\alpha ; \kappa}^{* 2} \leq & C(\varepsilon) N_{\kappa}^{\max (1-2 \alpha+2 \varepsilon, 0)} \\
& \times\left\{N_{\kappa}\left\|u^{*}-u_{N, h}^{*}\right\|_{1, \omega_{\kappa}}^{2}+N_{\kappa}^{2 \varepsilon} \frac{h_{\kappa}^{2}}{N_{\kappa}^{2}}\right. \\
& \left.\times\left\|\lambda_{N, h}^{*} u_{N, h}^{*}-\lambda^{*} u^{*}+L^{*} u_{N, h}^{*}-L_{\kappa}^{*} u_{N, h}^{*}\right\|_{0, \omega_{\kappa}}^{2}\right\} \tag{96}
\end{align*}
$$

In order to estimate bounds of $\left|\lambda-\lambda_{N, h}\right|$, we also need Lemma 18 (see [8, 10]).

Lemma 18. Let $(\lambda, u)$ be an eigenpair of (24), and let $\left(\lambda^{*}=\right.$ $\left.\bar{\lambda}, u^{*}\right)$ be the associated eigenpair of the adjoint problem (31). Then for all $w, w^{*} \in H_{0}^{1}(\Omega),\left(w, w^{*}\right) \neq 0$,

$$
\begin{align*}
& \frac{a\left(w, w^{*}\right)}{\left(w, w^{*}\right)}-\lambda  \tag{97}\\
& \quad=\frac{a\left(w-u, w^{*}-u^{*}\right)}{\left(w, w^{*}\right)}-\lambda \frac{\left(w-u, w^{*}-u^{*}\right)}{\left(w, w^{*}\right)}
\end{align*}
$$

Theorem 19. Under the assumptions of Theorem 7 , we assume that $D, \mathbf{b}$, and $c$ are smooth enough, and let $\alpha \in[0,1]$. Then there exists an eigenpair $(\lambda, u)$ of $(24)$, such that

$$
\begin{gather*}
\left\|u-u_{N, h}\right\|_{1, \Omega} \leq C\left(\sum_{\kappa \in K_{h}} N_{\kappa}^{2 \alpha} \eta_{\alpha ; \kappa}^{2}\right)^{1 / 2},  \tag{98}\\
\eta_{\alpha ; \kappa}^{2} \leq C(\varepsilon) N_{\kappa}^{\max (2-2 \alpha+2 \varepsilon, 1)}\left\|u-u_{N, h}\right\|_{1, \omega_{\kappa}}^{2} . \tag{99}
\end{gather*}
$$

Further let the ascent of $\lambda_{N, h}$ be $l=1$, and let $\left(\lambda_{N, h}^{*}, u_{N, h}^{*}\right)$ be the corresponding adjoint eigenpair of (32), then there exists an adjoint eigenpair $\left(\lambda^{*}, u^{*}\right)$ of (31), such that

$$
\begin{equation*}
\left|\lambda_{N, h}-\lambda\right| \leq C\left(\sum_{\kappa \in K_{h}} N_{\kappa}^{2 \alpha} \eta_{\alpha ; \kappa}^{2}\right)^{1 / 2}\left(\sum_{\kappa \in K_{h}} N_{\kappa}^{2 \alpha} \eta_{\alpha ; \kappa}^{* 2}\right)^{1 / 2} \tag{100}
\end{equation*}
$$

Particularly, if the eigenvalue problem (23) is symmetric (i.e., $\mathbf{b}=0$ ), then

$$
\begin{equation*}
C C(\varepsilon)^{-1} \sum_{\kappa \in K_{h}} N_{\kappa}^{\min (2 \alpha-2-2 \varepsilon,-1)} \eta_{\alpha ; \kappa}^{2} \leq\left|\lambda_{N, h}-\lambda\right| . \tag{101}
\end{equation*}
$$

Proof. We know from the assumption $D, c \in H^{t_{1}}(\kappa), \mathbf{b} \in$ $\left(H^{t_{1}}(\kappa)\right)^{2}$. By the interpolation error estimates (14) and (15), we have

$$
\begin{equation*}
\left\|L_{\kappa} u_{N, h}-L u_{N, h}\right\|_{0, \kappa} \leq C h_{\kappa}^{\min \left(N_{\kappa}+1, t_{1}\right)-1} N_{\kappa}^{-t_{1}+1} \tag{102}
\end{equation*}
$$

From $D \in H^{t_{1}}(\kappa)$, we know that $D \in H^{t_{1}-1 / 2}(e)$. By the interpolation error estimate on edge of element (see formula (5.4.42) in [2]), we get

$$
\begin{equation*}
\left\|D-D_{\kappa}\right\|_{0, e} \leq C h_{e}^{\min \left(N_{\kappa}+1, t_{1}-1 / 2\right)} N_{\kappa}^{-t_{1}+1 / 2} \tag{103}
\end{equation*}
$$

Note that the formula (51) gives the optimal orders of convergence; thus, we deduce that the second and third terms on the right side of (74) are higher order infinitesimals. We derive from (52) and (53), and $N=N_{\kappa}$, that

$$
\begin{align*}
\| \lambda u & -\lambda_{N, h} u_{N, h} \|_{0, \Omega} \\
\leq & \left|\lambda-\lambda_{N, h}\right|\|u\|_{0, \Omega} \\
& +\left|\lambda_{N, h}\right|\left\|u-u_{N, h}\right\|_{0, \Omega}  \tag{104}\\
\leq & \frac{C h^{\tau_{1}+\tau_{2}-2}}{N^{t_{1}+t_{2}-2}}+\frac{C h^{r_{2}+\tau_{1}-1}}{N^{r_{2}+t_{1}-1}} \leq \frac{C h^{r_{2}+\tau_{1}-1}}{N^{r_{2}+t_{1}-1}} .
\end{align*}
$$

Therefore, the fourth term on the right side of (74) is also a higher order infinitesimal. Up to higher order terms, we get (98). We ignore higher order infinitesimals in (95) and get (99). From Lemma 4 in [10], we know that $\left(u_{N, h}, u_{N, h}^{*}\right)=1$ and $u_{N, h}^{*}$ is uniformly bounded with $h$ and $N$. By the same argument of (98), we can deduce that

$$
\begin{equation*}
\left\|u^{*}-u_{N, h}^{*}\right\|_{1, \Omega}^{2} \leq C \sum_{\kappa \in K_{h}} N_{\kappa}^{2 \alpha} \eta_{\alpha ; \kappa}^{* 2} \tag{105}
\end{equation*}
$$

From (97), we have

$$
\begin{align*}
& \frac{a\left(u_{N, h}, u_{N, h}^{*}\right)}{\left(u_{N, h}, u_{N, h}^{*}\right)}-\lambda \\
& \quad=\frac{a\left(u_{N, h}-u, u_{N, h}^{*}-u^{*}\right)}{\left(u_{N, h}, u_{N, h}^{*}\right)}  \tag{106}\\
& \quad-\lambda \frac{\left(u_{N, h}-u, u_{N, h}^{*}-u^{*}\right)}{\left(u_{N, h}, u_{N, h}^{*}\right)}
\end{align*}
$$

that is,

$$
\begin{align*}
\lambda_{N, h}-\lambda= & a\left(u_{N, h}-u, u_{N, h}^{*}-u^{*}\right)  \tag{107}\\
& -\lambda\left(u_{N, h}-u, u_{N, h}^{*}-u^{*}\right) .
\end{align*}
$$

Substituting (98) and (105) into the above equality, we obtain (100).

If the eigenvalue problem (23) is symmetric (i.e., $\mathbf{b}=0$ ), then

$$
\begin{align*}
\lambda_{N, h}-\lambda= & a\left(u_{N, h}-u, u_{N, h}-u\right) \\
& -\lambda\left(u_{N, h}-u, u_{N, h}-u\right) . \tag{108}
\end{align*}
$$

Up to higher order term $\lambda\left(u_{N, h}-u, u_{N, h}-u\right)$, by (99) we get (101).

Remark 20. Babuška and Osborn [17] have discussed hp finite element approximation with simplex partition for eigenvalue problems. Obviously, the Interpolation estimates (14) and (15) hold for hp finite element with simplex partition (see [24]). Therefore, our theoretical results of spectral methods and spectral methods for eigenvalue problems, which have been discussed in Sections 3 and 4, hold for hp finite element with simplex partition.

Table 1: Errors of LGL-SM, modal, and Eq-SM for 1st eigenvalue.

| $N$ | DOF | LGL-SM <br> $\lambda_{1}$ | Modal-SM <br> $\lambda_{1}$ | Eq-SM <br> $\lambda_{1}$ |
| :---: | :---: | :---: | :---: | :---: |
| 4 | 9 | $5.19 E+00$ | $5.19 E+00$ | $5.19 E+00$ |
| 5 | 16 | $4.51 E-01$ | $4.51 E-01$ | $4.51 E-01$ |
| 6 | 25 | $7.68 E-03$ | $7.68 E-03$ | $7.68 E-03$ |
| 7 | 36 | $1.07 E-05$ | $1.07 E-05$ | $1.07 E-05$ |
| 8 | 49 | $1.21 E-05$ | $1.21 E-05$ | $1.21 E-05$ |
| 9 | 64 | $9.16 E-07$ | $9.16 E-07$ | $9.16 E-07$ |
| 10 | 81 | $2.46 E-08$ | $2.46 E-08$ | $2.48 E-08$ |
| 11 | 100 | $2.91 E-10$ | $2.91 E-10$ | $4.35 E-09$ |
| 12 | 121 | $9.31 E-13$ | $1.06 E-12$ | $2.79 E-08$ |
| 13 | 144 | $5.68 E-14$ | $1.28 E-13$ | $1.41 E-07$ |
| 14 | 169 | $2.84 E-14$ | $1.28 E-13$ | $2.28 E-06$ |
| 15 | 196 | $7.82 E-14$ | $2.13 E-14$ | $3.60 E-05$ |

## 5. Numerical Experiments

In this section, we simply denote spectral methods, spectral element methods, and finite element methods with SM, SEM, and FEM, respectively. And spectral methods with equidistant nodal basis, modal basis, and LGL nodal basis are replaced by Eq-SM, Modal-SM, and LGL-SM, respectively. Note that all these methods employ the tensorial basis.

In our experiment, we compute $1 /\left|\left(u_{N, h}, u_{N, h}^{*}\right)\right|$ as condition number for simple eigenvalue (see Remark 2.1 in [25]), where $u_{N, h}$ and $u_{N, h}^{*}$ are eigenfunctions of eigenvalue problem (25) and its adjoint problem (32) normalized with $\|\cdot\|_{0, \Omega}$, respectively.
5.1. Example 1. Consider the nonsymmetric eigenvalue problem

$$
\begin{gather*}
-\Delta u+10 u_{x}+u_{y}=\lambda u, \quad \text { in } \Omega=(0,1)^{2},  \tag{109}\\
u=0, \quad \text { on } \partial \Omega .
\end{gather*}
$$

The first eigenvalue of (109) $\lambda_{1}=101 / 4+2 \pi^{2}$ is a simple eigenvalue. And the corresponding eigenfunctions are sufficiently smooth.
5.1.1. Comparisons between LGL-SM, Modal, and Eq-SM. Figure 1 shows that the condition numbers of the first eigenvalue for LGL-SM, Modal-SM, and Eq-SM coincide with each other at the beginning but perform abnormally with $N>19$ for Eq-SM. Table 1 tells us that when $N>11$, the accuracy of first eigenvalue obtained by Eq-SM is not as good as obtained that by LGL-SM and Modal-SM. When $N=15$, the error of the first eigenvalues obtained by Eq-SM is greater than $1 \mathrm{E}-5$; however, the order of the magnitude of errors for LGL-SM and Modal-SM still keeps below 1E-13. The best result of first eigenvalue error for $\mathrm{Eq}-\mathrm{SM}$ is merely $\mathrm{IE}-9$ or so.
5.1.2. LGL-SM and Modal-SM versus hp-SEM. Tables 1 and 2 indicate that increasing the polynomial degree $N$ or


Figure 1: Condition number of first eigenvalue for SM.
decreasing the mesh fineness $h$ can decrease the errors of the first eigenvalue. But it is expensive to increase polynomial degree and decrease mesh fineness $h$ at the same time. For $h=1 / 4$ and $h=1 / 16$, we obtain from Table 2 the first eigenvalue errors $2.8 E-14$ and $1.3 E-13$ and the corresponding degree of freedom 1225 and 6241 for hp-SEM, respectively, Whereas from Table 1, to reach this accuracy, LGL-SM and Modal-SM should merely perform the interpolation approximations with polynomial degree bi-14 and bi-13 or so, and the corresponding degrees of freedom are merely 169 and 144, respectively. Therefore, we conclude that LGL-SM and Modal-SM are highly accurate and efficient for this kind of nonsymmetric eigenvalue problems.

In Figure 2 from [9], when the degree of freedom is up to 1000 , the error of linear FEM is about $1 \mathrm{E}-2$; the function value recovery techniques in [9] obviously improves the accuracy up to 1E-5. Comparing Tables 1 and 2 in this paper with Figure 2 in [9], we can also find the advantages of LGL-SM, ModalSM, and hp-SEM over the function value recovery techniques for FEM given by [9] from accuracy and degree of freedom.
5.1.3. $h p$-SEM versus $h p-F E M$. From Table 4, we find that the condition number of the first eigenvalue for hp -version methods (hp-SEM and hp-FEM) stays at 4.27. It is indicated from Tables 2 and 3 that, when $N$ is greater than 7 , compared with hp-SEM, the errors of hp-FEM tend to become large, whereas the errors of hp-SEM still keep stable or even stay a decreasing tendency; however, this phenomenon is not apparent for $h=1 / 2$.

Remark 21. Condition numbers of 1st eigenvalue for hp-FEM (not listed in Table 4) are almost the same to those for hpSEM.

TABLE 2: Errors and DOF of hp-SEM for the first eigenvalue.

| $h=1 / 2$ | $h=1 / 4$ |  |  |  |  |  |  |  |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  | DOF | Error | DOF | $h=1 / 8$ |  |  |  |  |
|  | Error | DOF | Error | DOF |  |  |  |  |
| 2 | $5.18 E+00$ | 9 | $2.54 E-01$ | 49 | $1.50 E-02$ | 225 | $9.00 E-04$ | 961 |
| 3 | $7.00 E-03$ | 25 | $6.10 E-04$ | 121 | $1.20 E-05$ | 529 | $1.90 E-07$ | 2209 |
| 4 | $8.40 E-03$ | 49 | $2.60 E-05$ | 225 | $9.70 E-08$ | 961 | $3.70 E-10$ | 3969 |
| 5 | $1.64 E-04$ | 81 | $1.60 E-07$ | 361 | $1.50 E-10$ | 1521 | $1.30 E-13$ | 6241 |
| 6 | $4.10 E-07$ | 121 | $2.30 E-11$ | 529 | $9.90 E-13$ | 2209 | $3.60 E-12$ | 9025 |
| 7 | $3.10 E-08$ | 169 | $1.70 E-12$ | 729 | $3.10 E-13$ | 3025 | $1.60 E-12$ | 12321 |
| 8 | $1.90 E-10$ | 225 | $1.90 E-13$ | 961 | $2.10 E-12$ | 3969 | $4.80 E-12$ | 16129 |
| 9 | $5.50 E-13$ | 289 | $2.80 E-14$ | 1225 | $6.00 E-13$ | 5041 | $1.10 E-12$ | 20449 |
| 10 | $3.80 E-13$ | 361 | $1.10 E-12$ | 1521 | $4.40 E-12$ | 6241 | $1.50 E-11$ | 25281 |

Table 3: Errors of hp-FEM for the first eigenvalue.

| $N$ | $h=1 / 2$ | $h=1 / 4$ | $h=1 / 8$ | $h=1 / 16$ |
| :--- | :---: | :---: | :---: | :---: |
| 3 | $7.00 E-03$ | $6.10 E-04$ | $1.20 E-05$ | $1.90 E-07$ |
| 4 | $8.40 E-03$ | $2.60 E-05$ | $9.70 E-08$ | $3.70 E-10$ |
| 5 | $1.60 E-04$ | $1.60 E-07$ | $1.50 E-10$ | $1.30 E-12$ |
| 6 | $4.10 E-07$ | $2.40 E-11$ | $3.60 E-13$ | $8.60 E-12$ |
| 7 | $3.10 E-08$ | $6.10 E-12$ | $1.30 E-11$ | $3.00 E-11$ |
| 8 | $1.80 E-10$ | $3.10 E-11$ | $2.30 E-10$ | $2.10 E-10$ |
| 9 | $7.50 E-11$ | $3.40 E-11$ | $6.80 E-10$ | $7.40 E-10$ |
| 10 | $2.50 E-11$ | $9.90 E-10$ | $8.70 E-09$ | $6.60 E-09$ |
| 11 | $2.00 E-09$ | $9.60 E-09$ | $8.90 E-09$ | $5.40 E-07$ |

Table 4: Condition number of first eigenvalue for hp-SEM.

| $N$ | $h=1 / 2$ | $h=1 / 4$ | $h=1 / 8$ | $h=1 / 16$ |
| :--- | :---: | :---: | :---: | :---: |
| 3 | 4.284381324 | 4.270132842 | 4.269625046 | 4.269615821 |
| 4 | 4.267343095 | 4.269607452 | 4.269615638 | 4.26961567 |
| 5 | 4.269636446 | 4.269615725 | 4.26961567 | 4.26961567 |
| 6 | 4.269619135 | 4.26961567 | 4.26961567 | 4.26961567 |
| 7 | 4.269615617 | 4.26961567 | 4.26961567 | 4.26961567 |
| 8 | 4.26961567 | 4.26961567 | 4.26961567 | 4.26961567 |
| 9 | 4.26961567 | 4.26961567 | 4.26961567 | 4.26961567 |

### 5.1.4. Validity of the Error Indicator. Denote

$$
\begin{equation*}
\psi_{\alpha}=\left(\sum_{\kappa \in K_{h}} N_{\kappa}^{2 \alpha} \eta_{\alpha ; \kappa}^{2}\right)^{1 / 2}\left(\sum_{\kappa \in K_{h}} N_{\kappa}^{2 \alpha} \eta_{\alpha ; \kappa}^{* 2}\right)^{1 / 2} \tag{110}
\end{equation*}
$$

From Theorem 19, we know that $\psi_{\alpha}$ is a reliable error indicator for $\lambda_{N, h}$. We choose $\psi_{0}$ (setting $\alpha=0$ in (110)) as a posteriorii error indicator.

In Figures 2 and 3, we denote the true error and est. error with $\left|\lambda_{N, h}-\lambda\right|$ and $\psi_{0}$, respectively.

As is depicted in Figure 2, when the polynomial degree $N \leq 12$, the error indicator $\psi_{0}$ can properly estimate the true errors of LGL-SM for the first eigenvalue, however, also slightly underestimate the true errors. It is easy to see that $\psi_{0}$ shows almost the same algebraic decay as the true error with the polynomial degree $N(\leq 12)$ increasing. Nevertheless, the error indicator $\psi_{0}$ cannot approximate the true errors if $N$

Table 5: The Approximate eigenvalues and indicator $\psi_{0}$ of P-SEM.

| $N$ | $\lambda_{N, h}$ | $\psi_{0}$ |
| :--- | :---: | :---: |
| 3 | 28.56900 | $2.72 E+01$ |
| 4 | 31.99175 | $3.49 E+00$ |
| 5 | 34.82082 | $2.25 E-01$ |
| 6 | 34.65087 | $1.31 E-02$ |
| 7 | 34.65057 | $3.32 E-03$ |
| 8 | 34.64765 | $1.92 E-03$ |
| 9 | 34.64567 | $1.22 E-03$ |
| 10 | 34.64432 | $8.11 E-04$ |
| 11 | 34.64335 | $5.62 E-04$ |
| 12 | 34.64265 | $4.02 E-04$ |
| 13 | 34.64212 | $2.95 E-04$ |
| 14 | 34.64171 | $2.22 E-04$ |
| 15 | 34.64139 | $1.71 E-04$ |
| 16 | 34.64114 | $1.33 E-04$ |
| 17 | 34.64094 | $1.06 E-04$ |
| 18 | 34.64078 | $8.49 E-05$ |

is large enough, which is caused by round-off errors derived from the bad condition number of eigenvalue. In Figure 3, we give the comparison between the error indicator $\psi_{0}$ and the true errors for hp-SEM.
5.2. Example 2. Consider the nonsymmetric eigenvalue problem

$$
\begin{align*}
&-\Delta u+10 u_{x}=\lambda u, \quad \text { in } \Omega=\frac{(-1,1)^{2}}{(0,1)^{2}}  \tag{111}\\
& u=0, \quad \text { on } \partial \Omega
\end{align*}
$$

A reference value for the first eigenvalue (simple eigenvalue) of (111) is 34.6397 given by [5]. And the corresponding eigenfunctions have the singularity at the origin. Next, we shall compare the relevant numerical results between P-SEM and the other methods adopted in this paper. Note that here and hereafter P-version methods are for the fixed mesh fineness $h=1$. Table 5 lists part data of the approximate eigenvalues computed by P-SEM and the corresponding error indicator $\psi_{0}$ for reference.


Figure 2: The Error indicator $\psi_{0}$ of LGL-SM.


Figure 3: The Error indicator $\psi_{0}$ of hp-SEM ( $h=1 / 2$ ).
5.2.1. Stability of $P$-Version Methods. Figure 4 indicates that the eigenvalues computed by P-FEM will not seriously deviate from the results computed by P-SEM until the interpolation polynomial degree $N$ is up to 19. This phenomenon coincides with the abnormity of condition number of first eigenvalue for P-FEM (see Figure 5). The reason is that the singularities of the eigenfunctions limit the accuracy of both kinds of methods; this is slightly different from the case of the eigenvalue problem with the sufficiently smooth eigenfunctions.
5.2.2. P-SEM versus Other Methods. By calculations, we find that, in the case of the linear FEM, for fixed mesh fineness $h=1 / 256$, the approximate eigenvalue is 34.6403 with degree of freedom up to 195585. But P-SEM with the polynomial degree bi-22 can reach this accuracy, and the corresponding


Figure 4: The Approximate 1st eigenvalue of P-SEM and P-FEM.


Figure 5: Condition number of first eigenvalue for P-SEM and PFEM.
degree of freedom is merely 1365 . Compared with the linear FEM, hp-SEM can obtain a higher accuracy with less degrees of freedom as follows: for fixed $h=1 / 16$ and $N=10$, the approximate eigenvalue is 34.63984 with degree of freedom 76161 but P-SEM with polynomial degree bi-44 can reach this accuracy. Therefore, P-SEM is more efficient for the eigenvalue problems with the singular solutions than the other methods.

## Acknowledgments

This work was supported by the National Natural Science Foundation of China (Grant no. 11161012) and the Educational

Administration and Innovation Foundation of Graduate Students of Guizhou Normal University (no. 2012(11)).

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## Research Article

# Nonlinear Hydroelastic Waves beneath a Floating Ice Sheet in a Fluid of Finite Depth 

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Received 21 May 2013; Revised 29 August 2013; Accepted 29 August 2013
Academic Editor: Rasajit Bera
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#### Abstract

The nonlinear hydroelastic waves propagating beneath an infinite ice sheet floating on an inviscid fluid of finite depth are investigated analytically. The approximate series solutions for the velocity potential and the wave surface elevation are derived, respectively, by an analytic approximation technique named homotopy analysis method (HAM) and are presented for the secondorder components. Also, homotopy squared residual technique is employed to guarantee the convergence of the series solutions. The present formulas, different from the perturbation solutions, are highly accurate and uniformly valid without assuming that these nonlinear partial differential equations (PDEs) have small parameters necessarily. It is noted that the effects of water depth, the ice sheet thickness, and Young's modulus are analytically expressed in detail. We find that, in different water depths, the hydroelastic waves traveling beneath the thickest ice sheet always contain the largest wave energy. While with an increasing thickness of the sheet, the wave elevation tends to be smoothened at the crest and be sharpened at the trough. The larger Young's modulus of the sheet also causes analogous effects. The results obtained show that the thickness and Young's modulus of the floating ice sheet all greatly affect the wave energy and wave profile in different water depths.


## 1. Introduction

In recent decades, the ice cover in the polar region has attracted more and more attention in the field of ocean engineering and polar engineering in view of their practical importance and theoretical investigations. The motivations for the research work are to study damage to offshore constructions by floating ice sheets, the transportation systems in the cold region where the ice cover can be considered as roads and aircraft runways and air-cushioned vehicles are used to break the ice, for example. One of the important problems in this field would appear to be the accurate measurement of the characteristics of waves traveling beneath a floating ice sheet. And such wave may have been generated in the ice cover itself by the wind, or it may have originated by a moving load on the ice sheets. Considerable work has been done since the first theoretical model of wave propagation in sea ice was
proposed by Greenhill [1] in 1887. A comprehensive summary on mathematical method and modeling for the problem can be found in some review articles such as Squire et al. [2, 3]. In addition to ice sheets, this work can apply to very large floating structures (VLFSs) such as floating airports, mobile offshore bases, offshore port facilities, offshore storage and waste disposal provisions, energy islands including some wave power configurations, and ultralarge ships, where there is an extensive complementary literature [4-6].

Most theoretical works on the problem are still in the scope of linear theory based on the assumption that the wave amplitudes generated are very small in comparison with the wave lengths. So such models are not appropriate to describe waves of arbitrary amplitude considered here. According to hydrodynamics and elasticity, we can construct the nonlinear partial differential equations (PDEs) (1)-(5) to describe nonlinear hydroelastic waves of arbitrary amplitude
traveling through water covered by an ice sheet in finite water depth. Unfortunately, it is very difficult to solve analytically the coupled nonlinear PDEs mathematically. Further, most of the most works literature on the nonlinear theory of sea waves ice sheet interaction are necessarily in the context of weakly nonlinear analysis due to the limitation of present mathematical tools. Now the main analytical study on such complex nonlinear PDEs still follows the well-known perturbation technique. For example, Forbes [7] derived nonlinear PDEs to describe two-dimensional periodic waves beneath an elastic sheet floating on the surface of an infinitely deep fluid. The periodic solutions are sought using the Fourier series and perturbation expansions for the Fourier coefficients. And it is found that the solutions have certain features in common with capillary-gravity waves. Following the framework in [7], Forbes continued his study of finite-amplitude surface waves beneath a floating elastic sheet in infinitely deep water [8], and optimized their previous perturbation technology directly by developing the Fourier coefficients as expansions in the wave height. Waves of extremely large amplitude are found to exist, and results are presented for waves belonging to several different nonlinear solution branches. Recently, Vanden-Broeck and Părău [9] further extended the results of Forbes for periodic waves to the arbitrary-amplitude waves. It is noted that perturbation and asymptotic approximations of nonlinear PDEs often break down as nonlinearity becomes strong. So the weakly nonlinear solutions of small-amplitude waves are derived by the perturbation approach, while fully nonlinear solutions of large-amplitude waves have to be calculated numerically by means of the numerical series truncation method in Vanden-Broeck's study.

Furthermore, perturbation and asymptotic techniques depend extremely on the small/large parameters in general, while our nonlinear PDEs have no any small/large parameters. Thus the perturbation techniques are not applicable to the nonlinear problem under consideration. In this paper, we apply a new analytic approximation method known as the homotopy analysis method (HAM) to effectively solve the nonlinear PDEs presented here. Based on the concept of homotopy in algebraic topology, the HAM was proposed by Liao [10] in 1992. Unlike the perturbation method, the HAM is entirely independent of any small/large parameters. Moreover, it provides us with extremely large freedom to choose base functions and initial approximations (16) and (17) of solutions and auxiliary linear operators (21)-(23) only under some basic rules [11, 12]. More importantly, in contrast to all other previous analytic techniques, the HAM provides us a convenient way to control and adjust the convergence of the approximate series solutions by means of introducing an auxiliary parameter $c_{0}$. The method has been systematically described by Liao [11, 12]. Recently the HAM has been successfully applied to the study of a number of classical nonlinear differential equations including nonlinear equations arising in fluid mechanics [13-18], heat transfer [19, 20], solitons and integrable models [21-24], and finance [25, 26]. These aforementioned studies show the validity and generality of the HAM for some highly nonlinear PDEs with multiple solutions, singularity, and unknown boundary conditions.

The objective of the present work is to analytically study the nonlinear hydroelastic waves under an ice sheet lying over an incompressible inviscid fluid of finite uniform depth by means of the HAM. According to the potential theory in hydrodynamics and elasticity, the nonlinear partial differential equations (PDEs) (1)-(5) are composed of the Laplace equation taken as the governing equation for inviscid flows, the kinematic and dynamic boundary conditions on the unknown ice sheet-water interface with a zero draft, a simple linear model for the thin sheet that includes the effects of flexural rigidity and vertical inertia, and a bottom boundary condition. The convergent homotopy-series solutions for the velocity potential and the wave surface elevation are formally derived by applying the HAM with the consideration of the minimum of the squared residual, respectively. It should be mentioned that we study the effects of the water depth and two important physical parameters including Young's modulus and the thickness of the ice sheet on the wave energy and its elevation in detail. Discussion and conclusions are made in Sections 4 and 5, respectively. All of results obtained will help enrich our understanding of nonlinear hydroelastic waves propagating under a floating ice sheet on a fluid of finite depth.

## 2. Mathematical Description

The problem under consideration is a train of nonlinear hydroelastic waves propagating beneath a two-dimensional infinite elastic plate floating on a fluid of finite depth $h$ and a zero draft. A Cartesian coordinate oxz is used in which the $z$-axis points vertically upward, while $z=0$ represents the undisturbed surface. We follow Greenhill in [1] assuming that this problem is capable of modeling ocean waves in the presence of sea ice when the fluid is inviscid and incompressible and the flow is irrotational, and the ice sheet is mathematically idealized as a thin elastic plate. Then the governing equations for a velocity potential $\phi(x, z, t)$ can be written as

$$
\begin{equation*}
\frac{\partial^{2} \phi}{\partial x^{2}}+\frac{\partial^{2} \phi}{\partial z^{2}}=0, \quad(-h \leq z \leq \zeta(x, t)) \tag{1}
\end{equation*}
$$

where $\zeta(x, t)$ is the wave surface elevation. The bottom boundary condition reads

$$
\begin{equation*}
\frac{\partial \phi}{\partial z}=0, \quad(z=-h) \tag{2}
\end{equation*}
$$

The motion of the fluid and the plate is coupled through the dynamic free-surface condition. We also assume that any particle which is once between the elastic plate and the water surface remains on it. So the kinematic and dynamic boundary conditions on the unknown surface $z=\zeta(x, t)$ are, respectively, modeled as

$$
\begin{gather*}
\frac{\partial \zeta}{\partial t}+\frac{\partial \phi}{\partial x} \frac{\partial \zeta}{\partial x}-\frac{\partial \phi}{\partial z}=0  \tag{3}\\
\frac{\partial \phi}{\partial t}+\frac{1}{2}|\nabla \phi|^{2}+\frac{p_{e}}{\rho}+g \zeta=0 \tag{4}
\end{gather*}
$$

where $p_{e}$ is the water-plate interface pressure, $\rho$ is the fluid density, and $g$ is the gravitational acceleration, for a thin homogeneous elastic plate with uniform mass density $\rho_{e}$ and constant thickness $d$.

Since we are considering long waves here, the linear Kirchhoff (Euler-Bernoulli) beam theory is applied to the floating elastic plate as follows:

$$
\begin{equation*}
p_{e}=D \frac{\partial^{4} \zeta}{\partial x^{4}}+m_{e}\left(\frac{\partial^{2} \zeta}{\partial t^{2}}+g\right) \tag{5}
\end{equation*}
$$

where $m_{e}=\rho_{e} d, D=E d^{3} /\left[12\left(1-v^{2}\right)\right]$ is the flexural rigidity of the plate, $E$ is the effective Young's modulus of the plate, and $\nu$ Poisson's ratio. We substitute (5) into (4) to derive a new form of the dynamic boundary condition as follows:

$$
\begin{equation*}
\frac{\partial \phi}{\partial t}+\frac{1}{2}|\nabla \phi|^{2}+g \zeta+\frac{1}{\rho}\left[D \frac{\partial^{4} \zeta}{\partial x^{4}}+m_{e}\left(\frac{\partial^{2} \zeta}{\partial t^{2}}+g\right)\right]=0 \tag{6}
\end{equation*}
$$

Here, we consider a train of nonlinear waves traveling beneath an elastic plate with constant wave number $k$ and constant angular frequency $\omega$ of the incident wave. For a general case it should be emphasized that, by means of the traveling-wave method directly, the progressive waves are transferred from the temporal differentiation into the spatial one, which is very different from the mathematical model obtained by simply eliminating the time-dependent terms from the kinematic and dynamic boundary conditions on the unknown free surface [7-9]. Namely, we introduce an independent variable transformation

$$
\begin{equation*}
X=k x-\omega t \tag{7}
\end{equation*}
$$

where the angular frequency $\omega$ and the wave number $k$ are given. Thus, we can express the potential function $\phi(x, z, t)=$ $\phi(X, z)$ and the traveling wave profile $\zeta(x, t)=\zeta(X)$.

Then the governing equation and the bottom boundary condition for the velocity potential are transformed, respectively, by

$$
\begin{gather*}
k^{2} \frac{\partial^{2} \phi}{\partial X^{2}}+\frac{\partial^{2} \phi}{\partial z^{2}}=0, \quad(-h \leq z \leq \zeta(X))  \tag{8}\\
\frac{\partial \phi}{\partial z}=0, \quad(z=-h) \tag{9}
\end{gather*}
$$

With the transformation (7), (3), and (6) on $z=\zeta(X)$ are given by

$$
\begin{gather*}
-\omega \frac{\mathrm{d} \zeta}{\mathrm{~d} X}+k^{2} \frac{\partial \phi}{\partial X} \frac{\mathrm{~d} \zeta}{\mathrm{~d} X}-\frac{\partial \phi}{\partial z}=0  \tag{10}\\
-\omega \frac{\partial \phi}{\partial X}+f+g \zeta+\frac{1}{\rho}\left[D k^{4} \frac{\mathrm{~d}^{4} \zeta}{\mathrm{~d} X^{4}}+m_{e}\left(\omega^{2} \frac{\mathrm{~d}^{2} \zeta}{\mathrm{~d} X^{2}}+g\right)\right]=0 \tag{11}
\end{gather*}
$$

respectively, where

$$
\begin{equation*}
f=\frac{1}{2}\left[k^{2}\left(\frac{\partial \phi}{\partial X}\right)^{2}+\left(\frac{\partial \phi}{\partial z}\right)^{2}\right] . \tag{12}
\end{equation*}
$$

We combine partially (10) and (11) to gain the boundary conditions on $z=\zeta(X)$ as follows:

$$
\begin{align*}
\omega^{2} \frac{\partial^{2} \phi}{\partial X^{2}} & +g \frac{\partial \phi}{\partial z}-\omega \frac{\partial f}{\partial X}-\frac{\omega}{\rho}\left(D k^{4} \frac{\mathrm{~d}^{5} \zeta}{\mathrm{~d} X^{5}}+m_{e} \omega^{2} \frac{\mathrm{~d}^{3} \zeta}{\mathrm{~d} X^{3}}\right)  \tag{13}\\
& -k^{2} g \frac{\partial \phi}{\partial X} \frac{\mathrm{~d} \zeta}{\mathrm{~d} X}=0
\end{align*}
$$

Now the corresponding unknown potential function $\phi(X, z)$ and the wave surface elevation $\zeta(X)$ are governed by (8), (9), (11), and (13).

## 3. Analytic Approach Based on the Homotopy Analysis Method

3.1. Solution Expression and Initial Approximation. Using the homotopy analysis method, we should first of all start from a set of base functions and solution expression which are very important to approximate the unknown solutions of the nonlinear boundary problem under consideration. Mathematically, it seems impossible to guess the expression forms of the unknown potential function and the wave vertical displacement. Fortunately, considering the physical background of our problem, we may gain proper solution expressions of it. From viewpoints of the physical considerations here, our problem is composed of a train of progressive waves cause by a load moving on the ice sheet, an infinite elastic plate acting as an ice sheet floating on an fluid of finite depth. As is well known, in case of the pure water waves, the progressive wave elevation can be expressed as

$$
\begin{equation*}
\zeta(X)=\sum_{n=0}^{+\infty} \beta_{n} \cos (n X) \tag{14}
\end{equation*}
$$

by a set of base functions $\{\cos (n X), n \geq 0\}$, where $\beta_{n}$ are unknown coefficients. In the case of plate-covered surface, since we assume that there is no gap between the bottom surface of the thin elastic plate and the top surface of the fluid layer and a zero draft, the vertical displacement of the thin plate is still periodic in the $X$ direction. Therefore, we clearly know that $\zeta(X)$ can be expressed in the above form (14) too.

Besides, according to the linear wave theory, we can find the solutions of the Laplace equation (8) by the separation of variables method. To acquire those solutions, we have to use kinematic and dynamic boundary conditions of the free surface and the boundary condition in finite water depth, and we consider the solution derived here as the solution expression of potential function

$$
\begin{equation*}
\phi(X, z)=\sum_{n=1}^{+\infty} \alpha_{n} \frac{\cosh [n k(z+h)]}{\cosh (n k h)} \sin (n X) \tag{15}
\end{equation*}
$$

by a set of base functions $\{\cosh [n k(z+h)] / \cosh (n k h) \sin (n X)$, $n \geq 0\}$, where $\alpha_{n}$ are unknown coefficients. Note that the potential function $\phi(X, z)$ defined by (15) automatically satisfies the governing equation (8) and the bottom boundary condition (9). The above expressions (14) and (15) are called the solution expressions of $\phi(X, z)$ and $\zeta(X)$, respectively,
which play important roles in the method of homotopy analysis.

According to the solution expression (15) and the boundary condition (9), we construct the initial approximation of the potential function:

$$
\begin{equation*}
\phi_{0}(X, z)=\alpha_{0,1} \frac{\cosh [k(z+h)]}{\cosh (k h)} \sin (X) \tag{16}
\end{equation*}
$$

where $\alpha_{0,1}$ is an unknown coefficient. We choose

$$
\begin{equation*}
\zeta_{0}(X)=0 . \tag{17}
\end{equation*}
$$

as the initial approximation of wave profile $\zeta(X)$ to simplify the subsequent solution procedure [18, 20]. It should be emphasized that higher order terms can hold the corrections of the analytic series solutions due to the nonlinearity inherent in (11) and (13) although the initial guess $\zeta_{0}(X)$ is zero.
3.2. Continuous Variation. The HAM is based on a kind of continuous mapping of an initial approximation to the exact solution through a series of deformation equations. For simplicity, based on the nonlinear boundary condition (13) and (11), we define the two following nonlinear operators $\mathcal{N}_{1}$ and $\mathcal{N}_{2}$ as follows

$$
\begin{align*}
\mathcal{N}_{1}[ & \Phi(X, z ; q), \eta(X ; q)] \\
= & \omega^{2} \frac{\partial^{2} \Phi(X, z ; q)}{\partial X^{2}}+g \frac{\partial \Phi(X, z ; q)}{\partial z}-\omega \frac{\partial F}{\partial X} \\
& -\frac{\omega}{\rho}\left(D k^{4} \frac{\partial^{5} \eta(X ; q)}{\partial X^{5}}+\omega^{2} m_{e} \frac{\partial^{3} \eta(X ; q)}{\partial X^{3}}\right)  \tag{18}\\
& -k^{2} g \frac{\partial \Phi(X, z ; q)}{\partial X} \frac{\partial \eta(X ; q)}{\partial X}, \\
\mathcal{N}_{2}[ & \eta(X ; q), \Phi(X, z ; q)] \\
= & -\omega \frac{\partial \Phi(X, z ; q)}{\partial X}+F+g \eta(X ; q) \\
& +\frac{1}{\rho}\left[D k^{4} \frac{\partial^{4} \eta(X ; q)}{\partial X^{4}}+m_{e}\left(\omega^{2} \frac{\partial^{2} \eta(X ; q)}{\partial X^{2}}+g\right)\right], \tag{19}
\end{align*}
$$

where

$$
\begin{equation*}
F=\frac{1}{2}\left[k^{2}\left(\frac{\partial \Phi}{\partial X}\right)^{2}+\left(\frac{\partial \Phi}{\partial z}\right)^{2}\right] \tag{20}
\end{equation*}
$$

and $q \in[0,1]$ is the embedding parameter of the HAM.
Here, it should be emphasized that, as mentioned by Liao and Cheung and Tao et al. [14, 15], the HAM provides us with extremely large freedom to choose the auxiliary linear operators and the initial guess. Note that both linear terms of $\Phi(X, z ; q)$ and linear terms of $\eta(X ; q)$ are all contained in (18). If we choose all linear terms, the subsequent iterative procedure will become very complex. Fortunately, based on
the HAM, we can completely forget the linear terms in (13) and choose proper auxiliary linear operator of $\Phi(X, z ; q)$ by means of the solution expression (15) which is obtained under the physical considerations as

$$
\begin{equation*}
\overline{\mathscr{L}}_{1}[\Phi(X, z ; q)]=\omega^{2} \frac{\partial^{2} \Phi(X, z ; q)}{\partial X^{2}}+g \frac{\partial \Phi(X, z ; q)}{\partial z} \tag{21}
\end{equation*}
$$

In particular, if the angular frequency $\omega$ is given, we can choose such an approximation based on the linear wave theory to simplify the subsequent resolution of the nonlinear PDEs as follows:

$$
\begin{equation*}
\omega \approx \sqrt{g k \tanh (k h)} \tag{22}
\end{equation*}
$$

So we simplify the auxiliary linear operator in (21) as follows:

$$
\begin{align*}
\mathscr{L}_{1}[\Phi(X, z ; q)]= & g k \tanh (k h) \frac{\partial^{2} \Phi(X, z ; q)}{\partial X^{2}}  \tag{23}\\
& +g \frac{\partial \Phi(X, z ; q)}{\partial z}
\end{align*}
$$

where $\mathscr{L}_{1}[0]=0$. Note that, due to the weakly nonlinear effects, the actual frequency $\omega$ is often slightly different from the linear dispersion relation $\omega_{0}=\sqrt{g k \tanh (k h)}$. In Section $4, \omega / \omega_{0}=1.01$ is chosen so that the perturbation theory is valid and corresponding results are highly accurate, and then we can compare our results with those obtained by the perturbation method.

Based on the linear operator of the wave profile function $\eta(X ; q)$ in the nonlinear operator $\mathcal{N}_{2}$, for simplicity, we may choose another auxiliary linear operator:

$$
\begin{equation*}
\mathscr{L}_{2}[\eta(X ; q)]=\frac{\partial^{4} \eta(X ; q)}{\partial X^{4}}+\frac{\partial^{2} \eta(X ; q)}{\partial X^{2}}+\eta(X ; q) \tag{24}
\end{equation*}
$$

where $\mathscr{L}_{2}[0]=0$.
We let $c_{0}$ be an nonzero convergence-control parameter. It is noted that both $c_{0}$ and $q$ in the HAM are auxiliary parameters without any physical meaning. Instead of the nonlinear PDEs (8), (9), (11), and (13), we reconstruct the socalled zeroth-order deformation equations as follows:

$$
\begin{equation*}
k^{2} \frac{\partial^{2} \Phi(X, z ; q)}{\partial X^{2}}+\frac{\partial^{2} \Phi(X, z ; q)}{\partial z^{2}}=0, \quad(-h \leq z \leq \eta(X ; q)) \tag{25}
\end{equation*}
$$

$$
\begin{gather*}
\frac{\partial \Phi(X, z ; q)}{\partial z}=0, \quad(z=-h),  \tag{26}\\
(1-q) \mathscr{L}_{1}\left[\Phi(X, z ; q)-\phi_{0}(X, z)\right] \\
=q c_{0} \mathscr{N}_{1}[\Phi(X, z ; q), \eta(X ; q)], \quad(z=\eta(X ; q)),  \tag{27}\\
(1-q) \mathscr{L}_{2}\left[\eta(X ; q)-\zeta_{0}(X)\right] \\
=q c_{0} \mathcal{N}_{2}[\eta(X ; q), \Phi(X, z ; q)], \quad(z=\eta(X ; q)) . \tag{28}
\end{gather*}
$$

Then, from (27) and (28), two mapping functions $\Phi(X, z ; q)$ and $\eta(X ; q)$ vary respectively continuously from their initial approximation $\phi_{0}(X, z)$ and $\zeta_{0}(X)$ to the exact solutions $\phi(X, z)$ and $\zeta(X)$ of the original problem. The Taylor series of $\Phi(X, z ; q)$ and $\eta(X ; q)$ at $q=0$ are

$$
\begin{gather*}
\Phi(X, z ; q)=\phi_{0}(X, z)+\sum_{m=1}^{+\infty} \phi_{m}(X, z) q^{m}  \tag{29}\\
\eta(X ; q)=\zeta_{0}(X)+\sum_{m=1}^{+\infty} \zeta_{m}(X) q^{m} \tag{30}
\end{gather*}
$$

where

$$
\begin{equation*}
\left\{\phi_{m}(X, z), \zeta_{m}(X)\right\}=\left.\frac{1}{m!} \frac{\partial^{m}}{\partial q^{m}}\{\Phi(X, z ; q), \eta(X ; q)\}\right|_{q=0} . \tag{31}
\end{equation*}
$$

Assume that $c_{0}$ is so properly chosen that the series in (29) and (30) converge at $q=1$; then we have the so-called homotopy-series solutions as follows:

$$
\begin{gather*}
\phi(X, z)=\Phi(X, z ; 1)=\phi_{0}(X, z)+\sum_{m=1}^{+\infty} \phi_{m}(X, z) \\
\zeta(X)=\eta(X ; 1)=\zeta_{0}(X)+\sum_{m=1}^{+\infty} \zeta_{m}(X) \tag{32}
\end{gather*}
$$

At the $n$ th-order of approximations, we have

$$
\begin{align*}
\phi(X, z) & \approx \phi_{0}(X, z)+\sum_{m=1}^{+n} \phi_{m}(X, z)  \tag{33}\\
\zeta(X) & \approx \zeta_{0}(X)+\sum_{m=1}^{+n} \zeta_{m}(X)
\end{align*}
$$

As shown later in the following section, the unknown terms $\phi_{m}(X, z)$ and $\zeta_{m}(X)$ are governed by the linear PDEs (34)-(36).
3.3. High-Order Deformation Equations. High-order deformation equations for the unknown $\phi_{m}(X, z), \zeta_{m}(X)$ can be derived directly from the zeroth-order deformation equations. Firstly, substituting the homotopy-Maclaurin series (29) and (30) into the governing equation (25) and the boundary condition in finite water depth (26) and then equating the like-power of the embedding parameter $q$, we have

$$
\begin{gather*}
k^{2} \frac{\partial^{2} \phi_{m}(X, z)}{\partial X^{2}}+\frac{\partial^{2} \phi_{m}(X, z)}{\partial z^{2}}=0, \quad(-h \leq z \leq 0)  \tag{34}\\
\frac{\partial \phi_{m}(X, z)}{\partial z}=0, \quad(z=-h)
\end{gather*}
$$

where $m \geq 1$.
Note that, $\Phi(X, z ; q)$ at the unknown surface $z=\eta(X ; q)$ may be expressed in terms of the Taylor expansion at $z=$ 0 instead of $z=\eta(X ; q)$. The detailed derivation of the
expansion of $\Phi(X, z ; q)$ at the unknown surface is given in Appendices (A.1)-(A.5). Upon the substitution of appropriate series (A.5) and (30) into the boundary conditions (27) and (28), we have two linear boundary conditions on $z=0$ as follows:

$$
\begin{gather*}
\left.\mathscr{L}_{1}\left(\phi_{m}\right)\right|_{z=0}=c_{0} \Delta_{m-1}^{\phi}+\chi_{m} S_{m-1}-\bar{S}_{m}  \tag{35}\\
\mathscr{L}_{2}\left(\zeta_{m}\right)=c_{0} \Delta_{m-1}^{\zeta}+\chi_{m}\left(\frac{\mathrm{~d}^{4} \zeta_{m-1}}{\mathrm{~d} X^{4}}+\frac{\mathrm{d}^{2} \zeta_{m-1}}{\mathrm{~d} X^{2}}+\zeta_{m-1}\right) \tag{36}
\end{gather*}
$$

where

$$
\chi_{m}= \begin{cases}0, & m \leqslant 1  \tag{37}\\ 1, & m>1\end{cases}
$$

The detailed derivation of the above equations and the expression for $\phi_{m}$ and $\zeta_{m}$ are given in Appendix A. It should be noted that (27) and (28) holds on the unknown boundary $z=\eta(X ; q)$, while (35) and (36) hold on $z=0$. Furthermore, the original nonlinear DPEs (1)-(5) are transferred into an infinite number of linear decoupled high-order deformation equations (34)-(36). Namely, given $\phi_{m-1}$ and $\zeta_{m-1}, \phi_{m}$ and $\zeta_{m}$ can be obtained easily by means of the inverse operators of the right-hand sides of (35) and (36), respectively, and a computer algebra system such as Mathematica. The resulting expressions for $\phi_{m}$ and $\zeta_{m}$ are presented to the second order in the coming subsection.
3.4. First-Order and Second-Order Approximations. Substituting initial approximations (16) and (17) into (36), we can get $\zeta_{1}(X)$ using the inverse linear operator $\mathscr{L}_{2}$ in (36) as follows:

$$
\begin{align*}
\zeta_{1}(X)= & \frac{1}{4}\left[4 d g c_{0}+c_{0} a_{0,1}^{2}+k^{2} c_{0} a_{0,1}^{2} \tanh ^{2}(h k)\right] \\
& -\omega c_{0} \alpha_{0,1} \cos (X)  \tag{38}\\
& +\frac{1}{52}\left[c_{0} a_{0,1}^{2}-k^{2} c_{0} a_{0,1}^{2} \tanh ^{2}(h k)\right] \cos (2 X)
\end{align*}
$$

But now the coefficient $\alpha_{0,1}$ in the initial approximation of $\phi_{0}(X, z)$ in (16) is still unknown. So we introduce an additional equation to relate the solutions with the wave height:

$$
\begin{equation*}
\zeta_{1}(m \pi)-\zeta_{1}(n \pi)=H \tag{39}
\end{equation*}
$$

in which $m$ is an even integer, $n$ is an odd integer, and $H$ is the wave height to the first order based on the HAM. The relation (39) for the wave height and its vertical displacement can determine the solution of $\alpha_{0,1}$.

Further, in the analogous manner as for the first-order approximation, by using the inverse linear operator $\mathscr{L}_{1}$ in (35), it is easy to get the solution of $\phi_{1}(X, z)$, especially
by means of the symbolic computation software such as Mathematica:

$$
\begin{gather*}
\alpha_{0,1}=-\frac{H}{2 \omega c_{0}}, \\
\phi_{1}(X, z)=\alpha_{1,1} \frac{\cosh [k(h+z)]}{\cosh (k h)} \sin (X) \\
+\frac{-H^{2}+H^{2} k^{2} \tanh ^{2}(h k)}{16 g k \omega c_{0}[2 \tanh (h k)-\tanh (2 h k)]}  \tag{40}\\
\times \frac{\cosh [2 k(h+z)]}{\cosh (2 k h)} \sin (2 X) .
\end{gather*}
$$

We find the common solution $\phi_{1}(X, z)$ has one unknown coefficient $\alpha_{1,1}$ which can be determined by avoiding the "secular" term $\sin (X)$ in $\phi_{2}(X, z)$. We note that all subsequent functions occur recursively. Utilizing the linear equations (35) and (36) to continue with the first-order approximations we have

$$
\begin{align*}
\zeta_{2}(X)= & \beta_{2,0}+\beta_{2,1} \cos (X)+\beta_{2,2} \cos (2 X) \\
& +\beta_{2,3} \cos (3 X)+\beta_{2,4} \cos (4 X) \\
\phi_{2}(X, z)= & \alpha_{2,1} \frac{\cosh [k(h+z)]}{\cosh (k h)} \sin (X) \\
& +\alpha_{2,2} \frac{\cosh [2 k(h+z)]}{\cosh (2 k h)} \sin (2 X)  \tag{41}\\
& +\alpha_{2,3} \frac{\cosh [3 k(h+z)]}{\cosh (3 k h)} \sin (3 X) \\
& +\alpha_{2,4} \frac{\cosh [4 k(h+z)]}{\cosh (4 k h)} \sin (4 X) \\
& +\alpha_{2,5} \frac{\cosh [5 k(h+z)]}{\cosh (5 k h)} \sin (5 X)
\end{align*}
$$

where $\alpha_{i, j}$ is the $j$ th unknown coefficient of $\phi_{i}(X, z)$ and $\beta_{i, j}$ is the $j$ th unknown coefficient of $\zeta_{i}(X)$. The detailed expressions of these coefficients for $\phi_{2}$ and $\zeta_{2}$ are given in Appendix B.

In order to obtain higher-order functions $\phi_{m}(X, z)$ and $\zeta_{m}(X)$, we need only to continue this approach. In principle, we can acquire infinite-order solutions for our physical model. It is also worthwhile to mention that these solutions will retain model parameters and the convergence control parameter $c_{0}$.
3.5. Optimal Convergence-Control Parameter. If we fix all model parameters in our approximate series solutions, there is still an unknown convergence control parameter $c_{0}$ in them, which is used to guarantee the convergence of our approximation solutions. According to Liao [12], it is the convergence control parameter $c_{0}$ that essentially differs the HAM from all other analytic methods. And the optimal value of $c_{0}$ is determined by the minimum of the total squaredresidual $\varepsilon_{m}^{T}$ of our nonlinear problem, defined by

$$
\begin{equation*}
\varepsilon_{m}^{T}=\varepsilon_{m}^{\phi}+\varepsilon_{m}^{\zeta}, \tag{42}
\end{equation*}
$$



Figure 1: Residual squares of $\log _{10} \varepsilon_{m}^{T}$ versus $c_{0}$. Solid line: firstorder approximation; dashed line: third-order approximation; dash-dotted line: fifth-order approximation; dash-dot-dotted line: seventh-order approximation.


Figure 2: Comparison of our present 3rd-order surface elevation $\zeta$ with those obtained by the perturbation method. Solid line: perturbation-series solution; dashed line: homotopy-series solution.
where

$$
\begin{align*}
& \varepsilon_{m}^{\phi}=\frac{1}{1+M} \sum_{i=0}^{M}\left(\left.\mathcal{N}_{1}[\phi(X, z), \zeta(X)]\right|_{X=i \Delta X}\right)^{2}, \\
& \varepsilon_{m}^{\zeta}=\frac{1}{1+M} \sum_{i=0}^{M}\left(\left.\mathcal{N}_{2}[\phi(X, z), \zeta(X)]\right|_{X=i \Delta X}\right)^{2}, \tag{43}
\end{align*}
$$

where $\varepsilon_{m}^{\phi}$ and $\varepsilon_{m}^{\zeta}$ are two residual square errors of boundary conditions (27) and (28), respectively. $M$ is the number of the discrete points, and $\Delta X=\pi / M$. In this paper, we choose $M=$ 10.


Figure 3: P.E. for (44) versus the water depth $h$ for different plate thicknesses $d$. Solid line: $d=0.001$; dashed line: $d=0.005$; dash-dot-dotted line: $d=0.01$.


Figure 4: P.E. for (44) versus the water depth $h$ for different Young's moduli of the plate $E$. Solid line: $E=10^{8}$; dashed line: $E=10^{9}$; dash-dot-dotted line: $E=10^{10}$.

Theorem 2.1 given by Liao in [12] can guarantee the rationality of (42). So we obtain the optimal convergence control parameter $c_{0}$ by the minimum of the squared-residual $\varepsilon_{m}^{T}$, generally corresponding to $d \varepsilon_{m}^{T} / d c_{0}=0$.

## 4. Results and Analysis

In order to show the convergence of the analytical series solution to our problems by means of the HAM, we consider the cases of $k=\pi / 5 \mathrm{~m}^{-1}, d=0.01 \mathrm{~m}, \rho_{e}=900 \mathrm{kgm}^{-3}, \nu=0.33$, $E=10^{10} \mathrm{Nm}^{-2}, h=5 \mathrm{~m}, H=0.1 \mathrm{~m}$, and $\omega / \omega_{0}=1.01$ and

TABLE 1: The total residual square error $\varepsilon_{m}^{T}$ for different approximation order $m$ with $c_{0}=-0.18$.

| $m$ | $\varepsilon_{m}^{T}$ |
| :--- | :---: |
| 1 | $3.497 \times 10^{-3}$ |
| 3 | $3.404 \times 10^{-4}$ |
| 5 | $3.700 \times 10^{-5}$ |
| 7 | $7.910 \times 10^{-6}$ |
| 10 | $4.803 \times 10^{-8}$ |
| 15 | $5.382 \times 10^{-11}$ |

take these data hereinafter for computation unless otherwise stated. The total residual square error $\varepsilon_{m}^{T}$ at several orders of approximation versus the convergence-control parameter $\mathcal{c}_{0}$ is shown in Figure 1. It is found that $\varepsilon_{m}^{T}$ at every order has the smallest values which corresponds to the optimal $c_{0}$. For example, as $m=7$, the optimal $c_{0}=-0.18$, and the smallest value of $\varepsilon_{7}^{T}=7.910 \times 10^{-6}$. So, let the optimal convergence-control parameter $c_{0}=-0.18$, the total residual square error $\varepsilon_{m}^{T}$ decreases quickly as the order $m$ increases, as shown in Table 1. It is also found that $\varepsilon_{15}^{T}$ is down to $5.382 \times$ $10^{-11}$ at the 15th-order of approximation, which indicates the convergence of our series solutions. In this way, we ensure that all our solutions are highly accurate.

Also, we compare our HAM solutions of waves propagating beneath an elastic plate floating on a fluid of finite depth with those results obtained by perturbation techniques, as shown in Figure 2. It should be noted that the perturbationseries solution is derived by substituting the series expansions (4.5) and (4.6) in [9] into the nonlinear PDEs (8)-(12), and equating power of small parameter $\epsilon$ leads to a succession of linear PDEs, and then the linear PDEs can be solved by the separation of variables. In Figure 2. It is seen that our homotopy-series approximation of the surface elevation $\zeta$ agrees well with the perturbation-series approximation, and only slight derivations occur at the trough of the wave profile as in Figure 2, which further indicates the validity of our present theory about nonlinear hydroelastic waves beneath a floating ice sheet.

We define quantities which measure how much energy there is in the wave propagating beneath an infinite elastic plate. Let P.E. be the mean potential density per unit length in the $X$-axis [27]. In terms of the wave surface elevation function, the energy density can be written as

$$
\begin{equation*}
\text { P.E. }=\frac{1}{4 \pi^{2}} \int_{0}^{2 \pi} \zeta^{2}(X) \mathrm{d} X . \tag{44}
\end{equation*}
$$

Different from all research objectives in [7-9], we firstly consider in this paper the effect of water depth on nonlinear hydroelastic waves beneath a floating elastic plate in detail. The energy of hydroelastic waves for different Young's moduli of the plate $E$ and different plate thicknesses $h$ in various water depths are as shown in Figures 3 and 4 and Tables 2 and 3, respectively. We find that, when water depth $h$ is about more than 2 , the hydroelastic waves traveling beneath the thickest plate always contain the largest wave energy in different water


Figure 5: Variation of the plate deflection $\zeta(X)$ near the crest versus $X$ for different Young's moduli of the plate $E$. Solid line: $E=10^{8}$; dashed line: $E=10^{9}$; dash-dot-dotted line: $E=10^{10}$.


Figure 6: Variation of the plate deflection $\zeta(X)$ near the trough versus $X$ for different Young's moduli of the plate $E$. Solid line: $E=10^{8}$; dashed line: $E=10^{9}$; dash-dot-dotted line: $E=10^{10}$.
depths. And with an increasing Young's modulus of the plate, the wave energy becomes large too.

The effect of Young's modulus $E$ of the plate on the wave elevation $\zeta(X)$ under a floating elastic plate is studied. Figures 5 and 6 show the differences in $\zeta(X)$ for $E=10^{8}, 10^{9}$, and $10^{10}$. According to Figures 5 and 6 , respectively, we can see that the nonlinear hydroelastic response of the waves becomes flatter at the crest and steeper at the trough due to the larger value of Young's modulus E. Finally, we consider the impact the plate thickness $d$ by increasing $d$ from 0.001 to 0.01 . In Figures 7 and 8, we show several displacements $\zeta(X)$ with $d=0.001, d=0.005$, and $d=0.01$, respectively. It


Figure 7: Variation of the plate deflection $\zeta(X)$ near the crest versus $X$ for different plate thicknesses $d$. Solid line: $d=0.001$; dashed line: $d=0.005$; dash-dot-dotted line: $d=0.01$.


Figure 8: Variation of the plate deflection $\zeta(X)$ near the trough versus $X$ for different plate thicknesses $d$. Solid line: $d=0.001$; dashed line: $d=0.005$; dash-dot-dotted line: $d=0.01$.
indicates that the results are very similar to the effects due to different Young's moduli $E$ of the plate.

## 5. Conclusions

In this paper, the nonlinear hydroelastic waves propagating beneath a two-dimensional infinite elastic plate floating on a fluid of finite depth are investigated analytically by the HAM. Mathematically, for a train of nonlinear hydroelastic waves traveling at a constant velocity in a fluid of finite or infinite depth, the PDEs in [7-9] were obtained by simply eliminating the time-dependent terms from the kinematic

Table 2: P.E. for (44) with different plate thicknesses and various water depths $h$.

| $h$ | P.E. <br> $(d=0.001)$ | P.E. <br> $(d=0.005)$ | P.E. <br> $(d=0.01)$ |
| :--- | :---: | :---: | :---: |
| 1 | 0.00067245 | 0.00100093 | 0.00040836 |
| 3 | 0.00031356 | 0.00055509 | 0.00069567 |
| 5 | 0.00030650 | 0.00053375 | 0.00074396 |
| 10 | 0.00030590 | 0.00053158 | 0.00074694 |
| 15 | 0.00030592 | 0.00053159 | 0.00074694 |
| 20 | 0.00030595 | 0.00053159 | 0.00074696 |
| 30 | 0.00030600 | 0.00053159 | 0.00074696 |
| $\infty$ | 0.00030600 | 0.00053159 | 0.00074696 |

Table 3: P.E. for (44) with different values of Young's modulus of the plate $E$ and various water depths $h$.

| $h$ | P.E. <br> $\left(E=10^{8}\right)$ | P.E. <br> $\left(E=10^{9}\right)$ | P.E. <br> $\left(E=10^{10}\right)$ |
| :--- | :---: | :---: | :---: |
| 1 | 0.00076484 | 0.0009946 | 0.00040836 |
| 3 | 0.00038884 | 0.00054698 | 0.00069567 |
| 5 | 0.00030138 | 0.00047932 | 0.00074396 |
| 10 | 0.00028886 | 0.00046970 | 0.00074694 |
| 15 | 0.00028884 | 0.00046969 | 0.00074694 |
| 20 | 0.00028884 | 0.00046969 | 0.00074696 |
| 30 | 0.00028884 | 0.00046969 | 0.00074696 |
| $\infty$ | 0.00028884 | 0.00046969 | 0.00074696 |

and dynamic boundary conditions on the unknown free surface in the frame of reference moving with the wave. Here, for a general case it should be noted that we construct the PDEs by directly applying the traveling-wave method to transfer the temporal differentiation into the spatial one in a fixed Cartesian coordinate oxz. Furthermore, the convergent homotopy-series solutions for the PDES are derived by the HAM with the optimal convergence control parameter.

Physically, we study the effect of the water depth on the nonlinear hydroelastic waves under an elastic plate in detail. It is found that, in different water depths, the wave energy density (P.E.) tends to become larger with an increasing thickness of the sheet. The same conclusions are obtained in various water depths by means of different values of Young's modulus of the plate. Additionally, the influences of Young's modulus and the thickness of the plate on the wave elevation $\zeta(X)$ are investigated, respectively. As Young's modulus of the plate increases, the wave elevation becomes lower. And the increasing thickness of the plate flattens the crest and sharpens the trough of the wave profile. The results obtained here demonstrate that Young's modulus and the thickness of the sheet have important effects on the energy and the profile of nonlinear hydroelastic waves under an ice sheet floating on a fluid of finite depth.

## Appendices

## A. The Detailed Derivation of (35) and (36) and the Expressions for $\phi_{m}$ and $\zeta_{m}$

Let

$$
\begin{equation*}
\eta^{n}=\left(\sum_{i=1}^{+\infty} \zeta_{i} q^{i}\right)^{n}=\sum_{i=n}^{+\infty} \mu_{n, i} q^{i} \tag{A.1}
\end{equation*}
$$

For any $z$, we have a Maclaurin series as follows:

$$
\begin{equation*}
\phi_{m}(X, z)=\left.\sum_{n=0}^{+\infty} \frac{1}{n!} \frac{\partial^{n} \phi_{m}}{\partial z^{n}}\right|_{z=0} z^{n} \tag{A.2}
\end{equation*}
$$

For $z=\eta(X ; q)$, it follows from (A.1) and (A.2) that

$$
\begin{align*}
\phi_{m}(X, \eta) & =\sum_{n=0}^{+\infty}\left(\left.\frac{1}{n!} \frac{\partial^{n} \phi_{m}}{\partial z^{n}}\right|_{z=0}\right)\left(\sum_{i=n}^{+\infty} \mu_{n, i} q^{i}\right)  \tag{A.3}\\
& =\sum_{i=0}^{+\infty} \psi_{m, i} q^{i}
\end{align*}
$$

where

$$
\begin{equation*}
\psi_{m, i}=\sum_{n=0}^{i}\left(\left.\frac{1}{n!} \frac{\partial^{n} \phi_{m}}{\partial z^{n}}\right|_{z=0}\right) \mu_{n, i} \tag{A.4}
\end{equation*}
$$

Thus we have, for $z=\eta(X ; q)$,

$$
\begin{align*}
\Phi(X, \eta ; q) & =\sum_{m=0}^{+\infty} \phi_{m}(X, \eta) q^{m}=\sum_{m=0}^{+\infty}\left(\sum_{n=0}^{+\infty} \psi_{m, i} q^{i}\right) q^{m}  \tag{A.5}\\
& =\sum_{m=0}^{+\infty} \varphi_{m} q^{m}
\end{align*}
$$

where

$$
\begin{equation*}
\varphi_{m}=\sum_{i=0}^{m} \psi_{m-i, i} . \tag{A.6}
\end{equation*}
$$

Substituting the series expansions (A.1) and (A.5) into the boundary conditions (27) and (28) and then equating the like-power of the embedding parameter $q$, we have two linear boundary conditions (35) and (36), respectively. And
the explicit expressions for $\Delta_{m-1}^{\phi}, S_{m-1}, \bar{S}_{m}$, and $\Delta_{m-1}^{\zeta}$ in these two conditions are given by

$$
\begin{align*}
& \Delta_{m-1}^{\phi}= \omega^{2} \frac{\mathrm{~d}^{2} \varphi_{m}}{\mathrm{~d} X^{2}}+g \bar{\varphi}_{m} \\
&- \omega \sum_{n=0}^{m}\left(\frac{\mathrm{~d} \varphi_{n}}{\mathrm{~d} X} \frac{\mathrm{~d}^{2} \varphi_{m-n}}{\mathrm{~d} X^{2}}+\bar{\varphi}_{n} \frac{\mathrm{~d} \bar{\varphi}_{m-n}}{\mathrm{~d} X}\right) \\
&- \frac{\omega}{\rho} D k^{4} \frac{\mathrm{~d}^{5} \zeta_{m}}{\mathrm{~d} X^{5}}-\frac{\omega^{3} m_{e}}{\rho} \frac{\mathrm{~d}^{3} \zeta_{m}}{\mathrm{~d} X^{3}}-k^{2} g \sum_{n=0}^{m} \frac{\mathrm{~d} \varphi_{n}}{\mathrm{~d} X} \frac{\mathrm{~d} \zeta_{m-n}}{\mathrm{~d} X}, \\
& S_{m-1}=\sum_{i=0}^{m-2}\left(\frac{\mathrm{~d}^{2} \psi_{m-1-i, i}}{\mathrm{~d} X^{2}}+\gamma_{m-1-i, i}\right), \\
& \bar{S}_{m}=\sum_{i=1}^{m-1}\left(\frac{\mathrm{~d}^{2} \psi_{m-i, i}}{\mathrm{~d} X^{2}}+\gamma_{m-i, i}\right), \\
& \Delta_{m-1}^{\zeta}=-\omega \frac{\mathrm{d} \varphi_{m-1}}{\mathrm{~d} X} \\
&+\frac{1}{2} \sum_{n=0}^{m-1}\left(\frac{\mathrm{~d} \varphi_{n}}{\mathrm{~d} X} \frac{\mathrm{~d} \varphi_{m-1-n}}{\mathrm{~d} X}+\bar{\varphi}_{n} \bar{\varphi}_{m-1-n}\right)+\zeta_{m-1} \\
&+\frac{D k^{4}}{\rho} \frac{\mathrm{~d}^{4} \zeta_{m-1}}{\mathrm{~d} X^{4}}+\frac{m_{e} \omega^{2}}{\rho} \frac{\mathrm{~d}^{2} \zeta_{m-1}}{\mathrm{~d} X^{2}}, \quad(m \geq 2), \\
& \Delta_{0}^{\zeta}= \frac{1}{2}\left[\left(\frac{\mathrm{~d} \varphi_{0}}{\mathrm{~d} X}\right)^{2}+\bar{\varphi}_{0}^{2}\right]-\omega \frac{\mathrm{d} \varphi_{0}}{\mathrm{~d} X}+\frac{m_{e} g}{\rho}, \tag{A.7}
\end{align*}
$$

where

$$
\begin{gather*}
\bar{\varphi}_{m}=\sum_{i=0}^{m} \gamma_{m-i, i} \\
\gamma_{m-i, i}=\sum_{n=0}^{i} \frac{1}{n!}\left(\left.\frac{\partial^{n+1} \phi_{m-i}}{\partial z^{n+1}}\right|_{z=0}\right) \mu_{n, i} . \tag{A.8}
\end{gather*}
$$

## B. Expressions of the Coefficients

$$
\begin{aligned}
\beta_{2,0}=\frac{1}{16 \omega^{2} c_{0}}[ & H^{2}+g H^{2} c_{0}+16 \mathrm{~d} g \omega^{2} c_{0}^{2} \\
& +16 \mathrm{~d} g^{2} \omega^{2} c_{0}^{3}-4 H \omega c_{0} \alpha_{1,1} \\
& +H^{2} k^{2} \tanh ^{2}(h k)+g H^{2} k^{2} c_{0} \tanh ^{2}(h k) \\
& \left.-4 H k^{2} \omega c_{0} \alpha_{1,1} \tanh ^{2}(h k)\right]
\end{aligned}
$$

$$
\begin{align*}
& \beta_{2,1}=\left(\left[H^{3} \rho+32 g H k \rho \omega^{2} c_{0} \tanh (h k)\right.\right. \\
& +32 D g H k^{5} \omega^{2} c_{0}^{2} \tanh (h k) \\
& +32 g^{2} H k \rho \omega^{2} c_{0}^{2} \tanh (h k) \\
& -32 \mathrm{~d} g H k \rho \omega^{4} c_{0}^{2} \tanh (h k) \\
& -64 g k \rho \omega^{3} c_{0}^{2} \alpha_{1,1} \tanh (h k) \\
& -16 g H k \rho \omega^{2} c_{0} \tanh (2 h k) \\
& -16 D g H k^{5} \omega^{2} c_{0}^{2} \tanh (2 h k) \\
& -16 g^{2} H k \rho \omega^{2} c_{0}^{2} \tanh (2 h k) \\
& +16 \mathrm{~d} g H k \rho \omega^{4} c_{0}^{2} \tanh (2 h k) \\
& +32 g k \rho \omega^{3} c_{0}^{2} \alpha_{1,1} \tanh (2 h k) \\
& +H^{3} k^{2} \rho \tanh (h k) \tanh (2 h k) \\
& -H^{3} k^{4} \rho \tanh ^{3}(h k) \tanh (2 h k) \\
& \left.\left.-H^{3} k^{2} \rho \tanh ^{2}(h k)\right]\right) \\
& \times\left(\left[32 g k \rho \omega^{2} c_{0}(2 \tanh (h k)-\tanh (2 h k))\right]\right)^{-1}, \\
& \beta_{2,2}=\frac{1}{13}-\frac{\mathrm{d} H^{2}}{52}+\frac{g H^{2}}{208 \omega^{2}}+\frac{D H^{2} k^{4}}{13 \rho \omega^{2}} \\
& +\frac{H^{2}}{16 \omega^{2} c_{0}}-\frac{H \alpha_{1,1}}{4 \omega}+\frac{1}{52} \mathrm{~d} H^{2} k^{2} \tanh ^{2}(h k) \\
& -\frac{g H^{2} k^{2} \tanh ^{2}(h k)}{208 \omega^{2}}-\frac{D H^{2} k^{6} \tanh ^{2}(h k)}{13 \rho \omega^{2}} \\
& -\frac{H^{2} k^{2} \tanh ^{2}(h k)}{16 \omega^{2} c_{0}} \\
& +\frac{H k^{2} \alpha_{1,1} \tanh ^{2}(h k)}{4 \omega} \\
& +\frac{H^{2}}{8 g k(2 \tanh (h k)-\tanh (2 h k))} \\
& -\frac{H^{2} k \tanh ^{2}(h k)}{8 g(2 \tanh (h k)-\tanh (2 h k))}, \\
& \beta_{2,3}=\left(H^{3}-H^{3} k^{2} \tanh ^{2}(h k)\right. \\
& -H^{3} k^{2} \tanh (h k) \tanh (2 h k) \\
& \left.+H^{3} k^{4} \tanh ^{3}(h k) \tanh (2 h k)\right) \\
& \times\left(2336 g k \omega^{2} c_{0}(2 \tanh (h k)-\tanh (2 h k))\right)^{-1}, \\
& \beta_{2,4}=0, \tag{B.1}
\end{align*}
$$

$$
\begin{aligned}
& \alpha_{1,1}=\left(\left[H \left(-13 H^{2}\left(1+2 g k^{2}\right) \rho \omega^{2}\right.\right.\right. \\
& -g k\left(g H^{2} k^{2} \rho-208\left(D k^{4} \omega^{4}-d \rho \omega^{6}\right) c_{0}^{2}\right) \\
& \times \tanh (2 h k) \\
& +H^{2} k^{2} \rho \tanh ^{2}(h k) \\
& \times\left(13\left(1+2 g k^{2}\right) \omega^{2}+g^{2} k^{3} \tanh (2 h k)\right) \\
& +H^{2} k^{4} \rho \tanh ^{3}(h k) \\
& \times\left(-2 g^{2} k-13(-1+g) \omega^{2} \tanh (2 h k)\right) \\
& +k \tanh (h k) \\
& \times\left(416 g \omega^{4}\left(-D k^{4}+\mathrm{d} \rho \omega^{2}\right) c_{0}^{2}+H^{2} k \rho\right. \\
& \times\left(2 g^{2} k+13(-1+g) \omega^{2} \tanh (2 h k)\right) \\
& +H^{2} k \rho\left(2 g^{2} k+13(-1+g)\right. \\
& \left.\left.\left.\left.\left.\times \omega^{2} \tanh (2 h k)\right)\right)\right)\right]\right) \\
& \times\left(\left[-416 \omega^{2} c_{0}^{2}\right.\right. \\
& \times\left(\mathrm{d} g^{2} k^{2}+\omega^{2} g k \rho \omega\right. \\
& \times\left(g H^{2} k^{2}\left(-1+\tanh ^{2}(h k)\right)\right. \\
& \times\left(25+27 k^{2} \tanh ^{2}(h k)\right) \\
& \left.\left.-g k \tanh (h k)-\mathrm{d} g^{2} k^{2} \tanh ^{2}(h k)\right)\right) \\
& \times(2 \tanh (h k)-\tanh (2 h k))])^{-1}, \\
& \alpha_{2,2}=\frac{1}{-4 g k \tanh (h k)+2 g k \tanh (2 h k)} \\
& \times\left[-\frac{g H^{2} k^{2}}{8 \omega}+\frac{2 D H^{2} k^{4}}{13 \rho \omega}-\frac{\mathrm{d} H^{2} \omega}{26}-\frac{H \alpha_{1,1}}{2}\right. \\
& -\frac{g H k^{2} \alpha_{1,1}}{4}-\frac{2 D H^{2} k^{6} \tanh ^{2}(h k)}{13 \rho \omega} \\
& +\frac{1}{26} \mathrm{~d} H^{2} k^{2} \omega \tanh ^{2}(h k) \\
& +\frac{1}{2} H k^{2} \alpha_{1,1} \tanh ^{2}(h k) \\
& +g H k^{2} \alpha_{1,1} \tanh ^{2}(h k) \\
& -\frac{H^{2} k \omega \tanh ^{2}(h k)}{8 g \tanh (h k)-4 g \tanh (2 h k)}
\end{aligned}
$$

$$
\begin{gather*}
+\left(H^{3} k \tanh (h k)(-\tanh (h k)-\tanh (2 h k)\right. \\
\left.\left.+k^{2} \tanh ^{2}(h k) \tanh (2 h k)\right)\right) \\
\times\left(2\left(16 g \omega c_{0} \tanh (h k)-8 g \omega c_{0} \tanh (2 h k)\right)\right)^{-1} \\
+\left(3 H^{3}\right)\left(4 \left(16 g k \omega c_{0} \tanh (h k)\right.\right. \\
\left.\left.\left.-8 g k \omega c_{0} \tanh (2 h k)\right)\right)^{-1}\right], \\
\alpha_{2,4} \\
=\frac{1}{-16 g k \tanh (h k)+4 g k \tanh (4 h k)} \\
\times\left[\frac{H^{4} k\left(1+k^{4} \tanh { }^{4}(h k)-2 k^{2} \tanh ^{2}(h k)\right)}{4\left(832 \omega^{3} c_{0}^{2} \tanh (h k)-416 \omega^{3} c_{0}^{2} \tanh (2 h k)\right)}\right. \\
\quad+\frac{H^{3} k \tanh (h k) \tanh (2 h k)\left(-1+k^{2} \tanh h^{2}(h k)\right)}{2\left(8 \omega c_{0} \tanh (h k)-4 \omega c_{0} \tanh (2 h k)\right)} \\
\times\left(H^{4} k \tanh (h k) \tanh (2 h k)\right. \\
\left.\times\left(-1-k^{4} \tanh { }^{4}(h k)+2 k^{2} \tanh ^{2}(h k)\right)\right) \\
\times\left(4 \left(104 \omega^{3} c_{0}^{2} \tanh (h k)\right.\right. \\
\left.\left.\left.\quad-52 \omega^{3} c_{0}^{2} \tanh (2 h k)\right)\right)^{-1}\right], \\
\alpha_{2,5}=0 . \tag{B.2}
\end{gather*}
$$

## Conflict of Interests

There is no conflict of interests in the paper. The authors themselves used the program of the symbolic computation software named Mathematica independently to gain the approximate analytical solutions of the PDEs considered here.

## Acknowledgments

This work was supported in part by China Postdoctoral Science Foundation funded Project 20100481088 and in part by the Natural Science Foundation of Shandong Province of China under Grant ZR2010FL016. The authors would like to thank the reviewer for his constructive comments.

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# Numerical Solution of Nonlinear Fredholm Integrodifferential Equations by Hybrid of Block-Pulse Functions and Normalized Bernstein Polynomials 

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Received 24 April 2013; Accepted 1 September 2013
Academic Editor: Santanu Saha Ray
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#### Abstract

A numerical method for solving nonlinear Fredholm integrodifferential equations is proposed. The method is based on hybrid functions approximate. The properties of hybrid of block pulse functions and orthonormal Bernstein polynomials are presented and utilized to reduce the problem to the solution of nonlinear algebraic equations. Numerical examples are introduced to illustrate the effectiveness and simplicity of the present method.


## 1. Introduction

Integrodifferential equations are often involved in mathematical formulation of physical phenomena. Fredholm integrodifferential equations play an important role in many fields such as economics, biomechanics, control, elasticity, fluid dynamics, heat and mass transfer, oscillation theory, and airfoil theory; for examples see [1-3] and references cited therein. Finding numerical solutions for Fredholm integrodifferential equations is one of the oldest problems in applied mathematics. Numerous works have been focusing on the development of more advanced and efficient methods for solving integrodifferential equations such as wavelets method [4, 5], Walsh functions method [6], sinc-collocation method [7], homotopy analysis method [8], differential transform method [9], the hybrid Legendre polynomials and blockpulse functions [10], Chebyshev polynomials method [11], and Bernoulli matrix method [12].

Block-pulse functions have been studied and applied extensively as a basic set of functions for signals and functions approximations. All these studies and applications show that block-pulse functions have definite advantages for solving problems involving integrals and derivatives due to their clearness in expressions and their simplicity in formulations; see [13]. Also, Bernstein polynomials play a prominent role in various areas of mathematics. Many authors have used these
polynomials in the solution of integral equations, differential equations, and approximation theory; see for instance [1417].

The purpose of this work is to utilize the hybrid functions consisting of combination of block-pulse functions with normalized Bernstein polynomials for obtaining numerical solution of nonlinear Fredholm integrodifferential equation:

$$
\begin{array}{r}
\sum_{i=0}^{s} p_{i}(x) y^{(i)}(x)=g(x)+\lambda \int_{0}^{1} k(x, t)[y(t)]^{q} d t, \\
0 \leq x, t<1,
\end{array}
$$

with the conditions

$$
\begin{equation*}
y^{(i)}(0)=\alpha_{i}, \quad 0 \leq i \leq s-1 \tag{2}
\end{equation*}
$$

where $y^{(i)}(x)$ is the $i$ th derivative of the unknown function that will be determined, $k(x, t)$ is the kernel of the integral, $g(x)$ and $p_{i}(x)$ are known analytic functions, $q$ is a positive integer, and $\lambda$ and $\alpha_{i}$ are suitable constants. The proposed approach for solving this problem uses few numbers of basis and benefits of the orthogonality of block-pulse functions and the advantages of orthonormal Bernstein polynomials properties to reduce the nonlinear integrodifferential equation to easily solvable nonlinear algebraic equations.

This paper is organized as follows. In the next section, we present Bernstein polynomials and hybrid of block-pulse functions. Also, their useful properties such as functions approximation, convergence analysis, operational matrix of product, and operational matrix of differentiation are given. In Section 3, the numerical scheme for the solution of (1) and (2) is described. In Section 4, the proposed method is applied to some nonlinear Fredholm integrodifferential equations, and comparisons are mad with the existing analytic or numerical solutions that were reported in other published works in the literature. Finally conclusions are given in Section 5.

## 2. Properties of Hybrid Functions

2.1. Hybrid of Block-Pulse Functions and Orthonormal Bernstein Polynomials. The Bernstein polynomials of $n$th degree are defined on the interval $[0,1]$ as $[16]$

$$
\begin{equation*}
B_{i, n}(x)=\binom{n}{i} x^{i}(1-x)^{n-i}, \quad \text { for } i=0,1,2, \ldots, n \tag{3}
\end{equation*}
$$

where

$$
\begin{equation*}
\binom{n}{i}=\frac{n!}{i!(n-i)!} \tag{4}
\end{equation*}
$$

There are $(n+1) n$th degree Bernstein polynomials. Using Gram-Schmidt orthonormalization process on $B_{i, n}(x)$, we obtain a class of orthonormal polynomials from the Bernstein polynomials. We call them orthonormal Bernstein polynomials of degree $n$ and denote them by $b_{i, n}(x), 0 \leq i \leq n$. For $n=3$, the four orthonormal Bernstein polynomials are given by

$$
\begin{gather*}
b_{0,3}(x)=-\sqrt{7}\left[x^{3}-3 x^{2}+3 x-1\right] \\
b_{1,3}(x)=\sqrt{5}\left[7 x^{3}-15 x^{2}+9 x-1\right]  \tag{5}\\
b_{2,3}(x)=-\sqrt{3}\left[21 x^{3}-33 x^{2}+13 x-1\right] \\
b_{3,3}(x)=35 x^{3}-45 x^{2}+15 x-1
\end{gather*}
$$

Hybrid functions $h_{j i}(x), j=1,2, \ldots, m$ and $i=0,1, \ldots, n$ are defined on the interval $[0,1)$ as

$$
h_{j i}(x)= \begin{cases}\sqrt{m} b_{i, n}(m x-j+1), & \frac{j-1}{m} \leq x<\frac{j}{m}  \tag{6}\\ 0, & \text { otherwise }\end{cases}
$$

where $j$ and $n$ are the order of block-pulse functions and degree of orthonormal Bernstein polynomials, respectively.

It is clear that these sets of hybrid functions in (6) are orthonormal and disjoint.
2.2. Functions Approximation. A function $y(x) \in L^{2}[0,1)$ may be approximated as

$$
\begin{equation*}
y(x) \approx \sum_{j=1}^{m} \sum_{i=0}^{n} c_{j i} h_{j i}(x)=\mathbf{C}^{T} \mathbf{H}(x) \tag{7}
\end{equation*}
$$

where

$$
\begin{gather*}
\mathbf{C}=\left[\mathbf{C}_{1}^{T}, \mathbf{C}_{2}^{T}, \ldots, \mathbf{C}_{j}^{T}, \ldots, \mathbf{C}_{m}^{T}\right]^{T}, \\
\mathbf{C}_{j}=\left[c_{j 0}, c_{j 1}, c_{j 2}, \ldots, c_{j n}\right]^{T}, \quad j=1,2, \ldots, m,  \tag{8}\\
\mathbf{H}(x)=\left[\mathbf{H}_{1}^{T}(x), \mathbf{H}_{2}^{T}(x), \ldots, \mathbf{H}_{j}^{T}(x), \ldots, \mathbf{H}_{m}^{T}(x)\right]^{T}, \tag{9}
\end{gather*}
$$

and $\mathbf{H}_{j}(x)=\left[h_{j 0}(x), h_{j 1}(x), \ldots, h_{j n}(x)\right]^{T}, j=1,2, \ldots, m$. The constant coefficients $c_{j i}$ are $\left(y(x), h_{j i}(x)\right), i=0,1,2, \ldots, n$, $j=1,2, \ldots, m$, and $(\cdot, \cdot)$ is the standard inner product on $L^{2}[0,1)$.

We can also approximate the function $k(x, t) \in L^{2}([0,1) \times$ $[0,1)$ ) by

$$
\begin{equation*}
k(x, t) \approx \sum_{i=1}^{m} \sum_{j=1}^{m} \sum_{l=0}^{n} \sum_{r=0}^{n} k_{l r}^{i j} h_{i l}(x) h_{j r}(t)=\mathbf{H}^{T}(x) \mathbf{K H}(t), \tag{10}
\end{equation*}
$$

where $\mathbf{K}=\left[\mathbf{K}^{i j}\right]$ is an $m(n+1) \times m(n+1)$ matrix, such that the elements of the sub matrix $\mathbf{k}^{i j}$ are

$$
\begin{gather*}
k_{l r}^{i j}=\int_{i-1 / m}^{i / m} \int_{j-1 / m}^{j / m} k(x, t) h_{i(l-1)}(x) h_{j(r-1)}(t) d x d t  \tag{11}\\
\quad l, r=1,2, \ldots, n+1, i, j=1,2, \ldots, m
\end{gather*}
$$

utilizing properties of block-pulse function and orthonormal Bernstein polynomials.
2.3. Convergence Analysis. In this section, the error bound and convergence is established by the following lemma.

Lemma 1. Suppose that $f \in C^{(n+1)}[0,1)$ is $n+1$ times continuously differentiable function such that $f=\sum_{j=1}^{m} f_{j}$, and let $Y_{j}=\operatorname{Span}\left\{h_{j 0}(x), h_{j 1}(x), \ldots, h_{j n}(x)\right\}, j=1,2, \ldots, m$. If $\mathbf{C}_{j}^{T} \mathbf{H}_{j}(x)$ is the best approximation to $f_{j}$ from $Y_{j}$, then $\mathbf{C}^{T} \mathbf{H}(x)$ approximates $f$ with the following error bound:

$$
\begin{gather*}
\left\|f-\mathbf{C}^{T} \mathbf{H}(x)\right\|_{2} \leq \frac{\gamma}{m^{n+1}(n+1)!\sqrt{2 n+3}}  \tag{12}\\
\gamma=\max _{x \in[0.1)}\left|f^{(n+1)}(x)\right|
\end{gather*}
$$

Proof. The Taylor expansion for the function $f_{j}(x)$ is

$$
\begin{array}{r}
\tilde{f}_{j}(x)=f_{j}\left(\frac{j-1}{m}\right)+f_{j}^{\prime}\left(\frac{j-1}{m}\right)\left(x-\frac{j-1}{m}\right) \\
+\cdots+f_{j}^{(n)}\left(\frac{j-1}{m}\right) \frac{(x-(j-1 / m))^{n}}{n!},  \tag{13}\\
\frac{j-1}{m} \leq x<\frac{j}{m}
\end{array}
$$

for which it is known that

$$
\begin{align*}
\left|f_{j}(x)-\tilde{f}_{j}(x)\right| \leq & \left|f^{(n+1)}(\eta)\right| \frac{(x-(j-1 / m))^{n+1}}{(n+1)!}  \tag{14}\\
& \eta \in\left[\frac{j-1}{m}, \frac{j}{m}\right), j=1,2, \ldots, m
\end{align*}
$$

Since $\mathbf{C}_{j}^{T} \mathbf{H}_{j}(x)$ is the best approximation to $f_{j}$ form $Y_{j}$ and $\widetilde{f}_{j} \in Y_{j}$, using (14) we have

$$
\begin{align*}
\left\|f_{j}-\mathbf{C}_{j}^{T} \mathbf{H}_{j}(x)\right\|_{2}^{2} & \leq\left\|f_{j}-\tilde{f}_{j}\right\|^{2} \\
& =\int_{j-1 / m}^{j / m}\left|f_{j}(x)-\tilde{f}_{j}(x)\right|^{2} d x \\
& \leq \int_{j-1 / m}^{j / m}\left[\frac{f^{(n+1)}(\eta)(x-(j-1 / m))^{n+1}}{(n+1)!}\right]^{2} d x \\
& \leq\left[\frac{\gamma}{(n+1)!}\right]^{2} \int_{j-1 / m}^{j / m}\left(x-\frac{j-1}{m}\right)^{2 n+2} d x \\
& =\left[\frac{\gamma}{(n+1)!}\right]^{2} \frac{1}{m^{2 n+3}(2 n+3)} . \tag{15}
\end{align*}
$$

Now,

$$
\begin{align*}
\left\|f-\mathbf{C}^{T} \mathbf{H}(x)\right\|_{2}^{2} & \leq \sum_{j=1}^{m}\left\|f_{j}-\mathbf{C}_{j}^{T} \mathbf{H}_{j}(x)\right\|_{2}^{2}  \tag{16}\\
& \leq \frac{\gamma^{2}}{m^{2 n+2}[(n+1)!]^{2}(2 n+3)} .
\end{align*}
$$

By taking the square roots we have the above bound.
2.4. The Operational Matrix of Product. In this section, we present a general formula for finding the $m(n+1) \times m(n+1)$ operational matrix of product $\widetilde{\mathbf{C}}$ whenever

$$
\begin{equation*}
\mathbf{C}^{T} \mathbf{H}(x) \mathbf{H}^{T}(x) \approx \mathbf{H}^{T}(x) \widetilde{\mathbf{C}} \tag{17}
\end{equation*}
$$

where

$$
\begin{equation*}
\widetilde{\mathbf{C}}=\operatorname{diag}\left[\widetilde{\mathbf{C}}_{1}, \widetilde{\mathbf{C}}_{2}, \ldots, \widetilde{\mathbf{C}}_{j}, \ldots, \widetilde{\mathbf{C}}_{m}\right] \tag{18}
\end{equation*}
$$

In (18), $\widetilde{\mathbf{C}}_{j}=\left[c_{l r}^{j}\right]$ are $(n+1) \times(n+1)$ symmetric matrices depending on $n$, where

$$
\begin{array}{r}
c_{l r}^{j}=\int_{j-1 / m}^{j / m}\left(h_{j(l-1)}(x) h_{j(r-1)}(x) \sum_{i=0}^{n} c_{j i} h_{j i}(x)\right) d x  \tag{19}\\
l, r=1,2, \ldots, n+1
\end{array}
$$

Furthermore, the integration of cross-product of two hybrid functions vectors is

$$
\begin{equation*}
\int_{0}^{1} \mathbf{H}(x) \mathbf{H}^{T}(x) d x=\mathbf{I} \tag{20}
\end{equation*}
$$

where $\mathbf{I}$ is the $m(n+1)$ identity matrix.
2.5. The Operational Matrix of Differentiation. The operational matrix of derivative of the hybrid functions vector $\mathbf{H}(x)$ is defined by

$$
\begin{equation*}
\frac{d}{d x} \mathbf{H}(x)=\mathbf{D H}(x) \tag{21}
\end{equation*}
$$

where $\mathbf{D}$ is the $m(n+1) \times m(n+1)$ operational matrix of derivative given as

$$
\mathbf{H}(x)=\left[\mathbf{H}_{1}^{T}(x), \mathbf{H}_{2}^{T}(x), \ldots, \mathbf{H}_{j}^{T}(x), \ldots, \mathbf{H}_{m}^{T}(x)\right]^{T}
$$

$$
\begin{equation*}
=\widetilde{\mathbf{A}} \widetilde{\mathbf{T}}(x) \tag{22}
\end{equation*}
$$

where $\widetilde{\mathbf{A}}=\operatorname{diag}\left[\mathbf{A}_{1}, \mathbf{A}_{2}, \ldots, \mathbf{A}_{j}, \ldots, \mathbf{A}_{m}\right]$ is the $m(n+1) \times$ $m(n+1)$ coefficient matrix of the $(n+1) \times(n+1)$ coefficient submatrix $\mathbf{A}_{j}$, and $\widetilde{\mathbf{T}}(x)=\left[\mathbf{t}_{1}(x), \mathbf{t}_{2}(x), \ldots, \mathbf{t}_{j}(x), \ldots, \mathbf{t}_{m}(x)\right]^{T}$ is the $m(n+1)$ vector with $\mathbf{t}_{j}(x)=\left[1, x, x^{2}, \ldots, x^{n}\right]^{T}$, such that $\mathbf{H}_{j}(x)=\mathbf{A}_{j} \mathbf{t}_{j}(x)$. Now

$$
\begin{equation*}
\frac{d}{d x} \mathbf{H}(x)=\widetilde{\mathbf{A}} \widetilde{\mathbf{Q}} \widetilde{\mathbf{T}}(x)=\widetilde{\mathbf{A}} \widetilde{\mathbf{Q}} \widetilde{\mathbf{A}}^{-1} \mathbf{H}(x) \tag{23}
\end{equation*}
$$

where $\widetilde{\mathbf{Q}}=\operatorname{diag}[\mathbf{Q}, \ldots, \mathbf{Q}]$ is the $m(n+1) \times m(n+1)$ matrix of the $(n+1) \times(n+1)$ sub-matrix $\mathbf{Q}$, such that

$$
\mathbf{Q}=\left[\begin{array}{cccccc}
0 & 0 & 0 & \cdots & 0 & 0  \tag{24}\\
1 & 0 & 0 & \cdots & 0 & 0 \\
0 & 2 & 0 & \cdots & 0 & 0 \\
\vdots & \vdots & \vdots & \cdots & \vdots & \vdots \\
0 & 0 & 0 & \cdots & n & 0
\end{array}\right]
$$

Hence,

$$
\begin{equation*}
\mathbf{D}=\widetilde{\mathbf{A}} \widetilde{\mathbf{Q}} \widetilde{\mathbf{A}}^{-1} \tag{25}
\end{equation*}
$$

In general, we can have

$$
\begin{equation*}
\frac{d^{k}}{d x^{k}} \mathbf{H}(x)=\mathbf{D}^{k} \mathbf{H}(x), \quad k=1,2,3, \ldots . \tag{26}
\end{equation*}
$$

## 3. Outline of the Solution Method

This section presents the derivation of the method for solving sth-order nonlinear Fredholm integrodifferential equation (1) with the initial conditions (2).

Step 1. The functions $y^{(i)}(x), i=0,1,2, \ldots, s$ are being approximated by

$$
\begin{equation*}
y^{(i)}(x)=\mathbf{C}^{T}(\mathbf{H}(x))^{(i)}=\mathbf{C}^{T} \mathbf{D}^{i} \mathbf{H}(x), \quad i=0,1,2, \ldots, s \tag{27}
\end{equation*}
$$

where $\mathbf{D}$ is given by (25).
Step 2. The function $k(x, t)$ is being approximated by (10).
Step 3. In this step, we present a general formula for approximate $y^{q}(x)$. By using (7) and (17), we can have

$$
\begin{align*}
y^{2}(x) & =\left[\mathbf{C}^{T} \mathbf{H}(x)\right]^{2}=\mathbf{C}^{T} \mathbf{H}(x) \mathbf{H}^{T}(x) \mathbf{C}=\mathbf{H}^{T}(x) \widetilde{\mathbf{C}} \mathbf{C}  \tag{28}\\
y^{3}(x) & =\mathbf{C}^{T} \mathbf{H}(x)\left[\mathbf{C}^{T} \mathbf{H}(x)\right]^{2}=\mathbf{C}^{T} \mathbf{H}(x) \mathbf{H}^{T}(x) \widetilde{\mathbf{C}} \mathbf{C}  \tag{29}\\
& =\mathbf{H}^{T}(x) \widetilde{\mathbf{C}} \widetilde{\mathbf{C}} \mathbf{C}=\mathbf{H}^{T}(x)(\widetilde{\mathbf{C}})^{2} \mathbf{C}
\end{align*}
$$

and so by use of induction, $y^{q}(x)$ will be approximated as

$$
\begin{equation*}
y^{q}(x)=\mathbf{H}^{T}(x)(\widetilde{\mathbf{C}})^{q-1} \mathbf{C} \tag{30}
\end{equation*}
$$

Table 1: Numerical comparison of absolute difference errors for Example 3.

| $x$ | Method of $[17]$ | The proposed method |  |
| :--- | :---: | :---: | :---: |
|  | $n=7$ | $n=2, m=30$ | $n=3, m=30$ |
| 0.0 | $3.2038 E-009$ | $3.1309 E-007$ | $4.0173 E-010$ |
| 0.2 | $7.1841 E-010$ | $3.8241 E-007$ | $4.9068 E-010$ |
| 0.4 | $1.4151 E-010$ | $4.6707 E-007$ | $5.9932 E-010$ |
| 0.6 | $4.0671 E-011$ | $5.7048 E-007$ | $7.3201 E-010$ |
| 0.8 | $9.1044 E-010$ | $6.9679 E-007$ | $8.9407 E-010$ |
| 1.0 | $3.7002 E-009$ | $8.2709 E-007$ | $1.4907 E-010$ |

Step 4. Approximate the functions $g(x)$ and $p_{i}(x)$ by

$$
\begin{gather*}
g(x) \approx \mathbf{G}^{T} \mathbf{H}(x)  \tag{31}\\
p_{i}(x) \approx \mathbf{P}_{i}^{T} H(x), \quad i=0,1,2, \ldots, s \tag{32}
\end{gather*}
$$

where $\mathbf{G}$ and $\mathbf{P}_{i}$ are constant coefficient vectors which are defined similarly to (7).

Now, using (27)-(32) and (10) to substitute into (1), we can obtain

$$
\begin{align*}
& \sum_{i=0}^{s} \mathbf{P}_{i}^{T} \mathbf{H}(x) \mathbf{H}^{T}(x)\left(\mathbf{D}^{i}\right)^{T} \mathbf{C} \\
& \quad=\mathbf{H}^{T}(x) \mathbf{G}+\lambda \int_{0}^{1} \mathbf{H}^{T}(x) \mathbf{K} \mathbf{H}(t) \mathbf{H}^{T}(t)(\widetilde{\mathbf{C}})^{q-1} \mathbf{C} d t . \tag{33}
\end{align*}
$$

Utilizing (17) and (20), we may have

$$
\begin{equation*}
\sum_{i=0}^{s} \mathbf{H}^{T}(x) \widetilde{\mathbf{P}}_{i}\left(\mathbf{D}^{i}\right)^{T} \mathbf{C}=\mathbf{H}^{T}(x) \mathbf{G}+\lambda \mathbf{H}^{T}(x) \mathbf{K}(\widetilde{\mathbf{C}})^{q-1} \mathbf{C} \tag{34}
\end{equation*}
$$

and hence we get

$$
\begin{equation*}
\sum_{i=0}^{s} \widetilde{\mathbf{P}}_{i}\left(\mathbf{D}^{i}\right)^{T} \mathbf{C}-\lambda \mathbf{K}(\widetilde{\mathbf{C}})^{q-1} \mathbf{C}=\mathbf{G} \tag{35}
\end{equation*}
$$

The matrix (35) gives a system of $m(n+1)$ nonlinear algebraic equations which can be solved utilizing the initial condition for the elements of $\mathbf{C}$. Once $\mathbf{C}$ is known, $y(x)$ can be constructed by using (7).

## 4. Applications and Numerical Results

In this section, numerical results of some examples are presented to validate accuracy, applicability, and convergence of the proposed method. Absolute difference errors of this method is compared with the existing methods reported in the literature $[5,6,17,18]$. The computations associated with these examples were performed using MATLAB 9.0.

Example 1. Consider the first-order nonlinear Fredholm integrodifferential equation $[17,18]$ as follows:

$$
\begin{equation*}
y^{\prime}(x)=1-\frac{1}{3} x+\int_{0}^{1} x y^{2}(t) d t, \quad 0 \leq x<1 \tag{36}
\end{equation*}
$$

with the initial condition

$$
\begin{equation*}
y(0)=0 \tag{37}
\end{equation*}
$$

In this example, we have $p_{0}=0, p_{1}=1, g(x)=1-(1 / 3) x$, $\lambda=1, k(x, t)=x$, and $q=2$.

The matrix (35) for this example is

$$
\begin{equation*}
\widetilde{\mathbf{P}}_{1} \mathbf{D}^{T} \mathbf{C}-\mathbf{K}(\widetilde{\mathbf{C}}) \mathbf{C}=\mathbf{G} \tag{38}
\end{equation*}
$$

where for $n=1$ and $m=2$ we have

$$
\begin{gathered}
\widetilde{\mathbf{P}}_{1}=\mathbf{I}, \quad \mathbf{D}^{T}=\left[\begin{array}{cccc}
-3 & 3 \sqrt{3} & 0 & 0 \\
-\sqrt{3} & 3 & 0 & 0 \\
0 & 0 & -3 & 3 \sqrt{3} \\
0 & 0 & -\sqrt{3} & 3
\end{array}\right], \quad \mathbf{C}=\left[\begin{array}{l}
c_{10} \\
c_{11} \\
c_{20} \\
c_{21}
\end{array}\right], \\
\mathbf{K}=\left[\begin{array}{cccc}
\frac{1}{16} & \frac{\sqrt{3}}{48} & \frac{1}{16} & \frac{\sqrt{3}}{48} \\
\frac{\sqrt{3}}{16} & \frac{1}{16} & \frac{\sqrt{3}}{16} & \frac{1}{16} \\
\frac{1}{4} & \frac{\sqrt{3}}{12} & \frac{1}{4} & \frac{\sqrt{3}}{12} \\
\frac{\sqrt{3}}{8} & \frac{1}{8} & \frac{\sqrt{3}}{8} & \frac{1}{8}
\end{array}\right], \\
\widetilde{\mathbf{C}}=\frac{1}{4}\left[\begin{array}{cccc}
3 \sqrt{6} c_{10}-\sqrt{2} c_{11} & -\sqrt{2} c_{10}+\sqrt{6} c_{11} & 0 \\
-\sqrt{2} c_{10}+\sqrt{6} c_{11} & \sqrt{6} c_{10}+5 \sqrt{2} c_{11} & 0 \\
0 & 0 & 3 \sqrt{6} c_{20}-\sqrt{2} c_{21} & -\sqrt{2} c_{20}+\sqrt{6} c_{21} \\
0 & 0 & -\sqrt{2} c_{20}+\sqrt{6} c_{21} & \sqrt{6} c_{20}+5 \sqrt{2} c_{21}
\end{array}\right],
\end{gathered}
$$

$$
\mathbf{G}=\left[\begin{array}{c}
\frac{17 \sqrt{6}}{72}  \tag{39}\\
\frac{5 \sqrt{2}}{24} \\
\frac{7 \sqrt{6}}{36} \\
\frac{\sqrt{2}}{6}
\end{array}\right] .
$$

Equation (38) gives a system of nonlinear algebraic equations that can be solved utilizing the initial condition (37); that is, $\sqrt{6} c_{10}-\sqrt{2} c_{11}=0$, we obtain

$$
\begin{array}{ll}
c_{10}=\frac{\sqrt{6}}{24}, & c_{11}=\frac{\sqrt{2}}{8}  \tag{40}\\
c_{20}=\frac{\sqrt{6}}{6}, & c_{21}=\frac{\sqrt{2}}{4} .
\end{array}
$$

Substituting these values into (7), the result will be $y(x)=x$, that is, the exact solution. It is noted that the result gives the exact solution as in [17], while in [18] using the sinc method the maximum absolute error is $1.52165 \times 10^{-3}$.

Example 2. Consider the first-order nonlinear Fredholm integrodifferential equation $[6,17]$ as follows:

$$
\begin{array}{r}
x y^{\prime}(x)-y(x)=-\frac{1}{6}+\frac{4}{5} x^{2}+\int_{0}^{1}\left(x^{2}+t\right) y^{2}(t) d t  \tag{41}\\
0 \leq x<1
\end{array}
$$

with the initial condition

$$
\begin{equation*}
y(0)=0 \tag{42}
\end{equation*}
$$

In this example, we have $p_{0}=-1, p_{1}=x, g(x)=-(1 / 6)+$ $(4 / 5) x^{2}, \lambda=1, k(x, t)=x^{2}+t$, and $q=2$.

The matrix (35) for this example is

$$
\begin{equation*}
\left(\widetilde{\mathbf{P}}_{0}+\widetilde{\mathbf{P}}_{1} \mathbf{D}^{T}\right) \mathbf{C}-\mathbf{K}(\widetilde{\mathbf{C}}) \mathbf{C}=\mathbf{G} \tag{43}
\end{equation*}
$$

where for $n=2$ and $m=2$ we have

$$
\begin{aligned}
& \widetilde{\mathbf{P}}_{0}=-\mathbf{I}, \quad \widetilde{\mathbf{P}}_{1}=\left[\begin{array}{cccccc}
\frac{1}{12} & \frac{\sqrt{15}}{60} & \frac{-\sqrt{5}}{120} & 0 & 0 & 0 \\
\frac{\sqrt{15}}{60} & \frac{1}{4} & \frac{\sqrt{3}}{24} & 0 & 0 & 0 \\
\frac{-\sqrt{5}}{120} & \frac{\sqrt{3}}{24} & \frac{5}{12} & 0 & 0 & 0 \\
0 & 0 & 0 & \frac{7}{12} & \frac{\sqrt{15}}{60} & \frac{-\sqrt{5}}{120} \\
0 & 0 & 0 & \frac{\sqrt{15}}{60} & \frac{3}{4} & \frac{\sqrt{3}}{24} \\
0 & 0 & 0 & \frac{-\sqrt{5}}{120} & \frac{\sqrt{3}}{24} & \frac{11}{12}
\end{array}\right], \\
& \mathbf{D}^{T}=\left[\begin{array}{cccccc}
-5 & \frac{7 \sqrt{15}}{3} & -2 \sqrt{5} & 0 & 0 & 0 \\
\frac{-\sqrt{15}}{3} & -3 & \frac{14 \sqrt{3}}{3} & 0 & 0 & 0 \\
0 & \frac{-8 \sqrt{3}}{3} & 8 & 0 & 0 & 0 \\
0 & 0 & 0 & -5 & \frac{7 \sqrt{15}}{3} & -2 \sqrt{5} \\
0 & 0 & 0 & \frac{-\sqrt{15}}{3} & \frac{-3}{} & \frac{14 \sqrt{3}}{3} \\
0 & 0 & 0 & 0 & \frac{-8 \sqrt{3}}{3} & 8
\end{array}\right], \quad \mathbf{C}=\left[\begin{array}{l}
c_{10} \\
c_{11} \\
c_{12} \\
c_{20} \\
c_{21} \\
c_{22}
\end{array}\right], \quad \mathbf{G}=\left[\begin{array}{c}
-\frac{11 \sqrt{10}}{450} \\
\frac{-\sqrt{6}}{90} \\
\frac{\sqrt{2}}{180} \\
\frac{23 \sqrt{10}}{900} \\
\frac{13 \sqrt{6}}{180} \\
\frac{19 \sqrt{2}}{180}
\end{array}\right],
\end{aligned}
$$

$$
\begin{align*}
& \mathbf{K}=\left[\begin{array}{cccccc}
\frac{1}{24} & \frac{\sqrt{15}}{45} & \frac{7 \sqrt{5}}{240} & \frac{13}{72} & \frac{\sqrt{15}}{20} & \frac{41 \sqrt{5}}{720} \\
\frac{\sqrt{15}}{72} & \frac{1}{12} & \frac{5 \sqrt{3}}{144} & \frac{\sqrt{15}}{24} & \frac{1}{16} & \frac{\sqrt{3}}{16} \\
\frac{\sqrt{5}}{48} & \frac{5 \sqrt{3}}{144} & \frac{1}{24} & \frac{7 \sqrt{5}}{144} & \frac{\sqrt{3}}{16} & \frac{5}{72} \\
\frac{7}{48} & \frac{31 \sqrt{15}}{720} & \frac{\sqrt{15}}{20} & \frac{41}{144} & \frac{17 \sqrt{15}}{240} & \frac{7 \sqrt{5}}{90} \\
\frac{7 \sqrt{15}}{144} & \frac{3}{16} & \frac{5 \sqrt{3}}{72} & \frac{11 \sqrt{15}}{144} & \frac{13}{48} & \frac{7 \sqrt{3}}{72} \\
\frac{\sqrt{5}}{16} & \frac{11 \sqrt{3}}{144} & \frac{1}{12} & \frac{13 \sqrt{5}}{144} & \frac{5 \sqrt{3}}{48} & \frac{1}{9}
\end{array}\right], \\
& \widetilde{\mathbf{c}}_{j}=\left[\begin{array}{ccc}
\frac{5 \sqrt{10}}{7} c_{j 0}-\frac{5 \sqrt{6}}{21} c_{j 1}+\frac{\sqrt{2}}{7} c_{j 2} & -\frac{5 \sqrt{6}}{21} c_{j 0}+\frac{11 \sqrt{10}}{35} c_{j 1}-\frac{8 \sqrt{30}}{105} c_{j 2} & \frac{\sqrt{2}}{7} c_{j 0}-\frac{8 \sqrt{30}}{105} c_{j 1}+\frac{3 \sqrt{10}}{35} c_{j 2} \\
-\frac{5 \sqrt{6}}{21} c_{j 0}+\frac{11 \sqrt{10}}{35} c_{j 1}-\frac{8 \sqrt{30}}{105} c_{j 2} & \frac{11 \sqrt{10}}{35} c_{j 0}+\frac{3 \sqrt{6}}{7} c_{j 1}+\frac{\sqrt{2}}{7} c_{j 2} & -\frac{8 \sqrt{30}}{105} c_{j 0}+\frac{\sqrt{2}}{7} c_{j 1}+\frac{5 \sqrt{6}}{21} c_{j 2} \\
\frac{\sqrt{2}}{7} c_{j 0}-\frac{8 \sqrt{30}}{105} c_{j 1}+\frac{3 \sqrt{10}}{35} c_{j 2} & -\frac{8 \sqrt{30}}{105} c_{j 0}+\frac{\sqrt{2}}{7} c_{j 1}+\frac{5 \sqrt{6}}{21} c_{j 2} & \frac{3 \sqrt{10}}{35} c_{j 0}+\frac{5 \sqrt{6}}{21} c_{j 1}+\frac{13 \sqrt{2}}{7} c_{j 2}
\end{array}\right], \\
& j=1,2 . \tag{44}
\end{align*}
$$

Equation (43) gives a system of nonlinear algebraic equations that can be solved utilizing the initial condition (42); that is, $\sqrt{10} c_{10}-\sqrt{6} c_{11}+\sqrt{2} c_{12}=0$, we obtain

$$
\begin{array}{ll}
c_{10}=\frac{\sqrt{10}}{240}, & c_{11}=\frac{\sqrt{6}}{48} \\
c_{12}=\frac{\sqrt{2}}{24}, & c_{20}=\frac{\sqrt{10}}{15}  \tag{45}\\
c_{21}=\frac{\sqrt{6}}{8}, & c_{22}=\frac{\sqrt{2}}{6}
\end{array}
$$

Substituting these values into (7), the result will be $y(x)=x^{2}$, that is, the exact solution. It is noted that the result gives the exact solution as in [17], while in [6] approximate solution is obtained with maximum absolute error $1.0000 \times 10^{-5}$.

Example 3. Consider the second-order nonlinear Fredholm integrodifferential equation [17] as follows:

$$
\begin{array}{r}
y^{\prime \prime}(x)+x y^{\prime}(x)-x y(x)=e^{x} \sin x+\int_{0}^{1} \sin x \cdot e^{-2 t} y^{2}(t) d t \\
0 \leq x<1 \tag{46}
\end{array}
$$

with the initial conditions

$$
\begin{equation*}
y(0)=y^{\prime}(0)=1 \tag{47}
\end{equation*}
$$

The exact solution is $y(x)=e^{x}$. We solve this example by using the proposed method with $n=2, m=30$ and $n=3$,
$m=30$. Comparison among the proposed method and methods in [17] is shown in Table 1. It is clear from this table that the results obtained by the proposed method, using few numbers of basis, are very promising and superior to that of [17].

Example 4. Consider the following nonlinear Fredholm integrodifferential equation [5, 17]:

$$
\begin{equation*}
y^{\prime}(x)+y(x)=\frac{1}{2}\left(e^{-2}-1\right)+\int_{0}^{1} y^{2}(t) d t, \quad 0 \leq x<1 \tag{48}
\end{equation*}
$$

with the initial conditions

$$
\begin{equation*}
y(0)=1 \tag{49}
\end{equation*}
$$

The exact solution of this problem is $y(x)=e^{-x}$. In Table 2 we have compared the absolute difference errors of the proposed method with the collocation method based on Haar wavelets in [5] and method in [17].

Maximum absolute errors of Example 4 for some different values of $n$ and $m$ are shown in Table 3. As it is seen from Table 3, for a certain value of $n$ as $m$ increases the accuracy increases, and for a certain value of $m$ as $n$ increases the accuracy increases as well. In case of $m=1$, the numerical solution obtained is based on orthonormal Bernstein polynomials only, while in case of $n=0$, the numerical solution obtained is based on block-pulse functions only.

Table 2: Numerical comparison of absolute difference errors for Example 4.

| Method of [5] <br> Number of collocation points <br> $N=128$ |  |  |  |  |  |  |  | Method of $[17]$ | The proposed method |  |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  | $n=7$ | $n=3, m=35$ | $n=4, m=15$ |  |  |  |  |  |  |  |
| 0.125 | $3.7591 E-007$ | $2.4509 E-010$ | $5.5200 E-011$ | $1.6710 E-011$ |  |  |  |  |  |  |
| 0.250 | $6.6413 E-007$ | $1.0202 E-010$ | $8.9982 E-011$ | $3.9705 E-012$ |  |  |  |  |  |  |
| 0.375 | $8.6917 E-007$ | $1.6139 E-010$ | $9.4606 E-011$ | $1.2126 E-011$ |  |  |  |  |  |  |
| 0.500 | $1.0020 E-006$ | $3.2362 E-010$ | $9.2457 E-011$ | $1.8312 E-012$ |  |  |  |  |  |  |
| 0.625 | $1.0757 E-006$ | $1.9197 E-010$ | $7.4991 E-011$ | $8.1299 E-012$ |  |  |  |  |  |  |
| 0.750 | $1.1029 E-006$ | $6.6120 E-011$ | $4.9442 E-011$ | $7.7237 E-012$ |  |  |  |  |  |  |
| 0.875 | $1.0944 E-006$ | $2.2417 E-010$ | $2.6083 E-011$ | $2.5547 E-012$ |  |  |  |  |  |  |

Table 3: Maximum absolute errors for different values of $n$ and $m$ for Example 4.

| $n$ | 1 | 5 | 10 | 15 | $m$ | 20 | 25 | 30 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  | 1 | $5.7735 E-01$ | $1.1547 E-01$ | $5.7735 E-02$ | $3.8490 E-02$ | $2.8868 E-02$ | $2.3094 E-02$ | $1.9245 E-02$ |
| $1.6496 E-02$ |  |  |  |  |  |  |  |  |
| 1 | $2.2361 E-01$ | $8.9443 E-03$ | $2.2361 E-03$ | $9.9381 E-04$ | $5.5902 E-04$ | $3.5777 E-04$ | $2.4845 E-04$ | $1.8254 E-04$ |
| 2 | $6.2994 E-02$ | $5.0395 E-04$ | $6.2994 E-05$ | $1.8665 E-05$ | $7.8743 E-06$ | $4.0316 E-06$ | $2.3331 E-06$ | $1.4693 E-06$ |
| 3 | $1.3889 E-02$ | $2.2222 E-05$ | $1.3889 E-06$ | $2.7435 E-07$ | $8.6806 E-08$ | $3.5556 E-08$ | $1.7147 E-08$ | $9.2554 E-09$ |
| 4 | $2.5126 E-03$ | $8.0403 E-07$ | $2.5126 E-08$ | $3.3088 E-09$ | $7.8519 E-10$ | $2.5729 E-10$ | $1.0340 E-10$ | $4.7839 E-11$ |
| 5 | $3.8521 E-04$ | $2.4653 E-08$ | $3.8521 E-10$ | $3.3818 E-11$ | $6.0189 E-12$ | $1.5778 E-12$ | $5.2841 E-13$ | $2.0955 E-13$ |
| 6 | $5.1230 E-05$ | $6.5574 E-10$ | $5.1230 E-12$ | $2.9984 E-13$ | $4.0023 E-14$ | $8.3935 E-15$ | $2.3425 E-15$ | $7.9625 E-16$ |

## 5. Conclusion

In this work, we present a numerical method for solving nonlinear Fredholm integrodifferential equations based on hybrid of block-pulse functions and normalized Bernstein polynomials. One of the most important properties of this method is obtaining the analytical solutions if the equation has an exact solution, that is, a polynomial function. Another considerable advantage is this method has high relative accuracy for small numbers of basis $n$. The matrices $\mathbf{K}$, $\widetilde{\mathbf{C}}$, and $\mathbf{D}$ in (10), (17), and (25), respectively, have large numbers of zero elements, and they are sparse; hence, the present method is very attractive and reduces the CPU time and computer memory. Moreover, satisfactory results of illustrative examples with respect to several other methods (e.g., Haar wavelets method, Walsh functions method, Bernstein polynomials method, and sinc collocation method) are included to demonstrate the validity and applicability of the proposed method.

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## Research Article

# Semi-Idealized Study on Estimation of Partly and Fully Space Varying Open Boundary Conditions for Tidal Models 

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Received 5 June 2013; Revised 1 September 2013; Accepted 1 September 2013
Academic Editor: Rasajit Bera
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#### Abstract

Two strategies for estimating open boundary conditions (OBCs) with adjoint method are compared by carrying out semi-idealized numerical experiments. In the first strategy, the OBC is assumed to be partly space varying and generated by linearly interpolating the values at selected feature points. The advantage is that the values at feature points are taken as control variables so that the variations of the curves can be reproduced by the minimum number of points. In the second strategy, the OBC is assumed to be fully space varying and the values at every open boundary points are taken as control variables. A series of semi-idealized experiments are carried out to compare the effectiveness of two inversion strategies. The results demonstrate that the inversion effect is in inverse proportion to the number of feature points which characterize the spatial complexity of open boundary forcing. The effect of illposedness of inverse problem will be amplified if the observations contain noises. The parameter estimation problems with more control variables will be much more sensitive to data noises, and the negative effects of noises can be restricted by reducing the number of control variables. This work provides a concrete evidence that ill-posedness of inverse problem can generate wrong parameter inversion results and produce an unreal "good data fitting."


## 1. Introduction

The tides and tidal currents are the basic motion forms of ocean water and play an important role in the research on other processes, such as the storm surge, the circulation and the estuarine dynamics [1, 2]. For tidal models, open boundary conditions (OBCs) are one of the most important parameters, which are determined by the physics of tides and tidal currents. Therefore, how to obtain reasonable and accurate OBCs for regional tidal models has been a subject of ongoing research. Data assimilation methods have been commonly used to optimize the open boundary conditions [3-7].

Data assimilation methods, especially the complex ones like four-dimensional variational (4DVAR), are developed on the base of rigorous mathematical theories, such as inverse problem theory and optimal control theory. The
ultimate purpose of applying data assimilation method is to reduce the data misfit between model results and various observations, by either improving the models or dynamically interpolating the observations. Among all the data assimilation methods, the 4DVAR is one of the most effective and powerful approaches. It is based on the optimal control methods and perturbation theory [8, 9]. This technique allows us to retrieve an optimal data for a given model from heterogeneous observation fields [9]. It is an advanced data assimilation method which involves the adjoint method and has the advantage of directly assimilating various observations distributed in time and space into numerical models while maintaining dynamical and physical consistency with the model. The adjoint method is a powerful tool for parameter estimation. Navon [10] presented an important overview on the state of the art of parameter estimation in meteorology and oceanography in view of application of

4DVAR data assimilation techniques to inverse parameter estimation problems. Zhang and Lu [7] studied the parameter estimation problems with a three-dimensional tidal model with $4 D V A R$ and also summarized relative works. More recently, Kazantsev [9] briefly revealed the history of data assimilation starting from Lorenz's pioneering work and then deeply studied the sensitivity of a shallow-water model to parameters by applying adjoint based technique.

For parameter estimation problems, it is of great importance to reasonably reduce the number of spatially varying control variables because of the ill-posedness of inverse problem. As noted by Yeh in the work of ground water flow parameter estimation, the inverse or parameter estimation problem is often ill-posed and beset by instability and nonuniqueness, particularly if one seeks parameters distributed in space and time domain [11]. The same viewpoint has been put forward by references [12-16]. Consequently, how to reduce the number of parameters to be estimated became an important aspect needing to draw attention to [13-17]. In this work two strategies for inverting the open boundary conditions with adjoint method are compared by carrying out semi-idealized numerical experiments. In the first strategy, the OBC is assumed to be partly space varying and generated by linearly interpolating the values at selected feature points. The feature points are selected by calculating the second-order derivatives of discrete curves and the values at selected feature points are taken as control variables to be estimated. The advantage is that most of the variations of the curves can be reproduced by the minimum number of points. In the second strategy, the $O B C$ is assumed to be fully space varying and the values at every open boundary points are taken as control variables.

This paper is organized as follows. The 2D tidal model with adjoint is briefly described in Section 2 . The two inversion strategies are developed in Section 3. A series of semiidealized numerical experiments are carried out and the results are analyzed and discussed in Section 4. Conclusions in Section 5 complete the paper.

## 2. The Adjoint Tidal Model

2.1. The $2 D$ Tidal Model. The governing equations for the tides used in the present study are the vertically integrated equations of continuity and momentum:

$$
\begin{gather*}
\frac{\partial \zeta}{\partial t}+\frac{1}{a} \frac{\partial[(h+\zeta) u]}{\partial \lambda}+\frac{1}{a} \frac{\partial[(h+\zeta) v \cos \phi]}{\partial \phi}=0 \\
\frac{\partial u}{\partial t}+\frac{u}{a} \frac{\partial u}{\partial \lambda}+\frac{v}{R} \frac{\partial u}{\partial \phi}-\frac{u v \tan \phi}{R}-f v-A \Delta u+\frac{g}{a} \frac{\partial(\zeta-\bar{\zeta})}{\partial \lambda}=F_{\lambda} \\
\frac{\partial v}{\partial t}+\frac{u}{a} \frac{\partial v}{\partial \lambda}+\frac{v}{R} \frac{\partial v}{\partial \phi}+\frac{u^{2} \tan \phi}{R}+f u-A \Delta v+\frac{g}{R} \frac{\partial(\zeta-\bar{\zeta})}{\partial \phi}=F_{\phi} \tag{1}
\end{gather*}
$$

where $t$ is time; $\lambda$ and $\phi$ are the east longitude and north latitude, respectively; $\zeta$ is the sea surface elevation above the undisturbed sea level; $u$ and $v$ are the east and north components of fluid velocity, respectively, $\bar{\zeta}$ is the adjusted
height of equilibrium tides; $R$ is the radius of the earth, $a=R \cos \phi ; f=2 \Omega \sin \phi$, where $\Omega$ represents the angular speed of earth rotation; $g$ is the acceleration due to gravity, $h$ is the undisturbed water depth and $H=h+\zeta$ denotes the total water depth; $A$ is the coefficient of horizontal eddy viscosity; $\Delta$ is the Laplace operator and $\Delta(u, v)=$ $a^{-1}\left[a^{-1} \partial_{\lambda}\left(\partial_{\lambda}(u, v)\right)+R^{-1} \partial_{\phi}\left(\cos \phi \partial_{\phi}(u, v)\right)\right] ; F_{\lambda}$ and $F_{\phi}$ are east and north components of bottom friction terms, respectively, and their expressions are given in quadratic form:

$$
\begin{equation*}
F_{\lambda}=-C_{Q} \frac{\sqrt{u^{2}+v^{2}}}{h+\zeta} u, \quad F_{\phi}=-C_{Q} \frac{\sqrt{u^{2}+v^{2}}}{h+\zeta} v . \tag{2}
\end{equation*}
$$

2.2. The Adjoint. The general idea of the adjoint method is described as follows. First, a model is defined by an algorithm and its independent variables such as initial conditions, boundary conditions, and empirical parameters. The cost function which measures the data misfit between the modeling results and observations is then minimized through optimizing the control variables. In detail, the cost function decreases along the opposite direction of the gradients with respect to the control variables, and this gradient is calculated by what has become known as the adjoint model. In order to construct the adjoint equations, the cost function is defined as

$$
\begin{equation*}
J(\zeta)=\frac{1}{2} K_{\zeta} \iint_{\Omega_{T, S}}(\zeta-\widehat{\zeta})^{2} d S d T \tag{3}
\end{equation*}
$$

and the Lagrangian function is defined as

$$
\begin{align*}
L=\iint_{\Omega_{T, S}}\left[\mu \left(\frac{\partial u}{\partial t}\right.\right. & +\frac{u}{a} \frac{\partial u}{\partial \lambda}+\frac{v}{R} \frac{\partial u}{\partial \phi}-\frac{u v \tan \phi}{R} \\
& \left.-f v-F_{\lambda}-A \Delta u+\frac{g}{a} \frac{\partial(\zeta-\bar{\zeta})}{\partial \lambda}\right) \\
& +v\left(\frac{\partial v}{\partial t}+\frac{u}{a} \frac{\partial v}{\partial \lambda}+\frac{v}{R} \frac{\partial v}{\partial \phi}+\frac{u^{2} \tan \phi}{R}\right. \\
& \left.+f u-F_{\phi}-A \Delta v+\frac{g}{R} \frac{\partial(\zeta-\bar{\zeta})}{\partial \phi}\right)  \tag{4}\\
& +\tau\left(\frac{\partial \zeta}{\partial t}+\frac{1}{a} \frac{\partial[(h+\zeta) u]}{\partial \lambda}\right. \\
& \left.\left.+\frac{1}{a} \frac{\partial[(h+\zeta) v \cos \phi]}{\partial \phi}\right)\right] d S d T \\
& +J(\zeta),
\end{align*}
$$

where $\hat{\zeta}$ is the observations of surface elevation; $\Omega_{T, S}$ stands for the whole integration area of time and space; $\mu, \nu$, and $\tau$ are the adjoint variables (namely, Lagrangian multipliers) of $u, v$, and $\zeta$, respectively. Based on the theory of Lagrangian


FIgure 1: Example of discrete curves and their feature points. GP stands for general points and FP indicates feature points.
multiplier method, we have the following first-order derivates of Lagrangian function with respect to all the model variables:

$$
\begin{array}{lll}
\frac{\partial L}{\partial \zeta}=0, & \frac{\partial L}{\partial u}=0, & \frac{\partial L}{\partial v}=0 \\
\frac{\partial L}{\partial \tau}=0, & \frac{\partial L}{\partial \mu}=0, & \frac{\partial L}{\partial v}=0 \\
\frac{\partial L}{\partial C_{Q}}=0, & \frac{\partial L}{\partial a}=0, & \frac{\partial L}{\partial b}=0 \tag{5c}
\end{array}
$$

Equations (5b) give the original governing (1) and the adjoint equations can be developed from (5a). In (5c), $a$ and $b$ are the Fourier coefficients along the open boundary and $C_{Q}$ denotes the bottom friction coefficients. From (5c) we can obtain the optimization formulae of model parameters.

Based on (5a) the adjoint equations can be obtained as

$$
\begin{aligned}
& \frac{\partial \tau}{\partial t}+ \frac{u}{a} \frac{\partial \tau}{\partial \lambda}+\frac{v}{a} \frac{\partial(\tau \cos \phi)}{\partial \phi}+\frac{g}{a} \frac{\partial \mu}{\partial \lambda}+\frac{g}{a} \frac{\partial v}{\partial \phi}-K_{\zeta}(\zeta-\widehat{\zeta}) \\
&=\Psi(1,1)+\Psi(1,2) \\
& \frac{\partial \mu}{\partial t}-f v-\frac{\mu}{a} \frac{\partial u}{\partial \lambda}-\frac{v}{a} \frac{\partial v}{\partial \lambda}+\frac{1}{a} \frac{\partial}{\partial \lambda}(\mu u)+\frac{1}{R} \frac{\partial}{\partial \phi}(\mu v) \\
&+\frac{h+\zeta}{a} \frac{\partial \tau}{\partial \lambda}+\frac{h+\zeta}{a} \frac{\partial \tau}{\partial \lambda}+A \Delta \mu+\frac{\mu v \tan \phi}{R} \\
&-\frac{2 v u \tan \phi}{R}=\Psi(2,1)+\Psi(2,2), \\
& \frac{\partial v}{\partial t}+ f \mu-\frac{\mu}{R} \frac{\partial u}{\partial \phi}-\frac{v}{R} \frac{\partial v}{\partial \phi}+\frac{1}{a} \frac{\partial}{\partial \lambda}(v u)+\frac{1}{R} \frac{\partial}{\partial \phi}(\nu v) \\
&+\frac{h+\zeta}{a} \frac{\partial(\tau \cos \phi)}{\partial \phi}+\frac{h+\zeta}{a} \frac{\partial(\tau \cos \phi)}{\partial \phi}+A \Delta v \\
&+\frac{\mu u \tan \phi}{R}=\Psi(3,1)+\Psi(3,2),
\end{aligned}
$$

where $\Psi(i, j)(1 \leq i \leq 3,1 \leq j \leq 2)$ is a matrix whose components denote the adjoint terms of bottom friction. The components of $\Psi$ for the quadratic parameterizations are given as

$$
\Psi=\left\{\begin{array}{cc}
-\mu \frac{C_{\mathrm{Q}} u \sqrt{u^{2}+v^{2}}}{(h+\zeta)^{2}} & -v \frac{C_{\mathrm{Q}} v \sqrt{u^{2}+v^{2}}}{(h+\zeta)^{2}}  \tag{7}\\
\mu \frac{C_{\mathrm{Q}}\left(2 u^{2}+v^{2}\right)}{(h+\zeta) \sqrt{u^{2}+v^{2}}} & v \frac{C_{\mathrm{Q}} u v}{(h+\zeta) \sqrt{u^{2}+v^{2}}} \\
\mu \frac{C_{\mathrm{Q}} u v}{(h+\zeta) \sqrt{u^{2}+v^{2}}} & v \frac{C_{\mathrm{Q}}\left(u^{2}+2 v^{2}\right)}{(h+\zeta) \sqrt{u^{2}+v^{2}}}
\end{array}\right\},
$$

The numerical schemes for the forward model and the adjoint model in this section are both based on Lu and Zhang [17] and Zhang et al. [18].

## 3. Methodology

3.1. Feature Points of a Curve. If the values of OBCs are plotted versus the location or index of grid points along open boundaries, they will form a discretized curve. Without loss of generality, the curve can be presented by Figure 1. Assume there are $N$ general (or, computational) points along open boundaries with index of $\operatorname{GP}(k), k=1,2, \ldots, N$. This type of curve can be approximately linearly expressed by a certain series of points which are defined as feature points in this paper. For the curve shown in Figure 1, one can easily obtain the feature points as indicated by symbol " + ." Assume the number of feature points is $M$ with index of $\operatorname{FP}(j), j=$ $1,2, \ldots, M$. Further assuming the feature point with index of $j$ is coincident with the general point with index of $\operatorname{II}(j)$, we can obtain the following relation: $\mathrm{II}(1)=1, \mathrm{II}(M)=N$, $\mathrm{II}(j)=k, 2<k<N-1$.

It is easy to conclude that any general point can be linearly expressed by two adjacent feature points. For example, as shown in Figure 1, an arbitrary general point $\operatorname{GP}(k)$ locates between two adjacent feature points $\mathrm{FP}(j-1)$ and $\mathrm{FP}(j)$, where $\mathrm{II}(j-1) \leq k \leq \mathrm{II}(j)$. Through linear interpolation, we can obtain the value of $\operatorname{GP}(k)$ as

GP ( $k$ )

$$
=\frac{\mathrm{II}(j)-k}{\mathrm{II}(j)-\mathrm{II}(j-1)} \mathrm{FP}(j-1)+\frac{k-\mathrm{II}(j-1)}{\mathrm{II}(j)-\mathrm{II}(j-1)} \mathrm{FP}(j) .
$$

For the whole curve (or the whole boundary), the relation between general points and feature points can be similarly expressed in matrix form as

$$
\begin{equation*}
\mathbf{V}_{\mathrm{GP}}=\mathbf{W}_{\mathrm{FG}} \times \mathrm{V}_{\mathrm{FP}} \tag{9}
\end{equation*}
$$

where $\mathbf{V}_{\mathbf{G P}}$ and $\mathbf{V}_{\mathbf{F P}}$ are both column vectors with dimensions of $N$ and $M$, respectively, and $\mathbf{W}_{\mathbf{F G}}$ is the weighting matrix of linear interpolation with dimensions of $N \times M$. The detailed forms of three matrixes are given as

$$
\begin{gathered}
\mathbf{V}_{\mathbf{G P}}=[\mathrm{GP}(1), \mathrm{GP}(2), \ldots, \mathrm{GP}(N)]^{T}, \\
\mathbf{W}_{\mathbf{F G}}=\left(\begin{array}{cccccccc}
w_{1,1} & 0 & 0 & 0 & 0 & 0 & 0 & 0 \\
w_{2,1} & w_{22} & 0 & 0 & 0 & 0 & 0 & 0 \\
\vdots & \vdots & 0 & 0 & 0 & 0 & 0 & 0 \\
w_{\mathrm{II}(2)-1,1} & w_{\mathrm{II}(2)-1,2} & 0 & 0 & 0 & 0 & 0 & 0 \\
0 & w_{\mathrm{II}(2), 2} & 0 & 0 & 0 & 0 & 0 & 0 \\
0 & 0 & \ddots & 0 & 0 & 0 & 0 & 0 \\
0 & 0 & 0 & w_{\mathrm{II}(j-1), j-1} & 0 & 0 & 0 & 0 \\
0 & 0 & 0 & w_{\mathrm{II}(j-1)+1, j-1} & w_{\mathrm{II}(j-1)+1, j} & 0 & 0 & 0 \\
0 & 0 & 0 & \vdots & \vdots & 0 & 0 & 0 \\
0 & 0 & 0 & w_{\mathrm{II}(j)-1, j-1} & w_{\mathrm{II}(j)-1, j} & 0 & 0 & 0 \\
0 & 0 & 0 & 0 & w_{\mathrm{III}(j), j} & 0 & 0 & 0 \\
0 & 0 & 0 & 0 & 0 & \ddots & 0 & 0 \\
0 & 0 & 0 & 0 & 0 & 0 & w_{\mathrm{II}(M-1), M-1} & 0 \\
0 & 0 & 0 & 0 & 0 & 0 & w_{\mathrm{II}(M-1)+1, M-1} & w_{\mathrm{II}(M-1)+1, M} \\
0 & 0 & 0 & 0 & 0 & 0 & \vdots & \vdots \\
0 & 0 & 0 & 0 & 0 & 0 & w_{\mathrm{II}(M)-1, M-1} & w_{\mathrm{II}(M)-1, M} \\
0 & 0 & 0 & 0 & 0 & 0 & & w_{\mathrm{II}(M), M}
\end{array}\right),
\end{gathered}
$$

where the nonzero components are the linear interpolation coefficients. Specifically, without loss of generality,

$$
\begin{gather*}
w_{\mathrm{II}(j), j}=1.0, \quad j=1,2, \ldots, M,  \tag{13a}\\
w_{\mathrm{II}(j-1)+m, j-1}=\frac{\mathrm{II}(j)-\mathrm{II}(j-1)-m}{\mathrm{II}(j)-\mathrm{II}(j-1)},  \tag{13b}\\
1 \leq m<\mathrm{II}(j)-\mathrm{II}(j-1), \\
w_{\mathrm{II}(j-1)+m, j}=\frac{m}{\mathrm{II}(j)-\mathrm{II}(j-1)},  \tag{13c}\\
1 \leq m<\mathrm{II}(j)-\mathrm{II}(j-1)
\end{gather*}
$$

Using (9), any general points along open boundaries can be highly approximated through the linear interpolation of
selected feature points. It indicates that the OBC identification problem can be transformed to seek the values of a few selected feature points, which reduces the number of control variables.

### 3.2. Selection of Feature Points for Periodic Tidal Open Bound-

 ary. Along a certain open boundary, we also assume that there are $N$ general grid points. The height of water level $\zeta$ at the $n$th time step is given by$$
\begin{equation*}
\zeta_{\mathrm{GP}(k)}^{n}=a_{0}+\left[a_{\mathrm{GP}(k)} \cos (\omega n \Delta t)+b_{\mathrm{GP}(k)} \sin (\omega n \Delta t)\right] \tag{14}
\end{equation*}
$$

where $\mathrm{GP}(k)$ stands for the general points of open boundaries and $1 \leqslant k \leqslant N, \omega$ is the frequency of $M_{2}$ constituent, $a_{\mathrm{GP}(k)}$ and $b_{\mathrm{GP}(k)}$ are the Fourier coefficients at $\mathrm{GP}(k), \Delta t$ is the time step of computation.

For regional tidal models the values of $a_{\mathrm{GP}(k)}$ and $b_{\mathrm{GP}(k)}$ can be obtained from large scale numerical models. It should be noted $a_{\mathrm{GP}(k)}$ and $b_{\mathrm{GP}(k)}$ are space dependent, and therefore the variations of their values versus the grids along the open boundary will constitute two curves (curve $a$ and curve $\_b$ ) similar to the one shown in Figure 1. The feature points for this type of curve can be selected by computing the secondorder differential of each general point. The detailed selection procedures are given as follows.
(1) Suppose the absolute values of second-order differentials of general points $\mathrm{GP}(k)$ are SD $\_a(k)$ for curve $\_a$ and SD $\_b(k)$ for curve $\_b$, respectively. For the general points locating in the middle of curve_a and curve_b, that is, $2 \leqslant k \leqslant N-1$, SD $\_a(k)$ and SD $\_b(k)$ can be computed as

$$
\begin{align*}
& \left.\mathrm{SD} \mathrm{\_a(k)=\mid} \mathrm{\frac{a} _{\mathrm{GP}(k+1)}-2 a_{\mathrm{GP}(k)}+a_{\mathrm{GP}(k-1)}}{2 \Delta d} \right\rvert\,, \\
& \mathrm{SD} \_b(k)=\left|\frac{b_{\mathrm{GP}(k+1)}-2 b_{\mathrm{GP}(k)}+b_{\mathrm{GP}(k-1)}}{2 \Delta d}\right|, \tag{15}
\end{align*}
$$

where $\Delta d$ is the size of computation grids and equals $\Delta x$ or $\Delta y$ according to the direction of open boundaries ( $\Delta x$ for west-east direction and $\Delta y$ for northsouth direction).
(2) Further define that the "maximum second-order differential" for point $\mathrm{GP}(k)$ is $\mathrm{SD}(k)$. The value of $\mathrm{SD}(k)$ is calculated as

$$
\begin{equation*}
\mathrm{SD}(k)=\max \left[\text { SD } \_a(k), \text { SD } \_b(k)\right] . \tag{16}
\end{equation*}
$$

(3) Define a threshold value of $\operatorname{SD}(k), 2 \leqslant k \leqslant N-1$, to be $T_{\mathrm{SD}}$. The points with larger values of $\mathrm{SD}(k)$ than $T_{\text {SD }}$ are selected as feature points. The value of $T_{\text {SD }}$ is problem dependent and should be determined according to the specific requirement on the number of control variables.
(4) It is easy to understand that the first and the last general points $\mathrm{GP}(1)$ and $\mathrm{GP}(N)$ are automatically selected as feature points indexed as $\mathrm{FP}(1)$ and $\mathrm{FP}(M)$.
3.3. Inversion Strategies and Gradients. In this work two strategies for inverting the open boundary conditions with adjoint method are compared by carrying out semi-idealized numerical experiments. In the first strategy the open boundary curves are assumed to be partly space varying and are generated by linearly interpolating the values at feature points. The feature points are selected by calculating the second-order derivatives of discrete curves and the values at selected feature points are taken as control variables to be estimated. The advantage is that most of the variations of the curves can be reproduced by the minimum number of points. In the second strategy, the OBC is assumed to be fully space varying and the values at every open boundary point are taken as control variables.

The Broyden-Fletcher-Goldfarb-Shanno (BFGS) method, which is a quasi-Newton conjugate-gradient algorithm, has been widely used in the unconstrained inverse problems and is famous for its efficiency [19, 20]. The limitedmemory BFGS (L-BFGS) algorithm is an adaptation of the BFGS method to large problem. Zou et al. [20] concluded that among the tested quasi-Newton methods, the L-BFGS method had the best performance. In this work L-BFGS method is employed to optimize the control variables, namely, the OBCs. In order to perform inversion with LBFGS, the gradients of cost function with respect to the control variables in two strategies have to be calculated.
3.3.1. Gradients for Partly Space Varying Inversion Strategy. In the first inversion strategy (partly space varying OBC), feature points for open boundary curves are selected and the OBCs at general points can be linearly interpolated from feature points. Consequently, the gradients of cost function with respect to the Fourier coefficients at feature points $a a_{\mathrm{FP}(j)}$ and $b b_{\mathrm{FP}(j)}\left(a a_{j}\right.$ and $b b_{j}$ for simplicity, $\left.1 \leqslant j \leqslant M\right)$ have to be computed in order to optimize the OBCs with LBFGS. The gradients are deduced from

$$
\begin{equation*}
\frac{\partial L}{\partial a a_{j}}=0, \quad \frac{\partial L}{\partial b b_{j}}=0, \quad 1 \leq j \leq M \tag{17}
\end{equation*}
$$

which yields

$$
\begin{array}{r}
\frac{\partial J}{\partial a a_{1}}+\sum_{k=1}^{i(2)-1} w_{k, 1} \sum_{n \in \Omega_{T}} T_{k}^{n} \cos (\omega n \Delta t)=0, \\
j=1, \\
\frac{\partial J}{\partial a a_{j}}+\sum_{k=i(j-1)+1}^{i(j+1)} w_{k, j} \sum_{n \in \Omega_{T}} T_{k}^{n} \cos (\omega n \Delta t)=0, \\
2 \leq j \leq M-1, \\
\frac{\partial J}{\partial a a_{M}}+\sum_{k=i(j-1)+1}^{N} w_{k, j} \sum_{n \in \Omega_{T}} T_{k}^{n} \cos (\omega n \Delta t)=0, \\
\frac{\partial J}{\partial b b_{1}}+\sum_{k=1}^{i(2)-1} w_{k, 1} \sum_{n \in \Omega_{T}} T_{k}^{n} \cos (\omega n \Delta t)=0, \\
j=1,  \tag{18}\\
\frac{\partial J}{\partial b b_{j}}+\sum_{k=i(j-1)+1}^{i(j+1)} w_{k, j} \sum_{n \in \Omega_{T}} T_{k}^{n} \cos (\omega n \Delta t)=0, \\
\frac{\partial J}{\partial b b_{M}}+\sum_{k=i(M-1)+1}^{N} w_{k, j} \sum_{n \in \Omega_{T}} T_{k}^{n} \cos (\omega n \Delta t)=0, \\
j=M \leq M-1, \\
j
\end{array}
$$



Figure 2: The bathymetric map of the Bohai, Yellow, and East China Seas (contour) and the position of $T / P$ satellite tracks (dot), tidal gauge stations (plus), and open boundaries (open circle). The numbers are the water depth in meter.
where

$$
T_{k}^{n}=-\frac{g \mu_{k}^{n}}{\Delta x}
$$

(for GP $(k)$ on the right of the area calculated),

$$
T_{k}^{n}=\frac{g \mu_{k_{l}}^{j}}{\Delta x}
$$

(for GP $(k)$ on the left of the area calculated),

$$
T_{k}^{n}=-\frac{g v_{k}^{n}}{\Delta y}
$$

(for GP $(k)$ under the area calculated),

$$
T_{k}^{n}=\frac{g v_{k}^{n}}{\Delta y}
$$

(for GP $(k)$ above the area calculated),
where $\mu$ and $\nu$ are the adjoint variables of west-east velocity component $u$ and north-south velocity component $v$, respectively. The values of $\mu$ and $\nu$ are computed by running the adjoint model.
3.3.2. Gradients for Fully Space Varying Inversion Strategy. In the second strategy, the OBC is assumed to be fully space varying and the values at every open boundary points (i.e., general points) are taken as control variables. Consequently, the gradients of cost function with respect to the Fourier coefficients at general points $a a_{\mathrm{GP}(k)}$ and $b b_{\mathrm{GP}(k)}\left(a a_{k}\right.$ and $b b_{k}$
for simplicity, $1 \leqslant k \leqslant N$ ) have to be computed. The gradients are deduced from

$$
\begin{equation*}
\frac{\partial L}{\partial a a_{k}}=0, \quad \frac{\partial L}{\partial b b_{k}}=0, \quad 1 \leq k \leq N \tag{20}
\end{equation*}
$$

which yields

$$
\begin{align*}
& \frac{\partial J}{\partial a a_{k}}+\sum_{n \in \Omega_{T}} T_{k}^{n} \cos (\omega n \Delta t)=0, \quad 1 \leq k \leq N \\
& \frac{\partial J}{\partial b b_{k}}+\sum_{n \in \Omega_{T}} T_{k}^{n} \cos (\omega n \Delta t)=0, \quad 1 \leq k \leq N \tag{21}
\end{align*}
$$

where $T_{k}^{n}$ can also be computed by using (19).

## 4. Numerical Experiments and Results Analysis

4.1. Model Settings. The computing area in the present study is the Bohai Sea, the Yellow Sea, and the East China Sea (BYECS), typical marginal shelf seas. The spatial resolution for the model is $1 / 12^{\circ} \times 1 / 12^{\circ} . T / P$ altimeter data and tidal gauge data are assimilated into the tidal model. The bathymetry map of the BYECS, the position of $T / P$ satellite tracks, tidal gauge stations, and the open boundaries are shown in Figure 2. Since the purpose of this paper is to discuss the inversion of OBCs, the bottom friction coefficients are fixed in all the experiments.

The numerical experiments in this work are semiidealized. Specifically, the coastline, the number, and location of the observations are real. On the contrary, the values of open boundary conditions and observations are artificial. The prescribed open boundary curves are generated by different number of feature points. Apparently, the complexity of open boundary curves is in direct proportion to the number of
feature points. For the semi-idealized experiments, only the location of real observations (satellite altimetry and tidal gauge stations) is used and the values of "observations" are obtained by running the dynamic forward model with prescribed open boundary conditions. The advantage of this kind of experiments is that we can obtain a thorough understanding of the "observations." The "observations" generated by the model can be accurate and we can control the quality of the "observations" by adding artificial error. In addition, because the other factors are real, the conclusions based on these semi-idealized experiments can be more useful for referring.

The semi-idealized numerical experiments are run as follows. First a distribution of artificial Fourier coefficients is prescribed and taken as "true values" of open boundary conditions. Then the forward tidal model is run using the "true values" and the simulation results recorded at grid points of $T / P$ satellite tracks and tidal gauge stations are taken as the "observations." Having obtained the "observations", an initial value (taken as zero in this work) of Fourier coefficients is assigned to run the forward model. The differences between simulated values and "observations" will function as the external force to drive the adjoint model. The optimized Fourier coefficients can be obtained through the backward integration of the adjoint equations. The inverse integral time of the adjoint equations is equal to a period of $M_{2}$ tide. With the procedures repeated above, the parameters will be optimized continuously and the difference between simulated values and "observations" will be diminished. Meanwhile, the difference between the prescribed and the inverted parameters will also be decreased.

The iteration of optimization will terminate once the following criterion is achieved [21]:

$$
\begin{equation*}
\|G\|<\mathrm{eps} \times \max (1,\|X\|), \tag{22}
\end{equation*}
$$

where $\|G\|$ is the $L_{2}$ norm of the gradients of cost function with respect to the control variables (i.e., the Fourier coefficients at feature points), eps is a positive variable that determines the accuracy with which the solution is to be found, and $\|X\|$ is the $L_{2}$ norm of control variables. Both the values of $\|G\|$ and $\|X\|$ vary along the iterations. For a correct adjoint model and a reasonable method, $\|G\|$ will gradually decrease versus the iteration steps and the inverted values of control variables must gradually approach the prescribed "true values". When using L-BFGS, the number of corrections used in the BFGS update is taken as 5 (usually between 3 and 7 , see Alekseev et al. [19]). In the minimization algorithm, the control variables should be scaled to similar magnitudes on the order of unity because within the optimization algorithm convergence, tolerances, and other criteria are based on an implicit definition of small and large [22]. Zou et al. [20] also proved that the efficiency could be greatly improved by a simple scaling. In twin experiments we use 10 to scale the Fourier coefficients [4].

### 4.2. Modeling Results

4.2.1. Effects of Complexity of Open Boundary Curves. In this section, the semi-idealized experiments (SE) are carried
out to calibrate the inversion ability of adjoint model and compare the effectiveness of two strategies developed in Section 3. The prescribed distributions of artificial Fourier coefficients at 173 grid points along the eastern open boundary are inverted. The prescribed distributions (PDs) are designed to be characterized by different numbers of feature points. PDs 1-7 are characterized by $2,6,10,14,18,22$, and 26 feature points, respectively. The twin experiments are correspondingly indexed with SEa 1-7 for inversion strategy 1 and SEb 1-7 for inversion strategy 2.

The prescribed and inverted distributions of open boundary curves in SEa 1-4 and SEb 1-4 are shown in Figure 3. The prescribed and inverted distributions of open boundary curves in SEa 5-6 and SEb 5-6 are shown in Figure 4. The feature points for prescribed distributions have also been indicated in Figures 3 and 4. Table 1 gives the error statistics for the experiments in this section. The $L_{2}$ norm of the gradients of cost function with respect to the control variables versus the iteration steps for the experiments using inversion strategies 1 and 2 are presented in Figures 4(c) and 4(d), respectively. The decrease in data misfit (i.e., cost function) calculated from (3) versus the iteration steps is shown in Figure 5. Note that the values of data misfit and $L_{2}$ norm of gradients have been normalized by their values at the first iteration step.

For strategy 1, the values of data misfit can sharply decrease by about 4 orders for all the experiments in about 30 iteration steps. For strategy 2, the values of data misfit can sharply decrease by about 5 orders for SEb 1-5 and by 4 orders for SEb 6-7 in about 60 iteration steps. The decrease in data misfit provides another proof for the inversion ability of the adjoint model and strategies in this work. Correspondingly, the $L_{2}$ norms of gradients also decrease by at least 2 orders for inversion strategy 1 and by 3 orders for inversion strategy 2, which demonstrates that the gradients calculated in Section 3.3 can work well with L-BFGS method.

From the decrease in data misfit and gradient it seems as if the effect of inversion strategy 2 is better than that of strategy 1 . However, the differences between prescribed and inverted distributions shown in Table 1 indicate that the inversion results of strategy 1 are much better than those of strategy 2. This inconsistency will be explained in Section 4.3. One can find that the adjoint model combined with inversion strategy 1 can reproduce the prescribed distributions of Fourier coefficients perfectly for SEa 1-2 or almost perfectly for SEa 3-4. For SEa 5-6 the inversion is acceptable but largely deviates from perfection. The major trend of the inversion is quite obvious that the effect of inversion is in inverse proportion to the number of feature points which characterizes the complexity of open boundary curves. The inverted open boundary curves shown in Figures 3 and 4 also prove that the inversion using strategy 1 is better than that using strategy 2.
4.2.2. Effects of Data Noises. As we know, the real observations either from satellite altimetry or from tidal gauge stations contain errors (or noises). In this section the effects


FIgure 3: The prescribed and inverted distributions of open boundary curves in SEa 1-4 and SEb 1-4. The feature points are indicated by open circles.
of the noises are studied. To do this, we replace each "observation" $\widehat{\zeta}_{i, j}^{n}$ by $\left(1+p r_{i, j}^{n}\right) \widehat{\zeta}_{i, j}^{n}$, where $r_{i, j}^{n}$ are uniform random numbers lying in $[-1,1]$ and $p$ is a factor determining the maximum percentage error. The maximum percentage errors for each prescribed distribution (PDs 1-7) are assigned to 5\%, $10 \%, 15 \%$, and $20 \%$. The corresponding inversion experiments are then indexed with $\mathrm{SE}_{x}$ i.1, $\mathrm{SE}_{x}$ i.2, $\mathrm{SE}_{x}$ i.3, and $\mathrm{SE}_{x}$ i.4, respectively, where $1 \leqslant i \leqslant 7$ and $x=a$ or $b$. The error statistics for the experiments with $P$ values of $5 \%, 10 \%, 15 \%$, and $20 \%$ are exhibited in Tables 2, 3, 4, and 5, respectively. The figures are omitted because they are similar to those in Section 4.2.1.

One can find the noises in artificial observations will significantly and negatively influence the inversion of open boundary conditions. It is clear that the inversion using strategy 2 is much more sensitive to the noise than that using strategy 1 . For example, when the simplest distribution PD 1 is inverted, the difference between prescribed and inverted values will sharply increase from 0.0101 (Table 1) to 0.0238 (Table 2) for strategy 2 even with a small value of error $5 \%$.

When $P$ was increased to $20 \%$, the value of this difference is also increased to 0.0562 (Table 5). However, for strategy 1 the values of this difference are just $0.0011,0.0011,0.0032$ and 0.0043 under $P$ value of $5 \%, 10 \%, 15 \%$, and $20 \%$. Similar results can be found from the inversion results of other distributions. This phenomenon indicates that the effect of ill-posedness of inverse problem will be amplified in the conditions that observations contain noises. In addition, the parameter estimation problems with more control variables will be much more sensitive to data noise and the negative effect of noises can be restricted by reducing the number of control variables.

### 4.3. Discussions

4.3.1. Rationality of the Adjoint Method (Suggested by an Anonymous Reviewer). The motivation of the present work is to take the open boundary condition as an example to investigate the performance of the adjoint method when applied to ocean modeling and the ill-posedness of relevant

(a)


| Gradient in SEa 1 | - Gradient in SEa 5 |
| :---: | :---: |
| Gradient in SEa 2 | -- Gradient in SEa 6 |
| Gradient in SEa 3 | Gradient in SEa 7 |
| Gradient in SEa 4 |  |

(c)

(b)

(d)

Figure 4: (a), (b) The prescribed and inverted distributions of open boundary curves in SEa 5-6 and SEb 5-6. The feature points are indicated by open circles. (c), (d) The $L_{2}$ norm of the gradients of cost function with respect to the control variables versus the iteration steps for strategies 1 and 2.


(a)


| Data misfit in SEb 1 | - Data misfit in SEb 5 |
| :---: | :---: |
| Data misfit in SEb 2 | - Data misfit in SEb 6 |
| Data misfit in SEb 3 | Data misfit in SEb 7 |

(b)

Figure 5: Data misfit versus the iteration steps for strategy 1 (a) and strategy 2 (b).

Table 1: Error statistics for SEa 1-7 and SEb 1-7.

| Exp. | $K_{1}^{\mathrm{a}}$ | $K_{2}^{\mathrm{a}}$ | Before | $K_{3}^{\mathrm{a}}$ | After | Before |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: |

${ }^{\text {a }} K_{1}$ is the number of feature points for PDs 1-7 prescribed in semi-idealized experiments. $K_{2}$ is the value of maximum percentage error. $K_{3}$ is the data misfit before and after assimilation. $K_{4}$ is the mean absolute difference between prescribed and inverted Fourier coefficients.

Table 2: Error statistics for SEa 1.1-7.1 and SEb 1.1-7.1.

| Exp. | $K_{1}^{\text {a }}$ | $K_{2}^{\text {a }}$ | $K_{3}^{\text {a }}$ |  | $K_{4}^{\mathrm{a}}$ (m) |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  |  |  | Before | After | Before | After |
| Inversion strategy 1 |  |  |  |  |  |  |
| SEa 1.1 | 2 | 0.05 | 5060.1284 | 4.3569 | 0.3500 | 0.0011 |
| SEa 2.1 | 6 | 0.05 | 4306.6660 | 3.5968 | 0.3332 | 0.0007 |
| SEa 3.1 | 10 | 0.05 | 4600.6445 | 3.9834 | 0.3055 | 0.0082 |
| SEa 4.1 | 14 | 0.05 | 4019.1911 | 3.2996 | 0.3121 | 0.0093 |
| SEa 5.1 | 18 | 0.05 | 3614.2876 | 4.0757 | 0.3014 | 0.0443 |
| SEa 6.1 | 22 | 0.05 | 3370.5825 | 3.4881 | 0.3066 | 0.0491 |
| SEa 7.1 | 26 | 0.05 | 3838.0024 | 4.3227 | 0.3124 | 0.0740 |
| Inversion strategy 2 |  |  |  |  |  |  |
| SEb 1.1 | 2 | 0.05 | 5060.1284 | 4.2224 | 0.3500 | 0.0238 |
| SEb 2.1 | 6 | 0.05 | 4306.6660 | 3.4525 | 0.3332 | 0.0250 |
| SEb 3.1 | 10 | 0.05 | 4600.6445 | 3.6353 | 0.3055 | 0.0332 |
| SEb 4.1 | 14 | 0.05 | 4019.1911 | 3.0429 | 0.3121 | 0.0337 |
| SEb 5.1 | 18 | 0.05 | 3614.2876 | 3.0501 | 0.3014 | 0.0482 |
| SEb 6.1 | 22 | 0.05 | 3370.5825 | 2.7539 | 0.3066 | 0.0736 |
| SEb 7.1 | 26 | 0.05 | 3838.0024 | 3.2047 | 0.3124 | 0.0833 |

${ }^{\mathrm{a}} K_{1}$ is the number of feature points for PDs 1-7 prescribed in semi-idealized experiments. $K_{2}$ is the value of maximum percentage error. $K_{3}$ is the data misfit before and after assimilation. $K_{4}$ is the mean absolute difference between prescribed and inverted Fourier coefficients.
inverse problem. The inverse problems in ocean models are often quite complex. The ocean modeling is not just to solve the partial differential equations which might also be solved by some simple methods like the method of characteristics. A reasonable ocean model should also be related to the field observations (satellite altimetry and tidal gauges in this work). In order to realize a more accurate simulation of ocean dynamics, how to organically combine the numerical ocean model with available observations has already become
a problem urgent to be solved. Data assimilation methods have been used widely to solve this problem. Among all data assimilation methods, the adjoint data assimilation method is one of the most effective and powerful approaches developed over the past three decades. It is an advanced data assimilation method and has the advantage of directly assimilating various observations distributed in time and space into the numerical model while maintaining dynamical and physical consistency with the model. The adjoint method

Table 3: Error statistics for SEa 1.2-7.2 and SEb 1.2-7.2.

| Exp. | $K_{1}^{\mathrm{a}}$ | $K_{2}^{\mathrm{a}}$ | Before | $K_{3}^{\mathrm{a}}$ | After | Before |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: |

${ }^{\text {a }} K_{1}$ is the number of feature points for PDs 1-7 prescribed in semi-idealized experiments. $K_{2}$ is the value of maximum percentage error. $K_{3}$ is the data misfit before and after assimilation. $K_{4}$ is the mean absolute difference between prescribed and inverted Fourier coefficients.

Table 4: Error statistics for SEa 1.3-7.3 and SEb 1.3-7.3.

| Exp. | $K_{1}^{\mathrm{a}}$ | $K_{2}^{\mathrm{a}}$ | Before | $K_{3}^{\mathrm{a}}$ | After | Before |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: |

${ }^{\mathrm{a}} K_{1}$ is the number of feature points for PDs 1-7 prescribed in semi-idealized experiments. $K_{2}$ is the value of maximum percentage error. $K_{3}$ is the data misfit before and after assimilation. $K_{4}$ is the mean absolute difference between prescribed and inverted Fourier coefficients.
might be complicated and expensive for some simple problems. However, the inverse problems in ocean modeling are often quite complex in contrast with those simple problems. As is known, one advantage of the numerical method over theoretical analysis lies in the disposal of nonlinear terms. The ocean numerical models are usually strongly nonlinear, increasing the complexity of the relevant inverse problem. Therefore, the increased complexity of the inverse problem makes the adjoint method effective. The adjoint method has been proved to be effective and powerful in ocean and atmosphere problems by many works (see the references listed
in Section 1). It has been widely applied to meteorological and oceanographic data assimilation, sensitivity studies, and parameter estimation.
4.3.2. Analysis on Ill-Posedness. From the statistics shown in Tables $1-5$, we can find an interesting phenomenon. Define the data misfits after assimilation to be $V 1_{\mathrm{dm}}$ for inversion strategy 1 and $V 2_{\mathrm{dm}}$ for inversion strategy 2 . Further define the differences between prescribed and inverted control variables to be $V 1_{\mathrm{cv}}$ for inversion strategy 1 and $V 2_{\mathrm{cv}}$ for

Table 5: Error statistics for SEa 1.4-7.4 and SEb 1.4-7.4.

| Exp. | $K_{1}^{\mathrm{a}}$ | $K_{2}^{\mathrm{a}}$ | Before | $K_{3}^{\mathrm{a}}$ | After | Before |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: |

${ }^{\text {a }} K_{1}$ is the number of feature points for PDs 1-7 prescribed in semi-idealized experiments. $K_{2}$ is the value of maximum percentage error. $K_{3}$ is the data misfit before and after assimilation. $K_{4}$ is the mean absolute difference between prescribed and inverted Fourier coefficients.
inversion strategy 2 . The values of $V i_{\mathrm{cv}}(i=1,2)$ and $V i_{\mathrm{dm}}(i=$ $1,2)$ for all the experiments are plotted in Figure 6. We can find $V 1_{\mathrm{dm}}$ are larger than or comparable with $V 2_{\mathrm{dm}}$ while $V 1_{\mathrm{cv}}$ are greatly smaller than $V 2_{\mathrm{cv}}$. Consequently, for all the experiments except SEa 1 and SEa 2, without loss of generality, we can obtain

$$
\begin{equation*}
V 1_{\mathrm{cv}}<V 2_{\mathrm{cv}}, \quad V 1_{\mathrm{dm}}>V 2_{\mathrm{dm}} . \tag{23}
\end{equation*}
$$

It is easy to understand that small values of $V i_{\mathrm{cv}}(i=1,2)$ indicate more accurate control variables, and small values of $V i_{\mathrm{dm}}(i=1,2)$ mean small differences between simulated and observed results. In this work, the open boundary conditions are the only parameters for estimation and other parameters are fixed all the time. Instead of formula (23), we should have expected

$$
\begin{equation*}
V 1_{\mathrm{cv}}<V 2_{\mathrm{cv}}, \quad \text { so } \quad V 1_{\mathrm{dm}}<V 2_{\mathrm{dm}} \tag{24}
\end{equation*}
$$

which means a better parameter estimation drives a more accurate simulation. In other words, what we want are small values of $V_{\mathrm{dm}}$ and what we need are small values of $V_{\mathrm{cv}}$. Formulas (23) and (24) exactly indicate an inconsistency between the effects of parameter estimation and observation restricted data reproduction.

For PDs 1-7 the numbers of feature points are $2,6,10$, $14,18,22$, and 26 , respectively. It should be noted that at each feature point the Fourier coefficients include $a$ and $b$. Therefore the numbers of control variables for inversion are doubled, that is, $4,12,20,28,36,44$, and 52 , respectively. There are a total of 35 semi-idealized experiments in this work. Among these experiments, only SEa 1 and SEa 2 can realize a perfect inversion of control variables. Here we define perfect inversion as follows: the data misfit between observed and simulated values can decrease to zero and the difference
between prescribed and inverted control variables can also reach a value of zero. With more control variables and larger data noises, the inversion results will not be exactly equal to the prescribed distributions. In the work of Smedstad and O'Brien [12] where the spatially distributed phase speed in an equatorial Pacific Ocean model was estimated, they could not produce the exact values either, even in the condition that perfect observations were available at every grid of the model. Zhang and Lu [4] put forward the similar viewpoint and it also occurs in the parameter estimation of internal tidal model [23-25]. With identical twin experiments, the "observations" are perfect in the sense that they are produced by the model and thus are consistent with the model physics. From the results of this paper and previous works, we can conclude that ill-posedness has happened in other 33 experiments and the effects of ill-posedness will be amplified by increasing the number of control variables and data noises. Formula (23) obtained in this work provides a concrete evidence that ill-posedness of inverse problem can generate poor parameter inversion results while producing an unreal "good data fitting". For a specific problem, it is necessary and helpful to perform identical semi-idealized experiments in order to find the optimal choices for the number of control variables and inversion strategy.

## 5. Conclusions

In this work, two strategies for inverting the open boundary conditions with adjoint method are compared by carrying out semi-idealized numerical experiments. In the first strategy, the open boundary curves are assumed to be partly space varying and are generated by linearly interpolating the values at feature points. The feature points are selected by calculating the second-order derivatives of discrete curves and the values

(b)

Figure 6: (a) The values of $V i_{\mathrm{dm}}(i=1,2)$ versus the index of experiments. (b) The values of $V i_{\mathrm{cv}}(i=1,2)$ versus the index of experiments.
at selected feature points are taken as control variables to be estimated. The advantage is that most of the variations of the curves can be reproduced by the minimum number of points. In the second strategy, the OBC is assumed to be fully space varying and the values at every open boundary points are taken as control variables.

A series of semi-idealized experiments are carried out to calibrate the inversion ability of adjoint model and compare the effectiveness of two inversion strategies. The results demonstrate that the effect of inversion is in inverse proportion to the number of feature points which characterize the complexity of open boundary curves. The effect of ill-posedness of inverse problem will be amplified in the conditions that observations contain noises. The parameter estimation problems with more control variables will be much more sensitive to data noises and the negative effects of noises can be restricted by reducing the number of control variables. This work provides a concrete evidence that illposedness of inverse problem can generate wrong parameter inversion results while producing an unreal "good data fitting". For a specific problem, it is necessary and helpful to perform identical semi-idealized experiments in order to find the optimal choices for the number of control variables and inversion strategy.

## Acknowledgments

The authors thank Professor Jorge Nocedal at Northwestern University for sharing the source codes of L-BFGS. Partial
support for this research was provided by the National Natural Science Foundation of China through Grants 41206001 and 41076006, the Major State Basic Research Development Program of China through Grant 2013CB956500, the Natural Science Foundation of Jiangsu Province through Grant BK2012315, the Priority Academic Program Development of Jiangsu Higher Education Institutions, and the Fundamental Research Funds for the Central Universities 201261006.

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## Research Article

# Decoupling the Stationary Navier-Stokes-Darcy System with the Beavers-Joseph-Saffman Interface Condition 

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Received 5 April 2013; Accepted 31 July 2013
Academic Editor: R. K. Bera
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#### Abstract

This paper proposes a domain decomposition method for the coupled stationary Navier-Stokes and Darcy equations with the Beavers-Joseph-Saffman interface condition in order to improve the efficiency of the finite element method. The physical interface conditions are directly utilized to construct the boundary conditions on the interface and then decouple the Navier-Stokes and Darcy equations. Newton iteration will be used to deal with the nonlinear systems. Numerical results are presented to illustrate the features of the proposed method.


## 1. Introduction

The Stokes-Darcy model has been extensively studied in the recent years due to its wide range of applications in many natural world problems and industrial settings, such as the subsurface flow in karst aquifers, oil flow in vuggy porous media, industrial filtrations, and the interaction between surface and subsurface flows [1-8]. Since the problem domain naturally consists of two different physical subdomains, several different numerical methods have been developed to decouple the Stokes and Darcy equations [6, 9-26]. For other works on the numerical methods and analysis of the StokesDarcy model, we refer the readers to [27-45].

Recently the more physically valid Navier-Stokes-Darcy model has attracted scientists' attention, and several coupled finite element methods have been studied for it [46-51]. On the other hand, the advantages of the domain decomposition methods (DDMs) in parallel computation and natural preconditioning have motivated the development of different DDMs for solving the Stoke-Darcy model [6, 10-18, 21, 22]. In this paper, we will develop a domain decomposition method for the Navier-Stokes-Darcy model based on Robin boundary conditions constructed from the interface conditions. This physics-based DDM is different from the traditional ones in the sense that they focus on decomposing different physical
domains by directly utilizing the given physical interface conditions.

The rest of paper is organized as follows. In Section 2, we introduce the Navier-Stokes-Darcy model with the Beavers-Joseph-Saffman interface condition. In Section 3, we recall the coupled weak formulation and the corresponding coupled finite element method for the Navier-Stokes-Darcy model. In Section 4, a parallel domain decomposition method and its finite element discretization are proposed to decouple the Navier-Stokes-Darcy system by using the Robin-type boundary conditions constructed from the physical interface conditions. Finally, in Section 5, we present a numerical example to illustrate the features of the proposed method.

## 2. Stationary Navier-Stokes-Darcy Model

In this section we introduce the following coupled Navier-Stokes-Darcy model on a bounded domain $\Omega=\Omega_{m} \cup \Omega_{c} \subset$ $\mathbb{R}^{d},(d=2,3)$; see Figure 1. In the porous media region $\Omega_{m}$, the flow is governed by the Darcy system

$$
\begin{gather*}
\vec{u}_{m}=-\mathbb{K} \nabla \phi_{m},  \tag{1}\\
\nabla \cdot \vec{u}_{m}=f_{m} .
\end{gather*}
$$



Figure 1: A sketch of the porous median domain $\Omega_{m}$, fluid domain $\Omega_{c}$, and the interface $\Gamma$.

Here, $\vec{u}_{m}$ is the fluid discharge rate in the porous media, $\mathbb{K}$ is the hydraulic conductivity tensor, $f_{m}$ is a sink/source term, and $\phi_{m}$ is the hydraulic head defined as $z+\left(p_{m} / \rho g\right)$, where $p_{m}$ denotes the dynamic pressure, $z$ the height, $\rho$ the density, and $g$ the gravitational acceleration. We will consider the following second-order formulation, which eliminates $\vec{u}_{m}$ in the Darcy system:

$$
\begin{equation*}
-\nabla \cdot\left(\mathbb{K} \nabla \phi_{m}\right)=f_{m} . \tag{2}
\end{equation*}
$$

In the fluid region $\Omega_{c}$, the fluid flow is assumed to be governed by the Navier-Stokes equations:

$$
\begin{gather*}
\vec{u}_{c} \cdot \nabla \vec{u}_{c}-\nabla \cdot \mathbb{T}\left(\vec{u}_{c}, p_{c}\right)=\vec{f}_{c},  \tag{3}\\
\nabla \cdot \vec{u}_{c}=0, \tag{4}
\end{gather*}
$$

where $\vec{u}_{c}$ is the fluid velocity, $p_{c}$ is the kinematic pressure, $\vec{f}_{c}$ is the external body force, $\nu$ is the kinematic viscosity of the fluid, $\mathbb{T}\left(\vec{u}_{c}, p_{c}\right)=2 \nu \mathbb{D}\left(\vec{u}_{c}\right)-p_{c} \mathbb{}$ is the stress tensor, and $\mathbb{D}\left(\vec{u}_{c}\right)=\left(\nabla \vec{u}_{c}+\nabla^{T} \vec{u}_{c}\right) / 2$ is the deformation tensor.

Let $\Gamma=\bar{\Omega}_{m} \cap \bar{\Omega}_{c}$ denote the interface between the fluid and porous media regions. On the interface $\Gamma$, we impose the following three interface conditions:

$$
\begin{gather*}
\vec{u}_{c} \cdot \vec{u}_{c}=-\vec{u}_{m} \cdot \vec{n}_{m},  \tag{5}\\
-\vec{u}_{c} \cdot\left(\mathbb{T}\left(\vec{u}_{c}, p_{c}\right) \cdot \vec{n}_{c}\right)=g\left(\phi_{m}-z\right),  \tag{6}\\
-\boldsymbol{\tau}_{j} \cdot\left(\mathbb{T}\left(\vec{u}_{c}, p_{c}\right) \cdot \vec{n}_{c}\right)=\frac{\alpha v \sqrt{\mathbf{d}}}{\sqrt{\operatorname{trace}\left(\prod\right)}} \boldsymbol{\tau}_{j} \cdot \vec{u}_{c}, \tag{7}
\end{gather*}
$$

where $\vec{n}_{c}$ and $\vec{n}_{m}$ denote the unit outer normal to the fluid and the porous media regions at the interface $\Gamma$, respectively, $\boldsymbol{\tau}_{j}(j=1, \ldots, d-1)$ denote mutually orthogonal unit tangential vectors to the interface $\Gamma$, and $\Pi=\mathbb{K} \nu / g$. The third condition (7) is referred to as the Beavers-Joseph-Saffman (BJS) interface condition [52-55].

In this paper, for simplification, we assume that the hydraulic head $\phi_{m}$ and the fluid velocity $\vec{u}_{c}$ satisfy the
homogeneous Dirichlet boundary condition except on $\Gamma$, that is, $\phi_{m}=0$ on the boundary $\partial \Omega_{m} / \Gamma$ and $\vec{u}_{c}=0$ on the boundary $\partial \Omega_{c} / \Gamma$.

## 3. Coupled Weak Formulation and Finite Element Method

In this section we will recall the coupled weak formulation and the corresponding coupled finite element method for the Navier-Stokes-Darcy model with Beavers-Joseph-Saffman condition. Let $(\cdot, \cdot)_{D}$ denote the $L^{2}$ inner product on the domain $D\left(D=\Omega_{c}\right.$ or $\left.\Omega_{m}\right)$ and let $\langle\cdot, \cdot\rangle$ denote the $L^{2}$ inner product on the interface $\Gamma$ or the duality pairing between $\left(H_{00}^{1 / 2}(\Gamma)\right)^{\prime}$ and $H_{00}^{1 / 2}(\Gamma)[5]$. Define the spaces

$$
\begin{gather*}
X_{c}=\left\{\vec{v} \in\left[H^{1}\left(\Omega_{c}\right)\right]^{d} \mid \vec{v}=0 \text { on } \frac{\partial \Omega_{c}}{\Gamma}\right\}, \\
Q_{c}=L^{2}\left(\Omega_{c}\right)  \tag{8}\\
X_{m}=\left\{\psi \in H^{1}\left(\Omega_{m}\right) \mid \psi=0 \text { on } \frac{\partial \Omega_{m}}{\Gamma}\right\},
\end{gather*}
$$

the bilinear forms

$$
\begin{gather*}
a_{m}\left(\phi_{m}, \psi\right)=\left(\mathbb{K} \nabla \phi_{m}, \nabla \psi\right)_{\Omega_{m}} \\
a_{c}\left(\vec{u}_{c}, \vec{v}\right)=2 v\left(\mathbb{D}\left(\vec{u}_{c}\right), \mathbb{D}(\vec{v})\right)_{\Omega_{c}}  \tag{9}\\
b_{c}(\vec{v}, q)=-(\nabla \cdot \vec{v}, q)_{\Omega_{c}}
\end{gather*}
$$

and the projection onto the tangent space on $\Gamma$ :

$$
\begin{equation*}
P_{\tau} \vec{u}=\sum_{j=1}^{d-1}\left(\vec{u} \cdot \boldsymbol{\tau}_{j}\right) \boldsymbol{\tau}_{j} \tag{10}
\end{equation*}
$$

With these notations, the weak formulation of the coupled Navier-Stokes-Darcy model with BJS interface condition is given as follows [46-51]: find $\left(\vec{u}_{c}, p_{c}, \phi_{m}\right) \in X_{c} \times Q_{c} \times X_{m}$ such that

$$
\begin{align*}
\left(\vec{u}_{c} \cdot\right. & \left.\nabla \vec{u}_{c}, \vec{v}\right)_{\Omega_{c}}+a_{c}\left(\vec{u}_{c}, \vec{v}\right)+b_{c}\left(\vec{v}, p_{c}\right) \\
& -b_{c}\left(\vec{u}_{c}, q\right)+a_{m}\left(\phi_{m}, \psi\right) \\
& +\left\langle g \phi_{m}, \vec{v} \cdot \vec{n}_{c}\right\rangle-\left\langle\vec{u}_{c} \cdot \vec{n}_{c}, \psi\right\rangle \\
& +\frac{\alpha v \sqrt{\mathbf{d}}}{\sqrt{\operatorname{trace}\left(\prod\right)}}\left\langle P_{\tau} \vec{u}_{c}, P_{\tau} \vec{v}\right\rangle  \tag{11}\\
= & \left(f_{m}, \psi\right)_{\Omega_{m}}+\left(\vec{f}_{c}, \vec{v}\right)_{\Omega_{c}} \\
& +\left\langle g z, \vec{v} \cdot \vec{n}_{c}\right\rangle, \quad \forall(\vec{v}, q, \psi) \in X_{c} \times Q_{c} \times X_{m}
\end{align*}
$$

Assume that we have in hand regular subdivisions of $\Omega_{m}$ and $\Omega_{c}$ into finite elements with mesh size $h$. Then one can define finite element spaces $X_{m h} \subset X_{m}, X_{c h} \subset X_{c}$ and
$Q_{c h} \subset Q_{c}$. We assume that $X_{c h}$ and $Q_{c h}$ satisfy the inf-sup condition $[56,57]$

$$
\begin{equation*}
\inf _{0 \neq q_{h} \in Q_{c h}} \sup _{0 \neq \vec{v}_{h} \in X_{c h}} \frac{b_{c}\left(\vec{v}_{h}, q_{h}\right)}{\left\|\vec{v}_{h}\right\|_{1}\left\|q_{h}\right\|_{0}}>\gamma \tag{12}
\end{equation*}
$$

where $\gamma>0$ is a constant independent of $h$. This condition is needed in order to ensure that the spatial discretizations of the Navier-Stokes equations used here are stable. See [56, 57] for more details of finite element spaces $X_{m h}, X_{c h}$, and $Q_{c h}$ that satisfy (12). One example is the Taylor-Hood element pair that we use in the numerical experiments; for that pair, $X_{c h}$ consists of continuous piecewise quadratic polynomials and $Q_{c h}$ consists of continuous piecewise linear polynomials.

Then a coupled finite element method with Newton iteration for the coupled Navier-Stokes-Darcy model is given as follows [46]: find $\left(\vec{u}_{c, h}, p_{c, h}, \phi_{m, h}\right) \in X_{c h} \times Q_{c h} \times X_{m h}$ in the following procedure.
(1) The initial value $\vec{u}_{c, h}^{0}$ is chosen.
(2) For $m=0,1,2, \ldots, M$, solve

$$
\begin{align*}
& \left(\vec{u}_{c, h}^{m+1} \cdot \nabla \vec{u}_{c, h}^{m}, \vec{v}_{h}\right)_{\Omega_{c}}+\left(\vec{u}_{c, h}^{m} \cdot \nabla \vec{u}_{c, h}^{m+1}, \vec{v}_{h}\right)_{\Omega_{c}} \\
& \quad+a_{c}\left(\vec{u}_{c, h}^{m+1}, \vec{v}_{h}\right)+b_{c}\left(\vec{v}_{h}, p_{c, h}^{m+1}\right) \\
& \quad-b_{c}\left(\vec{u}_{c, h}^{m+1}, q_{h}\right)+a_{m}\left(\phi_{m, h}^{m+1}, \psi_{h}\right) \\
& \quad+\left\langle g \phi_{m, h}^{m+1}, \vec{v}_{h} \cdot \vec{n}_{c}\right\rangle-\left\langle\vec{u}_{c, h}^{m+1} \cdot \vec{n}_{c}, \psi_{h}\right\rangle \\
& \quad+\frac{\alpha v \sqrt{\mathbf{d}}}{\sqrt{\operatorname{trace}\left(\prod\right)}}\left\langle P_{\tau} \vec{u}_{c, h}^{m+1}, P_{\tau} \vec{v}_{h}\right\rangle  \tag{13}\\
& =\left(\vec{u}_{c, h}^{m} \cdot \nabla \vec{u}_{c, h}^{m}, \vec{v}_{h}\right)_{\Omega_{c}}+\left(f_{m}, \psi_{h}\right)_{\Omega_{m}} \\
& \quad+\left(\vec{f}_{c}, \vec{v}_{h}\right)_{\Omega_{c}}+\left\langle g z, \vec{v}_{h} \cdot \vec{n}_{c}\right\rangle, \\
& \forall\left(\vec{v}_{h}, q_{h}, \psi_{h}\right) \in X_{c h} \times Q_{c h} \times X_{m h} .
\end{align*}
$$

(3) Set $\vec{u}_{c, h}=\vec{u}_{c, h}^{m+1}, p_{c, h}=\vec{p}_{c, h}^{m+1}$, and $\phi_{m, h}=\phi_{m, h}^{M+1}$.

## 4. Physics-Based Domain Decomposition Method

The coupled finite element method may end up with a huge algebraic system, which combines all parts from the Navier-Stokes equations, Darcy equation, and interface conditions together into one sparse matrix. Hence it is often impractical to directly apply this method to large-scale real world applications. In order to develop a more efficient numerical method in this section, we will directly utilize the three physical interface conditions to construct a physicsbased parallel domain decomposition method to decouple the Navier-Stokes and Darcy equations.

Let us first consider the following Robin condition for the Darcy system: for a given constant $\gamma_{p}>0$ and a given function $\eta_{p}$ defined on $\Gamma$,

$$
\begin{equation*}
\gamma_{p} \mathbb{K} \nabla \widehat{\phi}_{m} \cdot \vec{n}_{m}+g \widehat{\phi}_{m}=\eta_{p}, \quad \text { on } \Gamma . \tag{14}
\end{equation*}
$$

Then, the corresponding weak formulation for the Darcy part is given by the following: for $\eta_{p} \in L^{2}(\Gamma)$, find $\widehat{\phi}_{m} \in X_{m}$ such that

$$
\begin{align*}
& a_{m}\left(\hat{\phi}_{m}, \psi\right)+\left\langle\frac{g \widehat{\phi}_{m}}{\gamma_{p}}, \psi\right\rangle \\
& \quad=\left(f_{m}, \psi\right)_{\Omega_{m}}+\left\langle\frac{\eta_{p}}{\gamma_{p}}, \psi\right\rangle, \quad \forall \psi \in X_{m} \tag{15}
\end{align*}
$$

Second, we can propose the following two Robin-type conditions for the Navier-Stokes equations: for a given constant $\gamma_{f}>0$ and given functions $\eta_{f}$ and $\vec{\eta}_{f \tau}$ defined on Г,

$$
\begin{align*}
& \vec{n}_{c} \cdot\left(\mathbb{T}\left(\hat{\vec{u}}_{c}, \hat{p}_{c}\right) \cdot \vec{n}_{c}\right)+\gamma_{f} \hat{\vec{u}}_{c} \cdot \vec{n}_{c}=\eta_{f}, \quad \text { on } \Gamma, \\
& -P_{\tau}\left(\mathbb{T}\left(\hat{\vec{u}}_{c}, p_{c}\right) \cdot \vec{n}_{c}\right)=\frac{\alpha v \sqrt{\mathbf{d}}}{\sqrt{\operatorname{trace}\left(\prod\right)}} P_{\tau} \hat{\vec{u}}_{c}, \quad \text { on } \Gamma . \tag{16}
\end{align*}
$$

Then, the corresponding weak formulation for the Navier-Stokes equation is given by the following: for $\eta_{f} \in$ $L^{2}(\Gamma)$, find $\left(\widehat{\vec{u}}_{c}, \widehat{p}_{c}\right) \in X_{c} \times Q_{c}$ such that

$$
\begin{align*}
\left(\widehat{\vec{u}}_{c} \cdot\right. & \left.\nabla \widehat{\vec{u}}_{c}, \vec{v}\right)_{\Omega_{c}}+a_{c}\left(\hat{\vec{u}}_{c}, \vec{v}\right)+b_{c}\left(\vec{v}, \widehat{p}_{c}\right) \\
& -b_{c}\left(\hat{\vec{u}}_{c}, q\right)+\gamma_{f}\left\langle\hat{\vec{u}}_{c} \cdot \vec{n}_{c}, \vec{v} \cdot \vec{u}_{c}\right\rangle \\
& +\frac{\alpha v \sqrt{\mathbf{d}}}{\sqrt{\operatorname{trace}\left(\prod\right)}}\left\langle P_{\tau} \hat{\vec{u}}_{c}, P_{\tau} \vec{v}\right\rangle  \tag{17}\\
= & \left(\vec{f}_{c}, \vec{v}\right)_{\Omega_{c}}+\left\langle\eta_{f}, \vec{v} \cdot \vec{n}_{c}\right\rangle, \quad \forall(\vec{v}, q) \in X_{c} \times Q_{c} .
\end{align*}
$$

Our next step is to show that, for appropriate choices of $\gamma_{f}, \gamma_{\mathrm{p}}, \eta_{f}$, and $\eta_{p}$, the solutions of the coupled system (11) are equivalent to the solutions of the decoupled equations (15) and (17), and hence we may solve the latter system instead of the former.

Lemma 1. Let $\left(\phi_{m}, \vec{u}_{c}, p_{c}\right)$ be the solution of the coupled Navier-Stokes-Darcy system (11) and let ( $\widehat{\phi}_{m}, \widehat{\vec{u}}_{c}, \widehat{p}_{c}$ ) be the solution of the decoupled Navier-Stokes and Darcy equations (15) and (17) with Robin boundary conditions at the interface. Then, $\left(\hat{\phi}_{m}, \widehat{\vec{u}}_{c}, \widehat{p}_{c}\right)=\left(\phi_{m}, \vec{u}_{c}, p_{c}\right)$ if and only if $\gamma_{f}, \gamma_{p}, \eta_{f}, \vec{\eta}_{f \tau}$, and $\eta_{p}$ satisfy the following compatibility conditions:

$$
\begin{gather*}
\eta_{p}=\gamma_{p} \widehat{\vec{u}}_{c} \cdot \vec{n}_{c}+g \widehat{\phi}_{m},  \tag{18}\\
\eta_{f}=\gamma_{f} \widehat{\vec{u}}_{c} \cdot \vec{n}_{c}-g \widehat{\phi}_{m}+g z . \tag{19}
\end{gather*}
$$



FIGURE 2: Convergence for the velocity of the free flow (a) and the hydraulic head of the porous medium flow (b) versus the iteration counter $m$ for the parallel DDM with BJS interface condition.

TABLE 1: Errors of the finite element method for the steady Navier-Stokes-Darcy model with BJS interface condition.

| $h$ | $\left\\|\vec{u}_{h}-\vec{u}\right\\|_{0}$ | $\left\\|\vec{u}_{h}-\vec{u}\right\\|_{1}$ | $\left\\|p_{h}-p\right\\|_{0}$ | $\left\\|\phi_{h}-\phi\right\\|_{0}$ | $\left\|\phi_{h}-\phi\right\|_{1}$ |
| :--- | :---: | :---: | :---: | :---: | :---: |
| $1 / 8$ | $1.367 \times 10^{-3}$ | $6.147 \times 10^{-2}$ | $8.002 \times 10^{-3}$ | $6.940 \times 10^{-4}$ | $2.452 \times 10^{-2}$ |
| $1 / 16$ | $1.687 \times 10^{-4}$ | $1.577 \times 10^{-2}$ | $8.559 \times 10^{-4}$ | $8.687 \times 10^{-5}$ | $6.187 \times 10^{-3}$ |
| $1 / 32$ | $2.086 \times 10^{-5}$ | $3.978 \times 10^{-3}$ | $9.506 \times 10^{-5}$ | $1.089 \times 10^{-5}$ | $1.553 \times 10^{-3}$ |
| $1 / 64$ | $2.594 \times 10^{-6}$ | $9.974 \times 10^{-4}$ | $1.121 \times 10^{-5}$ | $1.363 \times 10^{-6}$ | $3.890 \times 10^{-4}$ |
| $1 / 128$ | $3.235 \times 10^{-7}$ | $2.496 \times 10^{-4}$ | $1.363 \times 10^{-6}$ | $1.705 \times 10^{-7}$ | $9.733 \times 10^{-5}$ |

Proof. Adding (15) and (17) together, we obtain the following: given $\eta_{p}, \eta_{f} \in L^{2}(\Gamma)$, find $\left(\widehat{\phi}_{m}, \widehat{u}_{f}, \widehat{p}_{c}\right) \in X_{\mathrm{m}} \times X_{c} \times Q_{c}$ such that

$$
\begin{aligned}
\left(\widehat{\vec{u}}_{c} \cdot\right. & \left.\nabla \widehat{\vec{u}}_{c}, \vec{v}\right)_{\Omega_{c}}+a_{c}\left(\widehat{\vec{u}}_{c}, \vec{v}\right)+b_{c}\left(\vec{v}, \hat{p}_{c}\right) \\
& -b_{c}\left(\widehat{\vec{u}}_{c}, q\right)+a_{m}\left(\widehat{\phi}_{m}, \psi\right)+\gamma_{f}\left\langle\hat{\vec{u}}_{c} \cdot \vec{n}_{c}, \vec{v} \cdot \vec{n}_{c}\right\rangle \\
& +\left\langle\frac{g \hat{\phi}_{m}}{\gamma_{p}}, \psi\right\rangle+\frac{\alpha v \sqrt{\mathbf{d}}}{\sqrt{\operatorname{trace}\left(\prod\right)}}\left\langle P_{\tau} \widehat{\vec{u}}_{c}, P_{\tau} \vec{v}\right\rangle \\
= & \left(f_{m}, \psi\right)_{\Omega_{m}}+\left(\vec{f}_{c}, \vec{v}\right)_{\Omega_{c}}+\left\langle\eta_{f}, \vec{v} \cdot \vec{n}_{c}\right\rangle \\
& +\left\langle\frac{\eta_{p}}{\gamma_{p}}, \psi\right\rangle, \quad \forall(\vec{v}, q, \psi) \in X_{m} \times X_{c} \times Q_{c} .
\end{aligned}
$$

For the necessity of the lemma, we pick $\psi=0$ and $\vec{v}$ such that $P_{\tau} \vec{v}=0$ in (11) and (20); then by subtracting (20) from (11), we get

$$
\begin{array}{r}
\left\langle\eta_{f}-\gamma_{f} \vec{v}_{f} \cdot \vec{n}_{c}+g \phi_{m}-g z, \vec{v} \cdot \vec{n}_{c}\right\rangle=0, \\
\forall \vec{v} \in X_{c} \quad \text { with } P_{\tau} \vec{v}=0, \tag{21}
\end{array}
$$

which implies (19). The necessity of (18) can be derived in a similar fashion.

As for the sufficiency of the lemma, by substituting the compatibility conditions (18)-(19), we easily see that $\left(\widehat{\phi}_{m}, \widehat{\vec{u}}_{c}, \widehat{p}_{c}\right)$ solves the coupled Navier-Stokes-Darcy system (11), which completes the proof.

Now we use the decoupled weak formulation constructed above to propose a physics-based parallel domain decomposition method with Newton iteration as follows.


$$
\begin{aligned}
& \cdots-\gamma_{p}=1 / 4 \gamma_{f} \\
& \cdots \circ \cdots \gamma_{p}=\gamma_{f} \\
& \rightarrow \gamma_{p}=4 \gamma_{f}
\end{aligned}
$$


$\cdots-\gamma_{p}=1 / 4 \gamma_{f}$
$\cdots \circ \cdot \gamma_{p}=\gamma_{f}$
$\leadsto \quad \gamma_{p}=4 \gamma_{f}$
(b)

Figure 3: Convergence for the pressure of the free flow (a) and $\eta_{f}(\mathrm{~b})$ versus the iteration counter $m$ for the parallel DDM with BJS interface condition.
(1) Initial values $\eta_{p}^{0}$ and $\eta_{f}^{0}$ are guessed. They may be taken to be zero.
(2) For $k=0,1,2, \ldots$, independently solve the Darcy and Navier-Stokes equations constructed above. More precisely, $\phi_{m}^{k} \in X_{m}$ is computed from

$$
\begin{equation*}
a_{m}\left(\phi_{m}^{k}, \psi\right)+\left\langle\frac{g \phi_{m}^{k}}{\gamma_{p}}, \psi\right\rangle=\left\langle\frac{\eta_{p}^{k}}{\gamma_{p}}, \psi\right\rangle+\left(f_{m}, \psi\right)_{\Omega_{m}} \tag{22}
\end{equation*}
$$

$\forall \psi \in X_{m}$,
and $\vec{u}_{c}^{k} \in X_{c}$ and $p_{c}^{k} \in Q_{c}$ are computed from the following Newton iteration.
(i) Initial value $\vec{u}_{c}^{k, 0}$ is chosen for the Newton iteration. For instance, it may be taken to be $\vec{u}_{c}^{0,0}=0$ and $\vec{u}_{c}^{k, 0}=\vec{u}_{c}^{k-1}$ for $k=1,2, \ldots$..
(ii) For $m=0,1,2, \ldots, M$, solve

$$
\begin{aligned}
& \left(\vec{u}_{c}^{k, m+1} \cdot \nabla \vec{u}_{c}^{k, m}, \vec{v}\right)_{\Omega_{c}}+\left(\vec{u}_{c}^{k, m} \cdot \nabla \vec{u}_{c}^{k, m+1}, \vec{v}\right)_{\Omega_{c}} \\
& \quad+a_{c}\left(\vec{u}_{c}^{k, m+1}, \vec{v}\right)+b_{c}\left(\vec{v}^{\prime}, p_{c}^{k, m+1}\right) \\
& \quad-b_{c}\left(\vec{u}_{c}^{k, m+1}, q\right)+\gamma_{f}\left\langle\vec{u}_{c}^{k, m+1} \cdot \vec{n}_{c}, \vec{v} \cdot \vec{n}_{c}\right\rangle \\
& \quad+\frac{\alpha v \sqrt{\mathbf{d}}}{\sqrt{\operatorname{trace}\left(\prod\right)}}\left\langle P_{\tau} \vec{u}_{c}^{k, m+1}, P_{\tau} \vec{v}\right\rangle
\end{aligned}
$$

$$
\begin{array}{r}
=\left(\vec{u}_{c}^{k, m} \cdot \nabla \vec{u}_{c}^{k, m}, \vec{v}\right)_{\Omega_{c}}+\left\langle\eta_{f}^{k}, \vec{v} \cdot \vec{n}_{c}\right\rangle+\left(\vec{f}_{c}, \vec{v}\right)_{\Omega_{c}}  \tag{23}\\
\forall(\vec{v}, q, \psi) \in X_{c} \times Q_{c} \times X_{m} .
\end{array}
$$

(iii) Set $\vec{u}_{c}^{k}=\vec{u}_{c}^{k, M+1}$ and $p_{c}^{k}=p_{c}^{k, M+1}$.
(3) $\eta_{p}^{k+1}$ and $\eta_{f}^{k+1}$ are updated in the following manner:

$$
\begin{align*}
& \eta_{f}^{k+1}=\frac{\gamma_{f}}{\gamma_{p}} \eta_{p}^{k}-\left(1+\frac{\gamma_{f}}{\gamma_{p}}\right) g \phi_{m}^{k}+g z,  \tag{24}\\
& \eta_{p}^{k+1}=-\eta_{f}^{k}+\left(\gamma_{f}+\gamma_{p}\right) \vec{u}_{c}^{k} \cdot \vec{n}_{c}+g z .
\end{align*}
$$

Then the corresponding domain decomposition finite element method is proposed as follows.
(1) Initial values $\eta_{p, h}^{0}$ and $\eta_{f, h}^{0}$ are guessed. They may be taken to be zero.
(2) For $k=0,1,2, \ldots$, independently solve the Darcy and Navier-Stokes equations with the Robin boundary conditions on the interface, which are constructed previously. More precisely, $\phi_{m, h}^{k} \in X_{m h}$ is computed from

$$
\begin{align*}
& a_{m}\left(\phi_{m, h}^{k}, \psi_{h}\right)+\left\langle\frac{g \phi_{m, h}^{k}}{\gamma_{p}}, \psi_{h}\right\rangle \\
& \quad=\left\langle\frac{\eta_{p, h}^{k}}{\gamma_{p}}, \psi_{h}\right\rangle+\left(f_{m}, \psi_{h}\right)_{\Omega_{m}}, \quad \forall \psi_{h} \in X_{m h} \tag{25}
\end{align*}
$$



FIgURe 4: Geometric convergence rate of the velocity of the free flow (a) and the hydraulic head of the porous medium flow (b) for the parallel DDM with BJS interface condition.
and $\vec{u}_{c, h}^{k} \in X_{c h}$ and $p_{c, h}^{k} \in Q_{c h}$ are computed from the following Newton iteration.
(i) Initial value $\vec{u}_{c, h}^{k, 0}$ is chosen for the Newton iteration. For instance, it may be taken to be $\vec{u}_{c, h}^{0,0}=0$ and $\vec{u}_{c, h}^{k, 0}=\vec{u}_{c, h}^{k-1}$ for $k=1,2, \ldots$..
(ii) For $m=0,1,2, \ldots, M$, solve

$$
\begin{align*}
& \left(\vec{u}_{c, h}^{k, m+1} \cdot \nabla \vec{u}_{c, h}^{k, m}, \vec{v}_{h}\right)_{\Omega_{c}}+\left(\vec{u}_{c, h}^{k, m} \cdot \nabla \vec{u}_{c, h}^{k, m+1}, \vec{v}_{h}\right)_{\Omega_{c}} \\
& +a_{c}\left(\vec{u}_{c, h}^{k, m+1}, \vec{v}_{h}\right)+b_{c}\left(\vec{v}_{h}, P_{c}^{k, m+1}\right)-b_{c}\left(\vec{u}_{c, h}^{k, m+1}, q_{h}\right) \\
& +\gamma_{f}\left\langle\vec{u}_{c, h}^{k, m+1} \cdot \vec{n}_{c}, \vec{v}_{h} \cdot \vec{n}_{c}\right\rangle \\
& +\frac{\alpha v \sqrt{\mathbf{d}}}{\sqrt{\operatorname{trace}\left(\prod\right)}}\left\langle P_{\tau} \vec{u}_{c, h}^{k, m+1}, P_{\tau} \vec{v}_{h}\right\rangle \\
& =\left(\vec{u}_{c, h}^{k, m} \cdot \nabla \vec{u}_{c, h}^{k, m}, \vec{v}_{h}\right)_{\Omega_{c}}+\left\langle\eta_{f, h}^{k} \vec{v}_{h} \cdot \vec{n}_{c}\right\rangle+\left(\vec{f}_{c}, \vec{v}_{h}\right)_{\Omega_{c}} \\
& \quad \forall\left(\vec{v}_{h}, q_{h}, \psi_{h}\right) \in X_{c h} \times Q_{c h} \times X_{m h} . \tag{26}
\end{align*}
$$

(iii) Set $\vec{u}_{c, h}^{k}=\vec{u}_{c, h}^{k, m+1}$ and $p_{c, h}^{k}=p_{c, h}^{k, M+1}$.
(3) $\eta_{p, h}^{k+1}$ and $\eta_{f, h}^{k+1}$ are updated in the following manner:

$$
\begin{align*}
& \eta_{f, h}^{k+1}=\frac{\gamma_{f}}{\gamma_{p}} \eta_{p, h}^{k}-\left(1+\frac{\gamma_{f}}{\gamma_{p}}\right) g \phi_{m, h}^{k}+g z,  \tag{27}\\
& \eta_{p, h}^{k+1}=-\eta_{f, h}^{k}+\left(\gamma_{f}+\gamma_{p}\right) \vec{u}_{c, h}^{k} \cdot \vec{n}_{c}+g z .
\end{align*}
$$

## 5. Numerical Example

Example 1. Consider the model problem (2)-(6) with the BJS interface condition (7) on $\Omega=[0, \pi] \times[-1,1]$ with $\Omega_{m}=[0, \pi] \times[0,1]$ and $\Omega_{c}=[0, \pi] \times[-1,0]$. Choose $\left(\alpha \nu \sqrt{\mathbf{d}} / \sqrt{\operatorname{trace}\left(\prod\right)}\right)=1, v=1, g=1, z=0$, and $\mathbb{K}=K \square$, where $\rrbracket$ is the identity matrix and $K=1$. The boundary condition data functions and the source terms are chosen such that the exact solution is given by

$$
\begin{align*}
\phi_{m}= & \left(e^{y}-e^{-y}\right) \sin (x) e^{t} \\
\vec{u}_{c}= & {\left[\frac{K}{\pi} \sin (2 \pi y) \cos (x) e^{t}\right.}  \tag{28}\\
& \left.\left(-2 K+\frac{K}{\pi^{2}} \sin ^{2}(\pi y)\right) \sin (x) e^{t}\right]^{T},
\end{align*}
$$

$$
p_{c}=0
$$

We divide $\Omega_{m}$ and $\Omega_{c}$ into rectangles of height $h=1 / N$ and width $\pi h$, where $N$ is a positive integer, and then subdivide each rectangle into two triangles by drawing a diagonal. The Taylor-Hood element pair is used for the Navier-Stokes equations, and the quadratic finite element is used for the second-order formulation of the Darcy equation.

For the coupled finite element method of the steady Navier-Stokes-Darcy model with BJS interface condition,


Figure 5: Geometric convergence rate of the pressure of the free flow (a) and $\eta_{f}$ (b) versus the iteration counter $m$ for the parallel DDM with BJS interface condition.

TABLE 2: $L^{2}$ errors in velocity and hydraulic head for the parallel DDM with BJS interface condition.

|  | $L^{2}$ velocity errors | $e(i) / e(i-4)$ | $L^{2}$ hydraulic head errors | $e(i) / e(i-4)$ |
| :--- | :---: | :---: | :---: | :---: |
| $e(0)$ | $2.342 \times 10^{-2}$ |  | $6.338 \times 10^{-1}$ |  |
| $e(4)(i=4)$ | $1.225 \times 10^{-3}$ | 0.0523 | $3.337 \times 10^{-2}$ | 0.0527 |
| $e(8)(i=8)$ | $6.450 \times 10^{-5}$ | 0.0527 | $1.756 \times 10^{-3}$ | 0.0526 |
| $e(12)(i=12)$ | $3.395 \times 10^{-6}$ | 0.0526 | $9.246 \times 10^{-5}$ | 0.0527 |
| $e(16)(i=16)$ | $1.787 \times 10^{-7}$ | 0.0526 | $4.868 \times 10^{-6}$ | 0.0527 |
| $e(20)(i=20)$ | $9.409 \times 10^{-9}$ | 0.0527 | $2.562 \times 10^{-7}$ | 0.0526 |

Table 1 provides errors for different choices of $h$. Using linear regression, the errors in Table 1 satisfy

$$
\begin{gather*}
\left\|\vec{u}_{c, h}-\vec{u}_{c}\right\|_{0} \approx 0.714 h^{3.011}, \quad\left|\vec{u}_{c, h}-\vec{u}_{c}\right|_{1} \approx 3.867 h^{1.987} \\
\left\|p_{c, h}-p_{c}\right\|_{0} \approx 5.123 h^{3.129} \\
\left\|\phi_{m, h}-\phi_{m}\right\|_{0} \approx 0.354 h^{2.998}, \quad\left|\phi_{m, h}-\phi_{m}\right|_{1} \approx 1.556 h^{1.995} \tag{29}
\end{gather*}
$$

These rates of convergence are consistent with the approximation capability of the Taylor-Hood element and quadratic element, which is third order with respect to $L^{2}$ norm of $\vec{u}_{c}$ and $\phi_{m}$, second order with respect to $H^{1}$ norm of $\vec{u}_{c}$ and $\phi_{m}$, and second order with respect to $L^{2}$ norms of $p_{c}$. In particular, the third-order convergence rate of $p_{c}$ observed above, which is better than the approximation capability of the linear element, is mainly due to the special analytic solution $p=0$.

For the parallel DDM with $\nu=1, K=1, \gamma_{f}=0.3$, and $h=1 / 32$, Figures 2 and 3 show the $L^{2}$ errors of hydraulic head, velocity, pressure, and $\eta_{f}$. We can see that the parallel domain decomposition method is convergent for $\gamma_{f} \leq \gamma_{p}$. Moreover, Figures 4 and 5 show that a smaller $\gamma_{f} / \gamma_{p}$ leads to faster convergence.

Then Tables 2 and 3 list some $L^{2}$ errors in velocity, hydraulic head, pressure, and $\eta_{f}$ for the parallel domain decomposition method with $\gamma_{f}=0.3$ and $\gamma_{p}=1.2$. The data in these two tables indicate the geometric convergence rate $\sqrt{\gamma_{f} / \gamma_{p}}$ since all the error ratios are less than $\left(\sqrt{\gamma_{f} / \gamma_{p}}\right)^{4}=$ $(\sqrt{1 / 4})^{4}=0.0625$.

Finally, for the preconditioning feature of the domain decomposition method, Table 4 shows the number of iterations $M$ is independent of the grid size $h$. Here, we set $\gamma_{S}=$ $0.3, \gamma_{D}=1.2, v=1$, and $K=1$. Let $\phi_{h}^{k}, \vec{u}_{h}^{k}$, and $p_{h}^{k}$ denote the finite element solutions of $\phi_{D}^{k}, \vec{u}_{S}^{k}$, and $p_{S}^{k}$ at the $k$ th step

Table 3: $L^{2}$ errors in pressure and $\eta_{f}$ for the parallel DDM with BJS interface condition.

|  | $L^{2}$ velocity errors | $e(i) / e(i-4)$ | $L^{2}$ hydraulic head errors | $e(i) / e(i-4)$ |
| :--- | :---: | :---: | :---: | :---: |
| $e(0)$ | $7.268 \times 10^{-1}$ |  | $5.668 \times 10^{-2}$ |  |
| $e(4)(i=4)$ | $3.826 \times 10^{-2}$ | 0.0526 | $2.752 \times 10^{-3}$ | 0.0486 |
| $e(8)(i=8)$ | $2.014 \times 10^{-3}$ | 0.0526 | $1.399 \times 10^{-4}$ | 0.0508 |
| $e(12)(i=12)$ | $1.060 \times 10^{-4}$ | 0.0526 | $7.233 \times 10^{-6}$ | 0.0517 |
| $e(16)(i=16)$ | $5.579 \times 10^{-6}$ | 0.0526 | $3.767 \times 10^{-7}$ | 0.0521 |
| $e(20)(i=20)$ | $2.937 \times 10^{-7}$ | 0.0526 | $1.969 \times 10^{-8}$ | 0.0523 |

Table 4: The iteration counter $M$ versus the grid size $h$ for both the parallel Robin-Robin domain decomposition method with BJS interface condition.

| $h$ | $1 / 8$ | $1 / 16$ | $1 / 32$ | $1 / 64$ |
| :--- | :---: | :---: | :---: | :---: |
| $M$ | 19 | 19 | 19 | 19 |

of the domain decomposition algorithm. The criterion used to stop the iteration, that is, to determine the value $M$, is $\left\|\vec{u}_{h}^{k}-\vec{u}_{h}^{k-1}\right\|_{0}+\left\|\phi_{h}^{k}-\phi_{h}^{k-1}\right\|_{0}+\left\|p_{h}^{k}-p_{h}^{k-1}\right\|_{0}<\varepsilon$, where the tolerance $\varepsilon=10^{-5}$.

## 6. Conclusions

In this paper, a parallel physics-based domain decomposition method is proposed for the stationary Navier-Stokes-Darcy model with the BJS interface condition. This method is based on the Robin boundary conditions constructed from the three physical interface conditions. Moreover, it is convergent with geometric convergence rates if the relaxation parameter is selected properly. The number of iteration steps is independent of the grid size due to the natural preconditioning advantage of the domain decomposition methods.

## Acknowledgments

This work is partially supported by DOE Grant DEFE0009843, National Natural Science Foundation of China (11175052).

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# A New Integro-Differential Equation for Rossby Solitary Waves with Topography Effect in Deep Rotational Fluids 

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Received 7 May 2013; Accepted 2 September 2013
Academic Editor: Rasajit Bera
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#### Abstract

From rotational potential vorticity-conserved equation with topography effect and dissipation effect, with the help of the multiplescale method, a new integro-differential equation is constructed to describe the Rossby solitary waves in deep rotational fluids. By analyzing the equation, some conservation laws associated with Rossby solitary waves are derived. Finally, by seeking the numerical solutions of the equation with the pseudospectral method, by virtue of waterfall plots, the effect of detuning parameter and dissipation on Rossby solitary waves generated by topography are discussed, and the equation is compared with KdV equation and BO equation. The results show that the detuning parameter $\alpha$ plays an important role for the evolution features of solitary waves generated by topography, especially in the resonant case; a large amplitude nonstationary disturbance is generated in the forcing region. This condition may explain the blocking phenomenon which exists in the atmosphere and ocean and generated by topographic forcing.


## 1. Introduction

Among the many wave motions that occur in the ocean and atmosphere, Rossby waves play one of the most important roles. They are largely responsible for determining the ocean's response to atmospheric and other climate changes [1]. In the past decades, the research on nonlinear Rossby solitary waves had been given much attention in the mathematics and physics, and some models had been constructed to describe this phenomenon. Based upon the pioneering work of Long [2] and Benney [3] on barotropic Rossby waves, there had been remarkably exciting developments [4-11] and formed classical solitary waves theory and algebraic solitary waves theory. The so-called classical solitary waves indicate that the evolution of solitary waves is governed by the Korteweg-de Vries (KdV) type model, while the behavior of solitary waves is governed by the Benjamin-Ono (BO) model, it is called algebraic solitary waves. After the KdV model and BO model, a more general evolution model for solitary waves in a finitedepth fluid was given by Kubota, and the model was called
intermediate long-wave (ILW) model [12, 13]. Many mathematicians solved the above models by all kinds of method and got a series of results [14-19]. We note that most of the previous researches about solitary waves were carried out in the zonal area and could not be applied directly to the spherical earth, and little attention had been focused on the solitary waves in the rotational fluids [20]. Furthermore, as everyone knows the real oceanic and atmospheric motion is a forced and dissipative system. Topography effect as a forcing factor has been studied by many researchers [21-25]; on the other hand, dissipation effect must be considered in the oceanic and atmospheric motion; otherwise, the motion would grow explosively because of the constant injecting of the external forcing energy. Our aim is to construct a new model to describe the Rossby solitary waves in rotational fluid with topography effect and dissipation effect. It has great difference from the previous researches.

In this paper, from rotational potential vorticity-conserved equation with topography effect and dissipation effect,
with the help of the multiple-scale method, we will first construct a new model to describe Rossby solitary waves in deep rotational fluids. Then we will analyse the conservation relations of the model and derive the conservation laws of Rossby solitary waves. Finally, the model is solved by the pseudospectral method [26]. Based on the waterfall plots, the effect of detuning parameter and dissipation on Rossby solitary waves generated by topography are discussed, the model is compared with KdV model and BO model, and some conclusions are obtained.

## 2. Mathematics Model

According to [27], taking plane polar coordinates $(r, \theta), r$ pointing to lower latitude is positive and the positive rotation is counter-clockwise, and then the rotational potential vortic-ity-conserved equation including topography effect and turbulent dissipation is, in the nondimensional form, given by

$$
\begin{align*}
{\left[\frac{\partial}{\partial t}+\right.} & \left.\left(\frac{1}{r} \frac{\partial \Psi}{\partial r} \frac{\partial}{\partial \theta}-\frac{1}{r} \frac{\partial \Psi}{\partial \theta} \frac{\partial}{\partial r}\right)\right] \\
& \times\left[\frac{1}{r} \frac{\partial}{\partial r}\left(r \frac{\partial \Psi}{\partial r}\right)+\frac{1}{r^{2}} \frac{\partial^{2} \Psi}{\partial \theta^{2}}+h(r, \theta)\right]+\frac{\beta}{r} \frac{\partial \Psi}{\partial \theta}  \tag{1}\\
= & -\lambda_{0}\left[\frac{1}{r} \frac{\partial}{\partial r}\left(r \frac{\partial \Psi}{\partial r}\right)+\frac{1}{r^{2}} \frac{\partial^{2} \Psi}{\partial \theta^{2}}\right]+Q
\end{align*}
$$

where $\Psi$ is the dimensionless stream function; $\beta=$ $\left(\omega_{0} / R_{0}\right) \cos \phi_{0}\left(L^{2} / U\right)$, in which $R_{0}$ is the Earth's radius, $\omega_{0}$ is the angular frequency of the Earth's rotation, $\phi_{0}$ is the latitude, $L$ and $U$ are the characteristic horizontal length and velocity scales, $h(r, \theta)$ expresses the topography effect, $\lambda_{0}\left[(1 / r)(\partial / \partial r)(r(\partial \Psi / \partial r))+\left(1 / r^{2}\right)\left(\partial^{2} \Psi / \partial \theta^{2}\right)\right]$ denotes the vorticity dissipation which is caused by the Ekman boundary layer and $\lambda_{0}$ is a dissipative coefficient, $Q$ is the external source, and the form of $Q$ will be given in the latter.

In order to consider weakly nonlinear perturbation on a rotational flow, we assume

$$
\begin{equation*}
\Psi=\int^{r}(\Omega(r)-c+\varepsilon \alpha) r d r+\varepsilon \psi(r, \theta, t) \tag{2}
\end{equation*}
$$

where $\alpha$ is a small disturbance in the basic flow and reflects the proximity of the system to a resonate state; $c$ is a constant, which is regarded as a Rossby waves phase speed; $\psi$ denotes disturbance stream function; $\Omega(r)$ expresses the rotational angular velocity. In order to consider the role of nonlinearity, we assume the following type of rotational angular velocity:

$$
\Omega(r)= \begin{cases}\omega(r) & r_{1} \leq r \leq r_{2}  \tag{3}\\ \omega_{1} & r>r_{2}\end{cases}
$$

where $\omega_{1}$ is constant and $\omega(r)$ is a function of $r$. For simplicity, $\omega(r)$ is assumed to be smooth across $r=r_{2}$.

In the domain $\left[r_{1}, r_{2}\right]$, in order to achieve a balance among topography effect, turbulent dissipation, and nonlinearity and to eliminate the derivative term of dissipation, we assume

$$
\begin{gather*}
h(r, \theta)=\varepsilon^{2} H(r, \theta), \quad \lambda_{0}=\varepsilon^{3 / 2} \lambda, \\
Q=\varepsilon^{3 / 2} \lambda \frac{1}{r} \frac{\partial}{\partial r}\left[r^{2}\left(\omega-c_{0}+\varepsilon \alpha\right)\right] . \tag{4}
\end{gather*}
$$

Substituting (2), (3), and (4) into (1) leads to the following equation for the perturbation stream function $\psi$ :

$$
\begin{align*}
{\left[\frac{\partial}{\partial t}+\right.} & \left.(\omega-c+\varepsilon \alpha) \frac{\partial}{\partial \theta}+\varepsilon\left(\frac{1}{r} \frac{\partial \psi}{\partial r} \frac{\partial}{\partial \theta}-\frac{1}{r} \frac{\partial \psi}{\partial \theta} \frac{\partial}{\partial r}\right)\right] \\
& \times\left[\frac{1}{r} \frac{\partial}{\partial r}\left(r \frac{\partial \psi}{\partial r}\right)+\frac{1}{r^{2}} \frac{\partial^{2} \psi}{\partial \theta^{2}}+\varepsilon H(r, \theta)\right] \\
& +\frac{1}{r}\left[\beta-\frac{d}{d r}\left(\frac{1}{r} \frac{d r^{2} \omega}{d r}\right)\right] \frac{\partial \psi}{\partial \theta}  \tag{5}\\
= & -\varepsilon^{3 / 2} \lambda\left[\frac{1}{r} \frac{\partial}{\partial r}\left(r \frac{\partial \psi}{\partial r}\right)+\frac{1}{r^{2}} \frac{\partial^{2} \psi}{\partial \theta^{2}}\right] .
\end{align*}
$$

In the domain $\left[r_{2}, \infty\right]$, the parameter $\beta$ is smaller than that in the domain $\left[r_{1}, r_{2}\right]$, and we assume $\beta=0$ for $\left[r_{2}, \infty\right]$. Furthermore, the turbulent dissipation and topography effect are absent in the domain and only consider the features of disturbances generated. Substituting (2) and (3) into (1), we have the following governing equations:

$$
\begin{align*}
{\left[\frac{\partial}{\partial t}\right.} & \left.+\left(\omega_{1}-c+\varepsilon \alpha\right) \frac{\partial}{\partial \theta}+\varepsilon\left(\frac{1}{r} \frac{\partial \psi}{\partial r} \frac{\partial}{\partial \theta}-\frac{1}{r} \frac{\partial \psi}{\partial \theta} \frac{\partial}{\partial r}\right)\right] \\
& \times\left[\frac{1}{r} \frac{\partial}{\partial r}\left(r \frac{\partial \psi}{\partial r}\right)+\frac{1}{r^{2}} \frac{\partial^{2} \psi}{\partial \theta^{2}}\right]=0 \tag{6}
\end{align*}
$$

For (5), we introduce the following stretching transformations:

$$
\begin{equation*}
\Theta=\varepsilon^{1 / 2} \theta, \quad r=r, \quad T=\varepsilon^{3 / 2} t, \tag{7}
\end{equation*}
$$

and the perturbation expansion of $\bar{\psi}$ is in the following form:

$$
\begin{equation*}
\bar{\psi}=\psi_{1}(\Theta, r, T)+\varepsilon \psi_{2}(\Theta, r, T)+\cdots \tag{8}
\end{equation*}
$$

Substituting (7) and (8) into (5), comparing the same power of $\varepsilon$ term, we can obtain the $\varepsilon^{1 / 2}$ equation:

$$
\begin{equation*}
\mathscr{L} \psi_{1}=0 \tag{9}
\end{equation*}
$$

where the operator $\mathscr{L}$ is defined as

$$
\begin{equation*}
\mathscr{L}=\frac{1}{r} \frac{\partial}{\partial \Theta}\left\{(\omega-c)\left[\frac{\partial}{\partial r}\left(r \frac{\partial}{\partial r}\right)\right]+\left[\beta-\frac{d}{d r}\left(\frac{1}{r} \frac{d r^{2} \omega}{d r}\right)\right]\right\} . \tag{10}
\end{equation*}
$$

Assume the perturbation at boundary $r=r_{1}$ does not exist, that is,

$$
\begin{equation*}
\psi_{1}=\psi_{2}=\cdots=0 \tag{11}
\end{equation*}
$$

and the perturbation at boundary $r=r_{2}$ is determined by (6). For the linear solution to be separable, assuming the solution of (9) in the form:

$$
\begin{equation*}
\psi_{1}=A(\Theta, T) \phi(r), \tag{12}
\end{equation*}
$$

thus $\phi(r)$ should satisfy the following equation:

$$
\begin{equation*}
(\omega-c) \frac{d}{d r}\left(r \frac{d \phi(r)}{d r}\right)+\left[\beta-\frac{d}{d r}\left(\frac{1}{r} \frac{d r^{2} \omega}{d r}\right)\right] \phi(r)=0 \tag{13}
\end{equation*}
$$

On the other hand, we proceed to the $\varepsilon^{3 / 2}$ equation:

$$
\begin{align*}
\mathscr{L} \psi_{2} & +\frac{\omega-c}{r^{2}} \frac{\partial^{3} \psi_{1}}{\partial \Theta^{3}} \\
& +\left(\frac{\partial}{\partial T}+\alpha \frac{\partial}{\partial \Theta}+\frac{1}{r} \frac{\partial \psi_{1}}{\partial r} \frac{\partial}{\partial \Theta}-\frac{1}{r} \frac{\partial \psi_{1}}{\partial \Theta} \frac{\partial}{\partial r}+\lambda\right)  \tag{14}\\
& \times\left[\frac{1}{r} \frac{\partial}{\partial r}\left(r \frac{\partial \psi_{1}}{\partial r}\right)\right]+(\omega-c) \frac{\partial H(r, \Theta)}{\partial \Theta}=0
\end{align*}
$$

Multiplying the both sides of (14) by $r \phi /(\omega-c)$ and integrating it with respect to $r$ from $r_{1}$ to $r_{2}$, employing the boundary conditions (11), we get

$$
\begin{array}{rl}
\frac{\partial}{\partial \Theta} r & \left.r\left(\phi \frac{\partial}{\partial r} \psi_{2}-\frac{d \phi}{d r} \psi_{2}\right)\right|_{r=r_{2}} \\
& +A \frac{\partial A}{\partial \Theta} \int_{r_{1}}^{r_{2}} \frac{\phi^{3}}{\omega-c} \frac{d}{d r} \\
& \quad \times\left[\frac{\beta-(d / d r)\left((1 / r)\left(d r^{2} \omega / d r\right)\right)}{\phi(\omega-c)}\right] d r \\
& +\left(\frac{\partial A}{\partial T}+\alpha \frac{\partial A}{\partial \Theta}+\lambda A\right) \int_{r_{1}}^{r_{2}} \frac{\phi}{\omega-c} \frac{d}{d r}\left(r \frac{d \phi}{d r}\right) d r \\
\quad+\frac{\partial^{3} A}{\partial \Theta^{3}} \int_{r_{1}}^{r_{2}} \frac{\phi^{2}}{r} d r+\int_{r_{1}}^{r_{2}} r \phi \frac{\partial H}{\partial \Theta} d r=0 \tag{15}
\end{array}
$$

In (15), if the boundary conditions on $\phi$ and $\psi_{2}$ are known, the equation governing the amplitude $A$ will be determined. Assuming the solution of (5) matches smoothly with the solution of (6) at $r=r_{2}$, we can solve (6) to seek the solution at $r=r_{2}$.

For (6), we adopt the transformations in the forms:

$$
\begin{equation*}
\rho=\theta, \quad r=r, \quad T=\varepsilon^{3 / 2} t \tag{16}
\end{equation*}
$$

and the perturbation function is shown $\widetilde{\psi}$; then by substituting (16) into (6), we can get the $\varepsilon^{0}$ equation:

$$
\begin{equation*}
\left(\omega_{1}-c\right) \frac{\partial}{\partial \rho}\left[\frac{1}{r} \frac{\partial}{\partial r}\left(r \frac{\partial \widetilde{\psi}}{\partial r}\right)+\frac{1}{r^{2}} \frac{\partial^{2} \widetilde{\psi}}{\partial \rho^{2}}\right]=0 \tag{17}
\end{equation*}
$$

It is easy to find that (17) can reduce to

$$
\begin{array}{r}
\frac{1}{r} \frac{\partial}{\partial r}\left(r \frac{\partial \widetilde{\psi}}{\partial r}\right)+\frac{1}{r^{2}} \frac{\partial^{2} \widetilde{\psi}}{\partial \rho^{2}}=0, \quad r \geq r_{2}  \tag{18}\\
\widetilde{\psi} \longrightarrow 0, \quad r \longrightarrow \infty
\end{array}
$$

Obviously, the solution of (18) is

$$
\begin{equation*}
\widetilde{\psi}(\rho, r, T, \varepsilon)=\frac{1}{2 \pi} \int_{0}^{2 \pi} \frac{\left(r^{2}-r_{2}^{2}\right) \widetilde{\psi}\left(\rho^{\prime}, r_{2}, T, \varepsilon\right)}{r_{2}^{2}-2 r_{2} r \cos \left(\rho-\rho^{\prime}\right)+r^{2}} d \rho^{\prime} \tag{19}
\end{equation*}
$$

Taking the derivative with respect to $r$ for both sides of (19) leads to

$$
\begin{align*}
& \frac{\partial \widetilde{\psi}}{\partial r} \\
& \quad=\frac{r_{2}}{\pi} \int_{0}^{2 \pi} \widetilde{\psi}\left(\rho^{\prime}, r_{2}, T, \varepsilon\right) \frac{\left[2 r_{2} r-\left(r^{2}+r_{2}^{2}\right) \cos \left(\rho-\rho^{\prime}\right)\right]}{\left[r_{2}^{2}-2 r_{2} r \cos \left(\rho-\rho^{\prime}\right)+r^{2}\right]^{2}} d \rho^{\prime} \tag{20}
\end{align*}
$$

Because the solution of (5) matches smoothly with the solution of (6) at $r=r_{2}$, we obtain

$$
\begin{align*}
& \psi_{1}\left(\Theta, r_{2}, T\right)+\varepsilon \psi_{2}\left(\Theta, r_{2}, T\right)=\widetilde{\psi}\left(\rho, r_{2}, T, \varepsilon\right)+O\left(\varepsilon^{2}\right),  \tag{21}\\
& \frac{\partial \psi_{1}}{\partial r}\left(\Theta, r_{2}, T\right)+\varepsilon \frac{\partial \psi_{2}}{\partial r}\left(\Theta, r_{2}, T\right)=\frac{\partial \widetilde{\psi}}{\partial r}\left(\rho, r_{2}, T, \varepsilon\right)+O\left(\varepsilon^{2}\right) \tag{22}
\end{align*}
$$

From (21), we have

$$
\begin{equation*}
A(\Theta, T) \phi\left(r_{2}\right)=\widetilde{\psi}\left(\rho, r_{2}, T, \varepsilon\right), \quad \psi_{2}\left(\Theta, r_{2}, T\right)=0 \tag{23}
\end{equation*}
$$

Substituting (23) into (20) leads to

$$
\begin{equation*}
\frac{\partial \widetilde{\psi}}{\partial r}\left(\rho, r_{2}, T, \varepsilon\right)=\varepsilon \phi\left(r_{2}\right) \frac{\partial^{2} \mathscr{F}(A(\Theta, T))}{\partial \Theta^{2}} \tag{24}
\end{equation*}
$$

where $\mathscr{J}(A(\Theta, T))=\left(r_{2} / 2 \pi\right) \int_{0}^{2 \pi} A\left(\Theta^{\prime}, T\right) \ln \mid \sin ((\Theta-$ $\left.\left.\Theta^{\prime}\right) / 2\right) \mid d \Theta^{\prime}$. Then, based on (22) and (24), we get

$$
\begin{equation*}
\phi^{\prime}\left(r_{2}\right)=0, \quad \frac{\partial \psi_{2}}{\partial r}\left(\Theta, r_{2}, T\right)=\phi\left(r_{2}\right) \frac{\partial^{2} \mathscr{F}(A(\Theta, T))}{\partial \Theta^{2}} \tag{25}
\end{equation*}
$$

Substituting the boundary conditions (23) and (25) into (15) yields

$$
\begin{align*}
\frac{\partial A}{\partial T} & +\alpha \frac{\partial A}{\partial \Theta}+a_{1} A \frac{\partial A}{\partial \Theta}+a_{2} \frac{\partial^{3} A}{\partial \Theta^{3}}  \tag{26}\\
& +a_{3} \frac{\partial^{3}}{\partial \Theta^{3}} \mathcal{J}(A(\Theta, T))+\lambda A=\frac{\partial G}{\partial \Theta}
\end{align*}
$$

Equation (26) can be rewritten as follows:

$$
\begin{align*}
\frac{\partial A}{\partial T} & +\alpha \frac{\partial A}{\partial \Theta}+a_{1} A \frac{\partial A}{\partial \Theta}+a_{2} \frac{\partial^{3} A}{\partial \Theta^{3}}  \tag{27}\\
& +a_{3} \frac{\partial^{2}}{\partial \Theta^{2}} \mathscr{H}(A(\Theta, T))+\lambda A=\frac{\partial G}{\partial \Theta}
\end{align*}
$$

where $\mathscr{H}(A(\Theta, T))=\left(r_{2} / 4 \pi\right) \int_{0}^{2 \pi} A\left(\Theta^{\prime}, T\right) \cot \left(\left(\Theta-\Theta^{\prime}\right) / 2\right) d \Theta^{\prime}$ and $a=\int_{r_{1}}^{r_{2}}(\phi /(\omega-c))(d / d r)(r(d \phi / d r)) d r, a_{1}=\int_{r_{1}}^{r_{2}}\left(\phi^{3} /(\omega-\right.$ $c))(d / d r)\left[\beta-(d / d r)\left((1 / r)\left(d r^{2} \omega / d r\right)\right) / \phi(\omega-c)\right] d r / a, a_{2}=$ $\int_{r_{1}}^{r_{2}}\left(\phi^{2} / r\right) d r / a, a_{3}=r_{2} \phi^{2}\left(r_{2}\right) / a, G=\int_{r_{1}}^{r_{2}} r \phi H d r / a$. Equation (27) is an integro-differential equation and $\lambda A$ expresses dissipation effect and has the same physical meaning with the term $\partial^{2} A / \partial \Theta^{2}$ in Burgers equation. When $a_{3}=\lambda=H=0$,
the equation degenerates to the KdV equation. When $a_{2}=$ $\lambda=H=0$, the equation degenerates to the so-called rotational BO equation. Here we call (27) forced rotational KdV-BO-Burgers equation. As we know, the forced rotational KdV-BO-Burgers equation as a governing model for Rossby solitary waves is first derived in the paper.

## 3. Conservation Laws

In this section, the conservation laws are used to explore some features of Rossby solitary waves. In [7], Ono presented four conservation laws of BO equation, and we extend Ono's work to investigate the following questions: Has the rotational KdV-BO-Burgers equation also conservation laws without dissipation effect? Has it four conservation laws or more? How to change of these conservation quantities in the presence of dissipation effect?

In this section, topography effect is ignored; that is, $H$ is taken zero in (27). Based on periodicity condition, we assume that the values of $A, A_{\Theta}, A_{\Theta \Theta}, A_{\Theta \Theta \Theta}$ at $\Theta=0$ equal that at $\Theta=2 \pi$. Then integrating (27) with respect to $\Theta$ over ( $0,2 \pi$ ), we are easy to obtain the following conservation relation:

$$
\begin{equation*}
Q_{1}=\int_{0}^{2 \pi} A d \Theta=\exp (-\lambda T) \int_{0}^{2 \pi} A(\Theta, 0) d \Theta \tag{28}
\end{equation*}
$$

From (28), it is obvious that $Q_{1}$ decreases exponentially with the evolution of time $T$ and the dissipation coefficient $\lambda$. By analogy with the KdV equation, $Q_{1}$ is regarded as the mass of the solitary waves. This shows that the dissipation effect causes the mass of solitary waves decrease exponentially. When dissipation effect is absent, the mass of the solitary waves is conserved.

In what follows, (27) has another simple conservation law, which becomes clear if we multiply (27) by $A(\Theta, T)$ and carry the integration; by using the property of the operator $\mathscr{H}: \int_{0}^{2 \pi} f(\Theta) \mathscr{H}(f(\Theta)) d \Theta=0$, then we get

$$
\begin{equation*}
Q_{2}=\int_{0}^{2 \pi} A^{2} d \Theta=\exp (-2 \lambda T) \int_{0}^{2 \pi} A^{2}(\Theta, 0) d \Theta \tag{29}
\end{equation*}
$$

Similar to the mass $Q_{1}, Q_{2}$ is regarded as the momentum of the solitary waves and is conserved without dissipation. The momentum of the solitary waves also decreases exponentially with the evolution of time $T$ and the increasing of dissipative coefficient $\lambda$ in the presence of dissipation effect. Furthermore, the rate of decline of momentum is faster than the rate of mass.

Next, we multiply (27) by $\left(A^{2}-\left(a_{3} / a_{1}\right) \mathscr{H}\left(A_{\Theta}\right)\right)$ and obtain

$$
\begin{aligned}
& \left(\frac{1}{3} A^{3}\right)_{T}-\frac{a_{3}}{a_{1}} \mathscr{H}\left(A_{\Theta}\right) A_{T} \\
& \quad+\left(\alpha+a_{1} A\right) A_{\Theta}\left(A^{2}-\frac{a_{3}}{a_{1}} \mathscr{H}\left(A_{\Theta}\right)\right) \\
& \quad+a_{2} A_{\Theta \Theta \Theta}\left(A^{2}-\frac{a_{3}}{a_{1}} \mathscr{H}\left(A_{\Theta}\right)\right)
\end{aligned}
$$

$$
\begin{align*}
& +a_{3}\left(A^{2}-\frac{a_{3}}{a_{1}} \mathscr{H}\left(A_{\Theta}\right)\right)(\mathscr{H}(A))_{\Theta \Theta} \\
& +\lambda\left(A^{2}-\frac{a_{3}}{a_{1}} \mathscr{H}\left(A_{\Theta}\right)\right) A=0 \tag{30}
\end{align*}
$$

Then taking the derivative of (27) with respect to $\Theta$ and multiplying $\left(-\left(2 a_{2} / a_{1}\right) A_{\Theta}+\left(a_{3} / a_{1}\right) \mathscr{H}(A)\right)$ lead to

$$
\begin{align*}
& \left(-\frac{2 a_{2}}{a_{1}} A_{\Theta}+\frac{a_{3}}{a_{1}} \mathscr{H}(A)\right) A_{\Theta T} \\
& \quad+\left[\alpha A_{\Theta \Theta}+a_{1}\left(A A_{\Theta}\right)_{\Theta}\right]\left(-\frac{2 a_{2}}{a_{1}} A_{\Theta}+\frac{a_{3}}{a_{1}} \mathscr{H}(A)\right) \\
& \quad+a_{2} A_{\Theta \Theta \Theta \Theta}\left(-\frac{2 a_{2}}{a_{1}} A_{\Theta}+\frac{a_{3}}{a_{1}} \mathscr{H}(A)\right)  \tag{31}\\
& \quad+a_{3}\left(-\frac{2 a_{2}}{a_{1}} A_{\Theta}+\frac{a_{3}}{a_{1}} \mathscr{H}(A)\right) \mathscr{H}(A)_{\Theta \Theta \Theta} \\
& \quad+\lambda A\left(-\frac{2 a_{2}}{a_{1}} A_{\Theta}+\frac{a_{3}}{a_{1}} \mathscr{H}(A)\right)=0 .
\end{align*}
$$

Adding (30) to (31), by virtue of the property of operator $\mathscr{H}$ :

$$
\begin{equation*}
\mathscr{H}(A)_{\Theta \Theta}=\mathscr{H}\left(A_{\Theta \Theta}\right), \quad \int_{0}^{2 \pi} u \mathscr{H} v d \Theta=-\int_{0}^{2 \pi} v \mathscr{H} u d \Theta \tag{32}
\end{equation*}
$$

we have

$$
\begin{aligned}
& \left(\frac{1}{3} A^{3}-\frac{a_{2}}{a_{1}} A_{\Theta}^{2}+\frac{a_{3}}{a_{1}} A_{\Theta} \mathscr{H}(A)\right)_{T} \\
& \quad+\alpha\left[\frac{1}{3} A^{3}-\frac{a_{2}}{a_{1}} A_{\Theta}^{2}+\frac{a_{3}}{a_{1}} \mathscr{H}(A) A_{\Theta}\right]_{\Theta} \\
& \quad+\left(\frac{a_{1}}{4} A^{4}\right)_{\Theta}+\frac{a_{3}^{2}}{2 a_{1}}\left[H(A) H(A)_{\Theta \Theta}\right]_{\Theta} \\
& \quad+\frac{a_{2} a_{3}}{a_{1}}\left(A_{\Theta \Theta \Theta} H(A)-2 A_{\Theta} H(A)_{\Theta \Theta}\right)_{\Theta} \\
& \quad-\frac{2 a_{2}}{a_{1}}\left(A_{\Theta} A_{\Theta \Theta \Theta}-\frac{1}{2} A_{\Theta \Theta}^{2}\right)_{\Theta}+a_{2}\left(A^{2} A_{\Theta \Theta}-2 A A_{\Theta}^{2}\right)_{\Theta} \\
& \quad+a_{3}\left[A(A \mathscr{H}(A))_{\Theta}\right]_{\Theta}+\lambda\left(A^{3}-\frac{a_{2}}{a_{1}} A_{\Theta}^{2}+\frac{a_{3}}{a_{1}} A_{\Theta} H(A)\right)
\end{aligned}
$$

$$
\begin{equation*}
=0 \tag{33}
\end{equation*}
$$

Taking $Q_{3}=\int_{0}^{2 \pi}\left((1 / 3) A^{3}-\left(a_{2} / a_{1}\right) A_{\Theta}^{2}+\left(a_{3} / a_{1}\right) A_{\Theta} \mathscr{H}(A)\right) d \Theta$, we are easy to see that when the dissipation effect is absent, that is, $\lambda=0, Q_{3}$ is a conserved quantity and regarded as the energy of the solitary waves. So we can conclude that the energy of solitary waves is conserved without dissipation. By analysing (33), we can find the decreasing trend of energy of solitary waves.

Finally, let us consider a quantity related to the phase of solitary waves:

$$
\begin{equation*}
\widetilde{\mathrm{Q}}_{4}=\frac{d}{d T} \int_{0}^{2 \pi} \Theta A d \Theta \tag{34}
\end{equation*}
$$

and we can get $d \widetilde{Q}_{4} / d T=0$ without dissipation. According [7], we present the velocity of the center of gravity for the ensemble of such waves $Q_{4}=\widetilde{Q}_{4} / Q_{1}$; by employing $d Q_{1} / d T=0$ and $d \widetilde{Q}_{4} / d T=0$, we have $d Q_{4} / d T=0$, which shows that the velocity of the center of gravity is conserved without dissipation.

After the four conservation relations are given, we can proceed to seek the fifth conservation quantity. In fact, after tedious calculation, we can also verify that

$$
\begin{equation*}
Q_{5}=\int_{0}^{2 \pi}\left(\frac{1}{4} A^{4}-\frac{3 a_{2}}{a_{1}} A A_{\Theta}^{2}+\frac{9 a_{2}}{a_{1}^{2}} A_{\Theta \Theta}^{2}+\frac{a_{3}}{4 a_{1}} A^{2} \mathscr{H}(A)\right) d \Theta \tag{35}
\end{equation*}
$$

is also conservation quantity. According the idea, we can obtain the sixth conservation quantity $Q_{6}$ and the seventh conservation quantity $Q_{7} \ldots$, so we can guess that, similar to the KdV equation, the rotational KdV -BO-Burgers equation without dissipation also owns infinite conservation laws, but it needs to be verified in the future.

## 4. Numerical Simulation and Discussion

In this section, we will take into account the generation and evolution feature of Rossby solitary waves under the influence of topography and dissipation, so we need to seek the solutions of forced rotational KdV-BO-Burgers equation. But we know that there is no analytic solution for (27), and here we consider the numerical solutions of (27) by employing the pseudospectral method.

The pseudo-spectral method uses a Fourier transform treatment of the space dependence together with a leap-forg scheme in time. For ease of presentation the spatial period is normalized to $[0,2 \pi]$. This interval is divided into $2 N$ points, and then $\Delta T=\pi / N$. The function $A(X, T)$ can be transformed to the Fourier space by

$$
\begin{array}{r}
\widehat{A}(v, T)=F A=\frac{1}{\sqrt{2 N}} \sum_{j=0}^{2 N-1} A(j \Delta X, T) e^{-\pi i j v / N},  \tag{36}\\
v=0, \pm 1, \ldots, \pm N .
\end{array}
$$

The inversion formula is

$$
\begin{equation*}
A(j \Delta X, T)=F^{-1} \widehat{A}=\frac{1}{\sqrt{2 N}} \sum_{v} \widehat{A}(v, T) e^{\pi i j v / N} \tag{37}
\end{equation*}
$$

These transformations can use Fast Fourier Transform algorithm to efficiently perform. With this scheme, $\partial A / \partial X$ can be evaluated as $F^{-1}\{i v F A\}, \partial^{3} A / \partial X^{3}$ as $-i F^{-1}\left\{v^{3} F A\right\}, \partial H / \partial X$ as
$F^{-1}\{i v F H\}$, and so on. Combined with a leap-frog time step, (27) would be approximated by

$$
\begin{align*}
& A(X, T+\Delta T)-A(X, T-\Delta T)+i \alpha F^{-1}\{v F A\} \Delta T \\
& \quad+i a_{1} A F^{-1}\{v F A\} \Delta T-a_{2} i F^{-1}\left\{v^{3} F A\right\} \Delta T  \tag{38}\\
& \quad-a_{3} F^{-1}\left\{v^{2} F \mathscr{H}(A)\right\} \Delta T+\lambda A=i F^{-1}\{v F G\} \Delta T .
\end{align*}
$$

The computational cost for (38) is six fast Fourier transforms per time step.

Once the zonal flow $\Omega(r)$ and the topography function $H(r, \Theta)$ as well as dissipative coefficient $\lambda$ are given, it is easy to get the coefficients of (27) by employing (13). In order to simplify the calculation and to focus attention on the time evolution of the solitary waves with topography effect and dissipation effect and to show the difference among the KdV model, BO model, and rotational KdV-BO model, we take $a_{1}=1, a_{2}=-1$, and $a_{3}=-1$. As an initial condition, we take $A(X, 0)=0$. In the present numerical computation, the topography forcing is taken as $G=e^{-[30(\Theta-\pi)]^{2} / 4}$.
4.1. Effect of Detuning Parameter $\alpha$ and Dissipation. In Figure 1, we consider the effect of detuning parameter $\alpha$ on solitary waves. The evolution features of solitary waves generated by topography are shown in the absence of dissipation with different detuning parameter $\alpha$. It is easy to find from these waterfall plots that the detuning parameter $\alpha$ plays an important role for the evolution features of solitary waves generated by topography.

When $\alpha>0$ (Figure $1(\mathrm{a})$ ), a positive stationary solitary wave is generated in the topographic forcing region, and a modulated cnoidal wave-train occupies the downstream region. There is no wave in the upstream region. A flat buffer region exists between the solitary wave in the forcing region and modulated cnoidal wave-train in the downstream. With the detuning parameter $\alpha$ decreasing, the amplitudes of both solitary wave in the forcing region and modulated cnoidal wave-train in the downstream region increase and the modulated cnoidal wave-train closes to the forcing region gradually and the flat buffer region disappears slowly.

Up to $\alpha=0$ (Figure 1(b)), the resonant case forms. In this case, a large amplitude nonstationary disturbance is generated in the forcing region. To some degree, this condition may explain the blocking phenomenon which exists in the atmosphere and ocean and generated by topographic forcing.

As $\alpha<0$, from Figure 1(c) we can easy to find that a negative stationary solitary wave is generated in the forcing region, and this is great difference with the former two conditions. Meanwhile, there are both wave-trains in the upstream and downstream region. The amplitude and wavelength of wave-train in the upstream region are larger than those in the downstream regions. Similar to Figure 1(b) and unlike Figure 1(a), the wave-trains in the upstream and downstream regions connect to the forcing region and the flat buffer region disappears.

Figure 2 shows the solitary waves generated by topography in the presence of dissipation with dissipative coefficient $\lambda=0.3$ and detuning parameter $\alpha=2.5$. The conditions of $\alpha=0$ and $\alpha<0$ are omitted. Compared to Figure 1(a), we will


Figure 1: Solitary waves generated by topography in the absence of dissipation.


Figure 2: Solitary waves generated by topography in the presence of dissipation ( $\lambda=0.3, \alpha=2.5$ ).
find that there is also a solitary wave generated in the forcing region, but because of dissipation effect the amplitude of solitary wave in the forcing region decreases as the dissipative coefficient $\lambda$ increases (Figures omitted) and time evolution. Meanwhile, the modulated cnoidal wave-train in the downstream region is dissipated. When $\lambda$ is big enough, the modulated cnoidal wave-train in the downstream region disappears.
4.2. Comparison of $K d V$ Model, BO Model, and $K d V-B O$ Model. We know that the rotation KdV-BO equation reduces to the KdV equation as $a_{3}=0$ and to the BO equation as $a_{2}=0$, so, in this subsection by comparing Figure 1(a) with Figure 3, we will look for the difference of solitary waves which is described by KdV-BO model, KdV model, and BO model. The role of detuning parameter $\alpha$ and dissipation effect has been studied in the former subsection, so here we only consider the condition of $\lambda=0, \alpha=2.5$.

At first, we can find that a positive solitary wave is all generated in the forcing region in Figures 1(a), 3(a) and 3(b), but it is stationary in Figures 1(a) and 3(a), and is nonstationary in Figure 3(b). By surveying carefully we find that the amplitude of stationary wave in the forcing region in Figure 1(a) is larger than that in Figure 3(a). Additionally, a modulated cnoidal wave-train is excited in the downstream region in Figures 1(a) and 3(a), and in both downstream and upstream region in Figure 3(b). The amplitude of modulated cnoidal wave-train in downstream region in Figure 3(b) is the largest and in Figure 1(a) is the smallest among the three models. Furthermore, in Figure 3(a) the wave number of modulated cnoidal wave-train is more than that in Figures 1(a) and 3(b). In a word, by the above analysis and comparison, it is easy to find that Figure 1(a) is similar to Figure 3(a) and has great difference with Figure 3(b). This indicates that the term $a_{2}\left(\partial^{3} A / \partial \Theta^{3}\right)$ plays more important role than the term $a_{3}\left(\partial^{2} / \partial \Theta^{2}\right) \mathscr{H}(A)$ in rotational KdV-BO equation.


Figure 3: Comparison of KdV model, Bo model, and KdV-BO model.

## 5. Conclusions

In this paper, we presented a new model: rotational KdV-BOBurgers model to describe the Rossby solitary waves generated by topography with the effect of dissipation in deep rotational fluids. By analysis and computation, five conservation quantities of KdV-BO-Burgers model were derived and corresponding four conservation laws of Rossby solitary waves were obtained; that is, mass, momentum, energy, and velocity of the center of gravity of Rossby solitary waves are conserved without dissipation effect. Further, we presented that the rotational KdV-BO-Burgers equation owns infinite conservation quantities in the absence of dissipation effect. Detailed numerical results obtained using pseudospectral method are presented to demonstrate the effect of detuning parameter $\alpha$ and dissipation. By comparing the KdV model, BO model, and KdV-BO model, we drew the conclusion that the term $a_{2}\left(\partial^{3} A / \partial \Theta^{3}\right)$ plays more important role than the term $a_{3}\left(\partial^{2} / \partial \Theta^{2}\right) \mathscr{H}(A)$ in rotational KdV -BO equation. More problems on KdV-BO-Burgers equation such as the analytical solutions, integrability, and infinite conservation quantities are not studied in the paper due to limited space. In fact, there are many methods carried out to solve some equations with special nonhomogenous terms [28] as well as multiwave solutions and other form solution [29,30] of homogenous equation. These researches have important value for understanding and realizing the physical phenomenon described by the equation and deserve to carry out in the future.

## Acknowledgments

This work was supported by Innovation Project of Chinese Academy of Sciences (no. KZCX2-EW-209), National Natural Science Foundation of China (nos. 41376030 and 11271107), Nature Science Foundation of Shandong Province of China (no. ZR2012AQ015), Science and Technology plan project of the Educational Department of Shandong Province of China (no. J12LI03), and SDUST Research Fund (no. 2012KYTD105).

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# A One Step Optimal Homotopy Analysis Method for Propagation of Harmonic Waves in Nonlinear Generalized Magnetothermoelasticity with Two Relaxation Times under Influence of Rotation 

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Received 1 May 2013; Revised 2 June 2013; Accepted 4 June 2013
Academic Editor: Santanu Saha Ray
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#### Abstract

The aim of this paper is to apply OHAM to solve numerically the problem of harmonic wave propagation in a nonlinear thermoelasticity under influence of rotation, thermal relaxation times, and magnetic field. The problem is solved in one-dimensional elastic half-space model subjected initially to a prescribed harmonic displacement and the temperature of the medium. The HAM contains a certain auxiliary parameter which provides us with a simple way to adjust and control the convergence region and rate of convergence of the series solution. This optimal approach has a general meaning and can be used to get fast convergent series solutions of the different type of nonlinear fractional differential equation. The displacement and temperature are calculated for the models with the variations of the magnetic field, relaxation times, and rotation. The results obtained are displayed graphically to show the influences of the new parameters.


## 1. Introduction

In the past recent years, much attention has been devoted to simulate some real-life problems which can be described by nonlinear coupled differential equations using reliable and more efficient methods. Nonlinear partial differential equations are useful in describing various phenomena in disciplines. The nonlinear coupled systems of partial differential equations often appear in the study of circled fuel reactor, high-temperature hydrodynamics, and thermoelasticity problems, see [1-4]. From the analytical point of view, a lot of work has been done for such systems. With the rapid development of nanotechnology, there appears an ever increasing interest of scientists and researchers in this field of science. Nanomaterials, because of their exceptional mechanical, physical, and chemical properties, have been
the main topic of research in many scientific publications. Wave generation in nonlinear thermoelasticity problems has gained a considerable interest for its utilitarian aspects in understanding the nature of interaction between the elastic and thermal fields as well as the system of PDEs for its applications. A lot of applications were paid on existence, uniqueness, and stability of the solution of the problem, see [5-7].

Recently, much attention has been devoted to numerical methods, which do not require discretization of space-time variables or linearization of the nonlinear equations, among the homotopy analysis methods. Since most of the nonlinear FDEs cannot be solved exactly, approximate and numerical methods must be used. Some of the recent analytical methods for solving nonlinear problems include the homotopy analysis method HAM [8-14]. The HAM, first proposed in 1992 by Liao [8], has been successfully applied to solve many
problems in physics and science. This method is applied to solve linear and nonlinear systems.

The homotopy perturbation method HPM has the merits of simplicity and easy execution. The homotopy perturbation method was first proposed by He [15]. Unlike the traditional numerical methods, the HPM does not need discretization and linearization. Most perturbation methods assume that a small parameter exists, but most nonlinear problems have no small parameter at all. Many new methods have been proposed to eliminate the small parameter. Recently, the applications of homotopy theory among scientists appeared, and the homotopy theory became a powerful mathematical tool, when it is successfully coupled with perturbation theory.

Recently, Gepreel et al. [16] investigated the homotopy perturbation method and variational iteration method for harmonic waves propagation in nonlinear magnetothermoelasticity with rotation. Abd-Alla and Abo-Dahab [17] investigated the effect of rotation and initial stress on an infinite generalized magnetothermoelastic diffusion body with a spherical cavity. Abo-Dahab and Mohamed [18] studied the influence of magnetic field and hydrostatic initial stress on reflection phenomena of P (Primary) and SV (Shear Vertical) waves from a generalized thermoelastic solid half space. Abd-Alla and Mahmoud [19] investigated the magnetothermoelastic problem in rotating non-homogeneous orthotropic hollow cylinder under the hyperbolic heat conduction model. Abd-Alla et al. [20] studied the thermal stresses effect in a non-homogeneous orthotropic elastic multilayered cylinder. Abd-Alla et al. [21] studied the generalized magnetothermoelastic Rayleigh waves in a granular medium under the influence of a gravity field and initial stress. Abd-Alla and Abo-Dahab [22] investigated the time-harmonic sources in a generalized magnetothermoviscoelastic continuum with and without energy dissipation.

In the present paper, investigation is devoted for solving numerically the problem of harmonic wave propagation in a nonlinear thermoelasticity under influence of magnetic field, thermal relaxation times, and rotation. The problem is solved in one-dimensional elastic half-space model subjected initially to a prescribed harmonic displacement and the temperature of the medium.

The HAM contains a certain auxiliary parameter which provides us with a simple way to adjust and control the convergence region and rate of convergence of the series solution. The $h$-curve of the third-order approximate solutions is displayed graphically to show the interval that the exact and approximate solutions take the same values. The displacement and temperature are calculated for the methods with the variations of the magnetic field and rotation. The results obtained are displayed graphically to show the influences of the new parameters.

## 2. A One-Step Optimal Homotopy Analysis Method for PDEs

To describe the basic ideas of the HAM, we consider the following general nonlinear differential equation:

$$
\begin{equation*}
N[u(x, t)]=0, \tag{1}
\end{equation*}
$$

where $N$ is a nonlinear operator for this problem, $x$ and $t$ denote independent variables, and $u(x, t)$ is an unknown function.

By means of the HAM, one first constructs the following zero-order deformation equation:

$$
\begin{equation*}
(1-q) \mathscr{L}\left(\phi(x, t ; q)-u_{0}(x, t)\right)=q h H(t) N[\phi(x, t ; q)] \tag{2}
\end{equation*}
$$

where $q \in[0,1]$ is the embedding parameter, $h \neq 0$ is an auxiliary parameter, $H(t) \neq 0$ is an auxiliary function, $\mathscr{L}$ is an auxiliary linear operator, and $u_{0}(x, t)$ is an initial guess. Obviously, when $q=0$ and $q=1$, it holds

$$
\begin{equation*}
\phi(x, t ; 0)=u_{0}(x, t), \quad \phi(x, t ; 1)=u(x, t) . \tag{3}
\end{equation*}
$$

Liao $[8,9]$ expanded $\phi(x, t ; q)$ in Taylor series with respect to the embedding parameter $q$, as follows:

$$
\begin{equation*}
\phi(x, t ; q)=u_{0}(x, t)+\sum_{m=1}^{\infty} u_{m}(x, t) q^{m} \tag{4}
\end{equation*}
$$

where

$$
\begin{equation*}
u_{m}(x, t)=\left.\frac{1}{m!} \frac{\partial^{m} \phi(x, t ; q)}{\partial q^{m}}\right|_{q=0} \tag{5}
\end{equation*}
$$

Assume that the auxiliary linear operator, the initial guess, the auxiliary parameter $h$, and the auxiliary function $H(t)$ are selected such that the series (4) is convergent at $q=1$; then we have from (4)

$$
\begin{equation*}
u(x, t)=u_{0}(x, t)+\sum_{m=1}^{\infty} u_{m}(x, t) \tag{6}
\end{equation*}
$$

Let us define the vector

$$
\begin{equation*}
\vec{u}_{n}(x, t)=\left\{u_{0}(x, t), u_{1}(x, t), u_{2}(x, t), \ldots, u_{n}(x, t)\right\} \tag{7}
\end{equation*}
$$

Differentiating (2) $m$ times with respect to $q$, then setting $q=0$ and dividing then by $m!$, we have the following $m$ thorder deformation equation:

$$
\begin{equation*}
\mathscr{L}\left(u_{m}(x, t)-\chi_{m} u_{m-1}(x, t)\right)=h H(t) \mathscr{R}_{m}\left(\vec{u}_{m-1}\right), \tag{8}
\end{equation*}
$$

where

$$
\begin{gather*}
\mathscr{R}_{m}\left(\vec{u}_{m-1}\right)=\left.\frac{1}{(m-1)!} \frac{\partial^{m-1} \mathcal{N}[\phi(x, t ; q)]}{\partial q^{m-1}}\right|_{q=0}  \tag{9}\\
\varkappa_{m}= \begin{cases}0, & m \leq 1 \\
1, & m>1\end{cases}
\end{gather*}
$$

Applying the integral operator on both sides of (8), we have

$$
\begin{equation*}
u_{m}(x, t)=\varkappa_{m} u_{m-1}(x, t)+h \int_{0}^{t} H(t) \mathscr{R}_{m}\left(\vec{u}_{m-1}\right) d t \tag{10}
\end{equation*}
$$

where the $m$ th-order deformation equation (8) can be easily solved, especially by means of symbolic computation software
such as Mathematica, Maple, and MathLab. The convergence of the homotopy analysis method for solving these equations is discussed in [23].

Abbasbandy and Jalili [24] and Turkyilmazoglu [25-29] applied the homotopy analysis method to nonlinear ODEs and suggested the so-called optimization method to find out the optimal convergence control parameters by minimum of the square residual error integrated in the whole region having physical meaning. Their approach is based on the square residual error.

Let $\Delta(h)$ denote the square residual error of the governing equation (1) and express it as

$$
\begin{equation*}
\Delta(h)=\int_{\Omega}\left(N\left[\tilde{u}_{n}(t)\right]\right)^{2} d \Omega \tag{11}
\end{equation*}
$$

where

$$
\begin{equation*}
\tilde{u}_{m}(t)=u_{0}(t)+\sum_{k=1}^{m} u_{k}(t) \tag{12}
\end{equation*}
$$

the optimal value of $h$ is given by a nonlinear algebraic equation:

$$
\begin{equation*}
\frac{d \Delta(h)}{d h}=0 \tag{13}
\end{equation*}
$$

## 3. Application of HAM on the Nonlinear Magnetothermoelastic with Rotation Equations

In this section, we use the homotopy analysis method to calculate the approximate solutions of the following nonlinear magnetothermoplastic model with rotation equations

$$
\begin{align*}
& \left(1+\sigma_{1}\right) u_{t t}+\Omega u_{t}-u_{x x}\left(1-\sigma_{2}+2 \gamma u_{x}+3 \delta u_{x}^{2}\right) \\
& \quad-\beta_{1}\left(1-i \omega \tau_{2}\right) \theta_{x}-\beta_{2}\left(\theta u_{x}\right)_{x}=0 \\
& \left(\theta\left(1-i \omega \tau_{1}\right)-a u_{x}\left(1-i \omega \delta \tau_{1}\right)-\frac{1}{2} b u_{x}^{2}\right)_{t}  \tag{14}\\
& \quad-\left[\left(1+\alpha u_{x}\right) \theta_{x}\right]_{x}=0
\end{align*}
$$

where $\sigma_{1}, \sigma_{2}, \Omega, \gamma, \beta_{1}, \beta_{2}, a, b$, and $\alpha$ are arbitrary constants with the initial conditions

$$
\begin{gather*}
u(x, 0)=\theta(x, 0)=A(1-\cos (x))  \tag{15}\\
u_{t}(x, 0)=\theta_{t}(x, 0)=0
\end{gather*}
$$

where $A$ is an arbitrary constant and the boundary conditions

$$
\begin{gather*}
u(0, t)=\theta(0, t)=0 \\
u_{t}(0, t)=\theta_{t}(0, t)=0 \tag{16}
\end{gather*}
$$

To demonstrate the effectiveness of the method, we consider the system of nonlinear initial-value problem (14) with the
initial conditions (15) and the boundary conditions (16) by choosing the linear operators

$$
\begin{align*}
& \mathscr{L}_{1}\left[\phi_{1}(x, t ; q)\right]=\frac{\partial^{2} \phi_{1}(x, t ; q)}{\partial t^{2}}  \tag{17}\\
& \mathscr{L}_{2}\left[\phi_{2}(x, t ; q)\right]=\frac{\partial \phi_{2}(x, t ; q)}{\partial t}
\end{align*}
$$

with the property $\mathscr{L}_{1}\left[c_{1}+c_{2} t\right]=0, \mathscr{L}_{2}\left[c_{3}\right]$, where $c_{i},(i=$ $1,2,3$ ) are the integral constants and the nonlinear operators are defined as

$$
\begin{align*}
N_{1}\left[\phi_{1}, \phi_{2}\right]= & \left(1+\sigma_{1}\right) \frac{\partial^{2} \phi_{1}}{\partial t^{2}}+\Omega \frac{\partial \phi_{1}}{\partial t}-\frac{\partial^{2} \phi_{1}}{\partial x^{2}} \\
& \times\left(1-\sigma_{2}+2 \gamma \frac{\partial \phi_{1}}{\partial x}+3 \delta\left(\frac{\partial \phi_{1}}{\partial x}\right)^{2}\right)-\beta_{1} \frac{\partial \phi_{2}}{\partial x} \\
& +i \omega \tau_{2} \beta_{1} \frac{\partial \phi_{2}}{\partial x}-\beta_{2} \frac{\partial}{\partial x}\left(\phi_{2} \frac{\partial \phi_{1}}{\partial x}\right) \\
N_{2}\left[\phi_{1}, \phi_{2}\right]= & \frac{\partial}{\partial t}\left(\left(1-i \omega \tau_{1}\right) \phi_{2}-a\left(1-i \omega \tau_{1} \delta\right) \frac{\partial \phi_{1}}{\partial x}\right. \\
& \left.-\frac{1}{2} b\left(\frac{\partial \phi_{1}}{\partial x}\right)^{2}-\frac{\partial}{\partial x}\left(1+\alpha \frac{\partial \phi_{1}}{\partial x}\right) \frac{\partial \phi_{2}}{\partial x}\right) \tag{18}
\end{align*}
$$

Choosing $H_{i}(t)=1$ for $i=1,2$, the zeroth-order deformation equations are

$$
\begin{align*}
(1-q) & \mathscr{L}_{1}\left[\phi_{1}(x, t ; q)-u_{0}(x, t)\right] \\
& =q h_{1} N_{1}\left[\phi_{1}(x, t ; q), \phi_{2}(x, t ; q)\right] \\
(1-q) & \mathscr{L}_{2}\left[\phi_{2}(x, t ; q)-v_{0}(x, t)\right]  \tag{19}\\
& =q h_{2} N_{2}\left[\phi_{1}(x, t ; q), \phi_{2}(x, t ; q)\right]
\end{align*}
$$

where

$$
\begin{array}{ll}
\phi_{1}(x, t ; 0)=u_{0}(x, t), & \phi_{1}(x, t ; 1)=u(x, t) \\
\phi_{2}(x, t ; 0)=v_{0}(x, t), & \phi_{2}(x, t ; 1)=\theta(x, t) \tag{20}
\end{array}
$$

Then, the $m$ th-order deformation equations become

$$
\begin{align*}
& \mathscr{L}_{1}\left[u_{m}(x, t)-\varkappa_{m} u_{m-1}(x, t)\right]=h_{1} \mathscr{R}_{1 m}\left(\vec{u}_{m-1}, \vec{\theta}_{m-1}\right), \\
& \mathscr{L}_{2}\left[\theta_{m}(x, t)-\varkappa_{m} \theta_{m-1}(x, t)\right]=h_{2} \mathscr{R}_{2 m}\left(\vec{u}_{m-1}, \vec{\theta}_{m-1}\right) \tag{21}
\end{align*}
$$

where

$$
\begin{align*}
& \mathscr{R}_{1 m}\left(\vec{u}_{m-1}, \vec{\theta}_{m-1}\right) \\
& =\frac{\partial^{2} u_{m-1}}{\partial t^{2}} \\
& +\frac{1}{1+\sigma_{1}}\left(\Omega\left(u_{m-1}\right)_{t}-\left(u_{m-1}\right)_{x x}+\sigma_{2}\left(u_{m-1}\right)_{x x}\right. \\
& -2 \gamma \sum_{j=0}^{m-1}\left(u_{j}\right)_{x x}\left(u_{m-1-j}\right)_{x} \\
& -3 \delta \sum_{i=0}^{m-1} \sum_{j=0}^{i}\left(u_{j}\right)_{x}\left(u_{i-j}\right)_{x}\left(u_{m-1-i}\right)_{x x} \\
& -\beta_{2} \sum_{j=0}^{m-1}\left(\theta_{j}\right)_{x}\left(u_{m-1-j}\right)_{x}-\beta_{1}\left(\theta_{m-1}\right)_{x} \\
& \left.-\beta_{2} \sum_{j=0}^{m-1}\left(\theta_{j}\right)\left(u_{m-1-j}\right)_{x x}+i \omega \tau_{2} \beta_{1} \theta_{m-1}\right), \\
& \mathscr{R}_{2 m}\left(\vec{u}_{m-1}, \vec{\theta}_{m-1}\right) \\
& =\frac{\partial \theta_{m-1}}{\partial t}+\frac{1}{1-i \omega \tau_{1}}\left(-a\left(u_{m-1}\right)_{x t}\right. \\
& -\frac{1}{2} b \sum_{j=0}^{m-1}\left(u_{j}\right)_{x t}\left(u_{m-1-j}\right)_{x t} \\
& -\left(\theta_{m-1}\right)_{x x} \\
& -\alpha \sum_{j=0}^{m-1}\left(\theta_{m-1-j}\right)_{x}\left(u_{j}\right)_{x x} \\
& -\alpha \sum_{j=0}^{m-1}\left(\theta_{m-1-j}\right)_{x x}\left(u_{j}\right)_{x} \\
& \left.-\left(\theta_{m-1}\right)_{x x}+\operatorname{ai\omega } \tau_{1} \delta\left(u_{m-1}\right)_{x t}\right) . \tag{22}
\end{align*}
$$

For simplicity, we suppose $h_{1}=h_{2}$; the system (21) has the following general solutions:

$$
\begin{gather*}
u_{m}(x, t)=\varkappa_{m} u_{m-1}(x, t)+h \iint_{0}^{t} \mathscr{R}_{1 m}\left(\vec{u}_{m-1}, \vec{\theta}_{m-1}\right) d t d t \\
\theta_{m}(x, t)=\varkappa_{m} \theta_{m-1}(x, t)+h \int_{0}^{t} \mathscr{R}_{2 m}\left(\vec{u}_{m-1}, \vec{\theta}_{m-1}\right) d t \tag{23}
\end{gather*}
$$

In this case, where $u_{0}$ and $\theta_{0}$ are constants, the general solution of (23) is taking the following form:

$$
\begin{align*}
& u(x, t)=u_{0}(x, t)+\sum_{m=1}^{\infty} u_{m}(x, t) \\
& \theta(x, t)=\theta_{0}(x, t)+\sum_{m=1}^{\infty} \theta_{m}(x, t) \tag{24}
\end{align*}
$$

The problems above can be readily solved by symbolic computation packages such as Mathematica. Thereupon, successive solving of these problems yields

$$
\begin{aligned}
& u_{0}(x, 0)=A(1-\cos (x)), \\
& \theta_{0}(x, 0)=A(1-\cos (x)) \text {, } \\
& u_{1}(x, t)=\frac{A h t^{2}}{4\left(1+\sigma_{1}\right)}\left(2 A \beta_{2}(-\cos (x)+\cos (2 x))\right. \\
& +\cos (x)\left(-2-3 A^{2} \delta\right. \\
& +3 A^{2} \delta \cos (2 x) \\
& \left.-4 A \gamma \sin (x)+2 \sigma_{2}\right) \\
& \left.+2 i \sin (x) \beta_{1}\left(i+\omega \tau_{2}\right)\right), \\
& \theta_{1}(x, t)=\frac{A h t \cos (x)(1+2 A \alpha \sin (x))}{-1+i \omega \tau_{1}}, \\
& u_{2}(x, t) \\
& =\frac{A h t^{2}}{4\left(1+\sigma_{1}\right)} \\
& \times\left(2 A \beta_{2}(-\cos (x)+\cos (2 x))+\cos (x)\right. \\
& \times\left(-2-3 A^{2} \delta+3 A^{2} \delta \cos (2 x)\right. \\
& \left.-4 A \gamma \sin (x)+2 \sigma_{2}\right) \\
& \left.+2 i \sin (x) \beta_{1}\left(i+\omega \tau_{2}\right)\right) \\
& +\frac{1}{384\left(1+\sigma_{1}\right)^{2}} \\
& \times\left(A h ^ { 2 } t ^ { 4 } \left(-135 A^{4} \delta^{2} \cos (5 x)\right.\right. \\
& +240 A^{3} \delta \gamma \sin (4 x)-16 A^{2} \\
& \times(2 \cos (x)-5 \cos 2(x)+3 \cos (3 x)) \\
& \times \beta_{2}^{2}-96 A \gamma \sin (2 x) \\
& \times\left(1+2 A^{2} \delta-\sigma_{2}\right)+3 A^{2} \cos (3 x) \\
& \times\left(48 \delta+63 A^{2} \delta^{2}+32 \gamma^{2}-48 \delta \sigma_{2}\right)
\end{aligned}
$$

$$
\begin{aligned}
& +8 A \beta_{2}\left(-2\left(2+3 A^{2} \delta\right) \cos (x)\right. \\
& +10 \cos (2 x) \\
& +A(3 A \delta(3 \cos (2 x) \\
& +6 \cos (3 x) \\
& -7 \cos (4 x)) \\
& -2 \gamma(\sin (x) \\
& +6 \sin (2 x) \\
& -9 \sin (3 x))) \\
& +2(2 \cos (x) \\
& \left.-5 \cos (2 x)) \sigma_{2}\right) \\
& -2 \cos (x)\left(8+A^{2}\right. \\
& \times\left(3 \delta\left(8+9 A^{2} \delta\right)\right. \\
& \left.+16 \gamma^{2}\right) \\
& \left.+8 \sigma_{2}\left(-2-3 A^{2} \delta+\sigma_{2}\right)\right) \\
& +4 i \beta_{1}(4 \sin (x) \\
& +A(-8 \gamma \cos (2 x) \\
& +3 A \delta(\sin (x) \\
& -3 \sin (3 x))) \\
& -4 \sin (x)(A(-1+2 \cos (x)) \\
& \left.\left.\times \beta_{2}+\sigma_{2}\right)\right) \\
& \left.\left.\times\left(i+\omega \tau_{0}\right)\right)\right) \\
& -\frac{1}{24\left(1+\sigma_{1}\right)^{2}\left(i+\omega \tau_{1}\right)} \\
& \times\left(A h ^ { 2 } t ^ { 3 } \left(-2-3 A^{2} \delta+3 A^{2} \delta \cos (2 x)\right.\right. \\
& \left.-4 A \gamma \sin (x)+2 \sigma_{2}\right) \\
& \times\left(i+\omega \tau_{1}\right)+2 A \beta_{2} \\
& \times\left(i A \alpha \sin (x)\left(1+\sigma_{1}\right)-3 A i \alpha \sin (3 x)\right. \\
& \times\left(1+\sigma_{1}\right)+2 \Omega \cos (x)\left(i+\omega \tau_{1}\right) \\
& \left.-2 i \cos (2 x)\left(1+\Omega+\sigma_{1}-i \Omega \omega \tau_{1}\right)\right) \\
& +4 \beta_{1}\left(i+\omega \tau_{2}\right) \\
& \times\left(-2 A \alpha \cos (2 x)\left(1+\sigma_{1}\right)\right. \\
& \left.\left.+\sin (x)\left(1+\Omega+\sigma_{1}-i \omega \Omega \tau_{1}\right)\right)\right),
\end{aligned}
$$

$$
\begin{aligned}
& \theta_{2}(x, t)=\frac{A h(1+h) t \cos (x)(1+2 A \alpha \sin (x))}{-1+i \omega \tau_{1}} \\
& +\frac{A^{2} h^{2} t^{3} \alpha}{6\left(1+\sigma_{1}\right)\left(i+\omega \tau_{1}\right)} \\
& \times\left(-\frac{1}{2} i(2 A \gamma \cos (x)-6 A \gamma \cos (3 x)\right. \\
& -9 A^{2} \delta \sin (4 x) \\
& +2 A(\sin (x)+\sin (2 x) \\
& -3 \sin (3 x)) \beta_{2} \\
& \left.+2 \sin (2 x)\left(1+3 A^{2} \delta-\sigma_{2}\right)\right) \\
& +\cos (2 x) \beta_{1}\left(i+\omega \tau_{2}\right) \\
& +\frac{1}{8\left(i+\omega \tau_{1}\right)} \\
& \times\left(i A h^{2} t^{2}\right. \\
& \times\left(\frac{a}{1+\sigma_{1}}\right. \\
& \times(-4 \sin (x) \\
& +A(8 v \cos (2 x) \\
& -3 A \delta(\sin (x) \\
& -3 \sin (3 x))) \\
& +4 \sin (x)(A(-1+4 \cos (x)) \\
& \left.\times \beta_{2}+\sigma_{2}\right) \\
& \left.+4 \cos (x) \beta_{1}\left(1-i \omega \tau_{2}\right)\right) \\
& +\frac{i a \omega \delta_{1}}{1+\sigma_{1}}(4 \sin (3 x) \\
& +A(-8 \nu \cos (x) \\
& +3 A \delta(\sin (x) \\
& -3 \sin (3 x))) \\
& -4 \sin (x)(A(-1 \\
& +4 \cos (x)) \\
& \left.\times \beta_{2}+\sigma_{2}\right) \\
& \left.+4 i \cos (x) \beta_{1}\left(i+\omega \tau_{2}\right)\right) \\
& \times \tau_{1}-\frac{1}{i+\omega \tau_{1}} \\
& \times 4 i\left(\left(1+A^{2}\right.\right. \\
& +\alpha^{2} \cos (x)+A \alpha
\end{aligned}
$$

Table 1: The optimal values of $h$ at third-order approximate solutions of (14) when $x=1.0, t=1.0$, for $\delta=1, \sigma_{1}=0.2, \sigma_{2}=$ $0.1, \Omega=0.1$, and $\omega=0.02$.

| $\tau_{1}$ | $\tau_{2}$ | Optimal value of $h$ | Minimum value |  |
| :--- | :---: | :---: | :---: | :---: |
|  |  |  |  |  |
| 1 | 0 | -0.95623 | $1.37312 \times 10^{-3}$ |  |
| 0 | 1 | -0.97543 | $7.9675 \times 10^{-7}$ |  |
|  | $\theta(1,1)$ |  |  |  |
| 1 | 0 | -0.95623 | $1.3242 \times 10^{-2}$ |  |
| 0 | 1 | -0.97543 | $2.1432 \times 10^{-5}$ |  |

$$
\begin{align*}
& \times(-3 A \times \alpha \cos (3 x) \\
& \quad+5 \sin (2 x))))) . \tag{25}
\end{align*}
$$

Now we make calculations for the results obtained by the HAM using the Mathematica software package with the following arbitrary constants:

$$
\begin{gather*}
a=0.5, \quad A=0.001, \quad b=0.5, \quad \alpha=1, \\
\beta_{1}=0.5, \quad \beta_{2}=0.05, \quad \gamma=1, \quad \delta=0.8  \tag{26}\\
\Omega=1, \quad \sigma_{1}=0.2, \quad \sigma_{2}=0.1 .
\end{gather*}
$$

To investigate the influence of $h$ on the convergence of the solution series given by the HAM, we first plot the so-called $h$-curves of $u(1,1)$ and $\theta(1,1)$. According to the $h$-curves, it is easy to discover the valid region of $h$. We used 3 terms in evaluating the approximate solution $u(x, t) \cong \sum_{i=0}^{2} u_{i}(x, t)$ and $\theta(x, t) \cong \sum_{i=0}^{2} \theta_{i}(x, t)$. Note that the solution series contains the auxiliary parameter $h$ which provides us with a simple way to adjust and control the convergence of the solution series. In general, by means of the so-called $h$-curve that is, a curve of a versus $h$. As pointed by Liao [8] and Turkyilmazoglu [25], the valid region of $h$ is a horizontal line segment. Therefore, it is straightforward to choose an appropriate range for $h$ which ensures the convergence of the solution series (Tables 1 and 2). We sketch the $h$-curve of $u(1,1)$ and $\theta(1,1)$ in Figure 1, which shows that the solution series is convergent when $-1.45<h<-0.5$.

## 4. Discussion

In order to gain physical insight, the temperature $T$ and radial displacement $u$ have been discussed by assigning numerical values to the parameter encountered in the problem in which the numerical results are displayed with the graphical illustrations in 2D and 3D formats. The variations are shown in Figures 1-15, with the view of illustrating the theoretical results obtained in the preceding sections; a numerical result is calculated for the homotopy analysis method.

Figures 1 and 2 display the $h$-curve of the third-order approximate solutions (14) when $x=1.0, t=1.0$; it

Table 2: The optimal values of $h$ at 5th-order approximate solutions of (14) when $x=1.0, t=1.0$, for $\delta=1, \sigma_{1}=0.2, \sigma_{2}=0.1, \Omega=$ 0.1 , and $\omega=0.02$.

| $\tau_{1}$ | $\tau_{2}$ | Optimal value of $h$ | Minimum value |  |
| :--- | :---: | :---: | :---: | :---: |
|  |  |  |  |  |
| 1 | 0 | -0.990321 | $1.37312 \times 10^{-7}$ |  |
| 0 | 1 | -0.995339 | $7.9675 \times 10^{-9}$ |  |
|  | $\theta(1,1)$ |  |  |  |
| 1 | 0 | -0.990321 | $1.3242 \times 10^{-4}$ |  |
| 0 | 1 | -0.995339 | $2.1432 \times 10^{-6}$ |  |

is concluded that the displacement $u(1,1)$ and temperature $\theta(1,1)$ increase with increasing the values of $h$ to their maxima and then decrease with the high values of $h$; also, it is shown that $u(1,1)$ is convergent when $-0.9<h<-0.6$ and $\theta(1,1)$ is convergent when $-0.6<h<-0.4$.

Figures $2-7$ show the variations of the radial displacement and temperature with respect to axial $x$, respectively, for different values of the time $t$, rotation $\Omega$, and sensitive parts of the magnetic fields $\sigma_{1}$ and $\sigma_{2}$. In both figures, it is clear that the radial displacement and temperature have a zero value only in a bounded region of space. It is observed from Figure 3 that the displacement $u$ and the temperature $\theta$ start from their maximum values, decrease, and increase periodically with an increase of the coordinate $x$; also, it is obvious that their values take the minimum values and increases with the increasing values of the time $t$. From Figure 4, one can see that $u$ and $\theta$ decrease with an increase of the rotation $\Omega$. It is shown that the components of the displacement $u$ and the temperature $\theta$ start from the minimum values near zero, increase, and then decrease periodically with the coordinate $x$; it is clear also that there is a slight increase with an increase of the sensitive parts of the magnetic field (see, Figures 5 and 6). It is shown that the increasing of the coordinate $x$ sensitive an increasing and decreasing on them periodically due to appearance of the pairs (cos, $\sin$ ) in the initial condition and the approximate solutions; it is also clear that the components begin from their minimum values and increase absolutely with the variation of the time $t$. The variations of the rotation and magnetic field tend to slightly affect the displacement and the temperature.

From Figures 7 and 8 (GL model), it is clear that the displacement component and temperature if the rotation and magnetic field are vanish, take larger values than the corresponding values with the rotation and magnetic field effects.

Figures 9, 10, and 11 show the variations of the displacement and temperature with respect to the time $t$ with LS and GL models; it is shown that the radial displacement and the temperature increase with an increase of $t$ that takes a slight change with the rotation $\Omega$ if $\sigma_{1}=\sigma_{2}=0, \sigma_{1}$ if $\Omega=\sigma_{2}=0$, and $\sigma_{2}$ if $\Omega=\sigma_{1}=0$.

Figures 12 and 13 show clearly the variations of the displacement and temperature in the presence and absence of the rotation and sensitive magnetic field; it is observed that $u$ and $\theta$ in presence of the parameters are smaller than the


Figure 1: The $h$-curve of the third-order approximate solutions of (14) when $x=1.0, t=1.0$; for LS model when $\tau_{2}=0, \tau_{1}=1, \delta=1$, $\sigma_{1}=0.2, \sigma_{2}=0.1, \Omega=0.1$, and $\omega=0.02$.


Figure 2: The $h$-curve of the third-order approximate solutions of (14) when $x=1.0, t=1.0$; for GL model when $\tau_{2}=0.2, \tau_{1}=0.1, \delta=0$, $\sigma_{1}=0.2, \sigma_{2}=0.1, \Omega=0.1$, and $\omega=0.02$.


Figure 3: Variations of the displacement $u$ and the temperature $\theta$ for various values of the $x$-axis and time $t$ when $\tau_{2}=0, \tau_{1}=0.1, \delta=1$, $\sigma_{1}=0.2, \sigma_{2}=0.1, \Omega=0.1$, and $\omega=0.02$.


Figure 4: Variations of the displacement $u$ and the temperature $\theta$ for various values of the $x$-axis and rotation $\Omega$ when $t=0.1, \tau_{2}=0, \tau_{1}=0.1$, $\delta=1, \sigma_{1}=0.2, \sigma_{2}=0.1$, and $\omega=0.02$.


Figure 5: Variations of the displacement $u$ and the temperature $\theta$ for various values of the $x$-axis and magnetic field $\sigma_{1}$ when $t=0.1, \tau_{2}=0$, $\tau_{1}=0.1, \delta=1, \sigma_{2}=0.1, \Omega=0.1$, and $\omega=0.02$.
corresponding values in the absence of $\sigma_{1}, \sigma_{2}$, and $\Omega$, but there is a slight change for LS and GL models.

Finally, Figures 14 and 15 show the variations of the displacement and the temperature with respect to the time $t$ for different values of $\beta=0.5,0.005$ and with and without rotation and magnetic field effects, respectively. It is obvious that the radial displacement and the temperature increase with an increase of $t$; the displacement increases with an increase of $\beta$ parameter, but the temperature is not affected by $\beta$. Is also seen that the radial displacement and the temperature take large values with the rotation and magnetic field effects. Also,
it is concluded that takes large values for LS comparing with those in GL model, vice versa for the temperature.

It is obvious that if the rotation and the sensitive part of the magnetic field are neglected, the approximate solutions obtained by HAM agree with the results obtained by Sweilam and Khader [1], taking into consideration VIM. Finally, it is obvious that the displacement takes large values if there are no rotation, thermal relaxation times, and sensitive part of the magnetic field parameters compared with the corresponding value with the rotation and magnetic fields parameters.


Figure 6: Variations of the displacement $u$ and temperature $\theta$ for various values of the $x$-axis and magnetic field $\sigma_{2}$ when $t=0.1, \tau_{2}=0$, $\tau_{1}=0.1, \delta=1$, and $\omega=0.02$.


Figure 7: Variations of the displacement $u$ and the temperature $\theta$ for various values of the $x$-axis and time $t$ when $\Omega=\sigma_{1}=\sigma_{2}=0, \tau_{2}=0.2$, $\tau_{1}=0.1, \delta=1$, and $\omega=0.02$.

The results indicate that the effect of the rotation and the magnetic field on the radial displacement and the temperature is very pronounced.

## 5. Conclusion

Due to the complicated nature of the governing equations of the magnetothermoelastic, the finished works in this field are unfortunately limited. The method used in this study
provides a quite successful in dealing with such problems. This method gives numerical solutions in the elastic medium without any restrictions on the actual physical quantities that appear in the governing equations of the considered problem. Important phenomena are observed in these computations.
(i) The homotopy analysis method has been successfully applied to obtain the numerical solutions of the nonlinear equation with initial conditions. The reliability of this method and reduction in computations give


Figure 8: Variations of the displacement $u$ and the temperature $\theta$ for various values of the $x$-axis and time $t$ when $\Omega=\sigma_{1}=\sigma_{2}=0, \tau_{2}=0.2$, $\tau_{1}=0.1, \delta=0$, and $\omega=0.02$.


FIGURE 9: Variations of the displacement $u$ and the temperature $\theta$ for various values of the $x$-axis and rotation $\Omega$ when $\sigma_{1}=\sigma_{2}=0, t=0.1$, $\tau_{2}=0.2, \tau_{1}=0.1, \delta=0$, and $\omega=0.02$.
this method a wider applicability. HAM contains a certain auxiliary parameter $h$ which provides us with a simple way to adjust and control the convergence region and rate of convergence of the series solution. It was also demonstrated that the Adomian decomposition method, homotopy perturbation method, and variational iteration method are specialcases of. HAM is clearly a very efficient and powerful technique for finding the numerical solutions of the proposed equation. It therefore provides more realistic series solutions that generally converge very rapidly in real physical problems. HAM provides us with a convenient way of controlling the convergence of approximation series, which is a fundamental qualitative difference in analysis between HAM and
other methods. The illustrative examples suggest that HAM is a powerful method for nonlinear problems in science and engineering. Mathematica has been used for computations in this paper.
(ii) It was found that for large values of time the large and the generalization give numerical results. The case is quite different when we consider small values of rotation and magnetic field. The coupled theory predicts infinite speeds of wave propagation. The solutions obtained in the context of generalized thermoelasticity theory, however, exhibit the behavior of finite speeds of wave propagation.
(iii) By comparing Figures 1-15 for thermoelastic medium with presence and absence of the rotation and


Figure 10: Variations of the displacement $u$ and the temperature $\theta$ for various values of the $x$-axis and magnetic field $\sigma_{1}$ when $\Omega=\sigma_{2}=0$, $t=0.1, \tau_{2}=0.2, \tau_{1}=0.1, \delta=0$, and $\omega=0.02$.


Figure 11: Variations of the displacement $u$ and the temperature $\theta$ for various values of the $x$-axis and magnetic field $\sigma_{2}$ when $\Omega=\sigma_{1}=0$, $t=0.1, \tau_{2}=0.2, \tau_{1}=0.1, \delta=0$, and $\omega=0.02$.
magnetic field, it was found that they have the same behavior in both media. The effect of rotation and sensitive parts of the magnetic field is strongly effective in the displacement and temperature of the propagation of the harmonic waves propagation in nonlinear thermoelasticity.
(iv) The results presented in this paper will be very helpful for researchers concerned with material science, designers of new materials, and low-temperature physicists, as well as for those working on the development of a theory of hyperbolic propagation of hyperbolic thermoelastic. The study of the phenomenon of


FIGURE 12: The displacement as a function of time and the temperature without and with rotation and magnetic field (LS) at $x=100, a=0.5$, $\beta_{1}=0.5, A=0.001, b=0.5, \alpha=1, \beta_{2}=0.05, \gamma=1, \delta=0.8, \tau_{2}=0, \tau_{1}=0.1, \delta=1, \omega=0.02\left(\Omega=\sigma_{1}=\sigma_{2}=0,-\right)$, and $\left(\Omega=0.1, \sigma_{1}=0.2\right.$, $\left.\sigma_{2}=0.1,---\right)$.


FIgURe 13: The displacement as a function of time and the temperature without and with rotation and magnetic field (GL) at $x=100, a=0.5$, $\beta_{1}=0.5, A=0.001, b=0.5, \alpha=1, \beta_{2}=0.05, \gamma=1, \delta=0.8, \tau_{2}=0.2, \tau_{1}=0.1, \delta=0, \omega=0.02\left(\Omega=\sigma_{1}=\sigma_{2}=0,-\right)$, and $\left(\Omega=0.1, \sigma_{1}=0.2\right.$, $\left.\sigma_{2}=0.1,---\right)$.


FIgure 14: The displacement as a function of time and the temperature for two values of $\beta_{1}(\mathrm{LS})$ at $x=100, a=0.5, A=0.001, b=0.5$, $\alpha=1, \beta_{2}=0.05, \gamma=1, \delta=0.8, \Omega=0.1, \sigma_{1}=0.2, \sigma_{2}=0.1, \tau_{2}=0, \tau_{1}=0.1, \delta=1, \omega=0.02,-\beta_{1}=0.5$, and $--\beta_{1}=0.005$.


Figure 15: The displacement as a function of time and the temperature for two values of $\beta_{1}$ (GL) at $x=100, a=0.5, A=0.001, b=0.5$, $\alpha=1, \beta_{2}=0.05, \gamma=1, \delta=0.8, \Omega=0.1, \sigma_{1}=0.2, \sigma_{2}=0.1, \tau_{2}=0.2, \tau_{1}=0.1, \delta=0, \omega=0.02,-\beta_{1}=0.5$, and $--\beta_{1}=0.005$.
rotation and magnetic field is also used to improve the conditions of oil extractions.

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# The Effect of Boundary Slip on the Transient Pulsatile Flow of a Modified Second-Grade Fluid 

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Received 20 May 2013; Accepted 9 August 2013
Academic Editor: Rasajit Bera
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We investigate the effect of boundary slip on the transient pulsatile fluid flow through a vessel with body acceleration. The FahraeusLindqvist effect, expressing the fluid behavior near the wall by the Newtonian fluid while in the core by a non-Newtonian fluid, is also taken into account. To describe the non-Newtonian behavior, we use the modified second-grade fluid model in which the viscosity and the normal stresses are represented in terms of the shear rate. The complete set of equations are then established and formulated in a dimensionless form. For a special case of the material parameter, we derive an analytical solution for the problem, while for the general case, we solve the problem numerically. Our subsequent analytical and numerical results show that the slip parameter has a very significant influence on the velocity profile and also on the convergence rate of the numerical solutions.

## 1. Introduction

In this paper, we study a fluid-structure interaction problem, namely, the effect of boundary slip on the flow of a nonNewtonian fluid through microchannels. This problem has many applications, and in this paper we particularly focus on blood flow in the cardiovascular system.

For the study of blood flow in arteries, two major types of constitutive models have been used. The first type of models is based on the microcontinuum or the structured continuum theories [1-6] in which the balance laws are used to determine the characteristics of blood motion. In the other type of models, blood is considered as a suspension, and its flow is modeled by the non-Newtonian fluid mechanics. Due to the red blood cells (RBC) migration as shown experimentally, blood has been modeled as a two-stage fluid by many researchers [7-9]. The first stage is a peripheral layer which is modeled as a Newtonian viscous fluid, while the other one is a centre core which is modeled as a non-Newtonian fluid. The effect of body acceleration and pulsatile conditions were
taken into account under the same problem by Majhi et al. [7, 10]. Later, Massoudi and Phuoc [11] used the (generalized) second-grade fluid constitutive model to describe the shear thinning and normal stress effect, and the behavior of blood flow near the wall is modeled by the Newtonian fluid model, while the behavior of the blood flow at the core is described by the second-grade fluid model.

In all of the above mentioned models, the so-called noslip boundary condition is used; namely, the velocity of flow relative to the solid is zero on the fluid-solid interface [12]. Although the no-slip condition is supported by many experimental results, the existence of slip of a fluid on the solid surface was also observed by many other researches [13-20]. The Navier slip condition has been used by various researchers to describe boundary slip and is a more general boundary condition, in which the fluid velocity component tangential to the solid surface, relative to the solid surface, is proportional to the shear stress on the fluid-solid interface and the slip length. The surface characteristics constant, slip length, describes the "slipperiness" of the surface. Recently,


Figure 1: The velocity profile in the small artery with radius 0.15 cm under two different slip parameter values: (a) $l_{b}=0$; (b) $l_{b}=2$. In the figure, the 3D graphs show the variation of velocity as a function of time and location, while the 2D graphs show the variation of velocity with time at three radial locations including the artery centre $(r=0)$, the interface of inner-outer layer $(r=0.6)$, and the arterial wall $(r=1)$.
we and many other researchers have investigated various flow problems of Newtonian fluids with the traditional no-slip and the Navier slip boundary conditions [12, 20-30], and it is found that the boundary slip and the slip parameter have significant influence on the flow of Newtonian fluids through microchannels and tubes.

Motivated by the above mentioned work, we extend previous work on slip flows of Newtonian fluids [21, 22] to the case involving both Newtonian and non-Newtonian fluid flow in the flow region. The new feature and contribution of
this work include establishment of the underlying boundary value problem for the problem, the derivation of an exact solution for a special case, and demonstration of the influence of the slip parameter on the flow profile and flow behavior. The rest of the paper is organized as follows. In Section 2, we present the underlying boundary value problem for the problem in dimensionless form. Then in Section 3, we derive an exact solution for a special case. In Section 4, we investigate numerically the effect of the slip parameter for the general case. Finally, a conclusion is given in Section 5.



$$
\begin{aligned}
& -r=0 \\
& -+-r=0.6 \\
& -0-r=1
\end{aligned}
$$

(a)


(b)

Figure 2: The velocity profile in the large artery with radius 0.50 cm under two different slip parameter values: (a) $l_{b}=0$; (b) $l_{b}=2$. In the figure, the 3D graphs show the variation of velocity as a function of time and location, while the 2D graphs show the variation of velocity with time at three radial locations including the artery centre $(r=0)$, the interface of inner-outer layer $(r=0.6)$, and the arterial wall ( $r=1$ ).

## 2. Mathematical Formulation

The flow of a fluid with no thermochemical and electromagnetic effects can be described by the conservation equations of mass and linear momentum; namely,

$$
\begin{gather*}
\frac{\partial \rho}{\partial t}+\operatorname{div}(\rho \mathbf{v})=0  \tag{1}\\
\rho\left(\frac{\partial \mathbf{v}}{\partial t}+\mathbf{v} \cdot \nabla \mathbf{v}\right)=\operatorname{div} \mathbf{T}+\rho \mathbf{b}
\end{gather*}
$$

where $\rho$ is the density of the fluid, $\partial / \partial t$ is the partial derivative with respect to time, $\mathbf{v}$ is the velocity vector, $\mathbf{b}$ is the body force vector, and $\mathbf{T}$ is the stress tensor.

The stress tensor is related to the velocity gradient by the constitutive equations. For a modified (generalized) secondgrade fluid [11, 31, 32], the constitutive equations can be expressed by

$$
\begin{equation*}
\mathbf{T}=-p \mathbf{I}+\Pi^{m / 2}\left(\mu \mathbf{A}_{1}+\alpha_{1} \mathbf{A}_{2}+\alpha_{2} \mathbf{A}_{1}^{2}\right) \tag{2}
\end{equation*}
$$

where $m$ is a material parameter, $\Pi=(1 / 2) \operatorname{tr} \mathbf{A}_{1}^{2}$ is the second invariant of $\mathbf{A}_{1}, p$ is the fluid pressure, $\mu$ is the


$$
\begin{aligned}
& -\times-l_{b}=0 \\
& -o-l_{b}=2 \\
& -+-l_{b}=4
\end{aligned}
$$

$-a-l_{b}=6$

- $l_{b}=8$
(a)

(b)

Figure 3: Diagrams showing the velocity profile on the arterial wall with five different slip parameters $l_{b}$ for two different artery radii (a) $r=0.15 \mathrm{~cm}$; (b) $r=0.5 \mathrm{~cm}$.
coefficient of viscosity, $\alpha_{i}$ are material moduli (the normal stress coefficients), and $\mathbf{A}_{\mathbf{i}}$ are the kinematical tensors given by

$$
\begin{gather*}
\mathbf{A}_{\mathbf{1}}=\mathbf{L}+\mathbf{L}^{T} \\
\mathbf{A}_{\mathbf{2}}=\frac{\partial \mathbf{A}_{\mathbf{1}}}{\partial t}+\left[\operatorname{grad}\left(\mathbf{A}_{\mathbf{1}}\right)\right] \mathbf{v}+\mathbf{A}_{\mathbf{1}} \mathbf{L}+(\mathbf{L})^{T} \mathbf{A}_{\mathbf{1}} \tag{3}
\end{gather*}
$$

in which $L$ is $\operatorname{grad} \mathbf{v}$ and the superscript $T$ refers to matrix transposition.

For the axially symmetrical blood flow through a circular tube of radius $b$, we can assume that $\mathbf{v}=v(r, t) \mathbf{e}_{z}$, where $z$ is the axial direction and $r$ is the radial direction. Under the periodic body acceleration and a unsteady pulsatile pressure gradient $[7,10]$, the momentum equation in the $z$-direction in the cylindrical polar coordinate $(r, \theta, z)$ is

$$
\begin{equation*}
\rho \frac{\partial v}{\partial t}=-\frac{\partial p}{\partial z}+\rho G+\frac{1}{r} \frac{\partial}{\partial r}\left(r T_{r z}\right) . \tag{4}
\end{equation*}
$$

The shear stress $T_{r z}$ for a generalized second-grade fluid can be expressed by

$$
T_{r z}= \begin{cases}\mu_{1}\left|\frac{\partial v_{1}}{\partial r}\right|^{m} \frac{\partial v_{1}}{\partial r} & 0 \leq r \leq a  \tag{5}\\ \mu_{2} \frac{\partial v_{2}}{\partial r} & a \leq r \leq b\end{cases}
$$

The approximate periodic form of the pressure gradient generated by the heart can be described by

$$
\begin{equation*}
-\frac{\partial p}{\partial z}=A_{0}+A_{1} \cos \omega_{p} t \tag{6}
\end{equation*}
$$

where $A_{0}, A_{1}, \omega_{p}=2 \pi f_{p}$, and $f_{p}$ are the constant component of the pressure gradient, the amplitude of the pressure fluctuation (establishing the systolic and diastolic pressures), the circular frequency, and the frequency of pulse rate, respectively.

The body acceleration $G$ can be approximated by

$$
\begin{equation*}
G=A_{g} \cos \left(\omega_{b} t+\phi\right) \tag{7}
\end{equation*}
$$

where $A_{g}$ is the amplitude, $f_{b}=\omega_{b} / 2 \pi$ is the frequency, and $\phi$ is the lead angle of $G$ with respect to the action of the heart.

Substituting (5)-(7) into (4), the blood flow equation for a modified second-grade fluid in the $z$-direction, in the inner and outer core, becomes

$$
\begin{align*}
\rho_{1} \frac{\partial v_{1}}{\partial t}= & A_{0}+A_{1} \cos \omega_{p} t+\rho A_{g} \cos \left(\omega_{b} t+\phi\right) \\
& +\frac{1}{r} \frac{\partial}{\partial r}\left(r \mu_{1}\left|\frac{\partial v_{1}}{\partial r}\right|^{m} \frac{\partial v_{1}}{\partial r}\right), \quad \text { for } 0 \leq r \leq a  \tag{8}\\
\rho_{2} \frac{\partial v_{2}}{\partial t}= & A_{0}+A_{1} \cos \omega_{p} t+\rho A_{g} \cos \left(\omega_{b} t+\phi\right) \\
& +\frac{1}{r} \frac{\partial}{\partial r}\left(r \mu_{2} \frac{\partial v_{2}}{\partial r}\right), \quad \text { for } a \leq r \leq b
\end{align*}
$$

In order to completely define the problem, boundary and initial conditions are required. In this work, the Navier slip condition is applied. That is, on the solid-fluid interface $r=b$, the axial fluid velocity, relative to the solid surface, is proportional to the shear stress on the interface. As the fluid layer near the wall is modeled as a Newtonian fluid in our model, the shear stress on the boundary is related to the shear


FIgure 4: Velocity profiles in arteries with different radii $r$ : (a) $r=0.15 \mathrm{~cm}$; (b) $r=0.5 \mathrm{~cm}$. In the figure, the graphs on the left column correspond to $l_{b}=0$, while the graphs on the right column correspond to $l_{b}=2$.


$$
\begin{array}{ll}
-\times-l_{b}=0 & --l_{b}=6 \\
-o-l_{b}=2 & -l_{b}=8 \\
-+-l_{b}=4 & \\
&
\end{array}
$$



$$
\begin{array}{ll}
-\star-l_{b}=0 & -\square-l_{b}=\epsilon \\
-o-l_{b}=2 & - \\
-+-l_{b}=4 &
\end{array}
$$

(b)

FIGURE 5: Diagrams showing the convergence of numerical solutions for different slip parameters and artery radii: (a) $r=0.15 \mathrm{~cm}$; (b) $r=0.50 \mathrm{~cm}$.


Figure 6: Velocity profiles in arteries with different slip parameters $l_{b}$ and radii $r$ : (a) $r=0.15 \mathrm{~cm}$; (b) $r=0.50 \mathrm{~cm}$. In the Figure, the graphs on the left column correspond to $l_{b}=0$, while the graphs on the right column correspond to $l_{b}=2$.
strain rate by $\sigma_{r z}=\mu_{2}(\partial v / \partial z)$. Thus, the Navier slip condition can be written as

$$
\begin{equation*}
v_{2}(b, t)+l \frac{\partial v_{2}}{\partial t}(b, t)=0 \tag{9}
\end{equation*}
$$

where $l$ is the slip parameter. Moreover, we assume that the slip parameter does not change along the axial direction.

On $r=0$, the symmetry condition is introduced:

On the interface between two different fluids, for continuous and smooth behavior of the velocity and shear stresses, we require

$$
\begin{gather*}
v_{1}(a, t)=v_{2}(a, t) \\
{\left[\mu_{1}\left|\frac{\partial v_{1}}{\partial r}\right|^{m} \frac{\partial v_{1}}{\partial r}\right](a, t)=\left[\mu_{2} \frac{\partial v_{2}}{\partial r}\right](a, t) .} \tag{11}
\end{gather*}
$$

The initial conditions are set to

$$
\begin{equation*}
v_{1}(r, 0)=0=v_{2}(r, 0), \tag{12}
\end{equation*}
$$



Figure 7: Velocity profiles at three arterial locations $\left(r_{1}, r_{2}, r_{3}\right)$ : for $m=-1 / 4$ and under different slip parameters $l_{b}$ and artery radii (a) $r=0.15 \mathrm{~cm}$; (b) $r=0.50 \mathrm{~cm}$. In the Figure, the graphs on the left column correspond to $l_{b}=0$, while the graphs on the right column correspond to $l_{b}=2$.
which is essential for the numerical scheme adopted to estimate the time at which the pulsatile steady state is achieved.

$$
\rho^{*}=\frac{\rho_{1}}{\rho_{2}}, \quad \mu^{*}=\frac{\mu_{2}}{\bar{\mu}}
$$

To simplify the equations, we introduce the following nondimensional variables and parameters:

$$
C_{1}=\frac{A_{0} b^{2}}{\bar{\mu} u_{0}}, \quad C_{2}=\rho_{1} A_{g} \frac{b^{2}}{\bar{\mu} u_{0}}=\frac{\rho_{1} A_{g}}{A_{0}} B_{1}
$$

$$
\begin{align*}
& \bar{r}=\frac{r}{b}, \quad \bar{v}=\frac{v}{v_{0}}, \quad \bar{t}=\frac{\omega_{p}}{2 \pi} t, \quad u_{0}=\frac{A_{0} b^{2}}{\mu_{2}}, \\
& e=\frac{A_{1}}{A_{0}}, \quad \omega_{r}=\frac{\omega_{b}}{\omega_{p}}, \quad r_{0}=\frac{a}{b}, \quad \overline{m u}=\mu\left(\frac{u_{0}}{b}\right)^{m},  \tag{13}\\
& \alpha=\frac{\rho_{1} \omega_{p} b^{2}}{2 \pi \bar{\mu}}, \quad \gamma=\frac{\rho_{2} \omega_{p} b^{2}}{2 \pi \bar{\mu} \mu^{*}}=\frac{\rho_{2} \omega_{p} b^{2} \rho_{1}}{2 \pi \bar{\mu} \mu^{*} \rho_{1}}=\alpha \frac{\rho^{*}}{\mu^{*}}, \\
& \widehat{C}_{1}=\frac{A_{0} b^{2}}{\bar{\mu} u_{0} \mu^{*}}=\frac{C_{1}}{\mu^{*}}, \quad \widehat{C}_{2}=\frac{\rho_{2} A_{g} b^{2} \rho_{1}}{\bar{\mu} u_{0} \mu^{*} \rho_{1}}=C_{2} \frac{\rho^{*}}{\mu^{*}} .
\end{align*}
$$


(a)


$$
\begin{aligned}
& -x-l_{b}=0 \\
& -o-l_{b}=2 \\
& --l_{b}=4
\end{aligned}
$$

$$
\begin{aligned}
& -l_{b}=6 \\
& -l_{b}=8
\end{aligned}
$$

(b)

Figure 8: Diagrams showing the convergence of numerical results of the fluid velocity on the wall to the steady state pulsatile velocity field under various slip parameters $l_{b}$ for two different artery radii: (a) $r=0.15 \mathrm{~cm}$; (b) $r=0.50 \mathrm{~cm}$.

In terms of the nondimensional variables and parameters, (8)-(12) can be written in the form of

$$
\begin{align*}
\alpha \frac{\partial \bar{v}_{1}}{\partial \bar{t}}= & C_{1}(1+e \cos 2 \pi \bar{t})+C_{2} \cos \left(2 \pi \omega_{r} \bar{t}+\phi\right) \\
& +\frac{1}{\bar{r}} \frac{\partial}{\partial \bar{r}}\left[\bar{r}\left|\frac{\partial \bar{v}_{1}}{\partial \bar{r}}\right|^{m} \frac{\partial \bar{v}_{1}}{\partial \bar{r}}\right], \quad \text { for } 0 \leq r \leq r_{0} \\
\gamma \frac{\partial \bar{v}_{2}}{\partial \bar{t}}= & \bar{C}_{1}(1+e \cos 2 \pi \bar{t})+\bar{C}_{2} \cos \left(2 \pi \omega_{r} \bar{t}+\phi\right) \\
& +\frac{1}{\bar{r}} \frac{\partial}{\partial \bar{r}}\left[\bar{r} \frac{\partial \bar{v}_{2}}{\partial \bar{r}}\right], \quad \text { for } r_{0} \leq r \leq 1 \tag{14}
\end{align*}
$$

The boundary conditions and initial conditions, in dimensionless form, can be expressed by

$$
\begin{gather*}
\frac{\partial \bar{v}_{1}}{\partial \bar{r}}(0, \bar{t})=0  \tag{15}\\
b \bar{v}_{2}(1, t)+l \frac{\partial \bar{v}_{2}}{\partial \bar{r}}(1, \bar{t})=0  \tag{16}\\
\bar{v}_{1}\left(r_{0}, \bar{t}\right)=\bar{v}_{2}\left(r_{0}, \bar{t}\right)  \tag{17}\\
{\left[\left|\frac{\partial \bar{v}_{1}}{\partial \bar{r}}\right|^{m} \frac{\partial \bar{v}_{1}}{\partial \bar{r}}\right]\left(r_{0}, \bar{t}\right)=\left[\mu^{*} \frac{\partial \bar{v}_{2}}{\partial \bar{r}}\right]\left(r_{0}, \bar{t}\right),}  \tag{18}\\
\bar{v}_{1}(r, 0)=0=\bar{v}_{2}(r, 0) \tag{19}
\end{gather*}
$$

## 3. Analytical Solution

For $m=0$, the model reduces to the linear model with different viscosity in the peripheral layer and the centre core. In this case, (14) have the same form:

$$
\begin{align*}
L(v) & =\beta \frac{\partial v}{\partial t}-\frac{1}{r} \frac{\partial v}{\partial r}-\frac{\partial^{2} v}{\partial r^{2}}  \tag{20}\\
& =B_{1}(1+e \cos (2 \pi t))+B_{2} \cos \left(2 \pi \omega_{r} t+\phi\right)
\end{align*}
$$

By the superposition principle, if $v_{0}, v_{1}$, and $v_{2}$ are the solution of $L(v)=f(t)$, respectively, for $f(t)=$ $B_{1} e^{0 t i}, B_{1} a e^{2 \pi t i}$, and $B_{2} e^{\left(2 \pi \omega_{r} t+\phi\right) i}$, then the complete solution of (20) is $v=\sum_{n=0}^{2} \operatorname{Re}\left(v_{n}\right)$.

To determine $v_{n}$, we solve

$$
\begin{equation*}
\beta \frac{\partial v_{n}}{\partial t}=D_{n} e^{g_{n}(t) i}+\frac{1}{r} \frac{\partial v_{n}}{\partial r}+\frac{\partial^{2} v_{n}}{\partial r^{2}} \tag{21}
\end{equation*}
$$

where $g_{0}(t)=0, g_{1}(t)=2 \pi t, g_{2}(t)=2 \pi \omega_{r} t+\phi, D_{0}=B_{1}$, $D_{1}=a B_{1}$, and $D_{2}=B_{2}$. As (21) admits solutions of the form $v_{n}=f_{n}(r) e^{g_{n}(t) i}$, we have from (21) that

$$
\begin{align*}
\beta g_{n}^{\prime} & (t) f_{n}(r) e^{g_{n}(t) i} i \\
& =D_{n} e^{g_{n}(t) i}+\frac{1}{r} f_{n}^{\prime}(r) e^{g_{n}(t) i}+f_{n}^{\prime \prime}(r) e^{g_{n}(t) i} \tag{22}
\end{align*}
$$

Dividing by $e^{g_{n}(t) i}$ on both sides of (22), we obtain

$$
\begin{equation*}
\beta g_{n}^{\prime}(t) f_{n}(r) i=D_{n}+\frac{1}{r} f_{n}^{\prime}(r)+f_{n}^{\prime \prime}(r) \tag{23}
\end{equation*}
$$

For $n=0$, we get

$$
\begin{equation*}
f_{0}^{\prime \prime}(r)+\frac{1}{r} f_{0}^{\prime}(r)=-B_{1} \tag{24}
\end{equation*}
$$

which has the general solution: $f_{0}(r)=\left(c_{1}+c_{2} \ln r\right)-\left(B_{1} / 4\right) r^{2}$.
For $n=1$, we have

$$
\begin{equation*}
f_{1}^{\prime \prime}(r)+\frac{1}{r} f_{1}^{\prime}(r)-2 \pi \beta i f_{1}(r)=-e B_{1} \tag{25}
\end{equation*}
$$

Let $\bar{\beta}_{1}^{2}=-2 \pi \beta i$; then,

$$
\begin{equation*}
\frac{1}{\bar{\beta}_{1}^{2}} f_{1}^{\prime \prime}(r)+\frac{1}{\bar{\beta}_{1}^{2} r} f_{1}^{\prime}(r)+f_{1}(r)=-\frac{e B_{1}}{\bar{\beta}_{1}^{2}} \tag{26}
\end{equation*}
$$

Let $\widehat{r}=\bar{\beta}_{1} r$; we have

$$
\begin{equation*}
\widehat{r}^{2} f_{1}^{\prime \prime}(\widehat{r})+\widehat{r} f_{1}^{\prime}(\widehat{r})+\widehat{r}^{2} f_{1}(\widehat{r})=-\frac{e B_{1}}{\bar{\beta}_{1}^{2}} \widehat{r}^{2} . \tag{27}
\end{equation*}
$$

The general solution of (27) is

$$
\begin{equation*}
f_{1}(r)=d_{1} J_{0}\left(\bar{\beta}_{1} r\right)+e_{1} Y_{0}\left(\bar{\beta}_{1} r\right)-\frac{e B_{1}}{2 \pi \beta} i \tag{28}
\end{equation*}
$$

where $d_{1}$ and $e_{1}$ are integration constants and $J_{0}$ and $Y_{0}$ denote the zero-order Bessel functions of the first kind and the second kind, respectively.

Similarly, for $n=2$, we have

$$
\begin{equation*}
f_{2}^{\prime \prime}(r)+\frac{1}{r} f_{2}^{\prime}(r)-2 \beta \pi \omega_{r} f_{2}(r) i=-B_{2} \tag{29}
\end{equation*}
$$

and the general solution is

$$
\begin{equation*}
f_{2}=d_{2} J_{0}\left(\bar{\beta}_{2} r\right)+e_{2} Y_{0}\left(\bar{\beta}_{2} r\right)-\frac{B_{2}}{2 \beta \omega_{r} \pi} i \tag{30}
\end{equation*}
$$

where $\bar{\beta}_{2}^{2}=-2 \pi \beta \omega_{r} i$.
Because the boundness of $v_{1}, v_{2}, c_{2}, e_{1}$, and $e_{2}$ are set to zero, hence, from (14) and the solutions for (20), we have

$$
\begin{aligned}
& \bar{v}_{1}=\operatorname{Re}\{ c_{1}-\frac{C_{1}}{4} \bar{r}^{2}+\left[d_{1} J_{0}\left(\beta_{1} \bar{r}\right)-\frac{e C_{1}}{2 \pi \alpha} i\right] e^{2 \pi \bar{t} i} \\
&\left.+\left[d_{2} J_{0}\left(\beta_{2} \bar{r}\right)-\frac{C_{2}}{2 \pi \omega_{r} \alpha} i\right] e^{\left(2 \pi \omega_{r} \bar{t}+\phi\right) i}\right\} \\
& \bar{v}_{2}=\operatorname{Re}\left\{\widehat{c}_{1}+\widehat{c}_{2} \ln \bar{r}-\frac{\bar{C}_{1}}{4} \bar{r}^{2}\right. \\
&+\left[\widehat{d}_{1} J_{0}\left(\widehat{\beta}_{1} \bar{r}\right)+\widehat{e}_{1} Y_{0}\left(\widehat{\beta}_{1} \bar{r}\right)-\frac{e \widehat{C}_{1}}{2 \pi \gamma} i\right] e^{2 \pi \bar{t} i} \\
&+\left[\widehat{d}_{2} J_{0}\left(\widehat{\beta}_{2} \bar{r}\right)+\widehat{e}_{2} Y_{0}\left(\widehat{\beta}_{2} \bar{r}\right)-\frac{\bar{c}_{2} i}{2 \pi \omega_{r} \gamma}\right] \\
&\left.\times e^{\left(2 \pi \omega_{r} \bar{t}+\phi\right) i}\right\}
\end{aligned}
$$

where $\widehat{\beta}_{1}^{2}=-2 \pi \gamma i, \widehat{\beta}_{2}^{2}=-2 \pi \omega_{r} \gamma i, \beta_{1}^{2}=-2 \pi \alpha i$, and $\beta_{2}^{2}=$ $-2 \pi \omega_{r} \alpha i$.

As $d J_{0}(x) / d x=-J_{1}(x)$ and $d Y_{0}(x) / d x=-Y_{0}(x)$, we have

$$
\begin{align*}
& \frac{\partial \bar{v}_{1}}{\partial \bar{r}}=\operatorname{Re}( -\frac{C_{1}}{2} \bar{r}^{2}-d_{1} \beta_{1} J_{1}\left(\beta_{1} \bar{r}\right) e^{2 \pi \bar{t} i} \\
&\left.-d_{2} \beta_{2} J_{1}\left(\beta_{2} \bar{r}\right) e^{\left(2 \pi \omega_{r} \bar{t}+\phi\right) i}\right) \\
& \frac{\partial \bar{v}_{2}}{\partial \bar{r}}=\operatorname{Re}\left(\widehat{c}_{2} \frac{1}{\bar{r}}-\frac{\bar{C}_{1}}{2} \bar{r}\right. \\
&+ {\left[-\widehat{d}_{1} \widehat{\beta}_{1} J_{1}\left(\widehat{\beta}_{1} \bar{r}\right)-\widehat{e}_{1} \widehat{\beta}_{1} Y_{1}\left(\widehat{\beta}_{1} \bar{r}\right)\right] e^{2 \pi \bar{t} i} } \\
&+ {\left.\left[-\widehat{d}_{2} \widehat{\beta}_{2} J_{1}\left(\widehat{\beta}_{2} \bar{r}\right)-\widehat{e}_{2} \widehat{\beta}_{2} Y_{1}\left(\widehat{\beta}_{2} \bar{r}\right)\right] e^{\left(2 \pi \omega_{r} \bar{t}+\phi\right) i}\right) } \tag{32}
\end{align*}
$$

Obviously, $v_{1}$ satisfies the boundary condition (15) automatically. We now consider the boundary condition (16); namely,

$$
\begin{align*}
& \operatorname{Re}\left[\left(b \widehat{c}_{1}+l \widehat{c}_{2}-\left(l+\frac{b}{2}\right) \frac{\bar{C}_{1}}{2}\right)\right. \\
& \quad+\left(\left(b J_{0}\left(\widehat{\beta}_{1}\right)-l \widehat{\beta}_{1} J_{1}\left(\widehat{\beta}_{1}\right)\right) \widehat{d}_{1}\right. \\
& \left.\quad+\left(b Y_{0}\left(\widehat{\beta}_{1}\right)-l \widehat{\beta}_{1} Y_{1}\left(\widehat{\beta}_{1}\right)\right) \widehat{e}_{1}-\frac{e b \bar{C}_{1} i}{2 \pi \gamma}\right) e^{2 \pi \bar{t} i}  \tag{33}\\
& \quad+\left(\widehat{d}_{2}\left(b J_{0}\left(\widehat{\beta}_{2}\right)-l \widehat{\beta}_{2} J_{1}\left(\widehat{\beta}_{2}\right)\right)\right. \\
& \quad+\widehat{e}_{2}\left(b Y_{0}\left(\widehat{\beta}_{2}\right)-l \widehat{\beta}_{2} Y_{1}\left(\widehat{\beta}_{2}\right)\right) \\
& \left.\left.\quad-\frac{b \bar{C}_{2}}{2 \pi \omega_{r} \gamma} i\right) e^{\left(2 \pi \omega_{r} \bar{t}+\phi\right) i}\right]=0
\end{align*}
$$

Further, from boundary conditions (17) and (18), we have

$$
\begin{aligned}
& \operatorname{Re}\left[\left(c_{1}-\widehat{c}_{1}-\widehat{c}_{2} \ln r_{0}-\left(C_{1}-\widehat{C}_{1}\right) \frac{r_{0}^{2}}{4}\right)\right. \\
& \quad+\left(d_{1} J_{0}\left(\beta_{1} r_{0}\right)-\widehat{d}_{1} J_{0}\left(\widehat{\beta}_{1} r_{0}\right)-\widehat{e}_{1} Y_{0}\left(\widehat{\beta}_{1} r_{0}\right)\right. \\
& \left.\quad-\left(\gamma C_{1}-\alpha \bar{C}_{1}\right) \frac{e i}{2 \pi \alpha \gamma}\right) e^{2 \pi \overline{t i}} \\
& \quad+\left(d_{2} J_{0}\left(\beta_{2} r_{0}\right)-\widehat{d}_{2} J_{0}\left(\widehat{\beta}_{2} r_{0}\right)\right. \\
& \left.\quad-\widehat{e}_{2} Y_{0}\left(\widehat{\beta}_{2} r_{0}\right)-\left(\gamma C_{2}-\alpha \bar{C}_{2}\right) \frac{i}{2 \pi \omega_{r} \gamma \alpha}\right) \\
& \left.\quad \times e^{\left(2 \pi \omega_{r} \bar{t}+\phi\right) i}\right]=0,
\end{aligned}
$$

$$
\begin{align*}
\operatorname{Re}[( & \left.\left(\mu^{*} \bar{C}_{1}-C_{1}\right) \frac{r_{0}}{2}-\mu^{*} \frac{\widehat{\underline{c}}_{2}}{r_{0}}\right) \\
& +\left(-d_{1} \beta_{1} J_{1}\left(\beta_{1} r_{0}\right)+\widehat{d}_{1} \mu^{*} \widehat{\beta}_{1} J_{1}\left(\widehat{\beta}_{1} r_{0}\right)\right. \\
& \left.\quad+\widehat{e}_{1} \mu^{*} \widehat{\beta}_{1} Y_{1}\left(\widehat{\beta}_{1} r_{0}\right)\right) e^{2 \pi \bar{i} i} \\
& +\left(-d_{2} \beta_{2} J_{1}\left(\beta_{2} r_{0}\right)+\widehat{d}_{2} \mu^{*} \widehat{\widehat{\beta}}_{2} J_{1}\left(\widehat{\beta}_{2} r_{0}\right)\right. \\
& \left.\quad+\widehat{e}_{2} \mu^{*} \widehat{\beta}_{2} Y_{1}\left(\widehat{\beta}_{2} r_{0}\right)\right) \\
& \left.\times e^{\left(2 \pi \omega_{r} \bar{t}+\phi\right) i}\right]=0 . \tag{34}
\end{align*}
$$

As (33)-(34) must be satisfied for any instant of time $t$, we require that the constant terms and the coefficients of the exponential terms all vanish; namely,

$$
\begin{gather*}
b \widehat{c}_{1}+l \widehat{\varepsilon}_{2}-\left(l+\frac{b}{2}\right) \frac{\bar{C}_{1}}{2}=0, \\
c_{1}-\widehat{c}_{1}-\widehat{c}_{2} \ln r_{0}-\left(C_{1}-\bar{C}_{1}\right) \frac{r_{0}^{2}}{4}=0, \\
\left(\mu^{*} \bar{C}_{1}-C_{1}\right) \frac{r_{0}}{2}-\mu^{*} \frac{\widehat{c}_{2}}{r_{0}}=0, \\
\widehat{d}_{1}\left(b J_{0}\left(\widehat{\beta}_{1}\right)-l \widehat{\beta}_{1} J_{1}\left(\widehat{\beta}_{1}\right)\right)+\widehat{e}_{1}\left(b Y_{0}\left(\widehat{\beta}_{1}\right)-l \widehat{\beta}_{1} Y_{1}\left(\widehat{\beta}_{1}\right)\right) \\
-\frac{e b \bar{C}_{1}}{2 \pi \gamma} i=0, \\
d_{1} J_{0}\left(\beta_{1} r_{0}\right)-\widehat{d}_{1} J_{0}\left(\widehat{\beta}_{1} r_{0}\right)-\widehat{e}_{1} Y_{0}\left(\widehat{\beta}_{1} r_{0}\right)-\frac{e C_{1}}{2 \pi \alpha} i \\
\quad+\frac{e \bar{C}_{1}}{2 \pi \gamma} i=0, \\
-d_{1} \beta_{1} J_{1}\left(\beta_{1} r_{0}\right)+\widehat{d}_{1} \mu^{*} \widehat{\widehat{\beta}}_{1} J_{1}\left(\widehat{\beta}_{1} r_{0}\right)+\widehat{e}_{1} \mu^{*} \widehat{\beta}_{1} Y_{1}\left(\widehat{\beta}_{1} r_{0}\right) \\
=0, \\
\widehat{d}_{2}\left(b J_{0}\left(\widehat{\beta}_{2}\right)-l \widehat{\beta}_{2} J_{1}\left(\widehat{\beta}_{2}\right)\right)+\widehat{e}_{2}\left(b Y_{0}\left(\widehat{\beta}_{2}\right)-l \widehat{\beta}_{2} Y_{1}\left(\widehat{\beta}_{2}\right)\right) \\
-\frac{b \bar{C}_{2}}{2 \pi \omega_{r} \gamma} i=0, \\
d_{2} J_{0}\left(\beta_{2} r_{0}\right)-\widehat{d}_{2} J_{0}\left(\widehat{\beta}_{2} r_{0}\right)-\widehat{e}_{2} Y_{0}\left(\widehat{\beta}_{2} r_{0}\right)-\frac{C_{1}}{2 \pi \omega_{r} \alpha} i \\
\quad+\frac{\bar{C}_{2}}{2 \pi \omega_{r} \gamma} i=0, \\
-d_{2} \beta_{2} J_{1}\left(\beta_{2} r_{0}\right)+\widehat{d}_{2} \mu^{*} \widehat{\beta}_{2} J_{1}\left(\widehat{\beta}_{2} r_{0}\right)+\widehat{e}_{2} \mu^{*} \widehat{\beta_{2}} Y_{1}\left(\widehat{\beta}_{2} r_{0}\right) \\
=0 . \tag{35}
\end{gather*}
$$

Solving the above system of equations yields

$$
\begin{aligned}
& c_{1}=\left(\ln r_{0}-\frac{l}{b}\right)\left(\left(\mu^{*} \bar{C}_{1}-C_{1}\right) \frac{r_{0}^{2}}{2 \mu^{*}}\right) \\
& +\left(\frac{l}{b}+\frac{1-r_{0}^{2}}{2}\right) \frac{\bar{C}_{1}}{2}+C_{1} \frac{r_{0}^{2}}{4}, \\
& \widehat{c}_{1}=-\frac{l}{b}\left(\left(\mu^{*} \bar{C}_{1}-C_{1}\right) \frac{r_{0}^{2}}{2 \mu^{*}}\right)+\left(\frac{l}{b}+\frac{1}{2}\right) \frac{\bar{C}_{1}}{2}, \\
& \widehat{c}_{2}=\left(\mu^{*} \bar{C}_{1}-C_{1}\right) \frac{r_{0}^{2}}{2 \mu^{*}}, \\
& d_{1}=\mu^{*}\left[\left(J_{1}\left(\widehat{\beta}_{1} r_{0}\right) Y_{0}\left(\widehat{\beta}_{1} r_{0}\right)-J_{0}\left(\widehat{\beta}_{1} r_{0}\right) Y_{1}\left(\widehat{\beta}_{1} r_{0}\right)\right)\right. \\
& \times \frac{e b \widehat{\beta}_{1} \bar{C}_{1} i}{2 \pi \gamma}+\left(\gamma C_{1}-\alpha \bar{C}_{1}\right) \\
& \times\left[J_{1}\left(\widehat{\beta}_{1} r_{0}\right)\left(b Y_{0}\left(\widehat{\beta}_{1}\right)-l \widehat{\beta_{1}} Y_{1}\left(\widehat{\beta}_{1}\right)\right)\right. \\
& \left.-Y_{1}\left(\widehat{\beta}_{1} r_{0}\right)\left(b J_{0}\left(\widehat{\beta}_{1}\right)-l \widehat{\beta}_{1} J_{1}\left(\widehat{\beta}_{1}\right)\right)\right] \\
& \left.\times \frac{\widehat{\beta}_{1} e i}{2 \pi \gamma \alpha}\right] /\left(b J_{0}\left(\widehat{\beta}_{1}\right)-l \widehat{\beta}_{1} J_{1}\left(\widehat{\beta}_{1}\right)\right) \\
& \times\left(\beta_{1} J_{1}\left(\beta_{1} r_{0}\right) Y_{0}\left(\widehat{\beta}_{1} r_{0}\right)-\mu^{*} \widehat{\beta}_{1} Y_{1}\left(\widehat{\beta}_{1} r_{0}\right) J_{0}\left(\beta_{1} r_{0}\right)\right) \\
& +\left(b Y_{0}\left(\widehat{\beta}_{1}\right)-l \widehat{\beta}_{1} Y_{1}\left(\widehat{\beta}_{1}\right)\right) \\
& \times\left(\mu^{*} \widehat{\beta}_{1} J_{0}\left(\beta_{1} r_{0}\right) J_{1}\left(\widehat{\beta}_{1} r_{0}\right)-\beta_{1} J_{1}\left(\beta_{1} r_{0}\right) J_{0}\left(\widehat{\beta}_{1} r_{0}\right)\right), \\
& \widehat{d}_{1}=\left[\left(\beta_{1} J_{1}\left(\beta_{1} r_{0}\right) Y_{0}\left(\widehat{\beta}_{1} r_{0}\right)\right.\right. \\
& \left.-\mu^{*} \widehat{\beta}_{1} Y_{1}\left(\widehat{\beta}_{1} r_{0}\right) J_{0}\left(\beta_{1} r_{0}\right)\right) \\
& \times \frac{e b \bar{C}_{1}}{2 \pi \gamma} i+\frac{\left(\gamma C_{1}-\alpha \bar{C}_{1}\right)}{2 \pi \gamma \alpha} \\
& \times\left(b Y_{0}\left(\widehat{\beta}_{1}\right)-l \widehat{\beta}_{1} Y_{1}\left(\widehat{\beta}_{1}\right)\right) \\
& \left.\times e \beta_{1} J_{1}\left(\beta_{1} r_{0}\right) i\right] /\left(b J_{0}\left(\widehat{\beta}_{1}\right)-l \widehat{\beta}_{1} J_{1}\left(\widehat{\beta}_{1}\right)\right) \\
& \times\left(\beta_{1} J_{1}\left(\beta_{1} r_{0}\right) Y_{0}\left(\widehat{\beta}_{1} r_{0}\right)\right. \\
& \left.-\mu^{*} \widehat{\beta}_{1} Y_{1}\left(\widehat{\beta}_{1} r_{0}\right) J_{0}\left(\beta_{1} r_{0}\right)\right) \\
& +\left(b Y_{0}\left(\widehat{\beta}_{1}\right)-l \widehat{\beta}_{1} Y_{1}\left(\widehat{\beta}_{1}\right)\right) \\
& \times\left(\mu^{*} \widehat{\beta}_{1} J_{0}\left(\beta_{1} r_{0}\right) J_{1}\left(\widehat{\beta}_{1} r_{0}\right)\right. \\
& \left.-\beta_{1} J_{1}\left(\beta_{1} r_{0}\right) J_{0}\left(\widehat{\beta}_{1} r_{0}\right)\right), \\
& \widehat{e}_{1}=\left[\left(\mu^{*} \widehat{\beta}_{1} J_{0}\left(\beta_{1} r_{0}\right) J_{1}\left(\widehat{\beta}_{1} r_{0}\right)\right.\right. \\
& \left.-\beta_{1} J_{0}\left(\widehat{\beta}_{1} r_{0}\right) J_{1}\left(\beta_{1} r_{0}\right)\right)
\end{aligned}
$$

$$
\begin{aligned}
& \times \frac{e b \overline{\mathrm{C}}_{1} i}{2 \pi \gamma}-\left(b J_{0}\left(\widehat{\beta}_{1}\right)-l \widehat{\beta}_{1} J_{1}\left(\widehat{\beta}_{1}\right)\right) \\
& \times\left(\gamma C_{1}-\alpha \bar{C}_{1}\right) \\
& \left.\times J_{1}\left(\beta_{1} r_{0}\right) \frac{\beta_{1} e i}{2 \pi \alpha \gamma}\right] /\left(b J_{0}\left(\widehat{\beta}_{1}\right)-l \widehat{\beta}_{1} J_{1}\left(\widehat{\beta}_{1}\right)\right) \\
& \times\left(\beta_{1} J_{1}\left(\beta_{1} r_{0}\right) Y_{0}\left(\widehat{\beta}_{1} r_{0}\right)\right. \\
& \left.-\mu^{*} \widehat{\beta}_{1} Y_{1}\left(\widehat{\beta}_{1} r_{0}\right) J_{0}\left(\beta_{1} r_{0}\right)\right) \\
& +\left(b Y_{0}\left(\widehat{\beta}_{1}\right)-l \widehat{\beta}_{1} Y_{1}\left(\widehat{\beta}_{1}\right)\right) \\
& \times\left(\mu^{*} \widehat{\beta}_{1} J_{0}\left(\beta_{1} r_{0}\right) J_{1}\left(\widehat{\beta}_{1} r_{0}\right)-\beta_{1} J_{1}\left(\beta_{1} r_{0}\right) J_{0}\left(\widehat{\beta}_{1} r_{0}\right)\right), \\
& d_{2}=\mu^{*}\left[\left(J_{1}\left(\widehat{\beta}_{2} r_{0}\right) Y_{0}\left(\widehat{\beta}_{2} r_{0}\right)-J_{0}\left(\widehat{\beta}_{2} r_{0}\right) Y_{1}\left(\widehat{\beta}_{2} r_{0}\right)\right)\right. \\
& \times \frac{b \widehat{\beta}_{2} \bar{C}_{1} i}{2 \pi \omega_{r} \gamma}+\left(\gamma C_{2}-\alpha \bar{C}_{2}\right) \\
& \times\left[J_{1}\left(\widehat{\beta}_{2} r_{0}\right)\left(b Y_{0}\left(\widehat{\beta}_{2}\right)-l \widehat{\beta_{2}} Y_{1}\left(\widehat{\beta}_{2}\right)\right)\right. \\
& \left.-Y_{1}\left(\widehat{\beta}_{2} r_{0}\right)\left(b J_{0}\left(\widehat{\beta}_{2}\right)-l \widehat{\beta}_{2} J_{1}\left(\widehat{\beta}_{2}\right)\right)\right] \\
& \left.\times \frac{\widehat{\beta}_{2} i}{2 \pi \omega_{r} \gamma \alpha}\right] /\left(b J_{0}\left(\widehat{\beta}_{2}\right)-l \widehat{\beta}_{2} J_{1}\left(\widehat{\beta}_{2}\right)\right) \\
& \times\left(\beta_{2} J_{1}\left(\beta_{2} r_{0}\right) Y_{0}\left(\widehat{\beta}_{2} r_{0}\right)\right. \\
& \left.-\mu^{*} \widehat{\beta}_{2} Y_{1}\left(\widehat{\beta}_{2} r_{0}\right) J_{0}\left(\beta_{2} r_{0}\right)\right) \\
& +\left(b Y_{0}\left(\widehat{\beta}_{2}\right)-l \widehat{\beta}_{2} Y_{1}\left(\widehat{\beta}_{2}\right)\right) \\
& \times\left(\mu^{*} \widehat{\beta}_{2} J_{0}\left(\beta_{2} r_{0}\right) J_{1}\left(\widehat{\beta}_{2} r_{0}\right)\right. \\
& \left.-\beta_{2} J_{1}\left(\beta_{2} r_{0}\right) J_{0}\left(\widehat{\beta}_{2} r_{0}\right)\right) \text {, } \\
& \widehat{d}_{2}=\left[\left(\beta_{2} I_{1}\left(\beta_{2} r_{0}\right) Y_{0}\left(\widehat{\beta}_{2} r_{0}\right)\right.\right. \\
& \left.-\mu^{*} \widehat{\beta}_{2} Y_{1}\left(\widehat{\beta}_{2} r_{0}\right) J_{0}\left(\beta_{2} r_{0}\right)\right) \\
& \times \frac{b \bar{C}_{1}}{2 \pi \omega_{r} \gamma} i+\frac{\left(\gamma C_{1}-\alpha \bar{C}_{1}\right)}{2 \pi \gamma \alpha \omega_{r}} \\
& \times\left(b Y_{0}\left(\widehat{\beta}_{2}\right)-l \widehat{\beta}_{2} Y_{1}\left(\widehat{\beta}_{2}\right)\right) \\
& \left.\times \beta_{2} J_{1}\left(\beta_{2} r_{0}\right) i\right] /\left(b J_{0}\left(\widehat{\beta}_{2}\right)-l \widehat{\beta}_{2} J_{1}\left(\widehat{\beta}_{2}\right)\right) \\
& \times\left(\beta_{2} J_{1}\left(\beta_{2} r_{0}\right) Y_{0}\left(\widehat{\beta}_{2} r_{0}\right)\right. \\
& \left.-\mu^{*} \widehat{\beta}_{2} Y_{1}\left(\widehat{\beta}_{2} r_{0}\right) J_{0}\left(\beta_{2} r_{0}\right)\right) \\
& +\left(b Y_{0}\left(\widehat{\beta}_{2}\right)-l \widehat{\beta}_{2} Y_{1}\left(\widehat{\beta}_{2}\right)\right) \\
& \times\left(\mu^{*} \widehat{\beta}_{2} J_{0}\left(\beta_{2} r_{0}\right) J_{1}\left(\widehat{\beta}_{2} r_{0}\right)\right. \\
& \left.-\beta_{2} J_{1}\left(\beta_{2} r_{0}\right) J_{0}\left(\widehat{\beta}_{2} r_{0}\right)\right) \text {, }
\end{aligned}
$$

$$
\begin{align*}
\widehat{e}_{2}= & {\left[\left(\mu^{*} \widehat{\beta}_{2} J_{0}\left(\beta_{2} r_{0}\right) J_{1}\left(\widehat{\beta}_{2} r_{0}\right)\right.\right.} \\
& \left.-\beta_{2} J_{0}\left(\widehat{\beta}_{2} r_{0}\right) J_{1}\left(\beta_{2} r_{0}\right)\right) \\
\times & \times \frac{b \overline{\mathrm{C}}_{1} i}{2 \pi \omega_{r} \gamma}-\left(b J_{0}\left(\widehat{\beta}_{2}\right)-l \widehat{\beta}_{2} J_{1}\left(\widehat{\beta}_{2}\right)\right) \\
\times & \left(\gamma C_{1}-\alpha \bar{C}_{1}\right) J_{1}\left(\beta_{2} r_{0}\right) \\
\times & \left.\frac{\beta_{2} i}{2 \pi \omega_{r} \alpha \gamma}\right] /\left(b J_{0}\left(\widehat{\beta}_{2}\right)-l \widehat{\beta}_{2} J_{1}\left(\widehat{\beta}_{2}\right)\right) \\
\times & \left(\beta_{2} J_{1}\left(\beta_{2} r_{0}\right) Y_{0}\left(\widehat{\beta}_{2} r_{0}\right)\right. \\
& \left.\quad-\mu^{*} \widehat{\beta}_{2} Y_{1}\left(\widehat{\beta}_{2} r_{0}\right) J_{0}\left(\beta_{2} r_{0}\right)\right) \\
\times & \left(b Y_{0}\left(\widehat{\beta}_{2}\right)-l \widehat{\beta}_{2} Y_{1}\left(\widehat{\beta}_{2}\right)\right) \\
\times & \left(\mu^{*} \widehat{\beta}_{2} J_{0}\left(\beta_{2} r_{0}\right) J_{1}\left(\widehat{\beta}_{2} r_{0}\right)\right. \\
& \left.\quad-\beta_{2} J_{1}\left(\beta_{2} r_{0}\right) J_{0}\left(\widehat{\beta}_{2} r_{0}\right)\right) . \tag{36}
\end{align*}
$$

To show the flow behavior and the effect of the slip parameter, we investigate the velocity profiles in the arteries with different values of the slip parameter under various different conditions. In the first example of investigation, the radius of the artery is taken as $r=b=0.15 \mathrm{~cm}$, and the other parameters are set to $A_{0}=698.65 \mathrm{dyne} / \mathrm{cm}^{3}, A_{g}=0.5 \mathrm{~g}$, $f_{b}=f_{p}=1.2, \phi=0, C_{1}=6.6, C_{2}=4.64, A_{1}=1.2 A_{0}$, and $\rho_{1} / \rho_{2}=1$. Figure 1 shows the 3 -dimensional velocity profile as a function of time and location and the 2 -dimensional velocity profile as a function of time at three different radial locations for two different slip parameters $l=0$ (no-slip) and $l=2$. The results show that boundary slip has a very dramatical effect on the fluid flow in the artery. It affects not only the magnitude of the flow velocity significantly, but also the flow pattern and velocity profile on the cross-section of the artery. For the no-slip flow $\left(l_{b}=0\right)$, the pulsatile flow nature gradually disappears toward the arterial wall, while with boundary slip, the flow near the arterial wall also displays a pulsatile nature.

We then investigate whether the above observed flow phenomena associated with boundary slip are affected or not by the radius of the artery, and for this purpose, we consider the fluid flow through an artery with a larger radius $r=0.5 \mathrm{~cm}$. The constant pressure gradient is set to $A_{0}=32$ dyne $/ \mathrm{cm}^{3}$ in order to achieve a mean velocity magnitude approximately equal to that in the smaller artery, while all other parameters are set to the same values as those used for the smaller radius. Figure 2 shows the velocity profile in the artery for two different slip parameter values including $l_{b}=0$ (no-slip) and $l_{b}=2$. The 3 -dimensional graph shows the variation of the flow velocity with time and radial position, while the 2-dimensional graphs demonstrate the variation of the flow velocity with time at three different radial locations including $r=0$ (centre), $r=0.6$ (inner-outer layers interface), and $r=1$ (arterial wall). From Figures 1 and 2, it is clear
that the boundary slip related flow phenomena and behavior observed for the smaller artery also appear in the artery with a larger radius, and further, a more significant pulsatile nature of fluid flow is observed for the larger artery.

To further investigate the effect of the slip parameter on the velocity profile near the artery wall, we show in Figure 3 the velocity of fluid on the artery wall for four different values of the slip parameter including $l_{b}=0,2,4,6$, and 8 . The results clearly demonstrate that the slip parameter has a very significant effect on the near-wall velocity and that the magnitude of the average wall velocity is proportional to the slip parameter.

## 4. Numerical Investigation

A numerical scheme, based on the finite different method, is established to solve the underlying boundary value problem for the general case $m \neq 0$, consisting of (14) and boundary condition (15)-(19). To validate the numerical technique, we apply the numerical scheme to generate a series of numerical solutions for the case $m=0$ and then compare the numerical results with the exact solution derived in Section 3.

Figure 4 presents the velocity profile in the small and large arteries for two different slip parameters $l_{b}=0$ (no-slip) and $l_{b}=2$ obtained by the numerical technique. The numerical errors between the exact solution and the numerical solution, $E_{r}=V-U$, are presented in Figure 5 in which $V$ is the exact solution and $U$ is the numerical solution. The results clearly indicate that the numerical solution converges to the exact solution. This shows that a larger slip length has a lower convergence rate.

We then investigate the flow phenomena for the general case $m \neq 0$, and here we consider $m=-1 / 4$ in the investigation. Figure 6 gives the 3D graph showing the convergence of the transient velocity field to a steady state pulsatile velocity field and also demonstrating the substantial influence of boundary slip on the steady state velocity profile in both magnitude and flow pattern. Figure 7 shows the variations of velocities with time at three arterial locations for different slip parameters and artery radii and also clearly demonstrates the significant effect of boundary slip on the flow through the artery. Figure 8 shows the variation of fluid velocity along the artery wall under different slip parameters and artery radii. The results show that as the slip parameter increases, the time required for achieving convergence results increases, and the magnitude of the average steady state velocity also increases.

## 5. Conclusion

In this paper, a mathematical model for the transient pulsatile flow of fluids through vessels, taking into account boundary slip and the Fahraeus-Lindqvist effect, is established. For a special case of the underlying boundary value problem, an exact solution for the velocity field has been derived in explicit form, which provides one with an exact analytical method for investigating the flow phenomena under the special case and also a mean for validating the subsequently developed numerical scheme for generating numerical results for the
general case. Our analytical and numerical studies show that for the flow of fluids with the Fahraeus-Lindqvist effect, boundary slip has a very significant influence on the magnitude of the mean flow velocity and on the flow pattern and velocity profile on the cross-section. With boundary slip, the boundary layer near the wall also displays significant pulsatile flow nature. The results also show that as the boundary slip length increases, the convergence rate of numerical results to the exact solutions decreases and the time required to achieve the steady state pulsatile flow increases.

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# Analytical Solutions of Boundary Values Problem of 2D and 3D Poisson and Biharmonic Equations by Homotopy Decomposition Method 

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Received 13 June 2013; Accepted 18 August 2013
Academic Editor: Santanu Saha Ray
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#### Abstract

The homotopy decomposition method, a relatively new analytical method, is used to solve the 2D and 3D Poisson equations and biharmonic equations. The method is chosen because it does not require the linearization or assumptions of weak nonlinearity, the solutions are generated in the form of general solution, and it is more realistic compared to the method of simplifying the physical problems. The method does not require any corrected function or any Lagrange multiplier and it avoids repeated terms in the series solutions compared to the existing decomposition method including the variational iteration method, the Adomian decomposition method, and Homotopy perturbation method. The approximated solutions obtained converge to the exact solution as $N$ tends to infinity.


## 1. Introduction

The numerical solution of Poisson equations and biharmonic equations is an important problem in numerical analysis. A vast arrangement of investigating effort has been published on the development of numerical solution of Poisson equations and biharmonic equations. The finite difference schemes of second and fourth order for the solution of Poisson's equation in polar coordinates have been derived by Mittal and Gahlaut [1]. A numerical method to interpolate the source terms of Poisson's equation by using B-spline approximation has been devised by Perrey-Debain and ter Morsche [2]. Sutmann and Steffen [3] proposed compact approximation schemes for the Laplace operator of fourth and sixth order; the schemes are based on a Pade approximation of the Taylor expansion for the discretized Laplace operator. Ge [4] used fourth-order compact difference discretization scheme with unequal mesh sizes in different coordinate directions to solve a 3D Poisson equation on a cubic domain. Gumerov and Duraiswami [5] developed a complete translation theory for
the biharmonic equation in three dimensions. Khattar et al. [6] derived a fourth-order finite difference approximation based on arithmetic average discretization for the solution of three-dimensional nonlinear biharmonic partial differential equations on a 19-point compact stencil using coupled approach. Altas et al. [7] used multigrid and preconditioned Krylov iterative methods to solve three-dimensional nonlinear biharmonic partial differential equations. Jeon [8] derived scalar boundary integral equation formulas for both interior and exterior biharmonic equations with the Dirichlet boundary data. A spectral collocation method for numerically solving two-dimensional biharmonic boundaryvalue problems has been reported in [9]. An indirect radial-basis-function collocation method for numerically solving biharmonic boundary-value problems has been reported in [10]. A high-order boundary integral equation method for the solution of biharmonic equations has been presented in [11]. A Galerkin boundary node method for solving biharmonic problems was developed in [12]. An integral collocation approach based on Chebyshev polynomials for numerically
solving biharmonic equations for the case of irregularly shaped domains has been developed by Mai-Duy et al. [13]. A numerical method, based on neural-network-based functions, for solving partial differential equations has been in [14]. Mai-Duy and Tanner [15] presented a collocation method based on a Cartesian grid and a 1D integrated radial basis function scheme for numerically solving partial differential equations in rectangular domains and Haar wavelet presented in [16]. The aim of this paper is to solve these problems via the homotopy decomposition method.

## 2. Method

In this study we follow the method of [17-20]. In order to illustrate the basic idea of this method we consider a general nonlinear nonhomogeneous partial differential equation with initial conditions of the following form

$$
\begin{array}{r}
\frac{\partial^{m} U(x, t)}{\partial t^{m}}=L(U(x, t))+N(U(x, t))+f(x, t)  \tag{1}\\
m=1,2,3, \ldots
\end{array}
$$

subject to the initial conditions

$$
\begin{array}{r}
\frac{\partial^{i} U(x, 0)}{\partial t^{i}}=f_{m}(x), \quad \frac{\partial^{m-1} U(x, 0)}{\partial t^{m-1}}=0  \tag{2}\\
\quad i=0,1,2, \ldots, m-2
\end{array}
$$

where $f$ is a known function, $N$ is the general nonlinear differential operator, and $L$ represents a linear differential operator. The method's first step here is to apply the inverse operator $\partial^{m} / \partial t^{m}$ of on both sides (1) to obtain

$$
\begin{align*}
U(x, t)= & \sum_{k=0}^{m-1} \frac{t^{k}}{k!} \frac{d^{k} u(x, 0)}{d t^{k}} \\
& +\int_{0}^{t} \int_{0}^{t_{1}} \cdots \int_{0}^{t_{m-1}} L(U(x, \tau))+N(U(x, \tau))  \tag{3}\\
& +f(x, \tau) d \tau \cdots d t
\end{align*}
$$

The multi-integrals in (3) can be transformed to

$$
\begin{align*}
& \int_{0}^{t} \int_{0}^{t_{1}} \cdots \int_{0}^{t_{m-1}} L(U(x, \tau))+N(U(x, \tau)) \\
& \quad+f(x, \tau) d \tau \cdots d t_{1} \\
& =\frac{1}{(m-1)!} \int_{0}^{t}(t-\tau)^{m-1} L(U(x, \tau))+N(U(x, \tau)) \\
& \quad+f(x, \tau) d \tau \tag{4}
\end{align*}
$$

So that (3) can be reformulated as

$$
\begin{align*}
& U(x, t) \\
& \begin{aligned}
= & \sum_{k=0}^{m-1} \frac{t^{k}}{k!}\left\{\frac{d^{k} u(x, 0)}{d t^{k}}\right\} \\
& +\frac{1}{(m-1)!} \int_{0}^{t}(t-\tau)^{m-1} L(U(x, \tau))+N(U(x, \tau)) \\
& +f(x, \tau) d \tau .
\end{aligned}
\end{align*}
$$

Using the homotopy scheme the solution of the previous integral equation is given in a series form as

$$
\begin{align*}
U(x, t, p) & =\sum_{n=0}^{\infty} p^{n} U_{n}(x, t)  \tag{6}\\
U(x, t) & =\lim _{p \rightarrow 1} U(x, t, p)
\end{align*}
$$

and the nonlinear term can be decomposed as

$$
\begin{equation*}
N U(r, t)=\sum_{n=1}^{\infty} p^{n} \mathscr{H}_{n}(U), \tag{7}
\end{equation*}
$$

where $p \in(0,1]$ is an embedding parameter. $\mathscr{H}_{n}(U)$ is He's polynomials [21] that can be generated by

$$
\begin{array}{r}
\mathscr{H}_{n}\left(U_{0}, \ldots, U_{n}\right)=\frac{1}{n!} \frac{\partial^{n}}{\partial p^{n}}\left[N\left(\sum_{j=0}^{n} p^{j} U_{j}(x, t)\right)\right],  \tag{8}\\
n=0,1,2 \ldots .
\end{array}
$$

The homotopy decomposition method is obtained by the graceful coupling of decomposition method with He's polynomials and is given by

$$
\begin{align*}
& \sum_{n=0}^{\infty} p^{n} U_{n}(x, t) \\
& \quad=T(x, t) \\
& \quad+p \frac{1}{(m-1)!} \int_{0}^{t}(t-\tau)^{m-1} \\
& \\
& \quad \times\left[f(x, \tau)+L\left(\sum_{n=0}^{\infty} p^{n} U_{n}(x, \tau)\right)\right.  \tag{9}\\
& \left.\quad+\sum_{n=0}^{\infty} p^{n} \mathscr{H}_{n}(U)\right] d \tau
\end{align*}
$$

with

$$
\begin{equation*}
T(x, t)=\sum_{k=0}^{m-1} \frac{t^{k}}{k!}\left\{\left.\frac{d^{k} u(x, t)}{d t^{k}} \right\rvert\, t=0\right\} \tag{10}
\end{equation*}
$$

Comparing the terms of the same power of $p$ gives the solutions of various orders. The initial guess of the approximation is $T(x, t)$. Some further related results can be seen in [22-25].

Lemma 1 (see [17]). The complexity of the homotopy decomposition method is of order $O(n)$.

Proof. The number of computations including product, addition, subtraction, and division are as follows.

In step 2
$U_{0}: 0$ because it is obtained directly from the initial conditions

$$
\begin{aligned}
& U_{1}: 3 \\
& \vdots \\
& U_{n}: 3 .
\end{aligned}
$$

Now in step 4 the total number of computations is equal to $\sum_{j=0}^{n} U_{j}(x, t)=3 n=O(n)$.

## 3. Solutions of the Main Problems

Problem 1. Consider the following equation

$$
\begin{equation*}
\frac{\partial^{2} u}{\partial x^{2}}+\frac{\partial^{2} u}{\partial y^{2}}=\sin (\pi x) \sin (\pi y) \tag{11}
\end{equation*}
$$

$u(x, y)=0$ along the boundaries, $0 \leq x, y \leq 1$;

$$
u_{x}(0, y)=-\frac{\sin (y \pi)}{2 \pi}
$$

The exact solution of the previous equation is given as

$$
\begin{equation*}
u(x, y)=\frac{\sin (x \pi) \sin (\pi y)}{-2 \pi^{2}} \tag{12}
\end{equation*}
$$

In the view of the homotopy decomposition method, (11) can be first transformed to

$$
\begin{align*}
u(x, y)= & u(0, y)-\frac{\sin (\pi y)}{2 \pi} x \\
+ & \int_{0}^{x}(x-\tau)\left[\sin (\pi \tau) \sin (\pi y)-u_{y y}(\tau, y)\right] \\
& u(x, y, p)=\sum_{n=0}^{\infty} p^{n} u_{n}(x, y) \tag{13}
\end{align*}
$$

Following the decomposition techniques, we obtain the following equation

$$
\begin{align*}
& \sum_{n=0}^{\infty} p^{n} u_{n}(x, y) \\
& \quad=T(x, y) \\
& \quad+p \int_{0}^{x}(x-\tau)[\sin (\pi \tau) \sin (\pi y)  \tag{14}\\
& \left.\quad-\frac{\partial^{2}}{\partial y^{2}}\left[\sum_{n=0}^{\infty} p^{n} u_{n}(x, y)\right]\right]
\end{align*}
$$

Comparing the terms of the same power of $p$ leads to

$$
\begin{aligned}
p^{0}: u_{0}(x, y)= & -\frac{\sin (\pi y)}{2 \pi} x, \\
p^{1}: u_{1}(x, y)= & \int_{0}^{x}(x-\tau)\left[\sin (\pi \tau) \sin (\pi y)-\frac{\partial^{2}}{\partial y^{2}}\left[u_{0}\right]\right] d \tau \\
& u_{1}(x, y)=0 \text { along the boundaries, } \\
p^{2}: u_{2}(x, y)= & \int_{0}^{x}(x-\tau)\left[-\frac{\partial^{2}}{\partial y^{2}}\left[u_{1}\right]\right] d \tau \\
p^{3}: u_{3}(x, y)= & \int_{0}^{x}(x-\tau)\left[-\frac{\partial^{2}}{\partial y^{2}}\left[u_{2}\right]\right] d \tau
\end{aligned}
$$

$p^{n}: u_{n}(x, y)=\int_{0}^{x}(x-\tau)\left[-\frac{\partial^{2} u_{n-1}}{\partial y^{2}}\right] d \tau$,
$u_{n}(x, y)=0$ along the boundaries.

The following solutions are obtained:

$$
\begin{align*}
& u_{0}(x, y)=-\frac{\sin (\pi y)}{2 \pi} x, \\
& u_{1}(x, y)=\left[\frac{x}{\pi}-\frac{\pi x^{3}}{2 \times 3!}\right] \sin (\pi y)-\frac{\sin (\pi \tau) \sin (\pi y)}{\pi^{2}}, \\
& u_{2}(x, y)=\left[-\frac{x}{\pi}+\frac{\pi x^{3}}{6}-\frac{\pi^{3} x^{5}}{240}\right] \sin (\pi y)+\frac{\sin (\pi \tau) \sin (\pi y)}{\pi^{2}}, \\
& u_{3}(x, y)=\left[\frac{x}{\pi}-\frac{\pi x^{3}}{6}+\frac{\pi^{3} x^{5}}{120}-\frac{\pi^{5} x^{7}}{10080}\right] \sin (\pi y) \\
& -\frac{\sin (\pi \tau) \sin (\pi y)}{\pi^{2}}, \\
& u_{4}(x, y) \\
& =\left[-\frac{x}{\pi}+\frac{\pi x^{3}}{6}-\frac{\pi^{3} x^{5}}{120}+\frac{\pi^{5} x^{7}}{5040}-\frac{\pi^{7} x^{9}}{725760}\right] \sin (\pi y) \\
& +\frac{\sin (\pi \tau) \sin (\pi y)}{\pi^{2}}, \\
& u_{5}(x, y) \\
& =\left[\frac{x}{\pi}-\frac{\pi x^{3}}{6}+\frac{\pi^{3} x^{5}}{120}-\frac{\pi^{5} x^{7}}{5040}+\frac{\pi^{7} x^{9}}{362880}-\frac{\pi^{9} x^{11}}{79833600}\right] \sin (\pi y) \\
& -\frac{\sin (\pi \tau) \sin (\pi y)}{\pi^{2}}, \\
& u_{6}(x, y) \\
& =\left[-\frac{x}{\pi}+\frac{\pi x^{3}}{6}-\frac{\pi^{3} x^{5}}{120}+\frac{\pi^{5} x^{7}}{5040}-\frac{\pi^{7} x^{9}}{362880}\right. \\
& \left.+\frac{\pi^{9} x^{11}}{39916800}-\frac{\pi^{11} x^{13}}{12454041600}\right] \sin (\pi y) \\
& +\frac{\sin (\pi \tau) \sin (\pi y)}{\pi^{2}}, \\
& u_{7}(x, y) \\
& =\left[\frac{x}{\pi}-\frac{\pi x^{3}}{6}+\frac{\pi^{3} x^{5}}{120}-\frac{\pi^{5} x^{7}}{5040}+\frac{\pi^{7} x^{9}}{362880}-\frac{\pi^{9} x^{11}}{39916800}\right. \\
& \left.+\frac{\pi^{11} x^{13}}{6227020800}-\frac{\pi^{13} x^{15}}{2615348736000}\right] \sin (\pi y) \\
& -\frac{\sin (\pi \tau) \sin (\pi y)}{\pi^{2}} \text {. } \tag{16}
\end{align*}
$$

TABLE 1: Evaluation of numerical errors for $N=4$.

| $x$ | $Y$ | $u(x, y)$ exact | $u(x, y) N=4$ | Error |
| :---: | :---: | :---: | :---: | :---: |
| 0.25 | 0.25 | -0.0253303 | -0.0253303 | $6.27007 \cdot 10^{-11}$ |
|  | 0.5 | -0.0358224 | -0.0358224 | $8.86722 \cdot 10^{-11}$ |
|  | 0.75 | -0.0253303 | -0.0253303 | $6.27007 \cdot 10^{-11}$ |
|  | 0.95 | -0.00560387 | -0.00560387 | $1.38714 \cdot 10^{-11}$ |
|  | 0.25 | -0.0358224 | -0.0358224 | $1.26904 \cdot 10^{-11}$ |
|  | 0.5 | -0.0506606 | -0.0506604 | $1.79469 \cdot 10^{-7}$ |
|  | 0.75 | -0.0358224 | -0.0358223 | $1.26904 \cdot 10^{-11}$ |
|  | 0.95 | -0.00792506 | -0.00792506 | $2.80752 \cdot 10^{-8}$ |
| 0.75 | 0.25 | -0.0253303 | -0.0253195 | $1.07646 \cdot 10^{-5}$ |
|  | 0.5 | -0.0358224 | -0.0358072 | $1.52235 \cdot 10^{-5}$ |
|  | 0.75 | -0.0253303 | -0.0253195 | $1.07646 \cdot 10^{-5}$ |
|  | 0.95 | -0.00560387 | -0.00560148 | $2.38148 \cdot 10^{-6}$ |
|  | 0.25 | -0.00560387 | -0.00546191 | 0.000141956 |
| 0.95 | 0.5 | -0.00792506 | -0.00772431 | 000200756 |
|  | 0.75 | -0.00560387 | -0.00546191 | 0.000141956 |
|  | 0.95 | -0.00123975 | -0.00120835 | $3.14051 \cdot 10^{-5}$ |

In the same manner one can obtain the rest of the components. But for eight terms were computed and the asymptotic solution is given by

$$
\begin{align*}
& u_{N=8}(x, y) \\
& =\left[\frac{x}{2 \pi}-\frac{\pi x^{3}}{2 \times 3!}+\frac{\pi^{3} x^{5}}{2 \times 5!}-\frac{\pi^{5} x^{7}}{2 \times 7!}+\frac{\pi^{7} x^{9}}{2 \times 9!}\right. \\
& \left.\quad-\frac{\pi^{9} x^{11}}{2 \times 11!}+\frac{\pi^{11} x^{13}}{2 \times 13!}-\frac{\pi^{13} x^{15}}{2 \times 15!}\right] \sin (\pi y)  \tag{17}\\
& \\
& \quad-\frac{1}{\pi^{2}} \sin (x \pi) \sin (y \pi)
\end{align*}
$$

Therefore in general for any $N>8$ we have

$$
\begin{align*}
u_{N=n}(x, y)= & {\left[\frac{1}{2 \pi^{2}} \sum_{n=0}^{N} \frac{(-1)^{n}(x \pi)^{2 n+1}}{(2 n+1)!}\right] \sin (\pi y) } \\
& -\frac{1}{\pi^{2}} \sin (x \pi) \sin (y \pi), \\
\lim _{N \rightarrow \infty} u_{N}(x, y)= & \frac{1}{2 \pi^{2}} \sin (\pi x) \sin (\pi y)  \tag{18}\\
& -\frac{1}{\pi^{2}} \sin (x \pi) \sin (y \pi) \\
= & -\frac{1}{2 \pi^{2}} \sin (x \pi) \sin (y \pi) .
\end{align*}
$$

This is the exact solution of the problem. Figures 1 and 2 show the comparison of the exact solution and the approximated one for $N=4$. The approximate solution and the exact solution are compared in Figures 1 and 2, respectively.

The numerical errors for $N=4$ are evaluated in Table 1.


Figure 1: Exact solution.


Figure 2: Approximated solution for the 4 first terms.

Problem 2. Consider 3D Poisson equation:

$$
\begin{gather*}
\frac{\partial^{2} u}{\partial x^{2}}+\frac{\partial^{2} u}{\partial y^{2}}+\frac{\partial^{2} u}{\partial z^{2}}=\sin (\pi x) \sin (\pi y) \sin (\pi z),  \tag{19}\\
u(x, y, z)=0 \text { along the boundaries, } 0 \leq x, y \leq 1 .
\end{gather*}
$$

Following the discussion presented earlier we obtain the following set of integral equations:

$$
\left.\begin{array}{l}
p^{0}: u_{0}(x, y)=-\frac{\sin (\pi y) \sin (\pi z)}{3 \pi} x \\
p^{1}: u_{1}(x, y) \\
=\int_{0}^{x}(x-\tau)\left[\sin (\pi \tau) \sin (\pi y) \sin (\pi z)-\frac{\partial^{2} u_{0}}{\partial y^{2}}\right] d \tau \\
p^{n}: u_{n}(x, y)=\int_{0}^{x}(x-\tau)\left[-\frac{\partial^{2} u_{n-1}}{\partial y^{2}}\right] d \tau, \\
u_{n}(x, y) \tag{20}
\end{array}\right)=0 \text { along the boundaries, } n \geq 2 .
$$

The following solutions are obtained:

$$
\begin{align*}
& u_{0}(x, y, z)=-\frac{\sin (\pi y) \sin (\pi z)}{3 \pi} x, \\
& u_{1}(x, y, z)=\left[\frac{1}{\pi} x-\frac{\pi x^{3}}{9}\right] \sin (\pi y) \sin (\pi z) \\
& -\frac{\sin (\pi \tau) \sin (\pi y) \sin (\pi z)}{\pi^{2}}, \\
& u_{2}(x, y, z)=\left[-\frac{2}{\pi} x+\frac{\pi x^{3}}{9}-\frac{\pi^{3} x^{5}}{90}\right] \sin (\pi y) \sin (\pi z) \\
& +\frac{2 \sin (\pi \tau) \sin (\pi y) \sin (\pi z)}{\pi^{2}}, \\
& u_{3}(x, y, z) \\
& =\left[\frac{4}{\pi} x-\frac{2 \pi x^{3}}{3}+\frac{\pi^{3} x^{5}}{30}-\frac{\pi^{5} x^{7}}{1890}\right] \sin (\pi y) \sin (\pi z) \\
& -\frac{4 \sin (\pi \tau) \sin (\pi y) \sin (\pi z)}{\pi^{2}}, \\
& u_{4}(x, y, z) \\
& =\left[-\frac{8}{\pi} x+\frac{4 \pi x^{3}}{3}-\frac{\pi^{3} x^{5}}{15}+\frac{\pi^{5} x^{7}}{630}\right] \sin (\pi y) \sin (\pi z) \\
& +\frac{8 \sin (\pi \tau) \sin (\pi y) \sin (\pi z)}{\pi^{2}}, \\
& u_{5}(x, y, z)=\left[\frac{16}{\pi} x-\frac{8 \pi x^{3}}{3}+\frac{2 \pi^{3} x^{5}}{15}-\frac{\pi^{5} x^{7}}{315}\right. \\
& \left.+\frac{\pi^{7} x^{9}}{22680}-\frac{\pi^{7} x^{11}}{3742200}\right] \sin (\pi y) \sin (\pi z) \\
& -\frac{16 \sin (\pi \tau) \sin (\pi y) \sin (\pi z)}{\pi^{2}} . \tag{21}
\end{align*}
$$

In the same manner one can obtain the rest of the components. But for six terms were computed and the asymptotic solution is given by

$$
\begin{gathered}
u(x, y, z)_{N=6}=\left[\frac{x}{3 \pi}-\frac{\pi x^{3}}{18}+\frac{\pi^{3} x^{5}}{360}-\frac{\pi^{5} x^{7}}{15120}+\frac{\pi^{7} x^{9}}{1088640}\right. \\
\left.\quad-\frac{\pi^{9} x^{11}}{119750400}\right] \sin (\pi y) \sin (\pi z) \\
\\
-\frac{2 \sin (\pi \tau) \sin (\pi y) \sin (\pi z)}{\pi^{2}}
\end{gathered}
$$

$$
\begin{align*}
& u(x, y, z)_{N=6}= \frac{1}{3 \pi^{2}}\left[\frac{x}{3 \pi}-\frac{\pi x^{3}}{18}+\frac{\pi^{3} x^{5}}{360}-\frac{\pi^{5} x^{7}}{15120}+\frac{\pi^{7} x^{9}}{1088640}\right. \\
&\left.-\frac{\pi^{9} x^{11}}{119750400}\right] \sin (\pi y) \sin (\pi z) \\
&-\frac{2 \sin (\pi \tau) \sin (\pi y) \sin (\pi z)}{3 \pi^{2}} \\
& \begin{aligned}
u(x, y, z)_{N=6}= & \frac{1}{3 \pi^{2}}\left[\pi x-\frac{(\pi x)^{3}}{3!}+\frac{(\pi x)^{5}}{5!}-\frac{(\pi x)^{7}}{7!}\right. \\
& \left.+\frac{(\pi x)^{9}}{9!}-\frac{(\pi x)^{3}}{11!}\right] \sin (\pi y) \sin (\pi z) \\
& -\frac{2 \sin (\pi \tau) \sin (\pi y) \sin (\pi z)}{3 \pi^{2}}
\end{aligned}
\end{align*}
$$

Therefore, for any $n \geq 6$, the partial sum is given as

$$
\begin{align*}
u_{N=n}(x, y, z)= & \frac{1}{3 \pi^{2}}\left[\sum_{k=1}^{N} \frac{(-1)^{k}(\pi x)^{2 k+1}}{(2 k+1)!}\right] \sin (\pi y) \sin (\pi z) \\
& -\frac{2 \sin (\pi \tau) \sin (\pi y) \sin (\pi z)}{3 \pi^{2}} \tag{23}
\end{align*}
$$

Thus

$$
\begin{align*}
u(x, y, z)= & \lim _{N \rightarrow \infty} u_{N=n}(x, y, z) \\
= & \frac{\sin (\pi \tau) \sin (\pi y) \sin (\pi z)}{3 \pi^{2}} \\
& -\frac{2 \sin (\pi \tau) \sin (\pi y) \sin (\pi z)}{3 \pi^{2}}  \tag{24}\\
= & -\frac{\sin (\pi \tau) \sin (\pi y) \sin (\pi z)}{3 \pi^{2}}
\end{align*}
$$

And this is the exact solution to the problem. One can evaluate error committed by choosing the $N$ first terms in the series solutions, in the same manner as in Table 1. The accuracy of the results is estimated by error function

$$
\begin{equation*}
R_{N}(x, y, z)=\left|u_{N}(x, y, z)-u(x, y, z)\right| \tag{25}
\end{equation*}
$$

Problem 3. Let us consider the following biharmonic equation

$$
\begin{equation*}
\frac{d^{4} u(x)}{d x^{4}}+4 u(x)=0 \tag{26}
\end{equation*}
$$

for which the exact solution is

$$
\begin{equation*}
u(x)=\frac{\operatorname{Exp}[1-x] \cos [x]}{\cos [1]} \tag{27}
\end{equation*}
$$

The aim of this part is to compare the numerical results obtained via HDM and the method used in [26].

Table 2: Comparison of the HDM and [1] results with the exaction solution for $N=6$.

| $x$ | HDM | Exact | ADM | Err for HDM | Err for ADM |
| :--- | :---: | :---: | :---: | :---: | :---: |
| -1.0 | 7.38906 | 7.38906 | 7.38906 | $6.78 E-16$ | $8.88 E-16$ |
| -0.6 | 7.56598 | 7.56598 | 7.56598 | $4.76 E-12$ | $7.96 E-12$ |
| -0.2 | 6.02244 | 6.02244 | 6.02244 | $0.015 E-11$ | $1.46 E-11$ |
| 0.2 | 4.03696 | 4.03696 | 4.03696 | $0.017 E-11$ | $1.80 E-11$ |
| 0.6 | 2.27883 | 2.27883 | 2.27883 | $0.015 E-11$ | $1.46 E-11$ |
| 1.0 | 1.0 | 1.0 | 1.0 | $1.24 E-15$ | $2.22 E-15$ |

Applying the steps involved in the HDM, we arrive at the following:

$$
\begin{gather*}
u_{0}(x)=e \operatorname{Sec}(1)\left(1-x+\frac{x^{3}}{3}\right), \\
u_{1}(x)=-e \operatorname{Sec}(1)\left[\frac{x^{4}}{4}-\frac{x^{5}}{20}+\frac{x^{7}}{420}\right], \\
u_{2}(x)=-e \operatorname{Sec}(1)\left[-\frac{x^{8}}{1120}+\frac{x^{9}}{10080}-\frac{x^{11}}{554400}\right], \\
u_{3}(x)=-e \operatorname{Sec}(1)\left[\frac{x^{12}}{2217600}-\frac{x^{13}}{64864800}+\frac{x^{15}}{1135134000}\right], \\
u_{4}(x) \\
=-e \operatorname{Sec}(1)\left[-\frac{x^{16}}{16144128000}+\frac{x^{17}}{274450176000}\right. \\
u_{5}(x) \\
=-e \operatorname{Sec}(1)\left[\frac{\left.-\frac{x^{19}}{46930980096000}\right],}{312873200640000}-\frac{x^{20}}{6570337213440000}\right. \\
\left.+\frac{x^{23}}{1662295315000320000}\right] .
\end{gather*}
$$

In the same manner, one can obtain the remaining term by using the following recursive formula:

$$
\begin{equation*}
u_{n+1}(x)=-\int_{0}^{x}(x-t)^{3} u_{n}(t) d t \tag{29}
\end{equation*}
$$

In this paper we consider only the first six terms of the series solution as follows:

$$
\begin{equation*}
u_{N=6}=\sum_{n=0}^{5} u_{n}(x) \tag{30}
\end{equation*}
$$

To access the accuracy of the method used in paper, we compare in Table 2 the numerical results of the above equation, the solution obtained in [26] with the exact solution.

Problem 4. We consider the 2D biharmonic equation

$$
\begin{equation*}
\frac{\partial^{4} u}{\partial x^{4}}+2 \frac{\partial^{4} u}{\partial x^{2} \partial y^{2}}+\frac{\partial^{4} u}{\partial x^{4}}=\sin (3 \pi x) \sin (3 \pi y), \quad 0 \leq x, y \leq 1, \tag{31}
\end{equation*}
$$

subject to the initial conditions:

$$
\begin{gather*}
\frac{\partial u(x, y)}{\partial x} \left\lvert\,(x=0)=\frac{\sin (3 \pi y)}{108 \pi^{3}}\right., \quad \partial_{x, x} u(0, y)=0,  \tag{32}\\
\partial_{x, x, x} u(0, y)=-\frac{\sin (3 \pi y)}{12 \pi} .
\end{gather*}
$$

In the view of the homotopy decomposition method, the following integral equations are obtained:

$$
\begin{align*}
& p^{0}: u_{0}(x, y)=\frac{\sin (3 \pi y)}{108 \pi^{3}} x-\frac{\sin (3 \pi y)}{12 \pi} \frac{x^{3}}{3!}, \\
& p^{1}: u_{1}(x, y) \\
& \quad=\int_{0}^{x}(x-\tau)\left[\sin (\pi \tau) \sin (\pi y)-\frac{\partial^{2} u_{0}}{\partial y^{2}}-2 \frac{\partial^{4} u_{0}}{\partial x^{2} \partial y^{2}}\right] d \tau \\
& p^{n}: u_{n}(x, y)=\int_{0}^{x}(x-\tau)\left[-\frac{\partial^{2} u_{n-1}}{\partial y^{2}}-2 \frac{\partial^{4} u_{0}}{\partial x^{2} \partial y^{2}}\right] d \tau \\
& \quad u_{n}(x, y)=0 \text { along the boundaries, } n \geq 2 . \tag{33}
\end{align*}
$$

It is worth noting that if the zeroth component $u_{0}(x, y)$ is defined, then the remaining components $n \geq 1$ can be completely determined such that each term is determined by using the previous terms, and the series solutions are thus entirely determined. Finally, the solution $u(x, y)$ is approximated for $n=4$ :

$$
\begin{align*}
& u(x, y) \\
& \begin{aligned}
&=\sin (3 \pi y)\left[\frac{x}{108 \pi^{3}}-\frac{x^{3}}{72 \pi}+\frac{\pi x^{5}}{160}-\frac{3 \pi^{3} x^{7}}{2240}+\frac{3 \pi^{5} x^{9}}{17920}\right. \\
&\left.-\frac{27 \pi^{7} x^{11}}{1971200}+\frac{81 \pi^{9} x^{13}}{102502400}-\frac{81 \pi^{11} x^{15}}{7175168000}\right], \\
& \begin{aligned}
\begin{aligned}
&(x, y) \\
&=\frac{\sin (3 \pi y)}{324 \pi^{4}}\left[3 \pi x-\frac{(3 \pi x)^{3}}{3!}+\frac{(3 \pi x)^{5}}{5!}-\frac{(3 \pi x)^{7}}{7!}+\frac{(3 \pi x)^{9}}{9!}\right. \\
&\left.-\frac{(3 \pi x)^{11}}{11!}+\frac{(3 \pi x)^{13}}{13!}-\frac{(3 \pi x)^{15}}{15!}\right] .
\end{aligned}
\end{aligned}
\end{aligned} .
\end{align*}
$$

Therefore for any $N \geq 4$ we have the following:

$$
\begin{equation*}
u_{N}(x, y)=\frac{\sin (3 \pi y)}{324 \pi^{4}} \sum_{n=0}^{N} \frac{(3 \pi x)^{2 n+1}}{(2 n+1)!} \tag{36}
\end{equation*}
$$



Figure 3: Analytical solution.


Figure 4: Absolute value of the solution.

Thus

$$
\begin{equation*}
\lim _{N \rightarrow \infty} u_{N}(x, y)=\frac{\sin (3 \pi y) \sin (3 \pi x)}{324 \pi^{4}} \tag{37}
\end{equation*}
$$

The exact solution of (31) is given by

$$
\begin{equation*}
\frac{\sin (3 \pi y) \sin (3 \pi x)}{324 \pi^{4}}=u(x, y) \tag{38}
\end{equation*}
$$

Figures 3 and 4 are the graphical representation of the previous solution. We have plotted the solution for (31) in Figure 3 and showed absolute value of the solution in Figure 4.

Theorem 2. Let $m$ be a nonzero natural number and let $(x, y) \in[0,1] \times[0,1]$; then two dimensional biharmonic equation of form

$$
\begin{equation*}
\frac{\partial^{4} u}{\partial x^{4}}+2 \frac{\partial^{4} u}{\partial x^{2} \partial y^{2}}+\frac{\partial^{4} u}{\partial y^{4}}=\sin (m \pi x) \sin (m \pi y) \tag{39}
\end{equation*}
$$

with $u(x, y)=0$ along the boundaries has an exact solution as follows

$$
\begin{equation*}
u(x, y)=\frac{\sin (m \pi x) \sin (m \pi y)}{4 m^{4} \pi^{4}} \tag{40}
\end{equation*}
$$

Proof. Use the step of the homotopy decomposition method.

Problem 5. We consider the 3D biharmonic equation:

$$
\begin{gather*}
\frac{\partial^{4} u}{\partial x^{4}}+2 \frac{\partial^{4} u}{\partial x^{2} \partial y^{2}}+2 \frac{\partial^{4} u}{\partial x^{2} \partial z^{2}}+2 \frac{\partial^{4} u}{\partial z^{2} \partial y^{2}}+\frac{\partial^{4} u}{\partial y^{4}}+\frac{\partial^{4} u}{\partial z^{4}} \\
=\sin (\pi x) \sin (\pi y) \sin (\pi z)  \tag{41}\\
0 \leq x, y, z \leq 1 \\
u=0, \quad u_{x, x}=u_{y, y}=u_{z, z}=0
\end{gather*}
$$

In the view of the homotopy decomposition method, the following integral equations are obtained:

$$
\begin{aligned}
& p^{0}: u_{0}(x, y)=\frac{\sin (3 \pi y)}{108 \pi^{3}} x-\frac{\sin (3 \pi y)}{12 \pi} \frac{x^{3}}{3!} \\
& \begin{aligned}
p^{1}: u_{1}(x, y)
\end{aligned} \\
& =\int_{0}^{x}(x-\tau)[\sin (\pi \tau) \sin (\pi y) \sin (\pi z) \\
& \\
& -2 \frac{\partial^{4} u_{0}}{\partial x^{2} \partial y^{2}}-2 \frac{\partial^{4} u_{0}}{\partial x^{2} \partial z^{2}} \\
& \\
& \left.\left.p^{n}: 2 \frac{\partial^{4} u_{0}}{\partial z^{2} \partial y^{2}}-\frac{\partial^{4} u_{0}}{\partial y^{4}}-\frac{\partial^{4} u_{0}}{\partial z^{4}}\right] d \tau, y\right)
\end{aligned}
$$

$$
=\int_{0}^{x}(x-\tau)\left[-2 \frac{\partial^{4} u_{n-1}}{\partial x^{2} \partial y^{2}}-2 \frac{\partial^{4} u_{n-1}}{\partial x^{2} \partial z^{2}}\right.
$$

$$
\left.-2 \frac{\partial^{4} u_{n-1}}{\partial z^{2} \partial y^{2}}-\frac{\partial^{4} u_{n-1}}{\partial y^{4}}-\frac{\partial^{4} u_{n-1}}{\partial z^{4}}\right] d \tau
$$

$$
\begin{equation*}
u_{n}(x, y)=0 \text { along the boundaries, } \quad n \geq 2 \tag{42}
\end{equation*}
$$

Solving the previous integral equations, the series solutions for the first $N$ terms are given as

$$
\begin{equation*}
u_{N}(x, y, z)=\frac{\sin (z \pi) \sin (\pi y)}{9 \pi^{4}} \sum_{n=0}^{N} \frac{(\pi x)^{2 n+1}}{(2 n+1)!} \tag{43}
\end{equation*}
$$

Therefore taking the limit at $N$ tending to infinity we obtained

$$
\begin{equation*}
u(x, y, z)=\lim _{N \rightarrow \infty} u_{N}(x, y, z)=\frac{\sin (x \pi) \sin (z \pi) \sin (\pi y)}{9 \pi^{4}} \tag{44}
\end{equation*}
$$

## 4. Conclusion

In this paper the recent homotopy decomposition [18-21] is used to solve the 2D and 3D Poisson equations and biharmonic equations. The method is chosen because it does not require the linearization or assumptions of weak nonlinearity,
the solutions are generated in the form of general solution, and it is more realistic compared to the method of simplifying the physical problems. The method does not require any corrected function any Lagrange multiplier and it avoids repeated terms in the series solutions compared to the existing decomposition method including the variational iteration method and the Adomian decomposition method. The approximated solutions obtained converge to the exact solution as $N$ tends to infinity. The numerical values are presented in Table 1 shows that the method is very efficient and accurate.

## Acknowledgment

The authors would like to thank referee(s) for very useful comments regarding the details and their remarks that improved the presentation and the contents of the paper.

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## Research Article

# Pattern Dynamics in a Spatial Predator-Prey System with Allee Effect 

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Received 9 May 2013; Accepted 22 August 2013
Academic Editor: Rasajit Bera
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#### Abstract

We investigate the spatial dynamics of a predator-prey system with Allee effect. By using bifurcation analysis, the exact Turing domain is found in the parameters space. Furthermore, we obtain the amplitude equations and determine the stability of different patterns. In Turing space, it is found that predator-prey systems with Allee effect have rich dynamics. Our results indicate that predator mortality plays an important role in the pattern formation of populations. More specifically, as predator mortality rate increases, coexistence of spotted and stripe patterns, stripe patterns, spotted patterns, and spiral wave emerge successively. The results enrich the finding in the spatial predator-prey systems well.


## 1. Introduction

The Allee effect, named after the ecologist Warder Clyde Allee, has been recognized as an important phenomenon of positive density dependence in low-density population [1-5]. Allee effect can occur whenever fitness of an individual in a small or sparse population decreases as the population size or density also declines $[6,7]$. Since the outstanding work of Allee [1], the Allee effect has been regarded as one of the central and highly important issues in the population and community ecology. And its critical importance has widely been realized in the conservation biology that Allee effect is most likely to increase the extinction risk of low-density populations. As a result, studies on Allee effect have received more and more attention from both mathematicians and ecologists.

Long time series of the density of both prey and predator is needed, so it is difficult to analyse their dynamics. As a result, it may provide useful information by constructing mathematical models to investigate the dynamical behaviors of predator-prey systems. There have been a large group of papers on predator-prey systems with Allee effect [8-13].

However, these previous works did not take into account the effect of space.

There are also some works done on spatial predatorprey systems with Allee effect [14-16]. Petrovskii et al. found that the deterministic system with Allee effect can induce patch invasion [14]. Morozov et al. found that the temporal population oscillations can exhibit chaotic dynamics even when the distribution of the species in the space was regular [15]. Moreover, they found that the chaos accompanied with patch invasion even though the environments were heterogeneous [16]. However, their results were obtained by choosing particular initial conditions. Then, it is natural to ask what kind of patterns can be obtained in predator-prey systems with Allee effect by using other initial conditions. To understand that mechanism well, we will investigate a predator-prey system with Allee effect.

Because of the insightful work of many scientists over recent years, we can make research on pattern selection by using the standard multiple scale analysis [17, 18], in which the control parameters and the derivatives are expanded in terms of a small enough parameter. In the neighborhood of
the bifurcation points (Hopf and Turing bifurcation points), the critical amplitudes follow the normal forms, and thus their general forms can be obtained from the methods of symmetry-breaking bifurcations.

The paper is organized as follows. In Section 2, we present a predator-prey system with Allee effect and give Turing region in parameters space. In Section 3, by using multiple scale analysis, we obtain amplitude equations. In Section 4, we show the spatial patterns by a series of numerical simulations. Finally, conclusions and discussions are presented in Section 5.

## 2. A Predator-Prey System with Allee Effect

We consider the following model of two-dimensional spatiotemporal system [14-16, 19]:

$$
\begin{align*}
& \frac{\partial H}{\partial T}=F(H)-f(H, P)+D_{1} \Delta H  \tag{1a}\\
& \frac{\partial P}{\partial T}=\kappa f(H, P)-D(P)+D_{2} \Delta P \tag{lb}
\end{align*}
$$

where $H=H(X, Y, T)$ and $P=P(X, Y, T)$ are densities of prey and predator, respectively, at time $T$ and position $(X, Y)$. The function $F(H)$ represents the intrinsic prey growth, $f(H, P)=f(H) P$ represents predation term, $\kappa$ is the food utilization coefficient, $D_{1}$ and $D_{2}$ are diffusion coefficients, and $D(P)$ describes predator mortality.

It is assumed that the predation term is a bilinear form of prey and predator density and predator mortality is a nonlinear function of predator density. As a result, we choose $f(H, P)=H P$ and $D(P)=M P^{2}[20]$.

When the prey population obeys Allee dynamics, its growth rate can be parameterized as follows [14, 15, 21]:

$$
\begin{equation*}
F(H)=\frac{4 \omega}{\left(K-H_{0}\right)^{2}} H\left(H-H_{0}\right)(K-H) \tag{2}
\end{equation*}
$$

where $K$ is the prey-carrying capacity, $\omega$ is the maximum per capita growth rate, and $H_{0}$ quantifies the intensity of the Allee effect. If $0<H_{0}<K, F(H)$ is a strong Allee effect; if $-K<H_{0}<0, F(H)$ is a weak Allee effect; if $H_{0} \leq-K$, the Allee effect is absent.

In order to minimize the number of parameters involved in the model system, it is extremely useful to write the system in a nondimensionalized form. Although there is no unique method of doing this, it is often a good idea to relate the variables to some key relevant parameters. Introducing dimensionless variables

$$
\begin{align*}
& u=\frac{H}{K}, \quad v=\frac{P}{\kappa K}, \quad t=a T  \tag{3}\\
& \bar{X}=X \sqrt{\frac{a}{D_{1}}}, \quad \bar{Y}=Y \sqrt{\frac{a}{D_{1}}},
\end{align*}
$$

we obtain the following equations:

$$
\begin{gather*}
\frac{\partial u}{\partial t}=\gamma u(u-\beta)(1-u)-u v+\Delta u  \tag{4a}\\
\frac{\partial v}{\partial t}=u v-\delta v^{2}+\varepsilon \Delta v \tag{4b}
\end{gather*}
$$

where

$$
\begin{array}{ll}
\beta=\frac{H_{0}}{K}, & \gamma=\frac{4 \omega K}{A \kappa\left(K-H_{0}\right)^{2}},  \tag{5}\\
\delta=\frac{M}{a}, & \varepsilon=\frac{D_{2}}{D_{1}} .
\end{array}
$$

First of all, we need to investigate the dynamics of nonspatial model of systems (4a) and (4b)

$$
\begin{gather*}
\frac{d u}{d t}=\gamma u(u-\beta)(1-u)-u v  \tag{6a}\\
\frac{d v}{d t}=u v-\delta v^{2} \tag{6b}
\end{gather*}
$$

Systems (6a) and (6b) have three boundary equilibrium named $E_{0}=(0,0), E_{1}=(1,0)$, and $E_{2}=(\beta, 0)$ and two interior equilibriums named $E_{3}$ and $E_{4}$, where

$$
\begin{align*}
& E_{3}=\left(\frac{\gamma \delta+\gamma \beta \delta-1+\sqrt{Q}}{2 \gamma \delta}, \frac{\gamma \delta+\gamma \beta \delta-1+\sqrt{Q}}{2 \gamma \delta^{2}}\right),  \tag{7a}\\
& E_{4}=\left(\frac{\gamma \delta+\gamma \beta \delta-1-\sqrt{Q}}{2 \gamma \delta}, \frac{\gamma \delta+\gamma \beta \delta-1-\sqrt{Q}}{2 \gamma \delta^{2}}\right) \tag{7b}
\end{align*}
$$

where $Q=(\gamma \delta)^{2}-2(\gamma \delta)^{2} \beta-2 \gamma \beta \delta+1$.
From a biological point of view, we are concerned with the dynamics of $E_{3}$ and $E_{4}$. The Jacobian matrix corresponding to the equilibrium point is that

$$
J=\left(\begin{array}{ll}
a_{11} & a_{12}  \tag{8}\\
a_{21} & a_{22}
\end{array}\right)
$$

where

$$
\begin{gather*}
a_{11}=2 \gamma u^{*}-\gamma \beta-3 \gamma\left(u^{*}\right)^{2}+2 \gamma \beta u^{*} \\
a_{12}=-u^{*}  \tag{9}\\
a_{21}=v^{*}-\delta \\
a_{22}=u^{*}
\end{gather*}
$$

Diffusion-driven instability requires the stable, homogeneous steady state is driven unstable by the interaction of the dynamics and diffusion of the species; and therefore

$$
\begin{gather*}
a_{11}+a_{22}<0  \tag{10}\\
a_{11} a_{22}-a_{12} a_{21}>0
\end{gather*}
$$

It is found from direct calculations that $E_{3}$ is unstable and $E_{4}$ is stable. Denote $E_{4}=\left(u^{*}, v^{*}\right)$.

Following the standard linear analysis of the reactiondiffusion equation [22], we consider a perturbation near the steady state:

$$
\begin{align*}
& u(\vec{r}, t)=u^{*}+\bar{u}(r, t), \\
& v(\vec{r}, t)=v^{*}+\bar{v}(r, t) \tag{11}
\end{align*}
$$

where $\bar{u}(r, t) \ll u^{*}, \bar{v}(r, t) \ll v^{*}$, and $r=(\bar{X}, \bar{Y})$. Assume that

$$
\begin{equation*}
\binom{\bar{u}(r, t)}{\bar{v}(r, t)}=\binom{\alpha_{1}}{\alpha_{2}} e^{\lambda t} e^{i\left(\kappa_{X} X+\kappa_{Y} Y\right)} \tag{12}
\end{equation*}
$$

where $\lambda$ is the growth rate of perturbation in time $t, \alpha_{1}$ and $\alpha_{2}$ represent the amplitudes, and $\kappa_{X}$ and $\kappa_{Y}$ are the wave number of the solutions.

The characteristic equation of the systems (4a) and (4b) is

$$
\begin{equation*}
(A-\lambda I)\binom{\bar{u}}{\bar{v}}=0 \tag{13}
\end{equation*}
$$

where

$$
A=\left(\begin{array}{cc}
a_{11}-\left(\kappa_{X}^{2}+\kappa_{Y}^{2}\right) & a_{12}  \tag{14}\\
a_{21} & a_{22}-\varepsilon\left(\kappa_{X}^{2}+\kappa_{Y}^{2}\right)
\end{array}\right)
$$

As a result, we have characteristic polynomial:

$$
\begin{gather*}
\lambda^{2}-t r_{\kappa} \lambda+\Delta_{\kappa}=0  \tag{15}\\
t r_{\kappa}=a_{11}+a_{22}-\kappa^{2}(1+\varepsilon) \stackrel{\Delta}{=} t r_{J}-\kappa^{2}(1+\varepsilon), \\
\Delta_{\kappa}=a_{11} a_{22}-a_{12} a_{21}-\kappa^{2}\left(a_{11} \varepsilon+a_{22}\right)+\kappa^{4} \varepsilon  \tag{16}\\
\stackrel{\Delta}{=} \Delta_{J}-\kappa^{2}\left(a_{11} \varepsilon+a_{22}\right)+\kappa^{4} \varepsilon
\end{gather*}
$$

where $\kappa^{2}=\kappa_{X}^{2}+\kappa_{Y}^{2}$.
The roots of (15) can be obtained by the following form:

$$
\begin{equation*}
\lambda_{\kappa}=\frac{1}{2}\left(t r_{\kappa} \pm \sqrt{t r_{\kappa}^{2}-4 \Delta_{\kappa}}\right) . \tag{17}
\end{equation*}
$$

When $\operatorname{Im}\left(\lambda_{\kappa}\right) \neq 0$ and $\operatorname{Re}\left(\lambda_{\kappa}\right)=0$, Hopf bifurcation will emerge. Then, we have that the critical value of Hopf bifurcation parameter- $\delta$ equals

$$
\begin{equation*}
\delta_{H}=\frac{\gamma(\gamma+\beta-1)}{\gamma^{2} \beta^{2}+\gamma^{2}-2 \gamma^{2} \beta-1} \tag{18}
\end{equation*}
$$

When $\kappa^{2}=\left(\kappa_{T}\right)^{2}=\sqrt{\Delta_{J} / \varepsilon}$ and $\operatorname{Im}\left(\lambda_{\kappa}\right)=0, \operatorname{Re}\left(\lambda_{\kappa}\right)=0$, Turing bifurcation will occur. Denote $\delta_{T}$ as the critical value of $\delta$ as Turing instability occurs. Since the expression is complicated, we omit it here.

In Figure 1, we show the two critical lines in the parameter space spanned by $\beta$ and $\delta$. The equilibria that can be found in the region, marked by $T$ (Turing space), are stable with


Figure 1: Bifurcation diagram for the systems (4a) and (4b). The green one is the Hopf bifurcation critical line and the red one, Turing bifurcation critical line. The figure shows the Turing space which is marked by $T$. Parameters values: $\gamma=1.5$ and $\varepsilon=0.15$.
respect to the homogeneous perturbations, but they lose their stability with respect to the perturbations of specific wave numbers $\kappa$. In this region, stationary patterns can be observed. To see the effect of parameter $\delta$ well, we plot in Figure 2 the dispersion relation corresponding to several values of $\delta$ while keeping the other parameters fixed. We see that the available Turing modes shift to higher wave numbers when $\beta$ decreases.

## 3. Spatial Dynamics of Systems (4a) and (4b)

In the following, we use multiple scale analysis to determine the amplitude equations when $|\kappa|=\kappa_{T}$. Denote $\delta$ as the controlled parameters. When the controlled parameter is larger than the critical value of Turing point, the solutions of systems (4a) and (4b) can be expanded as

$$
\begin{equation*}
c=c_{0}+\sum_{i=1}^{N}\left(A_{i} \exp \left(i \kappa_{i} \vec{r}\right)+\left(\bar{A}_{i} \exp \left(-i \kappa_{i} \vec{r}\right)\right)\right. \tag{19}
\end{equation*}
$$

with $|\kappa|=\kappa_{T} . A_{j}$ and the conjugate $\bar{A}_{j}$ are the amplitudes associated with the modes $\kappa_{j}$ and $-\kappa_{j}$.

Close to onset $\beta=\beta_{T}$, one has that

$$
\begin{equation*}
\frac{\partial A_{i}}{\partial t}=s_{i} A_{i}+F_{i}\left(A_{i}, A_{j}, \ldots\right) \tag{20}
\end{equation*}
$$

Based on the center manifold near the Turing bifurcation point, it can be concluded that amplitude $A_{j}$ satisfies

$$
\begin{equation*}
\frac{\partial A_{i}}{\partial t}=F_{i}\left(A_{i}, \bar{A}_{i}, A_{j}, \bar{A}_{j}, \ldots\right) \tag{21}
\end{equation*}
$$



Figure 2: Dispersion relation for different $\delta$. Parameters values: $\beta=$ $0.02, \gamma=1.5$, and $\varepsilon=0.15$. (a) $\delta=1.08$; (b) $\delta=1.04$; (c) $\delta=1$; (d) $\delta=0.96$; and (e) $\delta=0.92$.

In order to obtain the amplitude equations, we first need to investigate the linearized form of systems (4a) and (4b) at the equilibrium point $E_{4}$. By setting $u=u^{*}+x$ and $v=v^{*}+y$, we have the following equations:

$$
\begin{align*}
\frac{\partial x}{\partial t}= & {\left[2 \gamma u^{*}-3 \gamma\left(u^{*}\right)^{2}+2 \gamma\left(u^{*}\right)^{2} \beta-\gamma \beta-v^{*}\right] x }  \tag{22a}\\
& +\left(\beta \gamma-3 \gamma u^{*}+\gamma\right) x^{2}-\gamma x^{3}-x y+\Delta x \\
\frac{\partial y}{\partial t}= & v^{*} x+u^{*} y+x y-2 \delta v^{*} y-\delta y^{2}+\varepsilon \Delta y \tag{22b}
\end{align*}
$$

Close to onset $\delta=\delta_{T}$, the solutions of systems (4a) and (4b) can be expanded as series form:

$$
\begin{equation*}
U=U_{s}+\sum_{j=1}^{3} U_{0}\left[A_{j} \exp \left(i \kappa_{j} \vec{r}\right)+\bar{A}_{j} \exp \left(-i \kappa_{j} \vec{r}\right)\right] \tag{23}
\end{equation*}
$$

System (19) can be expanded as

$$
\begin{equation*}
U^{*}=\sum_{j=1}^{3} U_{0}\left[A_{j} \exp \left(i \kappa_{j} \vec{r}\right)+\bar{A}_{j} \exp \left(-i \kappa_{j} \vec{r}\right)\right] \tag{24}
\end{equation*}
$$

where $U_{0}=\left(\left(a_{11}^{*} \varepsilon+a_{11}^{*}\right) /\left(2 a_{21}^{*}\right), 1\right)^{T}$ is the eigenvector of the linearized operator.

From the standard multiple scale analysis, up to the third order in the perturbations, the spatiotemporal evolution of the amplitudes can be described as

$$
\begin{equation*}
\tau \frac{\partial A_{k}}{\partial t}=\mu A_{k}+\sum_{l m} h_{l m} A_{l} A_{m}+\sum_{l m n} g_{l m n} A_{l} A_{m} A_{n} \tag{25}
\end{equation*}
$$

Due to spatial translational symmetry, we have the following equation:

$$
\begin{align*}
& \tau \frac{\partial A_{k}}{\partial t} \exp \left(i \kappa_{k} r_{0}\right) \\
&=\mu A_{k} \exp \left(i \kappa_{k} r_{0}\right)+\sum_{l m} h_{l m} A_{l} A_{m} \exp \left[i\left(\kappa_{l}+\kappa_{m}\right) r_{0}\right] \\
&+\sum_{l m n} g_{l m n} A_{l} A_{m} A_{n} \exp \left[i\left(\kappa_{l}+\kappa_{m}+\kappa_{n}\right) r_{0}\right] \tag{26}
\end{align*}
$$

Comparing (25) with (26), one can find that the two equations hold only if $\kappa_{k}=\kappa_{l}+\cdots+\kappa_{m}$. From the center manifold theory, we know that amplitude equations do not include the amplitude with unstable mode. As a result, we have the following equations:

$$
\begin{align*}
\tau_{0} \frac{\partial A_{1}}{d t}= & \mu A_{1}+h \bar{A}_{2} \bar{A}_{3} \\
& -\left(g_{1}\left|A_{1}\right|^{2}+g_{2}\left(\left|A_{2}\right|^{2}+\left|A_{3}\right|^{2}\right)\right) A_{1} \\
\tau_{0} \frac{\partial A_{2}}{d t}= & \mu A_{2}+h \bar{A}_{1} \bar{A}_{3}  \tag{27}\\
& -\left(g_{1}\left|A_{2}\right|^{2}+g_{2}\left(\left|A_{1}\right|^{2}+\left|A_{3}\right|^{2}\right)\right) A_{2} \\
\tau_{0} \frac{\partial A_{3}}{d t}= & \mu A_{3}+h \bar{A}_{1} \bar{A}_{2} \\
& -\left(g_{1}\left|A_{3}\right|^{2}+g_{2}\left(\left|A_{1}\right|^{2}+\left|A_{2}\right|^{2}\right)\right) A_{3}
\end{align*}
$$

where $\mu=\left(\delta_{T}-\delta\right) / \delta_{T}$ and $\tau_{0}$ is a typical relaxation time.
In the following part, we will give the expressions of $\tau_{0}, h$, $g_{1}$, and $g_{2}$. Let

$$
\begin{align*}
X & =\binom{x}{y} \\
N & =\binom{N_{1}}{N_{2}} \tag{28}
\end{align*}
$$

Then systems (4a) and (4b) can be written as:

$$
\begin{equation*}
\frac{\partial X}{\partial t}=L X+N \tag{29}
\end{equation*}
$$

where

$$
\begin{gather*}
L=\left(\begin{array}{cc}
2 \gamma u^{*}-3 \gamma\left(u^{*}\right)^{2}+2 \gamma \beta\left(u^{*}\right)^{2}-\gamma \beta-v^{*}+\Delta & 0 \\
v^{*} & u^{*}-2 \delta v^{*}+\varepsilon \Delta
\end{array}\right) \\
N=\binom{\left(\beta \gamma-3 \gamma u^{*}+\gamma\right)-\gamma x^{3}-x y}{x y-\delta y^{2}} \tag{30}
\end{gather*}
$$

We need to investigate the dynamical behavior when $\delta$ is close to $\delta_{T}$, and thus we expand $\delta$ as:

$$
\begin{equation*}
\delta_{T}-\delta=\epsilon \delta_{1}+\epsilon^{2} \delta_{2}+\epsilon^{3} \delta_{3}+O\left(\epsilon^{4}\right) \tag{31}
\end{equation*}
$$

where $\epsilon$ is a small enough parameter. We expand $X$ and $N$ as the series form of $\epsilon$ :

$$
\begin{gather*}
X=\binom{x}{y}=\epsilon\binom{x_{1}}{y_{1}}+\epsilon^{2}\binom{x_{2}}{y_{2}}+\epsilon^{3}\binom{x_{3}}{y_{3}}+\cdots, \\
N=\binom{\left(\beta \gamma-3 \gamma u^{*}+\gamma\right)\left(x_{1}^{2} \epsilon^{2}+2 x_{1} x_{2} \epsilon^{3}\right)-\gamma x_{1}^{3} \epsilon^{3}-x_{1} y_{1} \epsilon^{2}-\left(x_{2} y_{1}+x_{1} y_{2}\right) \epsilon^{3}+o\left(\epsilon^{4}\right)}{x_{1} y_{1} \epsilon^{2}+\left(x_{2} y_{1}+x_{1} y_{2}\right) \epsilon^{3}+o\left(\epsilon^{4}\right)} . \tag{32}
\end{gather*}
$$

Linear operator $L$ can be expanded as

$$
\begin{equation*}
L=L_{T}+\left(\delta_{T}-\delta\right) M \tag{33}
\end{equation*}
$$

where

$$
L_{T}=\left(\begin{array}{cc}
a_{11}^{*}+\Delta & a_{12}^{*}  \tag{34}\\
a_{21}^{*} & a_{22}^{*}+\varepsilon \Delta
\end{array}\right), \quad M=\left(\begin{array}{ll}
b_{11} & b_{12} \\
b_{21} & b_{22}
\end{array}\right)
$$

Let

$$
\begin{equation*}
T_{0}=t, \quad T_{1}=\epsilon t, \quad T_{2}=\epsilon^{2} t \tag{35}
\end{equation*}
$$

and $T_{i}$ is a dependent variable. For the derivation of time, we have that

$$
\begin{equation*}
\frac{\partial}{\partial t}=\frac{\partial}{\partial T_{0}}+\epsilon \frac{\partial}{\partial T_{1}}+\epsilon^{2} \frac{\partial}{\partial T_{2}}+o\left(\epsilon^{3}\right) \tag{36}
\end{equation*}
$$

The solutions of systems (4a) and (4b) have the following form:

$$
\begin{equation*}
X=\binom{x}{y}=\sum_{i=1}^{3}\binom{A_{i}^{x}}{A_{i}^{y}} \exp \left(i \kappa_{i} \vec{r}\right)+\cdots \tag{37}
\end{equation*}
$$

This expression implies that the bases of the solutions have nothing to do with time and the amplitude $A$ is a variable that changes slowly. As a result, one has the following equation:

$$
\begin{equation*}
\frac{\partial A}{\partial t}=\epsilon \frac{\partial A}{\partial T_{1}}+\epsilon^{2} \frac{\partial A}{\partial T_{2}}+o\left(\epsilon^{3}\right) . \tag{38}
\end{equation*}
$$

Substituting the above equations into (29) and expanding (29) according to different orders of $\epsilon$, we can obtain three equations as follows:

$$
\begin{aligned}
& \epsilon: L_{T}\binom{x_{1}}{y_{1}}=0 \\
\epsilon^{2}: L_{T}\binom{x_{2}}{y_{2}}= & \frac{\partial}{\partial T_{1}}\binom{x_{1}}{y_{1}}-\delta_{1} M\binom{x_{1}}{y_{1}} \\
& -\binom{\left(\beta \gamma-3 \gamma u^{*}+\gamma\right) x_{1}^{2}-x_{1} y_{1}}{x_{1} y_{1}} ; \\
\epsilon^{3}: L_{T}\binom{x_{3}}{y_{3}}= & \frac{\partial}{\partial T_{1}}\binom{x_{2}}{y_{2}}+\frac{\partial}{\partial T_{2}}\binom{x_{1}}{y_{1}}-\delta_{1} M\binom{x_{2}}{y_{2}} \\
& -\delta_{2} M\binom{x_{1}}{y_{1}}-E,
\end{aligned}
$$

where

$$
\begin{equation*}
E=\binom{2 x_{1} x_{2}\left(\beta \gamma-3 \gamma u^{*}+\gamma\right)-\gamma x_{1}^{3}-\left(x_{2} y_{1}+x_{1} y_{2}\right)}{x_{2} y_{1}+x_{1} y_{2}} . \tag{40}
\end{equation*}
$$

We first consider the case of the first order of $\varepsilon$. Since $L_{T}$ is the linear operator of the system close to the onset, $\left(x_{1}, y_{1}\right)^{T}$ is the linear combination of the eigenvectors that corresponds to the eigenvalue zero. Since that

$$
\begin{equation*}
\binom{x}{y}=\sum_{i=1}^{3}\binom{A_{i}^{x}}{A_{i}^{y}} \exp \left(i \kappa_{i} \vec{r}\right)+\text { c.c. } \tag{41}
\end{equation*}
$$

we have that

$$
\begin{gather*}
\left(a_{11}^{*}+\Delta\right) x_{1}+a_{12}^{*} y_{1}=0  \tag{42a}\\
a_{21}^{*} x_{1}+\left(a_{22}^{*}+\varepsilon \Delta\right) y_{1}=0 . \tag{42b}
\end{gather*}
$$

As $\varepsilon a_{12}^{*}=\left(\left(a_{22}^{*}-\varepsilon a_{11}^{*}\right) / 2 a_{21}^{*}\right)^{2}$, we can obtain that $x_{1}=\left(a_{22}^{*}-\right.$ $\left.\varepsilon a_{11}^{*}\right) /\left(2 a_{21}^{*}\right)$ by assuming $y_{1}=1$.

Let $R=\left(a_{11}^{*} \varepsilon-a_{22}^{*}\right) / 2 a_{21}^{*}$ then

$$
\begin{gather*}
\binom{x_{1}}{y_{1}}=\binom{R}{1}\left(W_{1} \exp \left(i \kappa_{1} \vec{r}\right)+W_{2} \exp \left(i \kappa_{2} \vec{r}\right)\right.  \tag{43}\\
\left.+W_{3} \exp \left(i \kappa_{3} \vec{r}\right)\right)+ \text { c.c. }
\end{gather*}
$$

where $\left|\kappa_{j}\right|=\kappa_{T}^{*}$ and $W_{j}$ is the amplitude of the mode $\exp \left(i \kappa_{j} r\right)$.

Now, we consider the case of the second order of $\varepsilon$. Note that

$$
\begin{align*}
L_{T}\binom{x_{2}}{y_{2}}= & \frac{\partial}{\partial T_{1}}\binom{x_{1}}{y_{1}}-\delta_{T}\binom{b_{11} x_{1}+b_{12} y_{1}}{b_{21} x_{1}+b_{22} y_{1}} \\
& -\binom{\left(\beta \gamma-3 \gamma u^{*}+\gamma\right) x_{1}^{2}-x_{1} y_{1}}{x_{1} y_{1}}  \tag{44}\\
= & \binom{F_{x}}{F_{y}} .
\end{align*}
$$

According to the Fredholm solubility condition, the vector function of the right hand of the above equation must be orthogonal with the zero eigenvectors of operator $\mathbf{L}_{c}^{+}$. And the zero eigenvectors of operator $\mathbf{L}_{c}^{+}$are

$$
\begin{equation*}
\binom{1}{-\frac{1}{\varepsilon} R} \exp \left(i \kappa_{j} \vec{r}\right)+c . c \quad(j=1,2,3) . \tag{45}
\end{equation*}
$$

It can be found from the orthogonality condition that

$$
\begin{equation*}
\binom{1}{-\frac{1}{\varepsilon} R}\binom{F_{x}^{i}}{F_{y}^{i}}=0 \tag{46}
\end{equation*}
$$

where $F_{x}^{i}$ and $F_{y}^{i}$ represent the coefficients corresponding to $\exp \left(i \kappa_{j} r\right)$ in $F_{x}$ and $F_{y}$.

By investigating $\exp \left(i \kappa_{1} \vec{r}\right)$, one has

$$
\begin{align*}
\binom{F_{x}^{1}}{F_{y}^{1}}= & \binom{R \frac{\partial W_{1}}{\partial T_{1}}}{\frac{\partial W_{1}}{\partial T_{1}}}-\delta_{1}\binom{b_{11} R W_{1}+b_{12} W_{1}}{b_{21} R W_{1}+b_{22} W_{1}} \\
& -\binom{2 R^{2}\left(\beta \gamma-3 \gamma u^{*}+\gamma\right) \bar{W}_{2} \bar{W}_{3}+2 R \bar{W}_{2} \bar{W}_{3}}{2 R \bar{W}_{2} \bar{W}_{3}} \tag{47}
\end{align*}
$$

It can be obtained from the orthogonality condition that

$$
\begin{align*}
\frac{\varepsilon-1}{\varepsilon} R \frac{\partial W_{1}}{\partial T_{1}}= & \delta\left(R b_{11}+b_{22}-\frac{R}{\varepsilon}\left(R b_{21}+b_{22}\right) W_{1}\right) \\
& +2 R^{2}\left(\beta \gamma-3 \gamma u^{*}+\gamma+\frac{1}{R}-\frac{1}{\varepsilon}\right) \bar{W}_{2} \bar{W}_{3} . \tag{48}
\end{align*}
$$

By using the same methods, one has

$$
\begin{align*}
\binom{x_{2}}{y_{2}}= & \binom{X_{0}}{Y_{0}}+\sum_{j=1}^{3}\binom{X_{j}}{Y_{j}} \exp \left(i \kappa_{j} \vec{r}\right)  \tag{49}\\
& +\sum_{j=1}^{3}\binom{X_{j j}}{Y_{j j}} \exp \left(2 i \kappa_{j} \vec{r}\right)+Q+\text { c.c., }
\end{align*}
$$

where

$$
\begin{align*}
Q= & \binom{X_{12}}{Y_{12}} \exp \left(i\left(\kappa_{1}-\kappa_{2}\right) \vec{r}\right)+\binom{X_{23}}{Y_{23}} \exp \left(i\left(\kappa_{2}-\kappa_{3}\right) \vec{r}\right) \\
& +\binom{X_{31}}{Y_{31}} \exp \left(i\left(\kappa_{3}-\kappa_{1}\right) \vec{r}\right) \tag{50}
\end{align*}
$$

By solving the sets of the linear equations about $\exp (0)$, $\exp \left(i \kappa_{j} \vec{r}\right), \exp \left(2 i \kappa_{j} \vec{r}\right)$, and $\exp \left(i\left(\kappa_{j}-\kappa_{k}\right) \vec{r}\right)$, we obtain that

$$
\begin{aligned}
&\binom{X_{0}}{Y_{0}} \\
&=\binom{\frac{a_{22}^{*}\left[-2 R^{2}\left(\beta \gamma-3 \gamma u^{*}+\gamma\right)+2 R\right]+2 R a_{12}^{*}}{a_{11}^{*} a_{22}^{*}-a_{12}^{*} a_{21}^{*}}}{\frac{a_{21}^{*}\left[2 R^{2}\left(\beta \gamma-3 \gamma u^{*}+\gamma\right)-2 R\right]+2 R a_{11}^{*}}{a_{11}^{*} a_{22}^{*}-a_{12}^{*} a_{21}^{*}}} \\
& \times\left(\left|W_{1}\right|^{2}+\left|W_{2}\right|^{2}+\left|W_{3}\right|^{2}\right)
\end{aligned}
$$

$$
\binom{X_{j j}}{Y_{j j}}
$$

$$
=\binom{\frac{\left(a_{22}^{*}-4 \varepsilon \kappa_{T}^{2}\right)\left[-R^{2}\left(\beta \gamma-3 \gamma u^{*}+\gamma\right)+R\right]+R a_{12}^{*}}{\left(a_{11}^{*}-4 \kappa_{T}^{2}\right)\left(a_{22}^{*}-4 \varepsilon \kappa_{T}^{2}\right)-a_{12}^{*} a_{21}^{*}}}{\frac{a_{21}^{*}\left[R^{2}\left(\beta \gamma-3 \gamma u^{*}+\gamma\right)-2 R\right]-R\left(a_{11}^{*}-4 \varepsilon \kappa_{T}^{2}\right)}{\left(a_{11}^{*}-4 \kappa_{T}^{2}\right)\left(a_{22}^{*}-4 \varepsilon \kappa_{T}^{2}\right)-a_{12}^{*} a_{21}^{*}}}
$$

$$
\times W_{j}^{2}
$$

$$
\binom{X_{j k}}{Y_{j k}}
$$

$$
=\binom{\frac{\left(a_{22}^{*}-3 \varepsilon \kappa_{T}^{2}\right)\left[-2 R^{2}\left(\beta \gamma-3 \gamma u^{*}+\gamma\right)+2 R\right]+2 R a_{12}^{*}}{\left(a_{11}^{*}-3 \kappa_{T}^{2}\right)\left(a_{22}^{*}-3 \varepsilon \kappa_{T}^{2}\right)-a_{12}^{*} a_{21}^{*}}}{\frac{a_{21}^{*}\left[2 R^{2}\left(\beta \gamma-3 \gamma u^{*}+\gamma\right)-2 R\right]-2 R\left(a_{11}^{*}-3 \kappa_{T}^{2}\right)}{\left(a_{11}^{*}-3 \kappa_{T}^{2}\right)\left(a_{22}^{*}-3 \varepsilon \kappa_{T}^{2}\right)-a_{12}^{*} a_{21}^{*}}}
$$

$$
\begin{equation*}
\times W_{j} \bar{W}_{k} \tag{51}
\end{equation*}
$$

where $\kappa_{T}^{2}=\sqrt{\left(a_{11}^{*} a_{22}^{*}-a_{12}^{*} a_{21}^{*}\right) / \varepsilon}$.
For the third order of $\varepsilon$, we have that

$$
\begin{align*}
L_{T}\binom{x_{3}}{y_{3}}= & \frac{\partial}{\partial T_{1}}\binom{x_{2}}{y_{2}}+\frac{\partial}{\partial T_{2}}\binom{x_{1}}{y_{1}} \\
& -\delta_{1} M\binom{x_{2}}{y_{2}}-\delta_{2} M\binom{x_{1}}{y_{1}}-S \tag{52}
\end{align*}
$$

where

$$
\begin{equation*}
S=\binom{2 x_{1} x_{2}\left(\beta \gamma-3 \gamma u^{*}\right)-\gamma x_{1}^{3}-\left(x_{2} y_{1}+x_{1} y_{2}\right)}{x_{2} y_{1}+x_{1} y_{2}} \tag{53}
\end{equation*}
$$

Using the Fredholm solubility condition, we can obtain

$$
\begin{align*}
\frac{\varepsilon-1}{\varepsilon} & R \frac{\partial W_{1}}{\partial T_{2}}+\frac{\varepsilon-1}{\varepsilon} R \frac{\partial Y_{1}}{\partial T_{1}} \\
= & \delta_{2}\left[R b_{11}+b_{12}-\frac{1}{\varepsilon} R\left(R b_{21}+b_{22}\right)\right] W_{1}  \tag{54}\\
& +\delta_{1}\left[R b_{11}+b_{12}-\frac{1}{\varepsilon}\left(R b_{21}+b_{22}\right)\right] Y_{1}+Z
\end{align*}
$$

where

$$
\begin{aligned}
Z= & {\left[2 R^{2}\left(\beta \gamma-3 \gamma u^{*}+\gamma\right)-2 R-2 \frac{1}{\varepsilon}\right] } \\
& \times\left[W_{1} Y_{0}+W-2 Y_{12}+W_{3} Y_{13}+\bar{W}_{1} Y_{11}\right. \\
& \left.+\bar{W}_{2} \bar{Y}_{3}+\bar{W}_{3} \bar{Y}_{2}\right] \\
& -\left(G_{1}\left|W_{1}\right|^{2}+G_{2}\left|W_{2}\right|^{2}+G_{3}\left|W_{3}\right|^{2}\right) W_{1}
\end{aligned}
$$

Table 1: Coefficients for different parameter sets.

| $\beta$ | $\delta$ | $h$ | $g_{1}$ | $g_{2}$ | $\mu_{1}$ | $\mu_{2}$ | $\mu_{3}$ |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| 0.02 | 0.92 | -19.08604 | 7599.215 | 6906.578 | 0.0042531 | 0 | 5.770186 |
| 0.02 | 0.96 | 2.1329690 | -740.11 | -1429.72 | -0.000315 | 0 | -0.00708 |
| 0.02 | 1 | 8.4304106 | -207.521 | -474.186 | -0.015371 | 0 | -0.20741 |
| 0.02 | 1.12 | 11.304093 | -99.3194 | -193.856 | -0.0655924 | 0 | -0.00611 |



Figure 3: Spatial pattern of prey population at different time. Parameters set: $\gamma=1.5, \varepsilon=0.15$, and $\delta=0.92$. (a) $t=0$; (b) $t=100$; (c) $t=200$; (d) $t=500$; (e) $t=1000$; and (f) $t=2000$.

$$
\begin{align*}
G_{1}= & \left(\frac{1}{\varepsilon} R-1\right)\left[R\left(y_{11}+y_{0}\right)+x_{11}+x_{0}\right] \\
& -2 R\left(x_{11}+x_{0}\right)\left(\beta \gamma-3 \gamma u^{*}+\gamma\right)+3 \gamma R^{3}  \tag{55}\\
G_{2}= & \left(\frac{1}{\varepsilon} R-1\right)\left[R\left(y_{12}+y_{0}\right)+x_{12}+x_{0}\right] \\
& -2 R\left(x_{12}+x_{0}\right)\left(\beta \gamma-3 \gamma u^{*}+\gamma\right)+6 \gamma R^{3} .
\end{align*}
$$

By using the same methods, we can obtain the other two equations. The amplitude $A_{i}$ can be expanded as

$$
\begin{equation*}
A_{i}=\epsilon W_{i}+\epsilon^{2} V_{i}+o\left(\epsilon^{3}\right) \tag{56}
\end{equation*}
$$

As a result, we have

$$
\begin{equation*}
\tau_{0} \frac{\partial A_{1}}{\partial t}=\mu A_{1}+h \bar{A}_{2} \bar{A}_{3}-\left(g_{1}\left|A_{1}\right|^{2}+g_{2}\left|A_{2}\right|^{2}+\left|A_{3}\right|^{2}\right) A_{1} \tag{57}
\end{equation*}
$$

The other two equations can be obtained through the transformation of the subscript of $A$. By calculations, we obtain the expressions of the coefficients of $\tau_{0}, h, g_{1}$, and $g_{2}$ as follows:

$$
\begin{align*}
& \tau_{0}=R \frac{\varepsilon-1}{\delta_{T}\left[R b_{11}+b_{12}-(R / \varepsilon)\left(R b_{21}+b_{22}\right)\right]} \\
& h=\frac{\left[2 R^{2}\left(\beta \gamma-3 \gamma u^{*}+\gamma\right)-2 R-2\left(R^{2} / \varepsilon\right)\right]}{\delta_{T}\left[R b_{11}+b_{12}-(R / \varepsilon)\left(R b_{21}+b_{22}\right)\right]}  \tag{58}\\
& g_{1}=\frac{G_{1}}{\delta_{T}\left[R b_{11}+b_{12}-(R / \varepsilon)\left(R b_{21}+b_{22}\right)\right]} \\
& g_{2}=\frac{G_{2}}{\delta_{T}\left[R b_{11}+b_{12}-(R / \varepsilon)\left(R b_{21}+b_{22}\right)\right]}
\end{align*}
$$



Figure 4: Spatial pattern of prey population at different time. Parameters set: $\gamma=1.5, \varepsilon=0.15$, and $\delta=0.96$. (a) $t=0$; (b) $t=50$; (c) $t=100$; (d) $t=200$; (e) $t=500$; and (f) $t=1000$.
where $G_{1}=((R / \varepsilon)-1)\left[R\left(y_{0}+y_{11}\right)+x_{0}+x_{11}\right]-2 R\left(x_{0}+\right.$ $\left.x_{11}\right)\left(\beta \gamma-3 \gamma u^{*}+\gamma\right)+3 \gamma R^{3}$ and $G_{2}=((R / \varepsilon)-1)\left[R\left(y_{0}+y_{12}\right)+\right.$ $\left.x_{0}+x_{12}\right]-2 R\left(x_{0}+x_{12}\right)\left(\beta \gamma-3 \gamma u^{*}+\gamma\right)+6 \gamma R^{3}$.

By using substitutions, we have

$$
\begin{gather*}
\tau_{0} \frac{\partial \varphi}{d t}=-h \frac{\rho_{1}^{2} \rho_{2}^{2}+\rho_{1}^{2} \rho_{3}^{2}+\rho_{2}^{2} \rho_{3}^{2}}{\rho_{1} \rho_{2} \rho_{3}} \sin \varphi \\
\tau_{0} \frac{\partial \rho_{1}}{d t}=\mu \rho_{1}+h \rho_{2} \rho_{3} \cos \varphi-g_{1} \rho_{1}^{3}-g_{2}\left(\rho_{2}^{2} \rho_{3}^{2}\right) \rho_{1}  \tag{59}\\
\tau_{0} \frac{\partial \rho_{2}}{d t}=\mu \rho_{2}+h \rho_{1} \rho_{3} \cos \varphi-g_{1} \rho_{2}^{3}-g_{2}\left(\rho_{1}^{2} \rho_{3}^{2}\right) \rho_{2} \\
\tau_{0} \frac{\partial \rho_{3}}{d t}=\mu \rho_{3}+h \rho_{1} \rho_{2} \cos \varphi-g_{1} \rho_{3}^{3}-g_{2}\left(\rho_{1}^{2} \rho_{2}^{2}\right) \rho_{3}
\end{gather*}
$$

where $\varphi=\varphi_{1}+\varphi_{2}+\varphi_{3}$. In order to see the relationships between different parameters, we give the values of coefficients for different parameter sets in Table 1.

The dynamical systems (4a) and (4b) possess five kinds of solutions [23] as follows.
(1) The stationary state (O), given by

$$
\begin{equation*}
\rho_{1}=\rho_{2}=\rho_{3}=0 \tag{60}
\end{equation*}
$$

is stable for $\mu<\mu_{2}=0$ and unstable for $\mu>\mu_{2}$.
(2) Stripe patterns (S), given by

$$
\begin{equation*}
\rho_{1}=\sqrt{\frac{\mu}{g_{1}}} \neq 0, \quad \rho_{2}=\rho_{3}=0 \tag{61}
\end{equation*}
$$

are stable for $\mu>\mu_{3}=h^{2} g_{1} /\left(g_{2}-g_{1}\right)^{2}$, and unstable for $\mu<\mu_{3}$.
(3) Hexagon patterns $\left(H_{0}, H_{\pi}\right)$ are given by

$$
\begin{equation*}
\rho_{1}=\rho_{2}=\rho_{3}=\frac{|h| \pm \sqrt{h^{2}+4\left(g_{1}+2 g_{2} \mu\right)}}{2\left(g_{1}+2 g_{2}\right)} \tag{62}
\end{equation*}
$$

with $\varphi=0$ or $\pi$, and exist when

$$
\begin{equation*}
\mu>\mu_{1}=\frac{-h^{2}}{4\left(g_{1}+2 g_{2}\right)} \tag{63}
\end{equation*}
$$

The solution $\rho^{+}=|h|+\sqrt{h^{2}+4\left(g_{1}+2 g_{2} \mu\right)} / 2\left(g_{1}+\right.$ $\left.2 g_{2}\right)$ is stable only for

$$
\begin{equation*}
\mu<\mu_{4}=\frac{2 g_{1}+g_{2}}{\left(g_{2}-g_{1}\right)^{2}} h^{2} \tag{64}
\end{equation*}
$$



Figure 5: Spatial pattern of prey population at different time. Parameters set: $\gamma=1.5, \varepsilon=0.15$, and $\delta=1$. (a) $t=0$; (b) $t=150$; (c) $t=300$; (d) $t=500$; (e) $t=600$; and (f) $t=1000$.
and $\rho^{-}=\left(|h|-\sqrt{h^{2}+4\left(g_{1}+2 g_{2} \mu\right)}\right) / 2\left(g_{1}+2 g_{2}\right)$ is always unstable.
(4) The mixed states are given by

$$
\begin{equation*}
\rho_{1}=\frac{|h|}{g_{2}-g_{1}}, \quad \rho_{2}=\rho_{3}=\sqrt{\frac{\mu-g_{1} \rho_{1}^{2}}{g_{1}+g_{2}}} \tag{65}
\end{equation*}
$$

with $g_{2}>g_{1}$. They exist when $\mu>\mu_{3}$ and are always unstable.

## 4. Spatial Pattern of Systems (4a) and (4b)

In this section, we perform extensive numerical simulations of the spatially extended systems (4a) and (4b) in twodimensional spaces. All our numerical simulations employ the zero-flux boundary conditions with a system size of $200 \times$ 200. The space step is $\Delta H=1$, and the time step is $\Delta t=$ 0.00001 .

In Figure 3, we show the spatial pattern of prey population at different time. In the parameter set, $\gamma=1.5, \varepsilon=0.15$, and $\delta=0.92$, we find that $\mu \in\left(\mu_{3}, \mu_{4}\right)$, which means that there is coexistence of spotted and stripe patterns. As shown in this figure, our theoretical results are consistent with the numerical results.

By setting $\gamma=1.5, \varepsilon=0.15$, and $\delta=0.96$, one can obtain that $\mu>\mu_{4}$. In Figure 4, we show the spatial pattern of prey population when $t$ equals $0,50,100,200,500$, and 1000 . At the initial time, the prey population shows patched invasion. As time increases, stripe pattern appears and the structure does not change a lot. While keeping other parameters fixed and increasing $\delta$, we find that stripe pattern will occupy the whole space. However, some stripe patterns connect with each other and cause the emergence of spotted patterns which are shown in Figure 5.

Figure 6 shows the evolution of the spatial pattern of prey population at $t=0,100,300,500,1000$, and 2000 iterations, with small random perturbation of the stationary solution of the spatially homogeneous systems (4a) and (4b). The corresponding parameters values are $\gamma=1.5, \varepsilon=0.15$, and $\delta=1.04$. By the amplitude equations, we can conclude that there are spotted patterns of prey population for this parameter set. In this case, one can see that for the systems (4a) and (4b), the random initial distribution leads to the formation of an irregular transient pattern in the domain. After these forms, it grows slightly and spotted patterns emerge. When the time is large enough, the spotted patterns prevail over the two-dimensional space. As time further increases, the pattern structures of the prey population do not undergo any further changes.


Figure 6: Spatial pattern of prey population at different time. Parameters set: $\gamma=1.5, \varepsilon=0.15$, and $\delta=1.04$. (a) $t=0$; (b) $t=100$; (c) $t=300$; (d) $t=500$; (e) $t=1000$; and (f) $t=2000$.

## 5. Conclusion and Discussion

Allee effect has been paid much attention due to its strong potential impact on population dynamics [24]. In this paper, we investigated the pattern dynamics of a spatial predatorprey systems with Allee effect. Based on the bifurcation analysis, exact Turing pattern region is obtained. By using amplitude equations, the Turing pattern selection of the predator-prey system is well presented. It is found that the predator-prey systems with Allee effect have rich spatial dynamics by performing a series of numerical simulations.

It should be noted that our results were obtained under the assumption that predation is modeled by the bilinear function of the prey and predator densities. However, this function has limitations to describe many realistic phenomena in the biology. By numerical simulations, we find that the system exhibits similar behaviour when the functional response is of other types, such as Holling-II and Holling-III forms.

To compare the spatial dynamics for different parameters, we give the spatial patterns of population $u$ when the parameter values are out of the domain of Turing space. For this parameter set, systems (4a) and (4b) have Hopf
bifurcation, and spiral waves occupy the whole domain instead of stationary patterns, which is shown in Figure 7. The stability of spiral wave can be done by using the spectrum theory analysis [25,26]. In the further study, we will use the spectrum theory to show the stability of spiral wave.

In [15], they found that a spatial predator-prey model with Allee effect and linear death rate could increase the system's complexity and enhance chaos in population dynamics. However, in this paper, we showed that a spatial population model with Allee effect and nonlinear death rate can induce stationary patterns, which is different from the previous results.

From a biological point of view, our results show that predator mortality plays an important role in the spatial invasions of populations. More specifically, low predator mortality will induce stationary patterns (cf. Figures 36 ), and high predator mortality corresponds to travelling patterns (cf. Figure 7). When the populations exhibit wave distribution in space, the dynamics of populations may be accompanied with chaotic properties [27, 28]. If the chaotic behavior occurs, it may lead to the extinction of the population, or the population may be out of control [29, 30]. In that case, we need to find out the best way to control the chaos or change the chaotic behavior.


FIGURE 7: Spatial pattern of prey population at different time. Parameters set: $\gamma=1.5, \varepsilon=0.15$, and $\delta=1.2$. (a) $t=0$; (b) $t=100$; (c) $t=200$; (d) $t=300$; (e) $t=400$; and (f) $t=500$.

## Acknowledgments

This research was partially supported by the National Natural Science Foundation of China under Grant nos. 11301490, 11301491, 11331009, 11147015, 11171314, 11305043, and 11105024; Natural Science Foundation of Shan'xi Province Grant nos. 2012021002-1 and 2012011002-2, the Opening Foundation of Institute of Information Economy, Hangzhou Normal University, Grant no. PD12001003002003; and the specialized research fund for the doctoral program of higher education (preferential development) Grant no. 20121420130001.

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# The Analytical Solution of Some Fractional Ordinary Differential Equations by the Sumudu Transform Method 

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Received 16 March 2013; Accepted 31 July 2013
Academic Editor: Santanu Saha Ray
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#### Abstract

We introduce the rudiments of fractional calculus and the consequent applications of the Sumudu transform on fractional derivatives. Once this connection is firmly established in the general setting, we turn to the application of the Sumudu transform method (STM) to some interesting nonhomogeneous fractional ordinary differential equations (FODEs). Finally, we use the solutions to form two-dimensional (2D) graphs, by using the symbolic algebra package Mathematica Program 7.


## 1. Introduction

The Sumudu transform was first defined in its current shape by Watugala as early as 1993, which he used to solve engineering control problems. Although he might have had ideas for it sooner than that (1989) as some conference proceedings showed, he used it to control engineering problems [1, 2]. Later, Watugala extended in 2002 the Sumudu transform to two variables [3]. The first applications to differential equations and inversion formulae were done by Weerakoon in two papers in 1994 and 1998 [4, 5]. The Sumudu transform was also first defended by Weerakoon against Deakin's definition who claimed that there is no difference between the Sumudu and the Laplace and who reminded Weerakoon that the Sumudu transform is really the Carson or the S-multiplied transform disguised [6, 7]. The applications followed in three consecutive papers by Asiru dealing with the convolution-type integral equations and the discrete dynamic systems [8, 9]. At this point, Belgacem et al. using previous references and connections to the Laplace transform extended the theory and the applications of the Sumudu transform in [10-17] to various applications. In the meantime, subsequent to exchanges between Belgacem and other scholars, the following papers sprang up in the last decade [18-22]. Moreover, the Sumudu transform was
also used to solve many ordinary differential equations with integer order [23-29]. The application of STM turns out to be pragmatic in getting analytical solution of the fractional ordinary differential equations fast. Notably, implementations of difference methods such as in the differential transform method (DTM), the Adomian decomposition method (ADM) [30-33], the variational iteration method (VIM) [34-40] empowered us to achieve approximate solutions of various ordinary differential equations. STM [41-44] which is newly submitted to the literature is a suitable technique for solving various kinds of ordinary differential equations with fractional order (FODEs). In this sense, it is estimated that this novel approach that is used to solve homogeneous and nonhomogeneous problems will be particularly valuable as a tool for scientists and applied mathematicians.

## 2. Fundamental Properties of Fractional Calculus and STM

2.1. Fundamental Facts of the Fractional Calculus. Firstly, we mention some of the fundamental properties of the fractional calculus. Fractional derivatives (and integrals as well) definitions may differ, but the most widely used definitions are those of Abel-Riemann (A-R). Following the nomenclature
in [45], a derivative of fractional order in the A-R sense is defined by

$$
\begin{align*}
& D^{\alpha}[f(t)] \\
& = \begin{cases}\frac{1}{\Gamma[m-\alpha]} \frac{d}{d t^{m}} \int_{0}^{t} \frac{f(\tau)}{(t-\tau)^{\alpha-m+1}} d \tau, & m-1<\alpha \leq m \\
\frac{d^{m}}{d t^{m}} f(t), & \alpha=m,\end{cases} \tag{1}
\end{align*}
$$

where $m \in \mathbb{Z}^{+}$and $\alpha \in R^{+} . D^{\alpha}$ is a derivative operator here, and

$$
\begin{equation*}
D^{-\alpha}[f(t)]=\frac{1}{\Gamma[\alpha]} \int_{0}^{t}(t-\tau)^{\alpha-1} f(\tau) d \tau, \quad 0<\alpha \leq 1 . \tag{2}
\end{equation*}
$$

On the other hand, according to A-R, an integral of fractional order is defined by implementing the integration operator $J^{\alpha}$ in the following manner:

$$
\begin{equation*}
J^{\alpha}[f(t)]=\frac{1}{\Gamma[\alpha]} \int_{0}^{t}(t-\tau)^{\alpha-1} f(\tau) d \tau, \quad t>0, \alpha>0 \tag{3}
\end{equation*}
$$

When it comes to some of the fundamental properties of fractional integration and fractional differentiation, these have been introduced to the literature by Podlubny [46]. Among these, we mention

$$
\begin{align*}
J^{\alpha}\left[t^{n}\right] & =\frac{\Gamma[1+n]}{\Gamma[1+n+\alpha]} t^{n+\alpha}  \tag{4}\\
D^{\alpha}\left[t^{n}\right] & =\frac{\Gamma[1+n]}{\Gamma[1+n-\alpha]} t^{n-\alpha} .
\end{align*}
$$

Another main definition of the fractional derivative is that of Caputo $[46,47]$ who defined it by

$$
\begin{align*}
{ }^{C} D^{\alpha} & {[f(t)] } \\
& = \begin{cases}\frac{1}{\Gamma[m-\alpha]} \int_{0}^{t} \frac{f^{(m)}(\tau)}{(t-\tau)^{\alpha-m+1}} d \tau, & m-1<\alpha<m \\
\frac{d^{m}}{d t^{m}} f(t), & \alpha=m .\end{cases} \tag{5}
\end{align*}
$$

A fundamental feature of the Caputo fractional derivative is that [17]

$$
\begin{equation*}
J^{\alpha}\left[{ }^{C} D^{\alpha} f(t)\right]=f(t)-\sum_{k=0}^{\infty} f^{(k)}\left(0^{+}\right) \frac{t^{k}}{k!} . \tag{6}
\end{equation*}
$$

2.2. Fundamental Facts of the Sumudu Transform Method. The Sumudu transform is defined in [1,2] as follows. Over the set of functions

$$
\begin{align*}
A=\{ & f(t) \mid \exists M, \tau_{1}, \tau_{2}>0 \\
& \left.|f(t)|<M e^{|t| / \tau_{i}}, \text { if } t \in(-1)^{j} \times[0, \infty)\right\}, \tag{7}
\end{align*}
$$

the Sumudu transform of $f(t)$ is defined as

$$
\begin{equation*}
F(u)=S[f(t)]=\int_{0}^{\infty} f(u t) e^{-t} d t, \quad u \in\left(-\tau_{1}, \tau_{2}\right) \tag{8}
\end{equation*}
$$

Theorem 1. IfF $(u)$ is the Sumudu transform of $f(t)$, one knows that the Sumudu transform of the derivatives with integer order is given as follows [46-49]:

$$
\begin{equation*}
S\left[\frac{d f(t)}{d t}\right]=\frac{1}{u}[F(u)-f(0)] . \tag{9}
\end{equation*}
$$

Proof. Let us take the Sumudu transform [46-49] of $f^{\prime}(t)=$ $d f(t) / d t$ as follows:

$$
\begin{align*}
S\left[\frac{d f(t)}{d t}\right]= & \int_{0}^{\infty} \frac{d f(u t)}{d t} e^{-t} d t=\lim _{p \rightarrow \infty} \int_{0}^{p} \frac{d f(u t)}{d t} e^{-t} d t \\
= & \lim _{p \rightarrow \infty}\left[\left.\frac{1}{u} e^{-(t / u)} f(t)\right|_{0} ^{p}+\frac{1}{u^{2}} \int_{0}^{p} e^{-(t / u)} f(t) d t\right] \\
= & \lim _{p \rightarrow \infty}\left[\left.\frac{1}{u} e^{-(t / u)} f(t)\right|_{0} ^{p}\right. \\
& \left.+\frac{1}{u}\left(\frac{1}{u} \int_{0}^{p} e^{-(t / u)} f(t) d t\right)\right] \\
= & \lim _{p \rightarrow \infty}\left[-\frac{1}{u} f(0)+\frac{1}{u}\left(\frac{1}{u} \int_{0}^{p} e^{-(t / u)} f(t) d t\right)\right] \\
= & -\frac{1}{u} f(0)+\frac{1}{u} F(u) . \tag{10}
\end{align*}
$$

Equation (10) gives us the proof of Theorem 1. When we continue in the same manner, we get the Sumudu transform of the second-order derivative as follows [46-49]:

$$
\begin{equation*}
S\left[\frac{d^{2} f(t)}{d t^{2}}\right]=\frac{1}{u^{2}}\left[F(u)-f(0)-\left.u \frac{d f(t)}{d t}\right|_{t=0}\right] \tag{11}
\end{equation*}
$$

If we go on the same way, we get the Sumudu transform of the $n$-order derivative as follows:

$$
\begin{equation*}
S\left[\frac{d^{n} f(t)}{d t^{n}}\right]=u^{-n}\left[F(u)-\left.\sum_{k=0}^{n-1} u^{k} \frac{d^{n} f(t)}{d t^{n}}\right|_{t=0}\right] \tag{12}
\end{equation*}
$$

Theorem 2. If $F(u)$ is the Sumudu transform of $f(t)$, one can take into consideration the Sumudu transform of the RiemannLiouville fractional derivative as follows [17]:

$$
\begin{align*}
& S\left[D^{\alpha} f(t)\right]=u^{-\alpha}\left[F(u)-\sum_{k=1}^{n} u^{\alpha-k}\right. {\left.\left[D^{\alpha-k}(f(t))\right]_{t=0}\right] }  \tag{13}\\
&-1<n-1 \leq \alpha<n
\end{align*}
$$

Proof. Let us take the Laplace transform of $f^{\prime}(t)=d f(t) / d t$ as follows:

$$
\begin{align*}
L\left[D^{\alpha} f(t)\right] & =s^{\alpha} F(s)-\sum_{k=0}^{n-1} s^{k}\left[D^{\alpha-k-1}(f(t))\right]_{t=0}  \tag{14}\\
& =s^{\alpha} F(s)-\sum_{k=0}^{n} s^{k-1}\left[D^{\alpha-k}(f(t))\right]_{t=0} .
\end{align*}
$$

Therefore, when we substitute $1 / u$ for $s$, we get the Sumudu transform of fractional order of $f(t)$ as follows:

$$
\begin{equation*}
S\left[D^{\alpha} f(t)\right]=u^{-\alpha}\left[F(u)-\sum_{k=1}^{n} u^{\alpha-k}\left[D^{\alpha-k}(f(t))\right]_{t=0}\right] . \tag{15}
\end{equation*}
$$

Now, we will introduce the improvement form of STM for solving FODEs. We take into consideration a general linear ordinary differential equation with fractional order as follows:

$$
\begin{equation*}
\frac{\partial^{\alpha} U(t)}{\partial t^{\alpha}}=\frac{\partial^{2} U(t)}{\partial t^{2}}+\frac{\partial U(t)}{\partial t}+U(t)+c \tag{16}
\end{equation*}
$$

being subject to the initial condition

$$
\begin{equation*}
U(0)=f(0) . \tag{17}
\end{equation*}
$$

Then, we will obtain the analytical solutions of some of the fractional ordinary differential equations by using STM. When we take the Sumudu transform of (16) under the terms of (12) and (15), we obtain the Sumudu transform of (16) as follows:

$$
\begin{align*}
& S\left[\frac{\partial^{\alpha} U(t)}{\partial t^{\alpha}}\right]= S\left[\frac{\partial^{2} U(t)}{\partial t^{2}}\right]+S\left[\frac{\partial U(t)}{\partial t}\right]+S[U(t)]+S[c] \\
& u^{-\alpha}[ \left.F(u)-\sum_{k=1}^{n} u^{\alpha-k}\left[D^{\alpha-k}(U(t))\right]_{t=0}\right] \\
&=\frac{1}{u^{2}}\left[F(u)-f(0)-\left.u \frac{\partial f(t)}{\partial t}\right|_{t=0}\right] \\
&+\frac{1}{u}[F(u)-f(0)]+F(u)+c, \\
& F(u)-\sum_{k=1}^{n} u^{\alpha-k}\left[D^{\alpha-k}(U(t))\right]_{t=0} \\
&=u^{\alpha-2}\left[F(u)-f(0)-\left.u \frac{\partial U(t)}{\partial t}\right|_{t=0}\right] \\
&+u^{\alpha-1}[F(u)-f(0)] \\
&+u^{\alpha} F(u)+c u^{\alpha}, \\
& F(u)= u^{\alpha-2} F(u)-u^{\alpha-2} U(0) \\
&+\sum_{k=1}^{n} u^{\alpha-k}\left[D^{\alpha-k}(U(t))\right]_{t=0}-\left.u^{\alpha-1} \frac{\partial U(t)}{\partial t}\right|_{t=0} \\
&+u^{\alpha-1} F(u)-u^{\alpha-1} f(0) \\
&+u^{\alpha} F(u)+c u^{\alpha}, \\
& F(u)- u^{\alpha-2} F(u)-u^{\alpha-1} F(u)-u^{\alpha} F(u) \\
&=-u^{\alpha-2} f(0) \\
&+\sum_{k=1}^{n} u^{\alpha-k}\left[D^{\alpha-k}(U(t))\right]_{t=0}-\left.u^{\alpha-1} \frac{\partial U(t)}{\partial t}\right|_{t=0} \\
&-u^{\alpha-1} U(0)+c u^{\alpha}, \\
& F(u)= \frac{1}{1-u^{\alpha-2}-u^{\alpha-1}-u^{\alpha}} \\
& \times\left[g(u)-u^{\alpha-1} U(0)-u^{\alpha-2} U(0)+c u^{\alpha}\right] \tag{18}
\end{align*}
$$

where $g(u)$ is defined by $\sum_{k=1}^{n} u^{\alpha-k}\left[D^{\alpha-k}(U(t))\right]_{t=0}-$ $\left.u^{\alpha-1}(\partial U(t) / \partial t)\right|_{t=0}$. When we take the inverse Sumudu transform of (18) by using the inverse transform table in [11, 17], we get the solution of (16) by using STM as follows:

$$
\begin{align*}
U(t)=S^{-1}[ & \frac{1}{1-u^{\alpha-2}-u^{\alpha-1}-u^{\alpha}} \\
& \left.\times\left[g(u)-u^{\alpha-1} U(0)-u^{\alpha-2} U(0)+c u^{\alpha}\right]\right] \tag{19}
\end{align*}
$$

## 3. Applications of STM to Nonhomogeneous Fractional Ordinary Differential Equations

In this section, we have applied STM to the nonhomogeneous fractional ordinary differential equations as follows.

Example 3. Firstly, we consider the nonhomogeneous fractional ordinary differential equation as follows [50]:

$$
\begin{align*}
D^{\alpha}[U(t)]= & -U(t)+\frac{2}{\Gamma[3-\alpha]} t^{2-\alpha}-\frac{1}{\Gamma[2-\alpha]} t^{1-\alpha}  \tag{20}\\
& +t^{2}-t, \quad t>0, \quad 0<\alpha \leq 1
\end{align*}
$$

With the initial condition being

$$
\begin{equation*}
U(0)=0 \tag{21}
\end{equation*}
$$

In order to solve (20) by using STM, when we take the Sumudu transform of both sides of (20), we get the Sumudu transform of (20) as follows:

$$
\begin{gather*}
S\left[D^{\alpha} U(t)\right]+S[U(t)] \\
=S\left[\frac{2}{\Gamma[3-\alpha]} t^{2-\alpha}-\frac{1}{\Gamma[2-\alpha]} t^{1-\alpha}+t^{2}-t\right], \\
S\left[D^{\alpha} U(t)\right]+F(u)=S\left[\frac{2}{\Gamma[3-\alpha]} t^{2-\alpha}\right] \\
-S\left[\frac{1}{\Gamma[2-\alpha]} t^{1-\alpha}\right]+S\left[t^{2}\right]-S[t], \\
\frac{F(u)}{u^{\alpha}}-\left.\frac{D^{\alpha-1}[U(t)]}{u}\right|_{t=0}+F(u)=\frac{2}{\Gamma[3-\alpha]} S\left[t^{2-\alpha}\right] \\
-\frac{1}{\Gamma[2-\alpha]} S\left[t^{1-\alpha}\right] \\
\frac{F(u)}{u^{\alpha}}+F(u)=\frac{2}{\Gamma[3-\alpha]} u^{2-\alpha} \Gamma[3-\alpha] \\
\quad-\frac{1}{\Gamma[2-\alpha]} u^{1-\alpha} \Gamma[2-\alpha], 2 u^{2}-u, \\
\left(1+\frac{1}{u^{\alpha}}\right) F(u)=2 u^{2-\alpha}-u^{1-\alpha}+2 u^{2}-u, \\
\left(1+u^{\alpha}\right) F(u)=2 u^{2}-u+2 u^{2+\alpha}-u^{1+\alpha}, \\
\left(1+u^{\alpha}\right) F(u)=u(2 u-1)+u^{\alpha} u(2 u-1), \\
F(u)=(2 u-1) u, \\
F(u)=2 u^{2}-u .
\end{gather*}
$$

When we take the inverse Sumudu transform of (22) by using the inverse transform table in [11], we get the analytical solution of (20) by STM as follows:

$$
\begin{equation*}
U(t)=t^{2}-t . \tag{23}
\end{equation*}
$$

Remark 4. If we take the corresponding values for some parameters into consideration, then the solution of (20) is in full agreement with the solution of (30) mentioned in [50]. To our knowledge, the analytical solution of FODEs that we find in this paper has been newly submitted to the literature.

Example 5. Secondly, we consider the nonhomogeneous fractional ordinary differential equation as follows [51]:

$$
\begin{equation*}
D^{0.5} U(t)+U(t)=t^{2}+\frac{\Gamma[3]}{\Gamma[2.5]} t^{1.5}, \quad t>0 \tag{24}
\end{equation*}
$$

With the initial condition being

$$
\begin{equation*}
U(0)=0 . \tag{25}
\end{equation*}
$$

In order to solve (24) by using STM, when we take the Sumudu transform of both sides of (24), we get the Sumudu transform of (24) as follows:

$$
\begin{align*}
& S\left[D^{0.5} U(t)\right]+S[U(t)]=S\left[t^{2}\right]+\frac{\Gamma[3]}{\Gamma[2.5]} S\left[t^{1.5}\right] \\
& S\left[D^{0.5} U(t)\right]+S[U(t)]=S\left[t^{2}\right]+1.50451 S\left[t^{1.5}\right] \\
& \frac{F(u)}{u^{0.5}}-\left.\frac{D^{\alpha-1}[U(t)]}{u}\right|_{t=0}+F(u) \\
& =2 u^{2}+2 u^{1.5} \\
& \Longrightarrow \frac{F(u)}{u^{0.5}}+F(u)=2 u^{2}+2 u^{1.5}  \tag{26}\\
& \quad \Longrightarrow\left(1+u^{0.5}\right) F(u)=2 u^{2}+2 u^{1.5} \\
& \left(\frac{1+u^{0.5}}{u^{0.5}}\right) F(u)=2 u^{2}+2 u^{1.5} \\
& F(u)=\frac{2 u^{2.5}}{1+u^{0.5}}+\frac{2 u^{2}}{1+u^{0.5}}=\frac{2 u^{2}\left(1+u^{0.5}\right)}{1+u^{0.5}}=2 u^{2}
\end{align*}
$$

When we take the inverse Sumudu transform of (26) by using the inverse transform table in [48], we get the analytical solution of (24) by using STM as follows:

$$
\begin{equation*}
U(t)=t^{2} \tag{27}
\end{equation*}
$$

Remark 6. The solution (27) obtained by using the Sumudu transform method for (24) has been checked by the Mathematica Program 7. To our knowledge, the analytical solution that we find in this paper has been newly submitted to the literature.

## 4. Conclusion and Future Work

Prior to this study, various approaches have been performed to obtain approximate solutions of some fractional differential equations [50, 51]. In this paper, nonhomogeneous


Figure 1: The 2D surfaces of the obtained solution by means of STM for (23) when $0<t<3$.


Figure 2: The 2D surfaces of the obtained solution by means of STM for (27) when $0<t<3$.
fractional ordinary differential equations have been solved by using the Sumudu transform after giving the related formulae for the fractional integrals, the derivatives, and the Sumudu transform of FODEs. The Sumudu technique can be used to solve many types such as initial-value problems and boundary-value problems in applied sciences, engineering fields, aerospace sciences, and mathematical physics. The Sumudu transform method has been used for The discrete fractional calculus in [43]. This technique has been investigated in terms of the double Sumudu transform in [44]. Consequently, this new approach has been implemented with success on interesting fractional ordinary differential equations. As such and pragmatically so, it enriches the library of integral transform approaches. Without a doubt, and based on our findings such as Figures 1 and 2, the STM technique remains direct, robust and valuable tool for solving same fractional differential equations.

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## Research Article

# Improved $\left(G^{\prime} / G\right)$-Expansion Method for the Space and Time Fractional Foam Drainage and KdV Equations 

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Received 10 June 2013; Accepted 17 July 2013
Academic Editor: Santanu Saha Ray
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The fractional complex transformation is used to transform nonlinear partial differential equations to nonlinear ordinary differential equations. The improved $\left(G^{\prime} / G\right)$-expansion method is suggested to solve the space and time fractional foam drainage and KdV equations. The obtained results show that the presented method is effective and appropriate for solving nonlinear fractional differential equations.

## 1. Introduction

The soliton solutions of nonlinear evolution equations have made a major impact in the flesh. These solitons appear in various areas of physical and biological sciences. They show up in nonlinear optics, plasma physics, fluid dynamics, biochemistry, and mathematical chemistry. Fractional partial differential equations (FPDEs) have received considerable interest in recent years and have been extensively investigated. These equations were applied for many real problems which are modeled in various areas, for example, in mathematical physics [1], fluid and continuum mechanics [2], viscoplastic and viscoelastic flow [3], biology, chemistry, acoustics, and psychology $[4,5]$. Some FPDEs do not have exact solutions, so approximation and numerical techniques must be used. There are several approximation and numerical methods. The most commonly used ones are the homotopy perturbation method [6, 7], the Adomian decomposition method [ 8,9 ], the variational iteration method [10-12], the homotopy analysis method [13, 14], the generalized differential transform method [15], the finite difference method [16], and the finite element method [17]. In recent years, some authors have got exact solutions of FPDEs by using analytical methods. S. Zhang and H.-Q. Zhang [18] proposed to solve the nonlinear time fractional biological population model and $(4+1)$ dimensional space-time fractional Fokas equation by using
the fractional subequation method. Guo et al. [19] presented the improved subequation method to solve the space-time fractional Whitham-Broer-Kaup and the generalized HirotaSatsuma coupled KdV equations. Tang et al. [20] used the generalized fractional subequation method to obtain exact solutions of the space-time fractional Gardner equation with variable coefficients. Lu [21] investigated the exact solutions of the nonlinear fractional Klein-Gordon equation, the generalized time fractional Hirota-Satsuma coupled KdV system, and the nonlinear fractional Sharma-Tasso-Olver equation. Bin [22] solved the time-space fractional generalized HirotaSatsuma coupled KdV equations and the time fractional fifthorder Sawada-Kotera equation by using the $\left(G^{\prime} / G\right)$-expansion method. Omran and Gepreel [23] used the improved ( $G^{\prime} / G$ )-expansion method to calculate the exact solutions to the time-space fractional foam drainage and KdV equations. In this paper, we will apply the improved $\left(G^{\prime} / G\right)$-expansion method to obtain the exact solutions for the time-space fractional foam drainage and KdV equations with the modified Riemann-Liouville derivative defined by Jumarie [24-27]:

$$
\begin{array}{r}
\frac{\partial^{\alpha} u}{\partial t^{\alpha}}=\frac{1}{2} u \frac{\partial^{2 \beta} u}{\partial x^{2 \beta}}+2 u^{2} \frac{\partial^{\beta} u}{\partial x^{\beta}}+\left(\frac{\partial^{\beta} u}{\partial x^{\beta}}\right)^{2},  \tag{1}\\
t>0, \alpha>0, \beta \leq 1
\end{array}
$$

$$
\begin{array}{r}
\frac{\partial^{\alpha} u}{\partial t^{\alpha}}+\alpha u \frac{\partial^{\beta} u}{\partial x^{\beta}}+\frac{\partial^{3 \beta} u}{\partial x^{3 \beta}}=0,  \tag{2}\\
t>0, \alpha>0, \beta \leq 1,
\end{array}
$$

where $\alpha$ is arbitrary constant. This paper is organized as follows. In Section 2, we introduce some basic definitions of Jumarie's modified Riemann-Liouville derivative. In Section 3, the main steps of the improved $\left(G^{\prime} / G\right)$-expansion method are given. In Section 4, we construct the exact solutions of (1) and (2) by the proposed method. Some conclusions are given in Section 5.

## 2. Preliminaries

There are several definitions for fractional differential equations. These definitions include Riemann-Liouville, Weyl, Grünwald-Letnikov, Riesz, and Jumarie fractional derivatives. The Riemann-Liouville fractional derivative of a constant is not zero. So the fractional derivative is only defined for differentiable function. In order to deal with nondifferentiable functions, Jumarie [24-27] presented a modification of the Riemann-Liouville definition which appears to provide a framework for a fractional calculus. This modification was successfully applied in the probability calculus, fractional Laplace problem, exact solutions of the nonlinear fractional differential equations, and many other types of linear and nonlinear fractional differential equations [28-30].

Definition 1. The Riemann-Liouville fractional integral is defined [31] as

$$
\begin{align*}
{ }_{0} I_{x}^{\alpha} f(x)= & I^{\alpha} f(x) \\
= & \frac{1}{\Gamma(\alpha)} \int_{0}^{x} f(\xi)(x-\xi)^{\alpha-1} d \xi, \quad \alpha>0  \tag{3}\\
& \quad I^{0} f(x)=f(x)
\end{align*}
$$

Definition 2. Jumarie [24-27] defined the fractional derivative in the limit form by

$$
\begin{equation*}
f^{(\alpha)}(x)=\lim _{h \rightarrow 0} \frac{\Delta^{\alpha}[f(x)-f(0)]}{h^{\alpha}} \tag{4}
\end{equation*}
$$

where $f(x)$ should be a continuous (but not necessarily differentiable) function and $h>0$ denotes a constant discretization span. So, the modified form of the Riemann-Liouville derivative is defined as

$$
\begin{equation*}
{ }_{0} D_{x}^{\alpha} f(x)=\frac{1}{\Gamma(n-\alpha)} \frac{d^{n}}{d x^{n}} \int_{0}^{x}(x-\xi)^{(n-\alpha)}[f(\xi)-f(0)] \tag{5}
\end{equation*}
$$

where $x \in[0,1], n-1 \leq \alpha<n$ and $n \geq 1$.
Lemma 3. The integral with respect to $(d x)^{\alpha}$ is defined by Jumarie [24, 25] as follows:

$$
\begin{align*}
& \int_{0}^{x} f(\xi)(d \xi)^{\alpha}=\alpha \int_{0}^{x}(x-\xi) f(\xi) d \xi, 0<\alpha \leq 1, \\
& \frac{d^{\alpha}}{d x^{\alpha}} \int_{0}^{u(x)} f(\xi)(d \xi)^{\alpha}=\Gamma(\alpha+1) f[u(\xi)]\left[u^{\prime}(\xi)\right]^{\alpha},  \tag{6}\\
& 0<\alpha \leq 1 .
\end{align*}
$$

Theorem 4. Assume that the continuous function $f(x)$ has a fractional derivative of order $\alpha$; then

$$
\begin{gather*}
\frac{d^{\alpha}}{d x^{\alpha}} I^{\alpha} f(x)=f(x), \\
I^{\alpha} \frac{d^{\alpha}}{d x^{\alpha}} f(x)=f(x)-f(0), \quad 0<\alpha \leq 1, \tag{7}
\end{gather*}
$$

hold.

## 3. Description of the Improved $\left(G^{\prime} / G\right)$ Expansion Method

In this section, we give the description of the improved $\left(G^{\prime} / G\right)$-expansion method for solving the nonlinear FPDEs as

$$
\begin{gather*}
F\left(u, D_{t}^{\alpha} u, D_{x}^{\beta} u, D_{y}^{\gamma} u, D_{z}^{\delta} u, D_{t}^{\alpha} D_{t}^{\alpha} u,\right. \\
\left.D_{t}^{\alpha} D_{x}^{\beta} u, D_{x}^{\beta} D_{x}^{\beta} u, \ldots\right)=0  \tag{8}\\
0<\alpha, \beta, \gamma, \delta \leq 1
\end{gather*}
$$

where $u$ is an unknown function and $F$ is a polynomial of $u$ and its partial fractional derivatives, in which the highest order derivatives and nonlinear terms are involved. We offer an improved $\left(G^{\prime} / G\right)$-expansion method [32]. The essential steps of this method are described as follows.

Step 1. Li and He [33] and He and Li [34] presented a fractional complex transform to transform fractional differential equations into ordinary differential equations. So, all analytical methods devoted to advanced calculus can be easily dedicated to fractional calculus. The traveling wave variable is given as

$$
\begin{gather*}
u(x, y, z, t)=u(\xi) \\
\xi=\frac{K x^{\beta}}{\Gamma(\beta+1)}+\frac{N y^{\gamma}}{\Gamma(\gamma+1)}+\frac{L t^{\alpha}}{\Gamma(\alpha+1)} \tag{9}
\end{gather*}
$$

where $K, N$, and $L$ are nonzero arbitrary constants. So, (9) is reduced to (10):

$$
\begin{equation*}
p\left(u, u^{\prime}, u^{\prime \prime}, u^{\prime \prime \prime}, \ldots\right)=0 \tag{10}
\end{equation*}
$$

where $u=u(\xi)$.
Step 2. Suppose that (10) has the solution (11):

$$
\begin{equation*}
u(\xi)=\sum_{i=0}^{n} a_{i} F^{i}(\xi) \tag{11}
\end{equation*}
$$

where $a_{i}(i=0,1, \ldots, n)$ are real constants to be determined, the balancing number $n$ is a positive integer which can be determined by balancing the highest derivative terms with the highest power nonlinear terms in (10). More precisely, we define the degree of $u(\xi)$ as $D[u(\xi)]=m$, which gives rise to the degrees of other expressions, as follows:

$$
\begin{gather*}
D\left[\frac{d^{q} u}{d \xi^{q}}\right]=m+q, \\
D\left[u^{p}\left(\frac{d^{q} u}{d \xi^{q}}\right)^{s}\right]=m p+s(q+m) . \tag{12}
\end{gather*}
$$

Therefore, we can obtain the value of $m$ in (11).

Step 3. $F(\xi)$ is

$$
\begin{equation*}
F(\xi)=\frac{G^{\prime}(\xi)}{G(\xi)} \tag{13}
\end{equation*}
$$

where $G(\xi)$ expresses the solution of the following auxiliary ordinary differential equation

$$
\begin{equation*}
G(\xi) G^{\prime \prime}(\xi)=A G^{2}(\xi)+B G(\xi) G^{\prime}(\xi)+C[G(\xi)]^{2} \tag{14}
\end{equation*}
$$

where the prime denotes derivative with respect to $\xi$. $A, B$, and $C$ are real parameters.

Step 4. Substituting (13) into (10), using (14), collecting all terms with the same order of $\left(G^{\prime}(\xi) / G(\xi)\right)$ together, and then equating each coefficient of the resulting polynomial to zero, we obtain a set of algebraic equations for $a_{i}(i=0,1, \ldots, n)$, $A, B, C, K, N$, and $L$.

Step 5. Using the general solutions of (14), with the aid of Mathematica, we have the following four solutions of (13).

Case 1. If $B \neq 0$ and $\Delta=B^{2}+4 A-4 A C \geq 0$, then

$$
\begin{align*}
F(\xi)= & \frac{B}{2(1-C)}+\frac{B \sqrt{\Delta}}{2(1-C)} \\
& \times \frac{c_{1} \exp ((\sqrt{\Delta} / 2) \xi)+c_{2} \exp ((-\sqrt{\Delta} / 2) \xi)}{c_{1} \exp ((\sqrt{\Delta} / 2) \xi)-c_{2} \exp ((-\sqrt{\Delta} / 2) \xi)} \tag{15}
\end{align*}
$$

Case 2. If $B \neq 0$ and $\Delta=B^{2}+4 A-4 A C<0$, then

$$
\begin{align*}
F(\xi)= & \frac{B}{2(1-C)}+\frac{B \sqrt{-\Delta}}{2(1-C)} \\
& \times \frac{i c_{1} \cos ((\sqrt{-\Delta} / 2) \xi)-c_{2} \sin ((\sqrt{-\Delta} / 2) \xi)}{i c_{1} \sin ((\sqrt{-\Delta} / 2) \xi)+c_{2} \cos ((\sqrt{-\Delta} / 2) \xi)} \tag{16}
\end{align*}
$$

Case 3. If $B=0$ and $\Delta=A(C-1) \geq 0$, then

$$
\begin{equation*}
F(\xi)=\frac{\sqrt{\Delta}}{(1-C)} \frac{c_{1} \cos (\sqrt{\Delta} \xi)+c_{2} \sin (\sqrt{\Delta} \xi)}{c_{1} \sin (\sqrt{\Delta} \xi)-c_{2} \cos (\sqrt{\Delta} \xi)} \tag{17}
\end{equation*}
$$

Case 4. If $B=0$ and $\Delta=A(C-1)<0$, then

$$
\begin{equation*}
F(\xi)=\frac{\sqrt{-\Delta}}{(1-C)} \frac{i c_{1} \cosh (\sqrt{-\Delta} \xi)-c_{2} \sinh (\sqrt{-\Delta} \xi)}{i c_{1} \sinh (\sqrt{-\Delta} \xi)-c_{2} \cosh (\sqrt{-\Delta} \xi)} \tag{18}
\end{equation*}
$$

where

$$
\begin{equation*}
\xi=\frac{K x^{\beta}}{\Gamma(\beta+1)}+\frac{N y^{\gamma}}{\Gamma(\gamma+1)}+\frac{L t^{\alpha}}{\Gamma(\alpha+1)} \tag{19}
\end{equation*}
$$

and $A, B, C, c_{1}$, and $c_{2}$ are real parameters.

## 4. Applications

We use the improved $\left(G^{\prime} / G\right)$-expansion method on the timespace fractional nonlinear foam drainage equation and the time-space fractional nonlinear KdV equation in this section.
4.1. The Time and Space-Fractional Nonlinear Foam Drainage Equation. We apply the improved $\left(G^{\prime} / G\right)$-expansion method to construct the exact solutions for the time-space fractional nonlinear foam drainage equation in this subsection. Foams are of great importance in many technological processes and applications. Their properties are subject of intensive studies from practical and scientific points of view [27,35-37]. Liquid foam is an example of soft matter with a very well-defined structure, described by Joseph Plateau in the 19th century. Foams are common in foods and personal care products such as lotions and creams. They have important applications in food and chemical industries, mineral processing, fire fighting, and structural material sciences [27, 35-37]. This equation is numerically and analytically taken into account by different authors [38-40]. The space-time fractional nonlinear foam drainage equation is solved analytically only by Omran and Gepreel [23]. We can see the fractional complex transform as

$$
\begin{gather*}
u(x, t)=u(\xi) \\
\xi=\frac{K x^{\beta}}{\Gamma(\beta+1)}+\frac{L t^{\alpha}}{\Gamma(\alpha+1)} \tag{20}
\end{gather*}
$$

where $K$ and $L$ are constants. So, (20) reduces to (21):

$$
\begin{equation*}
-L u^{\prime}+\frac{1}{2} K^{2} u u^{\prime \prime}+2 K u^{2} u^{\prime}+K^{2}\left(u^{\prime}\right)^{2}=0 . \tag{21}
\end{equation*}
$$

Balancing the highest order nonlinear term and the highest order linear term, we get $n=1$. Thus, we obtain

$$
\begin{equation*}
u(\xi)=a_{0}+a_{1} F(\xi), \quad F(\xi)=\frac{G^{\prime}(\xi)}{G(\xi)} \tag{22}
\end{equation*}
$$

where $a_{0}$ and $a_{1}$ will be determined constants. Substituting (22) into (21), using (14), collecting all the terms of powers of $\left(G^{\prime} / G\right)$, and setting each coefficient to zero, we have the following system of algebraic equations:

$$
\begin{aligned}
\left(\frac{G^{\prime}}{G}\right)^{0} & : 2 A a_{0}^{2} a_{1} K+A^{2} a_{1}^{2} K^{2}+\frac{1}{2} A a_{0} a_{1} B K^{2}-A a_{1} L=0 \\
\left(\frac{G^{\prime}}{G}\right)^{1} & : 2 a_{0}^{2} a_{1} B K-A a_{0} a_{1} K^{2}+\frac{5}{2} A a_{1}^{2} B K^{2} \\
& +\frac{1}{2} a_{0} a_{1} B^{2} K^{2}+A a_{0} a_{1} C K^{2}-a_{1} B L=0
\end{aligned}
$$

$$
\left(\frac{G^{\prime}}{G}\right)^{2}:-2 a_{0}^{2} a_{1} K+2 A a_{1}^{3} K+4 a_{0} a_{1}^{2} B K
$$

$$
+2 a_{0}^{2} a_{1} C K-3 A a_{1}^{2} K^{2}+\frac{3}{2} a_{0} a_{1} B K^{2}
$$

$$
+\frac{3}{2} a_{1}^{2} B^{2} K^{2}+3 A a_{1}^{2} C K^{2}+\frac{3}{2} a_{0} a_{1} B C K^{2}
$$

$$
+\frac{3}{2} a_{1}^{2} B^{2} K^{2}+3 A a_{1}^{2} C K^{2}+\frac{3}{2} a_{0} a_{1} B C K^{2}
$$

$$
+a_{1} L-a_{1} C L=0
$$

$$
\begin{align*}
\left(\frac{G^{\prime}}{G}\right)^{3}: & -4 a_{0} a_{1}^{2} K+2 a_{1}^{3} B K+4 a_{0} a_{1}^{2} C K \\
& +a_{0} a_{1} K^{2}-\frac{7}{2} a_{1}^{2} B K^{2}+2 a_{0} a_{1} C K^{2} \\
& +\frac{7}{2} a_{1}^{2} B C K^{2}+a_{0} a_{1} C^{2} K^{2}=0 \\
\left(\frac{G^{\prime}}{G}\right)^{4}: & -2 a_{1}^{3} K+2 a_{1}^{3} C K+2 a_{1}^{2} K^{2} \\
& -4 a_{1}^{2} C K^{2}+2 a_{1}^{2} C^{2} K^{2}=0 \tag{23}
\end{align*}
$$

Solving the set of the above algebraic equations, we get the following result:

$$
\begin{gather*}
a_{0}=\frac{8 A B K(C-1)}{B^{2}+6 A-6 A C}, \quad a_{1}=K(C-1) \\
L=\frac{1}{2}\left(4 a_{0}^{2} K+a_{0} B K^{2}+2 A K^{3}-2 A C K^{3}\right)  \tag{24}\\
K B(C-1) \neq 0
\end{gather*}
$$

Substituting this value in (22) and by Cases 1-4, we obtain the following exponential, hyperbolic and triangular function solutions of (1).
(1) If we choose $B \neq 0$ and $\Delta=B^{2}+4 A-4 A C \geq 0$, then the exponential function solutions can be found as

$$
\begin{align*}
u(x, t)= & \frac{16 A B K(C-1)-K B[\Delta+2 A(C-1)]}{2 \Delta+2 A(C-1)}-\frac{K B \sqrt{\Delta}}{2} \\
& \times \frac{c_{1} \exp ((\sqrt{\Delta} / 2) \xi)+c_{2} \exp ((-\sqrt{\Delta} / 2) \xi)}{c_{1} \exp ((\sqrt{\Delta} / 2) \xi)-c_{2} \exp ((-\sqrt{\Delta} / 2) \xi)} \tag{25}
\end{align*}
$$

where

$$
\begin{align*}
\xi= & \frac{K x^{\beta}}{\Gamma(\beta+1)}+\frac{1}{2}\left(4 a_{0}^{2} K+a_{0} B K^{2}+2 A K^{3}-2 A C K^{3}\right) \\
& \times \frac{t^{\alpha}}{\Gamma(\alpha+1)} \tag{26}
\end{align*}
$$

(2) If we choose $B \neq 0$ and $\Delta=B^{2}+4 A-4 A C<0$, then the triangular function solution will be

$$
\begin{align*}
u(x, t)= & \frac{16 A B K(C-1)-K B[\Delta+2 A(C-1)]}{2 \Delta+2 A(C-1)}-\frac{K B \sqrt{-\Delta}}{2} \\
& \times \frac{i c_{1} \cos ((\sqrt{-\Delta} / 2) \xi)-c_{2} \sin ((\sqrt{-\Delta} / 2) \xi)}{i c_{1} \sin ((\sqrt{-\Delta} / 2) \xi)+c_{2} \cos ((\sqrt{-\Delta} / 2) \xi)}, \tag{27}
\end{align*}
$$

where

$$
\begin{align*}
\xi= & \frac{K x^{\beta}}{\Gamma(\beta+1)}+\frac{1}{2}\left(4 a_{0}^{2} K+a_{0} B K^{2}+2 A K^{3}-2 A C K^{3}\right)  \tag{28}\\
& \times \frac{t^{\alpha}}{\Gamma(\alpha+1)} .
\end{align*}
$$

(3) If we choose $B=0$ and $\Delta_{1}=A(C-1) \geq 0$, then we get another triangular function solution

$$
\begin{equation*}
u(x, t)=-K \sqrt{\Delta_{1}} \frac{c_{1} \cos \left(\sqrt{\Delta_{1}} \xi\right)+c_{2} \sin \left(\sqrt{\Delta_{1}} \xi\right)}{c_{1} \sin \left(\sqrt{\Delta_{1}} \xi\right)-c_{2} \cos \left(\sqrt{\Delta_{1}} \xi\right)} \tag{29}
\end{equation*}
$$

where $\xi=K x^{\beta} / \Gamma(\beta+1)-\Delta_{1}\left(K^{3} t^{\alpha} / \Gamma(\alpha+1)\right)$.
(4) If we choose $B=0$ and $\Delta_{1}=A(C-1)<0$, then we obtain the hyperbolic function solution

$$
\begin{equation*}
u(x, t)=-K \sqrt{\Delta_{1}} \frac{i c_{1} \cosh \left(\sqrt{-\Delta_{1}} \xi\right)-c_{2} \sinh \left(\sqrt{-\Delta_{1}} \xi\right)}{i c_{1} \sinh \left(\sqrt{-\Delta_{1}} \xi\right)-c_{2} \cosh \left(\sqrt{-\Delta_{1}} \xi\right)} \tag{30}
\end{equation*}
$$

$$
\text { where } \xi=K x^{\beta} / \Gamma(\beta+1)-\Delta_{1}\left(K^{3} t^{\alpha} / \Gamma(\alpha+1)\right)
$$

If we take $c_{1}=-c_{2}$ and $c_{1}=c_{2}$ in (25), respectively, then we get

$$
\begin{align*}
u(x, t)= & \frac{16 A B K(C-1)-K B[\Delta+2 A(C-1)]}{2 \Delta+2 A(C-1)} \\
& -\frac{K B \sqrt{\Delta}}{2} \tanh \left(\frac{\sqrt{\Delta}}{2} \xi\right) \\
u(x, t)= & \frac{16 A B K(C-1)-K B[\Delta+2 A(C-1)]}{2 \Delta+2 A(C-1)}  \tag{31}\\
& -\frac{K B \sqrt{\Delta}}{2} \operatorname{coth}\left(\frac{\sqrt{\Delta}}{2} \xi\right) .
\end{align*}
$$

4.2. The Nonlinear Space-Time Fractional KdV Equation. The KdV equation is the most popular soliton equation, and it has been largely investigated. In addition, the space and time fractional KdV equations with initial conditions were widely worked by [27, 38, 39]. Integrating (2) with respect to $u$ and ignoring the integral constants leads to

$$
\begin{equation*}
\frac{1}{2} L u^{2}+\frac{1}{6} a K u^{3}+\frac{1}{2} K^{3}\left(u^{\prime}\right)^{2}=0 . \tag{32}
\end{equation*}
$$

Considering the homogeneous balance between the highest order derivatives and the nonlinear term in (32), we get $n=2$. So, we can suppose that (32) has the following ansatz:

$$
\begin{equation*}
u(\xi)=a_{0}+a_{1} F(\xi)+a_{2} F^{2}(\xi) \tag{33}
\end{equation*}
$$

where $a_{0}, a_{1}, a_{2}, L$, and $K$ are arbitrary constants to be determined later. Substituting (33) and (14), along with (13),
into (32) and using Mathematica yields a system of Equations of $\left(G^{\prime} / G\right)$ :

$$
\begin{align*}
& \left(\frac{G^{\prime}}{G}\right)^{0}: \frac{1}{3} a a_{0}^{3} K+A^{2} a_{1}^{2} K^{3}+a_{0}^{2} L=0, \\
& \left(\frac{G^{\prime}}{G}\right)^{1}: \frac{1}{2} a a_{0}^{2} a_{1} K+2 A^{2} a_{1} a_{2} K^{3} \\
& +A a_{1}^{2} B K^{3}+a_{0} a_{1} L=0, \\
& \left(\frac{G^{\prime}}{G}\right)^{2}: \frac{1}{2} a a_{0} a_{1}^{2} K+\frac{1}{2} a a_{0}^{2} a_{2} K-A a_{1}^{2} K^{3} \\
& +2 A^{2} a_{2}^{2} K^{3}+4 A a_{1} a_{2} B K^{3}+\frac{1}{2 a_{1}^{2} B^{2} K^{3}} \\
& +A a_{1}^{2} C K^{3}+\frac{1}{2} a_{1}^{2} L+a_{0} a_{2} L=0, \\
& \left(\frac{G^{\prime}}{G}\right)^{3}: \frac{1}{6} a a_{1}^{3} K+a a_{0} a_{1} a_{2} K-4 A a_{1} a_{2} K^{3} \\
& -a_{1}^{2} B K^{3}+4 A a_{2}^{2} B K^{3}+2 a_{1} a_{2} B^{2} K^{3} \\
& +4 A a_{1} a_{2} C K^{3}+a_{1}^{2} B C K^{3}+a_{1} a_{2} L=0, \\
& \left(\frac{G^{\prime}}{G}\right)^{4}: \frac{1}{2} a_{1}^{2} a_{2} K+\frac{1}{2} a a_{0} a_{2}^{2} K+\frac{1}{2} a_{1}^{2} K^{3} \\
& -4 A a_{2}^{2} K^{3}-4 a_{1} a_{2} B K^{3}+2 a_{2}^{2} B^{2} K^{3} \\
& -a_{1}^{2} C K^{3}+4 a_{1} a_{2} B C K^{3}+\frac{1}{2} a_{1}^{2} C^{2} K^{3} \\
& +\frac{1}{2} a_{2}^{2} L=0, \\
& \left(\frac{G^{\prime}}{G}\right)^{5}: \frac{1}{2} a a_{1} a_{2}^{2} K+2 a_{1} a_{2} K^{3}-4 a_{2}^{2} B K^{3} \\
& -4 a_{1} a_{2} C K^{3}+4 a_{2}^{2} B C K^{3}+2 a_{1} a_{2} C^{2} K^{3}=0, \\
& \left(\frac{G^{\prime}}{G}\right)^{6}: \frac{1}{6} a_{1} a_{2}^{2} K+2 a_{2}^{2} K^{3}-4 a_{2}^{2} C K^{3} \\
& +2 a_{2}^{2} C^{2} K^{3}=0 . \tag{34}
\end{align*}
$$

Solving the set of the above algebraic equations by use of Mathematica, we get the following results:

$$
\begin{gathered}
a \neq 0, \quad a_{0}=-\frac{12 A K^{2}}{a}(C-1), \\
a_{1}=\frac{12 B K^{2}}{a}(C-1), \\
a_{2}=-\frac{12 K^{2}}{a}(C-1), \\
L=-K^{3}\left(B^{2}-4 A C+4 A\right) .
\end{gathered}
$$

Substituting (35) into (33) and according to (15)-(18), we obtain the following exponential function solutions, hyperbolic function solutions, and triangular function solutions of (2), respectively.
(1) If we choose $B \neq 0$ and $\Delta=B^{2}+4 A-4 A C \geq 0$, then the exponential function solution can be found as

$$
\begin{align*}
u(x, t)= & \frac{3 K^{2}(\Delta-2 C)}{a(C-1)}-\frac{6 C K^{2} \sqrt{\Delta}}{a(C-1)} \\
& \times \frac{c_{1} \exp (\sqrt{\Delta} \xi / 2)+c_{2} \exp ((-\sqrt{\Delta} / 2) \xi)}{c_{1} \exp ((\sqrt{\Delta} / 2) \xi)-c_{2} \exp ((-\sqrt{\Delta} / 2) \xi)} \\
& -\frac{6 C K^{2} \Delta}{a(C-1)} \\
& \times\left[\frac{c_{1} \exp ((\sqrt{\Delta} / 2) \xi)+c_{2} \exp (-(\sqrt{\Delta} / 2) \xi)}{c_{1} \exp ((\sqrt{\Delta} / 2) \xi)-c_{2} \exp (-(\sqrt{\Delta} / 2) \xi)}\right]^{2} \tag{36}
\end{align*}
$$

where $\xi=K x^{\beta} / \Gamma(\beta+1)-K^{3} \Delta\left(t^{\alpha} / \Gamma(\alpha+1)\right)$.
(2) If we choose $B \neq 0$ and $\Delta=B^{2}+4 A-4 A C<0$, then the triangular function solution will be

$$
\begin{align*}
u(x, t)= & \frac{3 K^{2}(\Delta-2 C)}{a(C-1)}+\frac{6 C K^{2} \sqrt{\Delta}}{a(C-1)} \\
& \times \frac{i c_{1} \cos (\sqrt{-\Delta} \xi / 2)-c_{2} \sin ((\sqrt{-\Delta} / 2) \xi)}{i c_{1} \sin ((\sqrt{-\Delta} / 2) \xi)+c_{2} \cos ((\sqrt{-\Delta} / 2) \xi)} \\
& +\frac{6 C K^{2} \Delta}{a(C-1)} \\
& \times\left[\frac{i c_{1} \cos ((\sqrt{-\Delta} / 2) \xi)-c_{2} \sin ((\sqrt{-\Delta} / 2) \xi)}{i c_{1} \sin ((\sqrt{-\Delta} / 2) \xi)+c_{2} \cos ((\sqrt{-\Delta} / 2) \xi)}\right]^{2} \tag{37}
\end{align*}
$$

where $\xi=K x^{\beta} / \Gamma(\beta+1)-K^{3} \Delta\left(t^{\alpha} / \Gamma(\alpha+1)\right)$.
(3) If we choose $B=0$ and $\Delta_{1}=A(C-1) \geq 0$, then the triangular function solution is given as

$$
\begin{align*}
u(x, t)= & -\frac{12 K^{2} \Delta_{1}}{a}-\frac{12 K^{2} \Delta_{1}}{a(C-1)} \\
& \times\left[\frac{c_{1} \cos \left(\sqrt{\Delta_{1}} \xi\right)+c_{2} \sin \left(\sqrt{\Delta_{1}} \xi\right)}{c_{1} \sin \left(\sqrt{\Delta_{1}} \xi\right)-c_{2} \cos \left(\sqrt{\Delta_{1}} \xi\right)}\right]^{2} \tag{38}
\end{align*}
$$

(4) If we choose $B=0$ and $\Delta_{1}=A(C-1)<0$, then the hyperbolic function solution is given as

$$
\begin{align*}
u(x, t)= & -\frac{12 K^{2} \Delta_{1}}{a}+\frac{12 K^{2} \Delta_{1}}{a(C-1)} \\
& \times\left[\frac{i c_{1} \cosh \left(\sqrt{-\Delta_{1}} \xi\right)-c_{2} \sinh \left(\sqrt{-\Delta_{1}} \xi\right)}{i c_{1} \sinh \left(\sqrt{-\Delta_{1}} \xi\right)-c_{2} \cosh \left(\sqrt{-\Delta_{1}} \xi\right)}\right]^{2} \tag{39}
\end{align*}
$$

where $\xi=K x^{\beta} / \Gamma(\beta+1)-4 K^{3} \Delta_{1}\left(t^{\alpha} / \Gamma(\alpha+1)\right)$. Equation (36) can be rewritten at $c_{1}=-c_{2}$; so we get the other hyperbolic function solution of (2):

$$
\begin{align*}
u(x, t)= & \frac{3 K^{2}(\Delta-2 C)}{a(C-1)}-\frac{6 C K^{2} \sqrt{\Delta}}{a(C-1)} \\
& \times \tanh \left[\frac{\sqrt{\Delta}}{2}\left(\frac{K x^{\beta}}{\Gamma(\beta+1)}-K^{3} \Delta \frac{t^{\alpha}}{\Gamma(\alpha+1)}\right)\right] \\
& +\frac{6 C K^{2} \sqrt{\Delta}}{a(C-1)} \\
& \times \tanh ^{2}\left[\frac{\sqrt{\Delta}}{2}\left(\frac{K x^{\beta}}{\Gamma(\beta+1)}-K^{3} \Delta \frac{t^{\alpha}}{\Gamma(\alpha+1)}\right)\right] . \tag{40}
\end{align*}
$$

Equation (36) becomes

$$
\begin{align*}
u(x, t)= & \frac{3 K^{2}(\Delta-2 C)}{a(C-1)}-\frac{6 C K^{2} \sqrt{\Delta}}{a(C-1)} \\
& \times \operatorname{coth}\left[\frac{\sqrt{\Delta}}{2}\left(\frac{K x^{\beta}}{\Gamma(\beta+1)}-K^{3} \Delta \frac{t^{\alpha}}{\Gamma(\alpha+1)}\right)\right] \\
& +\frac{6 C K^{2} \sqrt{\Delta}}{a(C-1)} \\
& \times \operatorname{coth}^{2}\left[\frac{\sqrt{\Delta}}{2}\left(\frac{K x^{\beta}}{\Gamma(\beta+1)}-K^{3} \Delta \frac{t^{\alpha}}{\Gamma(\alpha+1)}\right)\right] \tag{41}
\end{align*}
$$

at $c_{1}=c_{2}$.
Remark 5. Kudryashov et al. [41-44] have showed that every solution, which was obtained when soliton solutions have been found by some analytic methods, is not a new solution. They also showed that these methods are very similar. Furthermore, they mentioned that authors who used these methods should check the obtained results very carefully. The reason for using improved $\left(G^{\prime} / G\right)$-expansion method in this work is to use nonlinear equation (14) instead of linear equation

$$
\begin{equation*}
G^{\prime \prime}-\lambda G^{\prime}-\mu G=0 \tag{42}
\end{equation*}
$$

which was used in standard $\left(G^{\prime} / G\right)$ method and to obtain lots of different solutions.

## 5. Conclusion

In this paper, we introduced an improved $\left(G^{\prime} / G\right)$-expansion method and carried it out to obtain new travelling wave solutions of the space-time fractional foam drainage equation and the space-time fractional KdV equation. This method gives new exact solutions for nonlinear FPDEs. These solutions include the hyperbolic function solution, the exponential function solution, the triangular function solution, and the trigonometric function solution. These solutions are useful to understand the mechanisms of the complicated nonlinear physical phenomena.

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## Research Article

# The Solution to the BCS Gap Equation for Superconductivity and Its Temperature Dependence 

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Received 27 May 2013; Accepted 9 August 2013
Academic Editor: Santanu Saha Ray
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#### Abstract

From the viewpoint of operator theory, we deal with the temperature dependence of the solution to the BCS gap equation for superconductivity. When the potential is a positive constant, the BCS gap equation reduces to the simple gap equation. We first show that there is a unique nonnegative solution to the simple gap equation, that it is continuous and strictly decreasing, and that it is of class $C^{2}$ with respect to the temperature. We next deal with the case where the potential is not a constant but a function. When the potential is not a constant, we give another proof of the existence and uniqueness of the solution to the BCS gap equation, and show how the solution varies with the temperature. We finally show that the solution to the BCS gap equation is indeed continuous with respect to both the temperature and the energy under a certain condition when the potential is not a constant.


## 1. Introduction

We use the unit $k_{B}=1$, where $k_{B}$ stands for the Boltzmann constant. Let $\omega_{D}>0$ and $k \in \mathbb{R}^{3}$ stand for the Debye frequency and the wave vector of an electron, respectively. Let $h>0$ be Planck's constant, and set $\hbar=h /(2 \pi)$. Let $m>0$ and $\mu>0$ stand for the electron mass and the chemical potential, respectively. We denote by $T(\geq 0)$ the absolute temperature, and by $x$ the kinetic energy of an electron minus the chemical potential; that is, $x=\hbar^{2}|k|^{2} /(2 m)-\mu$. Note that $0<\hbar \omega_{D} \ll \mu$.

In the BCS model [1,2] of superconductivity, the solution to the BCS gap equation (1) is called the gap function. The gap function corresponds to the energy gap between the superconducting ground state and the superconducting first excited state. Accordingly, the value of the gap function (the solution) is nonnegative. We regard the gap function as a function of both $T$ and $x$ and denote it by $u$; that is, $u$ : $(T, x) \mapsto u(T, x)(\geq 0)$. The BCS gap equation is the following nonlinear integral equation $\left(0<\varepsilon \leq x \leq \hbar \omega_{D}\right)$ :

$$
\begin{align*}
u(T, x)=\int_{\varepsilon}^{\hbar \omega_{D}} & \frac{U(x, \xi) u(T, \xi)}{\sqrt{\xi^{2}+u(T, \xi)^{2}}}  \tag{1}\\
& \times \tanh \frac{\sqrt{\xi^{2}+u(T, \xi)^{2}}}{2 T} d \xi
\end{align*}
$$

where $U(\cdot, \cdot)>0$ is the potential multiplied by the density of states per unit energy at the Fermi surface and is a function of $x$ and $\xi$. In (1) we introduce $\varepsilon>0$, which is small enough and fixed $\left(0<\varepsilon \ll \hbar \omega_{D}\right)$. In the original BCS model, the integration interval is $\left[0, \hbar \omega_{D}\right]$; it is not $\left[\varepsilon, \hbar \omega_{D}\right]$. However, we introduce very small $\varepsilon>0$ for the following mathematical reasons. In order to show the continuity of the solution to the BCS gap equation with respect to the temperature and in order to show that the transition to a superconducting state is a second-order phase transition, we make the form of the BCS gap equation somewhat easier to handle. So we choose the closed interval $\left[\varepsilon, \hbar \omega_{D}\right]$ as the integration interval in (1).

The integral with respect to $\xi$ in (1) is sometimes replaced by the integral over $\mathbb{R}^{3}$ with respect to the wave vector $k$. Odeh [3] and Billard and Fano [4] established the existence and uniqueness of the positive solution to the BCS gap equation in the case $T=0$. For $T \geq 0$, Vansevenant [5] determined the transition temperature (the critical temperature) and showed that there is a unique positive solution to the BCS gap equation. Recently, Frank et al. [6] gave a rigorous analysis of the asymptotic behavior of the transition temperature at weak coupling. Hainzl et al. [7] proved that the existence of a positive solution to the BCS gap equation is equivalent to the existence of a negative eigenvalue of a certain linear operator to show the existence of a transition temperature. Moreover,

Hainzl and Seiringer [8] derived upper and lower bounds on the transition temperature and the energy gap for the BCS gap equation.

Since the existence and uniqueness of the solution were established for each fixed $T$ in the previous literature, the temperature dependence of the solution is not covered. It is well known that studying the temperature dependence of the solution to the BCS gap equation is very important in condensed matter physics. This is because, by dealing with the thermodynamical potential, this study leads to a mathematical proof of the statement that the transition to a superconducting state is a second-order phase transition. So, in condensed matter physics, it is highly desirable to study the temperature dependence of the solution to the BCS gap equation.

When the potential $U(\cdot, \cdot)$ in (1) is a positive constant, the BCS gap equation reduces to the simple gap equation (3). In this case, one assumes in the BCS model that there is a unique nonnegative solution to the simple gap equation (3) and that the solution is of class $C^{2}$ with respect to the temperature $T$ (see e.g., [1] and [9, (11.45), page 392]). In this paper, applying the implicit function theorem, we first show that this assumption of the BCS model indeed holds true; we show that there is a unique nonnegative solution to the simple gap equation (3) and that the solution is of class $C^{2}$ with respect to the temperature $T$. We next deal with the case where the potential is not a constant but a function. In order to show how the solution varies with the temperature, we then give another proof of the existence and uniqueness of the solution to the BCS gap equation (1) when the potential is not a constant. More precisely, we show that the solution belongs to the subset $V_{T}$ (see (12)). Note that the subset $V_{T}$ depends on $T$. We finally show that the solution to the BCS gap equation (1) is indeed continuous with respect to both $T$ and $x$ when $T$ satisfies (20) when the potential is not a constant.

Let

$$
\begin{equation*}
U(x, \xi)=U_{1} \quad \text { at all }(x, \xi) \in\left[\varepsilon, \hbar \omega_{D}\right]^{2} \tag{2}
\end{equation*}
$$

where $U_{1}>0$ is a constant. Then the gap function depends on the temperature $T$ only. So we denote the gap function by $\Delta_{1}$ in this case; that is, $\Delta_{1}: T \mapsto \Delta_{1}(T)$. Then (1) leads to the simple gap equation

$$
\begin{equation*}
1=U_{1} \int_{\varepsilon}^{\hbar \omega_{D}} \frac{1}{\sqrt{\xi^{2}+\Delta_{1}(T)^{2}}} \tanh \frac{\sqrt{\xi^{2}+\Delta_{1}(T)^{2}}}{2 T} d \xi \tag{3}
\end{equation*}
$$

The following is the definition of the temperature $\tau_{1}>0$.
Definition 1 (see [1]). Consider

$$
\begin{equation*}
1=U_{1} \int_{\varepsilon}^{\hbar \omega_{D}} \frac{1}{\xi} \tanh \frac{\xi}{2 \tau_{1}} d \xi \tag{4}
\end{equation*}
$$

## 2. The Simple Gap Equation

Set

$$
\begin{equation*}
\Delta=\frac{\sqrt{\left(\hbar \omega_{D}-\varepsilon e^{1 / U_{1}}\right)\left(\hbar \omega_{D}-\varepsilon e^{-1 / U_{1}}\right)}}{\sinh \left(1 / U_{1}\right)} \tag{5}
\end{equation*}
$$

Proposition 2 (see [10, Proposition 2.2]). Let $\Delta$ be as in (5). Then there is a unique nonnegative solution $\Delta_{1}:\left[0, \tau_{1}\right] \rightarrow$ $[0, \infty)$ to the simple gap equation (3) such that the solution $\Delta_{1}$ is continuous and strictly decreasing on the closed interval $\left[0, \tau_{1}\right]$ :

$$
\begin{align*}
\Delta_{1}(0) & =\Delta>\Delta_{1}\left(T_{1}\right)>\Delta_{1}\left(T_{2}\right)  \tag{6}\\
& >\Delta_{1}\left(\tau_{1}\right)=0, \quad 0<T_{1}<T_{2}<\tau_{1}
\end{align*}
$$

Moreover, the solution $\Delta_{1}$ is of class $C^{2}$ on the interval $\left[0, \tau_{1}\right)$ and satisfies

$$
\begin{equation*}
\Delta_{1}^{\prime}(0)=\Delta_{1}^{\prime \prime}(0)=0, \quad \lim _{T \uparrow \tau_{1}} \Delta_{1}^{\prime}(T)=-\infty \tag{7}
\end{equation*}
$$

Proof. Setting $Y=\Delta_{1}(T)^{2}$ turns (3) into

$$
\begin{equation*}
1=U_{1} \int_{\varepsilon}^{\hbar \omega_{D}} \frac{1}{\sqrt{\xi^{2}+Y}} \tanh \frac{\sqrt{\xi^{2}+Y}}{2 T} d \xi \tag{8}
\end{equation*}
$$

Note that the right side is a function of the two variables $T$ and $Y$. We see that there is a unique function $T \mapsto Y$ defined by (8) implicitly, that the function $T \mapsto Y$ is continuous and strictly decreasing on $\left[0, \tau_{1}\right]$, and that $Y=0$ at $T=\tau_{1}$. We moreover see that the function $T \mapsto Y$ is of class $C^{2}$ on the closed interval $\left[0, \tau_{1}\right]$.

Remark 3. We set $\Delta_{1}(T)=0$ for $T>\tau_{1}$.
Remark 4. In Proposition 2, $\Delta_{1}(T)$ is nothing but $\sqrt{f(T)}$ in [10, Proposition 2.2].

We introduce another positive constant $U_{2}>0$. Let $0<$ $U_{1}<U_{2}$. We assume the following condition on $U(\cdot, \cdot)$ :

$$
\begin{align*}
U_{1} & \leq U(x, \xi) \\
& \leq U_{2} \quad \text { at all }(x, \xi) \in\left[\varepsilon, \hbar \omega_{D}\right]^{2}, U(\cdot, \cdot) \in C\left(\left[\varepsilon, \hbar \omega_{D}\right]^{2}\right) \tag{9}
\end{align*}
$$

When $U(x, \xi)=U_{2}$ at all $(x, \xi) \in\left[\varepsilon, \hbar \omega_{D}\right]^{2}$, an argument similar to that in Proposition 2 gives that there is a unique nonnegative solution $\Delta_{2}:\left[0, \tau_{2}\right] \rightarrow[0, \infty)$ to the simple gap equation

$$
\begin{align*}
1=U_{2} \int_{\varepsilon}^{\hbar \omega_{D}} & \frac{1}{\sqrt{\xi^{2}+\Delta_{2}(T)^{2}}}  \tag{10}\\
& \times \tanh \frac{\sqrt{\xi^{2}+\Delta_{2}(T)^{2}}}{2 T} d \xi, \quad 0 \leq T \leq \tau_{2}
\end{align*}
$$

Here, $\tau_{2}>0$ is defined by

$$
\begin{equation*}
1=U_{2} \int_{\varepsilon}^{\hbar \omega_{D}} \frac{1}{\xi} \tanh \frac{\xi}{2 \tau_{2}} d \xi \tag{11}
\end{equation*}
$$

We again set $\Delta_{2}(T)=0$ for $T>\tau_{2}$. A straightforward calculation gives the following.


Figure 1: The graphs of the functions $\Delta_{1}$ and $\Delta_{2}$.

Lemma 5 ([11, Lemma 1.5]). (a) The inequality $\tau_{1}<\tau_{2}$ holds. (b) If $0 \leq T<\tau_{2}$, then $\Delta_{1}(T)<\Delta_{2}(T)$. If $T \geq \tau_{2}$, then $\Delta_{1}(T)=\Delta_{2}(T)=0$.

Note that Proposition 2 and Lemma 5 point out how $\Delta_{1}$ and $\Delta_{2}$ depend on the temperature and how $\Delta_{1}$ and $\Delta_{2}$ vary with the temperature; see Figure 1.

Remark 6. On the basis of Proposition 2, the present author [10, Theorem 2.3] proved that the transition to a superconducting state is a second-order phase transition under the restriction (2).

## 3. The BCS Gap Equation

Let $0 \leq T \leq \tau_{2}$ and fix $T$, where $\tau_{2}$ is that in (11). We consider the Banach space $C\left[\varepsilon, \hbar \omega_{D}\right]$ consisting of continuous functions of $x$ only and deal with the following subset $V_{T}$ :

$$
\begin{align*}
& V_{T}=\left\{u(T, \cdot) \in C\left[\varepsilon, \hbar \omega_{D}\right]: \Delta_{1}(T)\right. \\
&\left.\leq u(T, x) \leq \Delta_{2}(T) \text { at } x \in\left[\varepsilon, \hbar \omega_{D}\right]\right\} . \tag{12}
\end{align*}
$$

Remark 7. The subset $V_{T}$ depends on $T$. So we denote each element of $V_{T}$ by $u(T, \cdot)$; see Figure 1.

As it is mentioned in the introduction, the existence and uniqueness of the solution to the BCS gap equation were established for each fixed $T$ in the previous literature, and the temperature dependence of the solution is not covered. We therefore give another proof of the existence and uniqueness of the solution to the BCS gap equation (1) so as to show how the solution varies with the temperature. More precisely, we show that the solution belongs to $V_{T}$.

Theorem 8 (see [11, Theorem 2.2]). Assume condition (9) on $U(\cdot, \cdot)$. Let $T \in\left[0, \tau_{2}\right]$ be fixed. Then there is a unique


Figure 2: For each $T$, the solution $u_{0}(T, x)$ lies between $\Delta_{1}(T)$ and $\Delta_{2}(T)$.
nonnegative solution $u_{0}(T, \cdot) \in V_{T}$ to the BCS gap equation (1) $\left(x \in\left[\varepsilon, \hbar \omega_{D}\right]\right)$ :

$$
\begin{align*}
u_{0}(T, x)=\int_{\varepsilon}^{\hbar \omega_{D}} & \frac{U(x, \xi) u_{0}(T, \xi)}{\sqrt{\xi^{2}+u_{0}(T, \xi)^{2}}}  \tag{13}\\
& \times \tanh \frac{\sqrt{\xi^{2}+u_{0}(T, \xi)^{2}}}{2 T} d \xi
\end{align*}
$$

Consequently, the solution is continuous with respect to $x$ and varies with the temperature as follows:

$$
\begin{align*}
\Delta_{1}(T) & \leq u_{0}(T, x)  \tag{14}\\
& \leq \Delta_{2}(T) \quad \text { at }(T, x) \in\left[0, \tau_{2}\right] \times\left[\varepsilon, \hbar \omega_{D}\right]
\end{align*}
$$

Proof. We define a nonlinear integral operator $A$ on $V_{T}$ by

$$
\begin{align*}
A u(T, x)=\int_{\varepsilon}^{\hbar \omega_{D}} & \frac{U(x, \xi) u(T, \xi)}{\sqrt{\xi^{2}+u(T, \xi)^{2}}}  \tag{15}\\
& \times \tanh \frac{\sqrt{\xi^{2}+u(T, \xi)^{2}}}{2 T} d \xi
\end{align*}
$$

where $u(T, \cdot) \in V_{T}$. Clearly, $V_{T}$ is a bounded, closed, and convex subset of the Banach space $C\left[\varepsilon, \hbar \omega_{D}\right]$. A straightforward calculation gives that the operator $A: V_{T} \rightarrow V_{T}$ is compact. Therefore, the Schauder fixed point theorem applies, and hence the operator $A: V_{T} \rightarrow V_{T}$ has at least one fixed point $u_{0}(T, \cdot) \in V_{T}$. Moreover, we can show the uniqueness of the fixed point; see Figure 2.

The existence of the transition temperature $T_{c}$ is pointed out in the previous papers [5-8]. In our case, it is defined as follows.

Definition 9. Let $u_{0}(T, \cdot) \in V_{T}$ be as in Theorem 8. The transition temperature $T_{c}$ stemming from the BCS gap equation (1) is defined by

$$
\begin{equation*}
T_{c}=\inf \left\{T>0: u_{0}(T, x)=0 \text { at all } x \in\left[\varepsilon, \hbar \omega_{D}\right]\right\} \tag{16}
\end{equation*}
$$

Remark 10. Combining Definition 9 with Theorem 8 implies that $\tau_{1} \leq T_{c} \leq \tau_{2}$. For $T>T_{c}$, we set $u_{0}(T, x)=0$ at all $x \in\left[\varepsilon, \hbar \omega_{D}\right]$.

## 4. Continuity of the Solution with respect to the Temperature

Let $U_{0}>0$ be a constant satisfying $U_{0}<U_{1}<U_{2}$. An argument similar to that in Proposition 2 gives that there is a unique nonnegative solution $\Delta_{0}:\left[0, \tau_{0}\right] \rightarrow[0, \infty)$ to the simple gap equation

$$
\begin{align*}
1=U_{0} \int_{\varepsilon}^{\hbar \omega_{D}} & \frac{1}{\sqrt{\xi^{2}+\Delta_{0}(T)^{2}}}  \tag{17}\\
& \times \tanh \frac{\sqrt{\xi^{2}+\Delta_{0}(T)^{2}}}{2 T} d \xi, \quad 0 \leq T \leq \tau_{0} .
\end{align*}
$$

Here, $\tau_{0}>0$ is defined by

$$
\begin{equation*}
1=U_{0} \int_{\varepsilon}^{\hbar \omega_{D}} \frac{1}{\xi} \tanh \frac{\xi}{2 \tau_{0}} d \xi \tag{18}
\end{equation*}
$$

We set $\Delta_{0}(T)=0$ for $T>\tau_{0}$. A straightforward calculation gives the following.

Lemma 11. (a) $\tau_{0}<\tau_{1}<\tau_{2}$.
(b) If $0 \leq T<\tau_{0}$, then $0<\Delta_{0}(T)<\Delta_{1}(T)<\Delta_{2}(T)$.
(c) If $\tau_{0} \leq T<\tau_{1}$, then $0=\Delta_{0}(T)<\Delta_{1}(T)<\Delta_{2}(T)$.
(d) If $\tau_{1} \leq T<\tau_{2}$, then $0=\Delta_{0}(T)=\Delta_{1}(T)<\Delta_{2}(T)$.
(e) If $\tau_{2} \leq T$, then $0=\Delta_{0}(T)=\Delta_{1}(T)=\Delta_{2}(T)$.

Remark 12. Let the functions $\Delta_{k}(k=0,1,2)$ be as above. For each $\Delta_{k}$, there is the inverse $\Delta_{k}^{-1}:\left[0, \Delta_{k}(0)\right] \rightarrow\left[0, \tau_{k}\right]$. Here,

$$
\begin{equation*}
\Delta_{k}(0)=\frac{\sqrt{\left(\hbar \omega_{D}-\varepsilon e^{1 / U_{k}}\right)\left(\hbar \omega_{D}-\varepsilon e^{-1 / U_{k}}\right)}}{\sinh \left(1 / U_{k}\right)} \tag{19}
\end{equation*}
$$

and $\Delta_{0}(0)<\Delta_{1}(0)<\Delta_{2}(0)$.
We introduce another temperature. Let $T_{1}$ satisfy $0<$ $T_{1}<\Delta_{0}^{-1}\left(\Delta_{0}(0) / 2\right)$ and

$$
\begin{align*}
& \frac{\Delta_{0}(0)}{4 \Delta_{2}^{-1}\left(\Delta_{0}\left(T_{1}\right)\right)} \tanh \frac{\Delta_{0}(0)}{4 \Delta_{2}^{-1}\left(\Delta_{0}\left(T_{1}\right)\right)} \\
& \quad>\frac{1}{2}\left(1+\frac{4 \hbar^{2} \omega_{D}^{2}}{\Delta_{0}(0)^{2}}\right) \tag{20}
\end{align*}
$$

Remark 13. Numerically, the temperature $T_{1}$ is very small.
Consider the following subset $V$ of the Banach space $C\left(\left[0, T_{1}\right] \times\left[\varepsilon, \hbar \omega_{D}\right]\right)$ consisting of continuous functions of both the temperature $T$ and the energy $x$ :

$$
\begin{align*}
V=\{u & \in C\left(\left[0, T_{1}\right] \times\left[\varepsilon, \hbar \omega_{D}\right]\right): \Delta_{1}(T) \leq u(T, x) \\
& \left.\leq \Delta_{2}(T) \text { at }(T, x) \in\left[0, T_{1}\right] \times\left[\varepsilon, \hbar \omega_{D}\right]\right\} \tag{21}
\end{align*}
$$



Figure 3: The solution $u_{0}$ is continuous on $\left[0, T_{1}\right] \times\left[\varepsilon, \hbar \omega_{D}\right]$.

Theorem 14 (see [12, Theorem 1.2]). Assume (9). Let $u_{0}$ be as in Theorem 8 and $V$ as in (21). Then $u_{0} \in V$. Consequently, the gap function $u_{0}$ is continuous on $\left[0, T_{1}\right] \times\left[\varepsilon, \hbar \omega_{D}\right]$.

Proof. We define a nonlinear integral operator $B$ on $V$ by

$$
\begin{align*}
B u(T, x)=\int_{\varepsilon}^{\hbar \omega_{D}} & \frac{U(x, \xi) u(T, \xi)}{\sqrt{\xi^{2}+u(T, \xi)^{2}}}  \tag{22}\\
& \quad \times \tanh \frac{\sqrt{\xi^{2}+u(T, \xi)^{2}}}{2 T} d \xi
\end{align*}
$$

where $u \in V$.
Clearly, $V$ is a closed subset of the Banach space $C\left(\left[0, T_{1}\right] \times\left[\varepsilon, \hbar \omega_{D}\right]\right)$. A straightforward calculation gives that the operator $B: V \rightarrow V$ is contractive as long as (20) holds true. Therefore, the Banach fixed-point theorem applies, and hence the operator $B: V \rightarrow V$ has a unique fixed point $u_{0} \in V$. The solution $u_{0} \in V$ to the BCS gap equation is thus continuous both with respect to the temperature and with respect to the energy $x$; see Figure 3.

## Acknowledgment

Shuji Watanabe is supported in part by the JSPS Grant-in-Aid for Scientific Research (C) 24540112.

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## Research Article

# Numerical Solution for IVP in Volterra Type Linear Integrodifferential Equations System 

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Received 23 May 2013; Accepted 9 July 2013
Academic Editor: Santanu Saha Ray
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A method is proposed to determine the numerical solution of system of linear Volterra integrodifferential equations (IDEs) by using Bezier curves. The Bezier curves are chosen as piecewise polynomials of degree $n$, and Bezier curves are determined on $\left[t_{0}, t_{f}\right]$ by $n+1$ control points. The efficiency and applicability of the presented method are illustrated by some numerical examples.

## 1. Introduction

Integrodifferential equations (IDEs) have been found to describe various kinds of phenomena, such as glass forming process, dropwise condensation, nanohydrodynamics, and wind ripple in the desert (see $[1,2]$ ).

There are several numerical and analytical methods for solving IDEs. Some different methods are presented to solve integral and IDEs in [3, 4]. Maleknejad et al. [5] used rationalized Haar functions method to solve the linear IDEs system. Linear IDEs system has been solved by using Galerkin methods with the hybrid Legendre and blockPulse functions on interval [0,1) (see [6]). Yusufoğlu [7] presented an application of He's homotopy perturbation (HPM) method to solve the IDEs system. He's variational iteration method has been used for solving two systems of Volterra integrodifferential equations (see [8]). Arikoglu and Ozkol [9] presented differential transform method (DTM) for integrodifferential and integral equation systems. He's homotopy perturbation (HPM) method was proposed for system of integrodifferential equations (see [10]). A numerical method based on interpolation of unknown functions at distinct interpolation points has been introduced for solving linear IDEs system with initial values (see [11]). Recently, Biazar introduced a new modification of homotopy perturbation method (NHPM) to obtain the solution of linear IDEs system (see [12]). Taylor expansion method has been used for solving

IDEs (see [13, 14]). Rashidinia and Tahmasebi developed and modified Taylor series method (TSM) introduced in [15] to solve the system of linear Volterra IDEs.

In the present work, we suggest a technique similar to the one which was used in [16] for solving the system of linear Volterra IDEs in the following form:

$$
\begin{gather*}
\sum_{i=1}^{n} \sum_{j=0}^{\alpha_{m i}} p_{m i j}(t) y_{i}^{(j)}(t)+\sum_{i=1}^{n} \int_{t_{0}}^{t}\left(k_{m i}(t, x) \sum_{j=0}^{\alpha_{m i}} y_{i}^{(j)}(x)\right) d x  \tag{1}\\
=f_{m}(t), \quad m=1,2, \ldots, n, t_{0} \leq t \leq t_{f}
\end{gather*}
$$

with the initial conditions

$$
\begin{equation*}
y_{i}^{(0)}\left(t_{0}\right)=c_{i 0}, \quad y_{i}^{(1)}\left(t_{0}\right)=c_{i 1}, \ldots, y_{i}^{\left(\alpha_{m i}-1\right)}\left(t_{0}\right)=c_{i\left(\alpha_{m i}-1\right)} \tag{2}
\end{equation*}
$$

where $y_{i}^{(j)}(t)$ stands for $j$ th-order derivative of $y_{i}(t) . f_{m}(t)$, $k_{m i}(t, x)$, and $p_{m i j}(t)$ are known functions $(m, i=1,2, \ldots$, $\left.n ; j=0,1, \ldots, \alpha_{m i}\right)$, and $t_{0}, t_{f}$, and $c_{i j}(i=1,2, \ldots, n ; j=$ $\left.0,1, \ldots, \alpha_{m i}-1\right)$ are appropriate constants.

The current paper is organized as follows. In Section 2, function approximation will be introduced. Numerical examples will be stated in Section 3. Finally, Section 4 will give a conclusion briefly.


Figure 1: The graph of approximated $y_{1}(t)$ for Example 1.

## 2. Function Approximation

Our strategy is to use Bezier curves to approximate the solutions $y_{i}(t)$, for $1 \leq i \leq n$, which are given below. Define the Bezier polynomials of degree $N$ that approximate, respectively, the actions of $y_{i}(t)$ over the interval $\left[t_{0}, t_{f}\right]$ as follows:

$$
\begin{equation*}
y_{i}(t)=\sum_{r=0}^{N} a_{r}^{i} B_{r, N}\left(\frac{t-t_{0}}{h}\right), \tag{3}
\end{equation*}
$$

where $h=t_{f}-t_{0}$ and $a_{r}$ is the control point of Bezier curve, and

$$
\begin{equation*}
B_{r, N}\left(\frac{t-t_{0}}{h}\right)=\binom{N}{r} \frac{1}{h^{N}}\left(t_{f}-t\right)^{N-r}\left(t-t_{0}\right)^{r} \tag{4}
\end{equation*}
$$

is the Bernstein polynomial of degree $N$ over the interval [ $t_{0}, t_{f}$ ] (see [17]). By substituting (3) in (2), $R_{m}(t)$ can be defined for $t \in\left[t_{0}, t_{f}\right]$ as

$$
\begin{align*}
R_{m}(t)= & \sum_{i=1}^{n} \sum_{j=0}^{\alpha_{m i}} p_{m i j}(t) y_{i}^{(j)}(t) \\
& +\sum_{i=1}^{n} \int_{0}^{t}\left(k_{m i}(t, x) \sum_{j=0}^{\alpha_{m i}} y_{i}^{(j)}(x)\right) d x-f_{m}(t),  \tag{5}\\
& m=1,2, \ldots, n
\end{align*}
$$

where (2) is satisfied. The convergence was proved in the approximation with Bezier curves when the degree of the approximate solution, $N$, tends to infinity (see [18]).

Now, the residual function is defined over the interval [ $t_{0}, t_{f}$ ] as follows:

$$
\begin{equation*}
R(t)=\int_{t_{0}}^{t_{f}} \sum_{m=1}^{n}\left\|R_{m}(t)\right\|^{2} d t \tag{6}
\end{equation*}
$$



Figure 2: The graph of approximated $y_{2}(t)$ for Example 1.
where $\|\cdot\|$ is the Euclidean norm. Our aim is to solve the following problem over the interval $\left[t_{0}, t_{f}\right]$ :

$$
\begin{array}{ll}
\min & R(t) \\
\text { s.t. } & y_{i}^{(0)}\left(t_{0}\right)=c_{i 0}  \tag{7}\\
& y_{i}^{(1)}\left(t_{0}\right)=c_{i 1}, \ldots, y_{i}^{\left(\alpha_{m i}-1\right)}\left(t_{0}\right)=c_{i\left(\alpha_{m_{i}}-1\right)}
\end{array}
$$

The mathematical programming problem (7) can be solved by many subroutine algorithms, and we used Maple 12 to solve this optimization problem.

Remark 1. In Chapter 1 of [19], it was proved that $N$ satisfies

$$
\begin{equation*}
N>\frac{S}{\delta^{2} \epsilon} \tag{8}
\end{equation*}
$$

where $S=\left\|y_{i}(t)\right\|$, and because of this reason that $y_{i}(t)$ is uniformly continuous on $\left[t_{0}, t_{f}\right]$, we have $s, t \in\left[t_{0}, t_{f}\right]$ that $|t-s|<\delta$ and $-(\epsilon / 2)<y_{i}(t)-y_{i}(s)<\epsilon / 2$, for more details see [19].

## 3. Applications and Numerical Results

Consider the following examples which can be solved by using the presented method.

Example 1. Consider a system of third-order linear Volterra IDEs on the interval $[0,1]$ (see [4]):

$$
\begin{align*}
& y_{1}^{\prime \prime}(t)+t^{2} y_{1}(t)-y_{2}^{\prime \prime}(t) \\
& \quad+\int_{0}^{t}\left((t-x) y_{1}(x)+y_{2}(x)\right) d x=g_{1}(t) \\
& 4 t^{3} y_{1}^{\prime}(t)+6 t^{2} y_{1}(t)+y_{2}^{\prime \prime \prime}(t)  \tag{9}\\
& \quad+\int_{0}^{t}\left(y_{1}(x)+(t+x) y_{2}(x)\right) d x=g_{2}(t)
\end{align*}
$$

Table 1: Computed errors for Example 1.

| $t$ | Absolute error for $y_{1}(t)$ | Absolute error for $y_{2}(t)$ |
| :--- | :---: | :---: |
| 0.0 | 0.000000 | 0.0000000000 |
| 0.2 | $1.4801 \times 10^{-10}$ | $2.2475 \times 10^{-11}$ |
| 0.4 | $0.162735585 \times 10^{-5}$ | $3.12780502 \times 10^{-7}$ |
| 0.6 | $0.251133963 \times 10^{-5}$ | $0.1536077787 \times 10^{-5}$ |
| 0.8 | $1.8337 \times 10^{-10}$ | $0.8864238659 \times 10^{-5}$ |
| 1.0 | $4.5905 \times 10^{-10}$ | $7.897 \times 10^{-12}$ |

Table 2: Computed errors for Example 2.

| $t$ | Absolute error for $y_{1}(t)$ | Absolute error for $y_{2}(t)$ |
| :--- | :---: | :---: |
| 0.0 | 0.000000 | 0.0000000000 |
| 0.2 | $3.840 \times 10^{-11}$ | $1.5360 \times 10^{-11}$ |
| 0.4 | $0.5791064832 \times 10^{-3}$ | $0.156041748480 \times 10^{-3}$ |
| 0.6 | $0.17373195072 \times 10^{-2}$ | $0.156041748480 \times 10^{-3}$ |
| 0.8 | $0.69492781056 \times 10^{-2}$ | $1.5360 \times 10^{-11}$ |
| 1.0 | 0.000 | 0.000 |

with the initial conditions $y_{1}(0)=y_{1}^{\prime}(0)=1, y_{2}(0)=y_{2}^{\prime \prime}(0)=$ 0 , and $y_{2}^{\prime}(0)=1$, where

$$
\begin{gather*}
g_{1}(t)=\left(2+t^{2}\right) e^{t}-t-\cos (t)+\sin (t) \\
g_{2}(t)=7 \sin (t)-(1+2 t) \cos (t)+e^{t}\left(1+4 t^{2}+4 t^{3}\right)+t-1 \tag{10}
\end{gather*}
$$

The exact solution of this system is $y_{1}(t)=e^{t}, y_{2}(t)=$ $\sin (t)$.

With $N=5$, the approximated solutions for $y_{1}(t)$ and $y_{2}(t)$ are shown, respectively, in Figures 1 and 2, and the computed errors are shown in Table 1 which show the high accuracy of the proposed method.

Example 2. Consider the following system of linear Volterra IDEs equations as follows (see [4]):

$$
\begin{align*}
-y_{1}^{\prime} & -\frac{1}{2} t y_{1}+\frac{3}{2} y_{2} \\
= & \frac{5}{2}-t-\frac{27}{2} t^{2}+t^{4}+\frac{3}{2}\left(-1+2 t^{2}\right)-\frac{1}{2}\left(-3 t+4 t^{3}\right) \\
& +\int_{-1}^{\mathrm{t}}\left(y_{1}-3 t y_{2}\right) d x  \tag{11}\\
t^{2} y_{1}+ & y_{2}^{\prime}-t y_{2} \\
= & \frac{2}{5}+3 t+3 t^{3}-\frac{8}{5} t^{5}+t^{2}\left(-3 t+4 t^{3}\right) \\
& +\int_{-1}^{t}\left((2 t+x) y_{1}+3 x^{2} y_{2}\right) d x
\end{align*}
$$

under the conditions $y_{1}(0)=0$ and $y_{2}(0)=-1$, with the exact solution $y_{1}(t)=4 t^{3}-3 t, y_{2}(t)=2 t^{2}-1$.

With $N=5$, the computed errors are shown in Table 2 which show the high accuracy of the proposed method.

## 4. Conclusions

In this paper, Bernstein's approximation is used to approximate the solution of linear Volterra IDEs. In this method, we approximate our unknown function with Bernstein's approximation. The present results show that Bernstein's approximation method for solving linear Volterra IDEs is very effective and simple, and the answers are trusty, and their accuracy is high, and we can execute this method in a computer simply. The numerical examples support this claim.

## Acknowledgment

The authors are very grateful to the referees for their valuable suggestions and comments that improved the paper.

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## Research Article

# Analytical and Multishaped Solitary Wave Solutions for Extended Reduced Ostrovsky Equation 

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Received 5 June 2013; Accepted 25 July 2013
Academic Editor: Santanu Saha Ray
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We present the analytical and multishaped solitary wave solutions for extended reduced Ostrovsky equation (EX-ROE). The exact solitary (traveling) wave solutions are expressed by three types of functions which are hyperbolic function solution, trigonometric function solution, and rational solution. These results generalized the previous results. Multishape solitary wave solutions such as loop-shaped, cusp-shaped, and hump-shaped can be obtained as well when the special values of the parameters are taken. The $\left(G^{\prime} / G\right)$-expansion method presents a wide applicability for handling nonlinear partial differential equations.

## 1. Introduction

The well-known Ostrovsky equation [1]

$$
\begin{equation*}
\left(u_{t}+c_{0} u_{x}+\alpha u u_{x}+\beta u_{x x x}\right)_{x}=\gamma u \tag{1}
\end{equation*}
$$

where $c_{0}$ is the velocity of dispersiveness linear waves, $\alpha$ is the nonlinear coefficient, and $\beta$ and $\gamma$ are dispersion coefficients, is a model for weakly nonlinear surface and internal waves in a rotating ocean.

For long waves, for which high-frequency dispersion is negligible, $\beta=0$, and (1) becomes the so-called reduced Ostrovsky equation (ROE) [2]

$$
\begin{equation*}
\left(u_{t}+c_{0} u_{x}+\alpha u u_{x}\right)_{x}=\gamma u . \tag{2}
\end{equation*}
$$

Parkes [3] has studied (2) and found its periodic and solitary traveling wave solutions.

In fact, by applying the following transformation [4]:

$$
\begin{equation*}
u \longrightarrow \frac{u}{\alpha}, \quad t \longrightarrow \frac{t}{\sqrt{|\gamma|}}, \quad x \longrightarrow \frac{\left(x+c_{0} t\right)}{\sqrt{|\gamma|}} \tag{3}
\end{equation*}
$$

to (2), we obtain the ROE in the neat form

$$
\begin{align*}
& \frac{\partial}{\partial x} \mathfrak{D u}+\delta u=0 \\
& \text { where } \mathfrak{D}:=\frac{\partial}{\partial t}+u \frac{\partial}{\partial x}, \quad \delta=\frac{\gamma}{|\gamma|}= \pm 1 \tag{4}
\end{align*}
$$

Just as it mentioned in [5] and the reference therein, when $\delta=$ $-1,(4)$ is referred to the Ostrovsky-Hunter equation (OHE). When $\delta=1$, (4) is referred to the Vakhnenko equation (VE), which is in order to model the propagation of waves in a relaxing medium [6, 7]. Parkes [3] pointed out that (4) is invariant under the transformation

$$
\begin{equation*}
u \longrightarrow-u, \quad t \longrightarrow-t, \quad \delta \longrightarrow-\delta \tag{5}
\end{equation*}
$$

so that the solutions of the OHE and VE are related in a simple way.

The purpose of this paper is to study the extended reduced Ostrovsky equation (EX-ROE):

$$
\begin{equation*}
\frac{\partial}{\partial x}\left(\mathfrak{D}^{2} u+\frac{1}{2} p u^{2}+\beta u\right)+q \mathfrak{D} u=0 \tag{6}
\end{equation*}
$$

where $\mathfrak{D}$ is defined previous, $p, q$, and $\beta$ are arbitrary nonzero constants. It is originally derived by Morrison and Parkes [8] which dubbed it as modified generalized Vakhnenko equation (mGVE) when $p=2 q$. They found that not only does it have loop soliton solutions, hump-like and cusp-like soliton solutions, but it also has N -soliton solutions.

In order to investigate mGVE's $N$-soliton solution, Morrison and Parkes [8] considered a Hirota-Satsuma-type shallow water wave equation [9] of the form

$$
\begin{equation*}
U_{X X T}+p U U_{T}-q U_{X} \int_{X}^{\infty} U_{T}\left(X^{\prime}, T\right) d X^{\prime}+\beta U_{T}+q U_{X}=0 \tag{7}
\end{equation*}
$$

where $p \neq 0, q \neq 0$, and $\beta$ is arbitrary constant. By using the transformation

$$
\begin{gather*}
x=T+\int_{-\infty}^{X} U_{T}\left(X^{\prime}, T\right) \mathrm{d} X^{\prime}+x_{0}  \tag{8}\\
t=X, \quad u(x, t)=U(X, T)
\end{gather*}
$$

where $x_{0}$ is a constant, (7) yields (6). So (6) and (7) are equivalent to each other under the transformation (8). Specifically, in (7), when $p=2 q$ and $\beta=-1$, it was discussed by Ablowitz et al. [10] and was shown to be integrable by inverse scatting method. When $p=q$ and $\beta=-1$, it was discussed by Hirota and Satsuma [11] and was shown to be integrable using Hirota's bilinear technique. In [12], the authors referred to (6) with $p=q=1$ and $\beta$ an arbitrary nonzero constant as the generalized Vakhnenko equation (GVE). In fact, when $p=q$ and $\beta=0$, (6) can be written as

$$
\begin{equation*}
\left(\frac{\partial u}{\partial x}+\mathfrak{D}\right)\left(\frac{\partial}{\partial x} \mathfrak{D} u+p u\right)=0 . \tag{9}
\end{equation*}
$$

Clearly, solutions of the ROE are also solutions of (9) with $p= \pm 1$. So for arbitrary $p, q$, and $\beta$, if we obtain the solutions of EX-ROE, then we can also obtain the solutions of VE, GVE, mGVE, ROE, and OHE by taking the special values of $p, q$, and $\beta$.

The EX-ROE has been studied by several researchers. For example, Liu et al. [13] used Jacobi elliptic function method to obtain exact double periodic wave solutions and solitary wave solutions. Parkes [4] constructed periodic and solitary wave solutions of EX-ROE and gave the categorization of the solutions. Xie and Cai [14] used the bifurcation method of dynamic systems and simulation method of differential equations to get exact compacton and generalized kink wave solutions of EX-ROE. Stepanyants [15] applied the qualitative theory of differential equations to give a full classification of its solutions.

Recently, there are many methods being proposed to study the traveling wave solutions of nonlinear partial differential equations which are derived from physics, for example, [16-27]. As well as these methods, there are still many other methods; we cannot list all of them. Here we will use modified $\left(G^{\prime} / G\right)$-expansion method to investigate EX-ROE. As a result, three types of traveling wave solutions are were obtained. When the special values of the parameters are taken, they are reduced to some previous results which obtained by an other method.

The rest of the paper is organized as follows. In Section 2, we present a methodology of the modified $\left(G^{\prime} / G\right)$-expansion method. In Section 3, we apply the method to the extended reduced Ostrovsky equation. In Section 4, some conclusions are given.

## 2. Description of the Modified $\left(G^{\prime} / G\right)$-Expansion Method

The $\left(G^{\prime} / G\right)$-expansion method is first proposed by Wang et al. [28]. The useful $\left(G^{\prime} / G\right)$-expansion method is then widely used by many authors [29-32]. Then it is modified in [33-35]. The main steps are as follows.

Suppose that a nonlinear equation is given by

$$
\begin{equation*}
P_{1}\left(u, u_{t}, u_{x}, u_{t t}, u_{x t}, u_{x x}, \ldots\right)=0 \tag{10}
\end{equation*}
$$

where $u=u(x, t)$ is an unknown function and $P$ is a polynomial in $u=u(x, t)$ and its partial derivatives, in which the highest-order derivatives and nonlinear terms are involved. In the following we give the main steps of the $\left(G^{\prime} / G\right)$ expansion method.

Step 1. The traveling wave variable $u(x, t)=u(\xi), \xi=x-c t$, where $c$ is a constant, permits us to reduce (10) to an ODE for $u=u(\xi)$ in the form

$$
\begin{equation*}
P_{2}\left(u,-c u^{\prime}, u^{\prime}, c^{2} u^{\prime \prime},-c u^{\prime \prime}, u^{\prime \prime}, \ldots\right)=0 \tag{11}
\end{equation*}
$$

Step 2. Suppose that the solution of (10) can be expressed by a polynomial in $\left(G^{\prime} / G\right)$ as follows:

$$
\begin{align*}
u(\xi)= & \alpha_{-m}\left(\frac{G^{\prime}}{G}\right)^{-m}+\alpha_{-(m-1)}\left(\frac{G^{\prime}}{G}\right)^{-(m-1)}+\cdots  \tag{12}\\
& +\alpha_{m-1}\left(\frac{G^{\prime}}{G}\right)^{m-1}+\alpha_{m}\left(\frac{G^{\prime}}{G}\right)^{m}
\end{align*}
$$

where $G=G(\xi)$ satisfies the second-order linear ordinary differential equation (LODE) in the form

$$
\begin{equation*}
G^{\prime \prime}+\lambda G^{\prime}+\mu G=0 \tag{13}
\end{equation*}
$$

where $\alpha_{-m}, \ldots, \alpha_{m}, \lambda$, and $\mu$ are constants to be determined later. The unwritten part in (12) is also a polynomial in $\left(G^{\prime} / G\right)$, but the degree of which is generally equal to or less than $m-1$. The positive integer $m$ can be determined by considering the homogeneous balance between the highest-order derivatives and nonlinear terms appearing in (11).

Step 3. Substituting (12) into (11) and using (13), collecting all terms with the same order of $\left(G^{\prime} / G\right)$ together, and then equating each coefficient of the resulting polynomial to zero yields a set of algebraic equations for $\alpha_{m}, \alpha_{m-1}, \ldots, \alpha_{-m}, c, \lambda$, and $\mu$.

Step 4. Since the general solutions of (13) have been well known for us, then substituting $\alpha_{m}, \alpha_{m-1}, \ldots, \alpha_{-m}$ and $c$ and the general solutions of (13) into (12) we have more traveling wave solutions of the nonlinear differential equation (10).

The main idea of $\left(G^{\prime} / G\right)$-expansion method is to use an integrable ODE to expand a solution to a nonlinear partial differential equation (PDE) as a polynomial or rational function of the solution of the ODE. However, such an idea was also presented in [36-38]. The method used in this paper can be also thought of as the application of transformed
rational function method used in [37] in some sense. Maybe the similar results can be obtained by using these very closely related methods. We plan to further study the EX-ROE in near future by using the methods proposed in [36-38]. We hope we can find much more interesting properties and new phenomenon of this equation.

## 3. Exact Traveling Wave Solutions of the Extended Reduced Ostrovsky Equation

In this section, we will use the $\left(G^{\prime} / G\right)$-expansion method to the extended reduced Ostrovsky equation to get exact traveling wave solutions.

First, in order to get traveling wave solutions, we need some transformation. Recall that in Section 1 we have stated that EX-ROE is equivalent to a Hirota-Satsuma-type shallow water wave equation (7) under the transformation of (8). So here we introduce a new variable $W$ defined by

$$
\begin{equation*}
U=W_{X} \tag{14}
\end{equation*}
$$

Substituting (14) into (7) yields

$$
\begin{equation*}
W_{X X X T}+p W_{X} W_{X T}+q W_{X X} W_{T}+\beta W_{X T}+q W_{X X}=0 \tag{15}
\end{equation*}
$$

Now giving the traveling wave transformation $W(X, T)=$ $W(\xi), \xi=X-c T$, where $c$ is wave speed. Substituting them into (15) and integrating once, we have

$$
\begin{equation*}
c_{1}+c W_{3 \xi}+\frac{1}{2} c(p+q) W_{\xi}^{2}+(c \beta-q) W_{\xi}=0 \tag{16}
\end{equation*}
$$

where $c_{1}$ is integral constant that is to be determined later.
Considering the homogeneous balance between $W_{3 \xi}$ and $W_{\xi}^{2}$, we have

$$
\begin{equation*}
m+3=2 m+2 \Longrightarrow m=1 \tag{17}
\end{equation*}
$$

We suppose that

$$
\begin{equation*}
W(\xi)=\alpha_{-1}\left(\frac{G^{\prime}}{G}\right)^{-1}+\alpha_{0}+\alpha_{1}\left(\frac{G^{\prime}}{G}\right) \tag{18}
\end{equation*}
$$

where the $G=G(\xi)$ satisfies the second-order LODE,

$$
\begin{equation*}
G^{\prime \prime}+\lambda G^{\prime}+\mu G=0 \tag{19}
\end{equation*}
$$

and $\alpha_{-1}, \alpha_{0}, \alpha_{1}, \lambda$, and $\mu$ are constants to be determined later.

By using (18) and (19), it is derived that

$$
\begin{align*}
W_{\xi}= & \mu \alpha_{-1}\left(\frac{G^{\prime}}{G}\right)^{-2}+\lambda \alpha_{-1}\left(\frac{G^{\prime}}{G}\right)^{-1}+\alpha_{-1} \\
& -\alpha_{1} \mu-\lambda \alpha_{1}\left(\frac{G^{\prime}}{G}\right)-\alpha_{1}\left(\frac{G^{\prime}}{G}\right)^{2},  \tag{20}\\
W_{\xi}^{2}= & \mu^{2} \alpha_{-1}^{2}\left(\frac{G^{\prime}}{G}\right)^{-4}+2 \lambda \alpha_{-1}^{2} \mu\left(\frac{G^{\prime}}{G}\right)^{-3} \\
& +\left(2 \alpha_{-1}^{2} \mu+\lambda^{2} \alpha_{-1}^{2}-2 \mu \alpha_{-1} \alpha_{-1}\right)\left(\frac{G^{\prime}}{G}\right)^{-2} \\
& +\left(2 \lambda \alpha_{-1}^{2}-4 \mu \lambda \alpha_{-1} \alpha_{1}\right)\left(\frac{G^{\prime}}{G}\right)^{-1}+\alpha_{-1}^{2}-4 \mu \alpha_{-1} \alpha_{1} \\
& -2 \lambda^{2} \alpha_{-1} \alpha_{1}+\alpha_{1}^{2} \mu^{2}+2 \lambda \alpha_{1}^{2} \mu\left(\frac{G^{\prime}}{G}\right) \\
& +\left(2 \alpha_{1}^{2} \mu+\lambda^{2} \alpha_{1}^{2}\right)\left(\frac{G^{\prime}}{G}\right)^{2}+2 \lambda \alpha_{1}^{2}\left(\frac{G^{\prime}}{G}\right)^{3}+\alpha_{1}^{2}\left(\frac{G^{\prime}}{G}\right)^{4} \tag{21}
\end{align*}
$$

$$
\begin{align*}
W_{3 \xi}= & 6 \alpha_{-1} \mu^{3}\left(\frac{G^{\prime}}{G}\right)^{-4}+12 \alpha_{-1} \lambda \mu^{2}\left(\frac{G^{\prime}}{G}\right)^{-3} \\
& -\left(8 \mu^{2} \alpha_{-1}+7 \alpha_{-1} \lambda^{2} \mu\right)\left(\frac{G^{\prime}}{G}\right)^{-2} \\
& -\left(8 \alpha_{-1} \lambda \mu+\alpha_{-1} \lambda^{3}\right)\left(\frac{G^{\prime}}{G}\right)^{-1} \\
& -\left(2 \alpha_{-1} \mu+\lambda^{2} \alpha_{-1}\right)-\left(2 \alpha_{1} \mu^{2}+\lambda^{2} \alpha_{1} \mu\right) \\
& -\left(8 \alpha_{1} \lambda \mu+\alpha_{1} \lambda^{3}\right)\left(\frac{G^{\prime}}{G}\right)-\left(8 \mu \alpha_{1}+7 \alpha_{1} \lambda^{2}\right)\left(\frac{G^{\prime}}{G}\right)^{2} \\
& -12 \alpha_{1} \lambda\left(\frac{G^{\prime}}{G}\right)^{3}-6 \alpha_{1}\left(\frac{G^{\prime}}{G}\right)^{4} . \tag{22}
\end{align*}
$$

By substituting (20)-(22) into (16) and collecting all terms with the same power of $\left(G^{\prime} / G\right)$ together, the left-hand sides of (16) are converted into the polynomials in $\left(G^{\prime} / G\right)$. Equating the coefficients of the polynomials to zero yields a set of simultaneous algebraic equations for $\alpha_{-1}, \alpha_{0}, \alpha_{1}, \lambda, c, c_{1}$, and $\mu$ as follows (denote $A$ for $\left(G^{\prime} / G\right)$ ):

$$
\begin{aligned}
& A^{-4}: 6 c \alpha_{-1} \mu^{3}+\frac{c(p+q) \mu^{2} \alpha_{-1}^{2}}{2}=0 \\
& A^{-3}: 12 c \lambda \mu^{2} \alpha_{-1}+c(p+q) \lambda \mu \alpha_{-1}^{2}=0
\end{aligned}
$$

$$
\begin{align*}
\begin{aligned}
& A^{-2}:(q-c \beta) \mu \alpha_{-1} \\
&-\frac{c(p+q)\left(\lambda^{2} \alpha_{-1}^{2}+2 \mu \lambda_{-1}^{2}-2 \mu^{2} \alpha_{1} \alpha_{-1}\right)}{2} \\
&-c\left(7 \lambda^{2} \mu \alpha_{-1}+8 \mu^{2} \alpha_{-1}\right)=0, \\
& A^{-1}:(c \beta-q) \lambda \alpha_{-1} \lambda+c(p+q)\left(\lambda \alpha_{-1}^{2}-2 \lambda \mu \alpha \alpha_{-1}\right) \\
&+c\left(\lambda^{3} \alpha_{-1}+8 \lambda \mu \alpha_{-1}\right)=0, \\
& A^{0}: c_{1}+ c\left(\lambda^{2} \alpha_{-1}+2 \mu \alpha_{-1}\right) \\
& \quad-\frac{c\left(2 \alpha_{1} \mu^{2}+\lambda^{2} \alpha_{1} \mu\right)+c(p+q) \alpha_{-1}^{2}}{2} \\
&-2 c \mu \alpha \alpha_{-1}(p+q)+\frac{c \alpha_{1}^{2} \mu^{2}(p+q)}{2}-(c \beta-q) \mu \alpha_{1} \\
&-c \lambda^{2} \alpha_{1} \alpha_{-1}(p+q)+\alpha_{-1}(c \beta-q)=0, \\
& A^{1}:(q-c \beta) \alpha_{1} \lambda+c(p+q)\left(\lambda \mu \alpha_{1}^{2}-2 \lambda \alpha_{1} \alpha_{-1}\right) \\
&-c\left(\lambda^{3} \alpha_{1}+8 \lambda \mu \alpha_{1}\right)=0, \\
& A^{2}:(q-c \beta) \alpha_{1}+c(p+q) \\
& \times\left(\lambda^{2} \alpha_{1}^{2}+2 a \alpha_{1}^{2} \mu-2 \alpha_{1} \alpha_{-1}-c\left(7 \lambda^{2} \alpha_{1}+8 \mu \alpha_{1}\right)\right) \\
&= 0,
\end{aligned} \\
A^{3}:-12 c \lambda \alpha_{1}+c(p+q) \lambda \alpha_{1}^{2}=0, \\
A^{4}:-6 c \alpha_{1}+\frac{c(p+q) \alpha_{1}^{2}}{2}=0 .
\end{align*}
$$

Solving the algebraic equations above yields

$$
\begin{gather*}
\alpha_{1}=\frac{12}{(p+q)}, \\
c=\frac{q}{\left(\beta+\lambda^{2}-4 \mu\right)}  \tag{24}\\
c_{1}=0, \quad \alpha_{-1}=0,
\end{gather*}
$$

or

$$
\begin{align*}
& \alpha_{-1}=-\frac{12 \mu}{(p+q)} \\
& c=\frac{q}{\left(\beta+\lambda^{2}-4 \mu\right)}  \tag{25}\\
& c_{1}=0, \quad \alpha_{1}=0
\end{align*}
$$

Substituting system (24) and (25) into (18), we have the formula of the solutions of (15) as follows:

$$
\begin{equation*}
W(X, T)=W(\xi)=\frac{12}{(p+q)}\left(\frac{G^{\prime}}{G}\right)+\alpha_{0} \tag{26}
\end{equation*}
$$

or

$$
\begin{equation*}
W(X, T)=W(\xi)=-\frac{12 \mu}{(p+q)}\left(\frac{G^{\prime}}{G}\right)^{-1}+\alpha_{0} \tag{27}
\end{equation*}
$$

where $G$ satisfies (19), $\xi=X-q T /\left(\beta+\lambda^{2}-4 \mu\right)$, and $\alpha_{0}$ is an arbitrary constant.

Since the general solutions $G=G(\xi)$ (hence $G^{\prime}=$ $\mathrm{d} G / \mathrm{d} \xi$ ) of ODE (19) have been known for us, substituting the solutions of (19) into (24) and (25), we have the general traveling wave solutions of (15) as follows.

Case 1. When $\lambda^{2}-4 \mu>0$, then we have the following exact traveling wave solution of (15):

$$
\begin{align*}
& W_{1}(X, T) \\
& \begin{array}{l}
=W_{1}(\xi)=\frac{6 \sqrt{\lambda^{2}-4 \mu}}{p+q} \\
\quad \times\left(\left(A_{1} \cosh \left(\frac{1}{2} \sqrt{\lambda^{2}-4 \mu} \xi\right)\right)\right. \\
\\
\left.+A_{2} \sinh \left(\frac{1}{2} \sqrt{\lambda^{2}-4 \mu} \xi\right)\right) \\
\\
\quad \times\left(A_{1} \sinh \left(\frac{1}{2} \sqrt{\lambda^{2}-4 \mu \xi}\right)\right. \\
\left.\left.\quad+A_{2} \cosh \left(\frac{1}{2} \sqrt{\lambda^{2}-4 \mu} \xi\right)\right)^{-1}\right) \\
-\frac{6 \lambda}{(p+q)+\alpha_{0}}
\end{array}
\end{align*}
$$

or

$$
\begin{align*}
& W_{2}(X, T) \\
& =W_{2}(\xi) \\
& =-24 \mu \times((p+q) \\
& \times\left(\sqrt{\lambda^{2}-4 \mu}\right. \\
& \times\left(\left(A_{1} \cosh \left(\frac{1}{2} \sqrt{\lambda^{2}-4 \mu} \xi\right)\right.\right. \\
& \left.+A_{2} \sinh \left(\frac{1}{2} \sqrt{\lambda^{2}-4 \mu} \xi\right)\right) \\
& \times\left(A_{1} \sinh \left(\frac{1}{2} \sqrt{\lambda^{2}-4 \mu} \xi\right)\right. \\
& \left.\left.+A_{2} \cosh \left(\frac{1}{2} \sqrt{\lambda^{2}-4 \mu} \xi\right)\right)^{-1}\right) \\
& -\lambda))^{-1}+\alpha_{0}, \tag{29}
\end{align*}
$$

where $\xi=X-q T /\left(\beta+\lambda^{2}-4 \mu\right)$ and $\alpha_{0}, A_{1}, A_{2}$ are arbitrary constants.

Case 2. When $\lambda^{2}-4 \mu<0$, then we have the following exact traveling wave solution of (15):

$$
\begin{align*}
& W_{3}(X, T) \\
& \begin{array}{l}
=W_{3}(\xi)=\frac{6 \sqrt{4 \mu-\lambda^{2}}}{p+q} \\
\quad \times\left(\left(-A_{1} \sin \left(\frac{1}{2} \sqrt{4 \mu-\lambda^{2}} \xi\right)+A_{2} \cos \left(\frac{1}{2} \sqrt{4 \mu-\lambda^{2}} \xi\right)\right)\right. \\
\quad \times\left(A_{1} \cos \left(\frac{1}{2} \sqrt{4 \mu-\lambda^{2}} \xi\right)\right. \\
\left.\left.\quad+A_{2} \sin \left(\frac{1}{2} \sqrt{4 \mu-\lambda^{2}} \xi\right)\right)^{-1}\right) \\
\\
-\frac{6 \lambda}{(p+q)+\alpha_{0}}
\end{array}
\end{align*}
$$

or

$$
\begin{align*}
& W_{4}(X, T) \\
& \qquad \begin{array}{r}
=W_{4}(\xi) \\
=-24 \mu
\end{array} \\
& \quad \times\left(( p + q ) \left(\sqrt{\lambda^{2}-4 \mu}\right.\right. \\
& \quad \times\left(\left(-A_{1} \sin \left(\frac{1}{2} \sqrt{4 \mu-\lambda^{2}} \xi\right)\right.\right. \\
& \\
& \left.\quad+A_{2} \cos \left(\frac{1}{2} \sqrt{4 \mu-\lambda^{2}} \xi\right)\right) \\
& \\
& \quad \times\left(A_{1} \cos \left(\frac{1}{2} \sqrt{4 \mu-\lambda^{2}} \xi\right)\right. \\
& +
\end{aligned} \begin{aligned}
& \left.\left.A_{2} \sin \left(\frac{1}{2} \sqrt{4 \mu-\lambda^{2}} \xi\right)\right)^{-1}\right)  \tag{31}\\
& \\
& \quad-\lambda))^{-1}+\alpha_{0}
\end{align*}
$$

where $\xi=X-q T /\left(\beta+\lambda^{2}-4 \mu\right)$ and $\alpha_{0}, A_{1}, A_{2}$ are arbitrary constants.

Case 3. When $\lambda^{2}-4 \mu=0$, then we have the following exact rational solution of (15):

$$
\begin{equation*}
W_{5}(X, T)=W_{5}(\xi)=\frac{12}{p+q}\left(\frac{A_{2}}{A_{1}+A_{2} \xi}\right)-\frac{6 \lambda}{(p+q)}+\alpha_{0} \tag{32}
\end{equation*}
$$

or

$$
\begin{equation*}
W_{6}(X, T)=W_{6}(\xi)=-\frac{24 \mu\left(A_{1}+A_{2} \xi\right)}{(p+q)\left[2 A_{2}-\lambda\left(A_{1}+A_{2} \xi\right)\right]}+\alpha_{0} \tag{33}
\end{equation*}
$$

where $\xi=X-q T /\left(\beta+\lambda^{2}-4 \mu\right)$ and $\alpha_{0}, A_{1}, A_{2}$ are arbitrary constants.

Now we will show how to get exact traveling wave solutions of (6). From (8) and (14), the solution of EX-ROE (6) is given in parametric form, with $T$ as the parameter, by

$$
\begin{equation*}
u(x, t)=U(t, T), \quad x=\theta(t, T) \tag{34}
\end{equation*}
$$

where

$$
\begin{equation*}
\theta(X, T)=T+W(X, T)+x_{0} \tag{35}
\end{equation*}
$$

So by using (8), (14), (34), (35), (28), and (29), we obtain a parameterized hyperbolic-function-type traveling wave solution of (6) as follows:

$$
\begin{align*}
& u_{1}(x, t) \\
& =3\left(A_{2}^{2}-A_{1}^{2}\right)\left(\lambda^{2}-4 \mu\right) \\
& \quad \times((p+q) \\
& \quad \times\left(A_{1} \sinh \left(\frac{1}{2} \sqrt{\lambda^{2}-4 \mu} \xi\right)\right.  \tag{36}\\
& \left.\left.+A_{2} \cosh \left(\frac{1}{2} \sqrt{\lambda^{2}-4 \mu} \xi\right)\right)^{2}\right)^{-1} \\
& x=T+W_{1}(t, T)+x_{0}
\end{align*}
$$

or

$$
\begin{align*}
& u_{2}(x, t) \\
& =-12 \mu\left(A_{2}^{2}-A_{1}^{2}\right)\left(\lambda^{2}-4 \mu\right) \\
& \times((p+q) \\
& \times\left[\left(A_{1} \sqrt{\lambda^{2}-4 \mu}-A_{2} \lambda\right) \cosh \left(\frac{1}{2} \sqrt{\lambda^{2}-4 \mu} \xi\right)\right. \\
& +\left(A_{2} \sqrt{\lambda^{2}-4 \mu}-A_{1} \lambda\right) \\
& \left.\left.\times \sinh \left(\frac{1}{2} \sqrt{\lambda^{2}-4 \mu} \xi\right)\right]^{2}\right)^{-1}, \\
& x=T+W_{2}(t, T)+x_{0}, \tag{37}
\end{align*}
$$

where $\xi=t-q T /\left(\beta+\lambda^{2}-4 \mu\right)$ and $x_{0}$ is an arbitrary constant.

By using (8), (14), (34), (35), (30), and (31), we obtain a parameterized trigonometric-function-type traveling wave solution of (6) as follows:

$$
\begin{align*}
& u_{3}(x, t) \\
& =-3\left(A_{2}^{2}+A_{1}^{2}\right)\left(\lambda^{2}-4 \mu\right) \\
& \times((p+q) \\
& \quad \times\left(A_{1} \cos \left(\frac{1}{2} \sqrt{\lambda^{2}-4 \mu} \xi\right)\right.  \tag{38}\\
& \left.\left.\quad+A_{2} \sin \left(\frac{1}{2} \sqrt{\lambda^{2}-4 \mu} \xi\right)\right)^{2}\right)^{-1} \\
& x=T+W_{3}(t, T)+x_{0}
\end{align*}
$$

or

$$
\begin{align*}
& u_{4}(x, t) \\
& =-12 \mu\left(A_{2}^{2}-A_{1}^{2}\right)\left(\lambda^{2}-4 \mu\right) \\
& \times((p+q) \\
& \times\left[\left(-A_{2} \sqrt{\lambda^{2}-4 \mu}+A_{1} \lambda\right) \cos \left(\frac{1}{2} \sqrt{\lambda^{2}-4 \mu} \xi\right)\right. \\
& \quad+\left(A_{1} \sqrt{\lambda^{2}-4 \mu}+A_{2} \lambda\right) \\
& \left.\left.\quad \times \sin \left(\frac{1}{2} \sqrt{\lambda^{2}-4 \mu} \xi\right)\right]^{2}\right)^{-1} \\
& x=T+W_{4}(t, T)+x_{0} \tag{39}
\end{align*}
$$

where $\xi=t-q T /\left(\beta+\lambda^{2}-4 \mu\right)$ and $x_{0}$ is an arbitrary constant.
By using (8), (14), (34), (35), (32), and (33), we obtain a parameterized rational-type traveling wave solution of (6) as follows:

$$
\begin{gather*}
u_{5}(x, t)=-\frac{12 A_{2}^{2}}{(p+q)\left(A_{1}+A_{2} \xi\right)^{2}}  \tag{40}\\
x=T+W_{5}(t, T)+x_{0}
\end{gather*}
$$

or

$$
\begin{gather*}
u_{6}(x, t)=-\frac{48 \mu A_{2}^{2}}{(p+q)\left(\lambda A_{1}+(-2+\lambda \xi) A_{2}\right)^{2}}  \tag{41}\\
x=T+W_{6}(t, T)+x_{0}
\end{gather*}
$$

where $\xi=t-q T /\left(\beta+\lambda^{2}-4 \mu\right)$ and $x_{0}$ is an arbitrary constant. To our knowledge, these solutions are presented for the first time; they are new exact solutions of EX-ROE.

If we take $A_{1}=0, \mu=0, A_{2} \neq 0$ and $\lambda>0$, then (36) yields the following solitary wave solution of (6):

$$
\begin{gather*}
u(x, t)=\frac{3 \lambda^{2}}{p+q} \operatorname{sech}^{2}\left[\frac{1}{2} \lambda\left(t-\frac{q}{\beta+\lambda^{2}} T\right)\right],  \tag{42}\\
x=T+\frac{6 \lambda}{(p+q)} \tanh \left[\frac{1}{2} \lambda\left(t-\frac{q}{\beta+\lambda^{2}} T\right)\right]+x_{0} .
\end{gather*}
$$

Now we will give some discussion of the solitary wave solution (42). Let $\lambda=2 k, x_{0}=0$; the solution (42) is reduced to the solution of (3.26) in [13] after correcting some minor errors [4]. Now from (35), we introduce a new variable:

$$
\begin{equation*}
\chi=x-v t=-v(X-c T)+\frac{6 \lambda}{p+q} \tanh \left[\frac{1}{2} \lambda(X-c T)\right]+x_{0} \tag{43}
\end{equation*}
$$

where $v=1 / c=\left(\lambda^{2}+\beta\right) / q$. In [8], the authors considered EX-ROE with $p=2 q, \beta \neq 0$ as mGVE and obtained 1 -soliton solution. In fact, if we take $p=2 q, \lambda=2 k$, the solitary wave solution (42) with (43) is reduced to the soliton solution (4.4) and (4.5) in [8]. From the above we can see that the solitary wave solution (3.26) in [4] and the 1 -soliton solution of mGVE are just a special case of the solution (42) in this paper.

## 4. Multishaped Solitary Wave Solutions

In $[8,13]$, the authors showed that the solutions of (4.4) and (4.5), (3.26) and (3.28) may be of different types, namely, loops, cusps, or humps for different values of parameters $\beta$, $k$, $p$. Here we also show that by choosing different values of the parameters $\beta, \lambda, p, q$, different shape wave solutions can be obtained. As it is stated in Section 1, (9) reduces to VE when $p=q=1, \beta=0$. Taking solution (42) with (43), for example, let $p=q=1, \beta=0, \lambda=2 k$; then it is reduced to one-loop soliton solution (3.4) and (3.5) in [39]. On the other hand, because the solutions of OHE and VE are connected in a particularly simple way, if we take $p=q=-1, \beta=0, \lambda=2 k$ in (42), we can obtain one-loop soliton solution of OHE.

From above analysis, one can clearly see that the solutions obtained in this paper are generalized for the previous results because here we only take the special case $A_{1}=0, \mu=0$, $A_{2} \neq 0, \lambda>0$ and give special discussion of solution (42). We conclude that if we take different values of the parameters $A_{1}, A_{2}, \lambda, \mu, p, q$, abundancy of types of exact solutions can be obtained from solutions (36), (38), and (40). Here we omit the detailed discussion.

Instead, we give some discussion about solution (37). Science from this solution, multishaped solitary wave solutions can be obtained. Suppose $A_{1} \neq 0, \mu<0, A_{2}=0, \lambda=0$; we reduce solution (37) to

$$
\begin{gather*}
u(x, t)=\frac{12 \mu}{p+q} \operatorname{sech}^{2}\left[\sqrt{-\mu}\left(t-\frac{q}{\beta-4 \mu} T\right)\right] \\
x=T+\frac{24 \sqrt{-\mu}}{(p+q)} \tanh \left[\sqrt{-\mu}\left(t-\frac{q}{\beta-4 \mu} T\right)\right]+x_{0} . \tag{44}
\end{gather*}
$$

We show that for different values of $\beta, \mu$, and $p$, the solution (44) may be of different types. It also owns the property of


Figure 1: The profile of solution (44) with $p=3, q=1.5, t=0$, and $x_{0}=0$. For (a) loop-shaped $\beta=0.005, \mu=-0.5$, (b) cusp-shaped $\beta=0.05, \mu=-1$, and (c) hump-shaped $\beta=0.5, \mu=-1$.
being loop-shaped, cusp-shaped and hump-shaped, as shown in Figure 1.

## 5. Conclusion

In this paper, we use $\left(G^{\prime} / G\right)$-expansion method to study extended reduced Ostrovsky equation. Several pairs of generalized traveling wave solutions are given directly. These solutions extend the previous results to more general cases. At the same time, multishaped wave solutions can be obtained if the different parameters values are chosen. These explicit solitary wave solutions own the property of being loop-shaped, cusp-shaped, and hump-shaped. These exact traveling wave solutions are also helpful to further study this nonlinear equation which has their physical meaning. The method used in this paper has more advantages. It is direct and concise. Much tedious algebraic calculations can be finished by computer program such as MATHEMATICA and MAPLE. Many well-known nonlinear wave equations can be handled by this method.

## Acknowledgment

The authors thank anonymous referees for valuable suggestions and comments which improve this paper readability and convincibility. This paper is supported by the Starting Research Founding of Wuhan Textile University.

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## Research Article

# New Exact Solutions for a Generalized Double Sinh-Gordon Equation 

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Received 26 June 2013; Revised 15 July 2013; Accepted 15 July 2013
Academic Editor: Santanu Saha Ray
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#### Abstract

We study a generalized double sinh-Gordon equation, which has applications in various fields, such as fluid dynamics, integrable quantum field theory, and kink dynamics. We employ the Exp-function method to obtain new exact solutions for this generalized double sinh-Gordon equation. This method is important as it gives us new solutions of the generalized double sinh-Gordon equation.


## 1. Introduction

It is well known that finding exact travelling wave solutions of nonlinear partial differential equations (NLPDEs) is useful in many scientific applications such as fluid mechanics, plasma physics, and quantum field theory. Due to these applications many researchers are investigating exact solutions of NLPDEs since they play a vital role in the study of nonlinear physical phenomena. Finding exact solutions of such NLPDEs provides us with a better understanding of the physical phenomena that these NLPDEs describe. Several techniques have been presented in the literature to find exact solutions of the NLPDEs. These include the homogeneous balance method, the Weierstrass elliptic function expansion method, the $F$-expansion method, the $\left(G^{\prime} / G\right)$-expansion method, the Exp-function method, the tanh function method, the extended tanh function method, and the Lie group method [1-10].

In this work, we study one such NLPDE, namely, the generalized double sinh-Gordon equation:

$$
\begin{equation*}
u_{t t}-k u_{x x}+2 \alpha \sinh (n u)+\beta \sinh (2 n u)=0, \quad n \geq 1, \tag{1}
\end{equation*}
$$

which appears in many scientific applications [11-13]. It should be noted that when $k=a, \alpha=(1 / 2) b$, and $\beta=0$,
(1) becomes the generalized sinh-Gordon equation [14, 15]. Furthermore, if $n=a=1$ and $b=2$, (1) reduces to the sinhGordon equation [16].

Many authors have studied the generalized double sinhGordon equation (1). Travelling waves solutions of (1) were obtained in [11] by using the tanh function method and the variable separable method. In [12] the method of bifurcation theory of dynamical system was used to prove the existence of periodic wave, solitary wave, kink and antikink wave, and unbounded wave solutions of (1). It should be noted that solutions obtained in [12] were different the ones obtained in [11]. Recently, solitary and periodic waves solutions of (1) were found in [13] by employing ( $\left.G^{\prime} / G\right)$-expansion method. It is further shown in [13] that solutions obtained by using the $\left(G^{\prime} / G\right)$-expansion method are more general than those given in [11], which were obtained by tanh function method.

In this paper, we employ an entirely different method, known as the Exp-function method, to obtain new exact solutions of the generalized sinh-Gordon equation (1). The paper is structured as follows. In Section 2, we obtain exact solutions of the generalized double sinh-Gordon equation (1) with the help of the Exp-function method. In Section 3 we present concluding remarks.

## 2. Exact Solutions of (1) Using

## Exp-Function Method

In this section we employ the Exp-function method to solve the generalized double sinh-Gordon equation (1). This method was introduced by He and Wu [17]. The Exp-function method results in the travelling wave solution based on the assumption that the solution can be expressed in the following form:

$$
\begin{equation*}
H(z)=\frac{\sum_{n=-c}^{d} a_{n} \exp (n z)}{\sum_{m=-p}^{q} b_{m} \exp (m z)}, \tag{2}
\end{equation*}
$$

where $c, d, p$, and $q$ are positive integers that can be determined and $a_{n}$ and $b_{m}$ are unknown constants. According to Exp-function method, we introduce the travelling wave substitution $u(x, t)=W(z)$, where $z=x-c t$. Then (1) transforms to the nonlinear ordinary differential equation:

$$
\begin{equation*}
\left(c^{2}-k\right) W^{\prime \prime}(z)+2 \alpha \sinh (n W(z))+\beta \sinh (2 n W(z))=0 \tag{3}
\end{equation*}
$$

Further, using the transformation $W(z)=(1 / n) \ln (H(z))$ on (3), we obtain

$$
\begin{gather*}
2\left(c^{2}-k\right) H(z) H^{\prime \prime}(z)-2\left(c^{2}-k\right) H^{\prime}(z)^{2}+2 \alpha n H(z)^{3} \\
-2 \alpha n H(z)+\beta n H(z)^{4}-\beta n=0 . \tag{4}
\end{gather*}
$$

We assume that the solution of (4) can be expressed as

$$
\begin{equation*}
H(z)=\frac{a_{c} \exp (c z)+\cdots+a_{-d} \exp (-d z)}{b_{p} \exp (p z)+\cdots+b_{-q} \exp (-q z)} \tag{5}
\end{equation*}
$$

The values of $c$ and $d, p$ and $q$ can be determined by balancing the linear term of the highest order with the highest order of nonlinear term in (4), that is, $H H^{\prime \prime}$ and $H^{4}$. By straight forward calculation, we have

$$
\begin{gather*}
H H^{\prime \prime}=\frac{c_{1} \exp [(2 c+3 p) z]+\cdots}{c_{2} \exp [5 p z]+\cdots}  \tag{6}\\
H^{4}=\frac{c_{3} \exp [4 c z]+\cdots}{c_{4} \exp [4 p z]+\cdots}=\frac{c_{3} \exp [(4 c+p) z]+\cdots}{c_{4} \exp [5 p z]+\cdots},
\end{gather*}
$$

where $c_{i}$ are coefficients only for simplicity. Balancing the highest order of Exp-function in (6), we have $2 c+3 p=4 c+p$, which yields $c=p$. Similarly, we balance the lowest order in (4) to determine values of $d$ and $q$. We have

$$
\begin{gather*}
H H^{\prime \prime}=\frac{\cdots+s_{1} \exp [-(2 d+3 q) z]}{\cdots+s_{2} \exp [-5 q z]} \\
H^{4}=\frac{\cdots+s_{3} \exp [4 d z]}{\cdots+s_{4} \exp [-4 q z]}=\frac{\cdots+s_{3} \exp [-(4 d+q) z]}{\cdots+s_{4} \exp [-5 q z]} \tag{7}
\end{gather*}
$$

where $s_{i}$ are coefficients only for simplicity. Balancing the lowest order of Exp-function in (7), we have $2 d+3 q=4 d+q$, which yields $d=q$. For simplicity, we first set $c=p=1$ and $d=q=1$. then (5) reduces to

$$
\begin{equation*}
H(z)=\frac{a_{1} \exp (z)+a_{0}+a_{-1} \exp (-z)}{b_{1} \exp (z)+b_{0}+b_{-1} \exp (-z)} \tag{8}
\end{equation*}
$$

Inserting (8) into (4) and using Maple, we obtain

$$
\begin{align*}
\frac{1}{B} & {\left[C_{4} \exp (4 z)+C_{3} \exp (3 z)+C_{2} \exp (2 z)\right.} \\
& +C_{1} \exp (z)+C_{0}+C_{-1} \exp (-z) \\
& \left.+C_{-2} \exp (-2 z)+C_{-3} \exp (-3 z)+C_{-4} \exp (-4 z)\right]=0 \tag{9}
\end{align*}
$$

where

$$
\begin{aligned}
& B=\left(b_{1} \exp (z)+b_{0}+b_{-1} \exp (-z)\right)^{4}, \\
& C_{4}=2 \alpha a_{1}^{3} b_{1} n-\beta b_{1}^{4} n+\beta a_{1}^{4} n-2 \alpha a_{1} b_{1}^{3} n, \\
& C_{3}=-2 a_{1}^{2} b_{0} b_{1} c^{2}+2 a_{1} a_{0} b_{1}^{2} c^{2}+6 \alpha a_{0} a_{1}^{2} b_{1} n \\
& -6 \alpha a_{1} b_{0} b_{1}^{2} n+2 a_{1}^{2} b_{0} b_{1} k-2 a_{0} a_{1} b_{1}^{2} k \\
& +2 \alpha a_{1}^{3} b_{0} n-2 a_{0} a_{1} b_{1}^{2} k+2 \alpha a_{1}^{3} b_{0} n \\
& +4 \beta a_{0} a_{1}^{3} n-2 \alpha a_{0} b_{1}^{3} n-4 \beta b_{0} b_{1}^{3} n, \\
& C_{2}=4 \beta a_{-1} a_{1}^{3} n-8 a_{1}^{2} b_{-1} b_{1} c^{2}+8 a_{-1} a_{1} b_{1}^{2} c^{2} \\
& +8 a_{1}^{2} b_{-1} b_{1} k-8 a_{-1} a_{1} b_{1}^{2} k+2 \alpha a_{1}^{3} b_{-1} n \\
& -2 \alpha a_{-1} b_{1}^{3} n-4 \beta b_{-1} b_{1}^{3} n+6 \alpha a_{0} a_{1}^{2} b_{0} n \\
& +6 \alpha a_{0}^{2} a_{1} b_{1} n-6 \alpha a_{1} b_{0}^{2} b_{1} n-6 \beta b_{0}^{2} b_{1}^{2} n \\
& +6 \alpha a_{-1} a_{1}^{2} b_{1} n-6 \alpha a_{1} b_{1}^{2} b_{-1} n+6 \beta a_{0}^{2} a_{1}^{2} n \\
& -6 \alpha a_{0} b_{0} b_{1}^{2} n, \\
& C_{1}=-2 a_{0}^{2} b_{0} b_{1} c^{2}+2 a_{0} a_{1} b_{0}^{2} c^{2}+2 a_{0}^{2} b_{0} b_{1} k \\
& -2 a_{0} a_{1} b_{0}^{2} k-2 a_{1}^{2} b_{0} b_{-1} c^{2}+2 a_{-1} a_{0} b_{1}^{2} c^{2} \\
& -2 a_{0} a_{-1} b_{1}^{2} k+2 \alpha a_{0}^{3} b_{1} n+4 \beta a_{0}^{3} a_{1} n-2 \alpha a_{1} b_{0}^{3} n \\
& -4 \beta b_{0}^{3} b_{1} n+12 a_{-1} a_{1} b_{0} b_{1} c^{2}-12 a_{-1} a_{1} b_{0} b_{1} k \\
& +12 a_{0} a_{1} b_{-1} b_{1} k+6 \alpha a_{0}^{2} a_{1} b_{0} n-6 \alpha a_{0} b_{0}^{2} b_{1} n \\
& +12 \alpha a_{-1} a_{0} a_{1} b_{1} n-12 \alpha a_{1} b_{-1} b_{0} b_{1} n \\
& +6 \alpha a_{-1} a_{1}^{2} b_{0} n-6 \alpha a_{0} b_{-1} b_{1}^{2} n-6 \alpha a_{-1} b_{0} b_{1}^{2} n \\
& +6 \alpha a_{0} a_{1}^{2} b_{-1} n+12 \beta a_{-1} a_{0} a_{1}^{2} n-12 \beta b_{-1} b_{0} b_{1}^{2} n \\
& +2 a_{1}^{2} b_{0} b_{-1} k-12 a_{0} a_{1} b_{-1} b_{1} c^{2},
\end{aligned}
$$

$$
\begin{align*}
& C_{0}=2 \alpha a_{0}^{3} b_{0} n-2 \alpha a_{0} b_{0}^{3} n+\beta a_{0}^{4} n \\
& +6 \alpha a_{-1} a_{1}^{2} b_{-1} n+6 \alpha a_{0}^{2} a_{1} b_{-1} n+6 \alpha a_{-1}^{2} a_{1} b_{1} n \\
& +6 \alpha a_{-1} a_{0}^{2} b_{1} n+12 \beta a_{-1} a_{0}^{2} a_{1} n-6 \alpha a_{1} b_{-1}^{2} b_{1} n \\
& -6 \alpha a_{1} b_{-1} b_{0}^{2} n-6 \alpha a_{-1} b_{-1} b_{1}^{2} n-\beta b_{0}^{4} n \\
& -6 \alpha a_{-1} b_{0}^{2} b_{1} n-12 \beta b_{-1} b_{0}^{2} b_{1} n+8 a_{-1} a_{1} b_{0}^{2} c^{2} \\
& -8 a_{0}^{2} b_{-1} b_{1} c^{2}-8 a_{-1} a_{1} b_{0}^{2} k+8 a_{0}^{2} b_{-1} b_{1} k \\
& +6 \beta a_{-1}^{2} a_{1}^{2} n-6 \beta b_{-1}^{2} b_{1}^{2} n+12 \alpha a_{-1} a_{0} a_{1} b_{0} n \\
& -12 \alpha a_{0} b_{-1} b_{0} b_{1} n \text {, } \\
& C_{-1}=12 \alpha a_{-1} a_{0} a_{1} b_{-1} n-12 \alpha a_{-1} b_{-1} b_{0} b_{1} n \\
& +2 a_{-1} a_{0} b_{0}^{2} c^{2}-2 a_{0}^{2} b_{-1} b_{0} c^{2}+2 a_{0}^{2} b_{-1} b_{0} k \\
& -2 a_{-1} a_{0} b_{0}^{2} k+2 a_{0} a_{1} b_{-1}^{2} c^{2}-2 a_{-1}^{2} b_{0} b_{1} c^{2} \\
& -2 a_{0} a_{1} b_{-1}^{2} k+2 a_{-1}^{2} b_{0} b_{1} k+2 \alpha a_{0}^{3} b_{-1} n \\
& +4 \beta a_{-1} a_{0}^{3} n-2 \alpha a_{-1} b_{0}^{3} n-4 \beta b_{-1} b_{0}^{3} n \\
& +12 a_{-1} a_{1} b_{-1} b_{0} c^{2}-12 a_{-1} a_{0} b_{-1} b_{1} c^{2} \\
& -12 a_{-1} a_{1} b_{-1} b_{0} k+12 a_{-1} a_{0} b_{-1} b_{1} k \\
& +6 \alpha a_{-1} a_{0}^{2} b_{0} n-6 \alpha a_{0} b_{-1} b_{0}^{2} n+6 \alpha a_{-1}^{2} a_{1} b_{0} n \\
& +6 \alpha a_{-1}^{2} a_{0} b_{1} n+12 \beta a_{-1}^{2} a_{0} a_{1} n-6 \alpha a_{1} b_{-1}^{2} b_{0} n \\
& -6 \alpha a_{0} b_{-1}^{2} b_{1} n-12 \beta b_{-1}^{2} b_{0} b_{1} n, \\
& C_{-2}=2 \alpha a_{-1}^{3} b_{1} n+8 a_{-1} a_{1} b_{-1}^{2} c^{2}+8 a_{-1}^{2} b_{-1} b_{1} k \\
& -8 a_{-1} a_{1} b_{-1}^{2} k+4 \beta a_{-1}^{3} a_{1} n-4 \beta b_{-1}^{3} b_{1} n \\
& -2 \alpha a_{1} b_{-1}^{3} n-8 a_{-1}^{2} b_{-1} b_{1} c^{2}+6 \alpha a_{-1} a_{0}^{2} b_{-1} n \\
& +6 \alpha a_{-1}^{2} a_{0} b_{0} n-6 \alpha a_{0} b_{-1}^{2} b_{0} n+6 \alpha a_{-1}^{2} a_{1} b_{-1} n \\
& -6 \alpha a_{-1} b_{-1}^{2} b_{1} n+6 \beta a_{-1}^{2} a_{0}^{2} n-6 \beta b_{-1}^{2} b_{0}^{2} n \\
& -6 \alpha a_{-1} b_{-1} b_{0}^{2} n, \\
& C_{-3}=6 \alpha a_{0} a_{-1}^{2} b_{-1} n-6 \alpha a_{-1} b_{-1}^{2} b_{0} n \\
& -2 a_{-1}^{2} b_{-1} b_{0} c^{2}+2 a_{-1} a_{0} b_{-1}^{2} c^{2} \\
& +2 a_{-1}^{2} b_{0} b_{-1} k-2 a_{-1} a_{0} b_{-1}^{2} k \\
& +2 \alpha a_{-1}^{3} b_{0} n+4 \beta a_{0} a_{-1}^{3} n \\
& -2 \alpha a_{0} b_{-1}^{3} n-4 \beta b_{0} b_{-1}^{3} n, \\
& C_{-4}=\beta a_{-1}^{4} n-\beta b_{-1}^{4} n+2 \alpha a_{-1}^{3} b_{-1} n-2 \alpha a_{-1} b_{-1}^{3} n . \tag{10}
\end{align*}
$$

Equating the coefficients of $\exp (z)$ in (9) to zero, we obtain a set of algebraic equations:
$C_{4}=0$,
$C_{3}=0$,
$C_{2}=0$,
$C_{1}=0$,
$C_{0}=0$,
$C_{-1}=0$,
$C_{-2}=0$,
$C_{-3}=0$,
$C_{-4}=0$.

Solving the system (11) with the help of Maple, we obtain the following three cases.

Case 1. We have the following:
$a_{-1}=b_{-1}, \quad a_{0}=-b_{0}, \quad a_{1}=b_{1}, \quad \beta=\frac{\alpha b_{0}^{2}-4 \alpha b_{1} b_{-1}}{4 b_{1} b_{-1}}$,

$$
\begin{equation*}
k=\frac{\alpha b_{0}^{2} n+2 b_{-1} b_{1} c^{2}}{2 b_{-1} b_{1}} \tag{12}
\end{equation*}
$$

Case 2. We have the following:

$$
\begin{gather*}
a_{-1}=\frac{b_{-1} b_{1}}{a_{1}}, \quad a_{0}=0, \quad b_{0}=0, \quad \alpha=\frac{-\beta\left(a_{1}^{2}+b_{1}^{2}\right)}{2 a_{1} b_{1}}, \\
k=\frac{-2 \beta a_{1}^{2} b_{1}^{2} n+\beta a_{1}^{4} n+\beta b_{1}^{4} n+8 a_{1}^{2} b_{1}^{2} c^{2}}{8 a_{1}^{2} b_{1}^{2}} . \tag{13}
\end{gather*}
$$

Case 3. We have the following:

$$
\begin{gather*}
a_{-1}=-\phi b_{1}, \quad b_{-1}=-\phi a_{1}, \quad \alpha=\frac{-\beta\left(a_{1}^{2}+b_{1}^{2}\right)}{2 a_{1} b_{1}}, \\
k=\frac{-2 \beta a_{1}^{2} b_{1}^{2} n+\beta a_{1}^{4} n+\beta b_{1}^{4} n+2 a_{1}^{2} b_{1}^{2} c^{2}}{2 a_{1}^{2} b_{1}^{2}} \tag{14}
\end{gather*}
$$

where $\phi=\left(-a_{0} a_{1}^{2} b_{0}+a_{0}^{2} a_{1} b_{1}+a_{1} b_{0}^{2} b_{1}-a_{0} b_{0} b_{1}^{2}\right) /\left(a_{1}-b_{1}\right)^{2}\left(a_{1}+\right.$ $\left.b_{1}\right)^{2}$.

Substituting values from (12) into (8), we obtain

$$
\begin{equation*}
H(z)=\frac{b_{1} \exp (z)-b_{0}+b_{-1} \exp (-z)}{b_{1} \exp (z)+b_{0}+b_{-1} \exp (-z)} \tag{15}
\end{equation*}
$$

As a result one of the solutions of (1) is given by

$$
\begin{equation*}
u_{1}(x, t)=\frac{1}{n} \ln \left(\frac{b_{1} \exp (z)-b_{0}+b_{-1} \exp (-z)}{b_{1} \exp (z)+b_{0}+b_{-1} \exp (-z)}\right) \tag{16}
\end{equation*}
$$

where $z=x-c t, \beta=\left(\alpha b_{0}^{2}-4 \alpha b_{1} b_{-1}\right) / 4 b_{1} b_{-1}$, and $k=\left(\alpha b_{0}^{2} n+\right.$ $\left.2 b_{-1} b_{1} c^{2}\right) / 2 b_{-1} b_{1}$.

As a special case, if we choose $b_{0}=2$ and $b_{-1}=b_{1}=1$ in (16), then we get $\beta=0, k=2 \alpha n+c^{2}$ and obtain the solution of the generalized sinh-Gordon equation as

$$
\begin{equation*}
u_{1}(x, t)=\frac{1}{n} \ln \left(\tanh ^{2}\left[\left(\frac{1}{2}\right)(x-c t)\right]\right) \tag{17}
\end{equation*}
$$

which is the solution obtained in $[14,15]$.

Now substituting the values from (13) (Case 2) into (8) results in the second solution of (1) as

$$
\begin{equation*}
u_{2}(x, t)=\frac{1}{n} \ln \left(\frac{a_{1} \exp (z)+\left(b_{-1} b_{1} / a_{1}\right) \exp (-z)}{b_{1} \exp (z)+b_{-1} \exp (-z)}\right) \tag{18}
\end{equation*}
$$

with $z=x-c t, \alpha=-\beta\left(a_{1}^{2}+b_{1}^{2}\right) / 2 a_{1} b_{1}$, and $k=\left(-2 \beta a_{1}^{2} b_{1}^{2} n+\right.$ $\left.\beta a_{1}^{4} n+\beta b_{1}^{4} n+8 a_{1}^{2} b_{1}^{2} c^{2}\right) / 8 a_{1}^{2} b_{1}^{2}$.

The third solution of (1) is obtained by using the values from (14) (Case 3) and substituting them into (8). Consequently, it is given by

$$
\begin{equation*}
u_{3}(x, t)=\frac{1}{n} \ln \left(\frac{a_{1} \exp (z)+a_{0}-b_{1} \phi \exp (-z)}{b_{1} \exp (z)+b_{0}-a_{-1} \phi \exp (-z)}\right) \tag{19}
\end{equation*}
$$

where $z=x-c t, \phi=\left(-a_{0} a_{1}^{2} b_{0}+a_{0}^{2} a_{1} b_{1}+a_{1} b_{0}^{2} b_{1}-a_{0} b_{0} b_{1}^{2}\right) /\left(a_{1}-\right.$ $\left.b_{1}\right)^{2}\left(a_{1}+b_{1}\right)^{2}, \alpha=-\beta\left(a_{1}^{2}+b_{1}^{2}\right) / 2 a_{1} b_{1}$, and $k=\left(-2 \beta a_{1}^{2} b_{1}^{2} n+\right.$ $\left.\beta a_{1}^{4} n+\beta b_{1}^{4} n+2 a_{1}^{2} b_{1}^{2} c^{2}\right) / 2 a_{1}^{2} b_{1}^{2}$.

To construct more solutions of (1), we now set $c=p=2$ and $d=q=2$. Then (5) reduces to

$$
\begin{align*}
H(z)= & \left(a_{2} \exp (2 z)+a_{1} \exp (z)+a_{0}+a_{-1} \exp (-z)\right. \\
& \left.+a_{-2} \exp (-2 z)\right) \\
& \times\left(b_{2} \exp (z)+b_{1} \exp (z)+b_{0}\right.  \tag{20}\\
& \left.\quad+b_{-1} \exp (-z)+b_{-2} \exp (-2 z)\right)^{-1}
\end{align*}
$$

Proceeding as above, we obtain the following three solutions of (1):

$$
\begin{gather*}
u_{4}(x, t)=\frac{1}{n} \ln \left(a_{2} \exp (2 z)+\left(\frac{a_{-1} b_{1}}{b_{-1}}\right) \exp (z)\right. \\
\left.\quad+\left(\frac{a_{-1} b_{0}}{b_{-1}}\right)+a_{-1} \exp (-z)\right) \\
\times\left(\frac{a_{2} b_{-1}}{a_{-1}} \exp (z)+b_{1} \exp (z)\right.  \tag{21}\\
\left.+b_{0}+b_{-1} \exp (-z)\right)^{-1}
\end{gather*}
$$

where $z=x-c t, \alpha=-\beta\left(a_{-1}^{2}+b_{-1}^{2}\right) / 2 a_{-1} b_{-1}$,

$$
\begin{equation*}
u_{5}(x, t)=\frac{1}{n} \ln \left(\frac{a_{2} \exp (2 z)+a_{1} \exp (z)+b_{0}}{-a_{2} \exp (z)+b_{1} \exp (z)+b_{0}}\right), \tag{22}
\end{equation*}
$$

with $z=x-c t, \beta=\alpha\left(b_{1}^{2}+4 a_{2} b_{0}\right) / 4 a_{2} b_{0}$, and $k=\left(\alpha n b_{1}^{2}+\right.$ $\left.2 a_{2} b_{0} c^{2}\right) / 2 a_{2} b_{0}$, and

$$
\begin{equation*}
u_{6}(x, t)=\frac{1}{n} \ln \left(\frac{a_{2} \exp (2 z)-b_{0}+b_{-2} \exp (-2 z)}{a_{2} \exp (2 z)+b_{0}+b_{-2} \exp (-2 z)}\right), \tag{23}
\end{equation*}
$$

where $z=x-c t, \alpha=-\left(8 a_{2} b_{-2}\left(c^{2}-k\right) / b_{0}^{2} n\right)$, and $\beta=$ $2\left(4 a_{2} b_{-2} c^{2}-4 a_{2} b_{-2} k-b_{0}^{2} c^{2}+b_{0}^{2} k\right) / b_{0}^{2} n$.


Figure 1: Profile of solution (16).


Figure 2: Profile of solution (23).

By taking $n=2, b_{-1}=-1, b_{0}=2, c=1$, and $b_{1}=-1$ in the solution (16), we have its profile given in Figure 1.

By taking $n=3, b_{-2}=1, b_{0}=2, c=1$, and $a_{1}=1$ in the solution (23), we have its profile given in Figure 2.

## 3. Concluding Remarks

In this paper we obtained new exact solutions of the generalized double sinh-Gordon equation (1) using the Expfunction method. We presented six different solutions of (1). Earlier, the tanh function, the bifurcation, and the $\left(G^{\prime} / G\right)$ expansion methods [11-13] were employed to obtain exact solutions of (1). The solutions obtained in this paper were new and were different from the ones obtained in [11-13]. By taking special values of the constants, we also retrieved the solution of the generalized sinh-Gordon equation, which was obtained in [14, 15]. The Exp-function method is very simple and straightforward method for solving nonlinear partial differential equations. Indeed this has some pronounced merit as compared to the other methods. The correctness of
the solutions obtained here has been verified by substituting them back into (1).

## Acknowledgments

Gabriel Magalakwe would like to thank SANHARP, NRF, and North-West University, Mafikeng Campus, South Africa, for their financial support.

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## Research Article

# Optimal Homotopy Asymptotic Method for Solving the Linear Fredholm Integral Equations of the First Kind 

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Received 20 April 2013; Accepted 16 June 2013
Academic Editor: Santanu Saha Ray
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#### Abstract

The aim of this study is to present the use of a semi analytical method called the optimal homotopy asymptotic method (OHAM) for solving the linear Fredholm integral equations of the first kind. Three examples are discussed to show the ability of the method to solve the linear Fredholm integral equations of the first kind. The results indicated that the method is very effective and simple.


## 1. Introduction

Integral equations of the first kind arise in several applications. These include applications in biology, chemistry, physics, and engineering. In recent years, much work has been carried out by researchers in mathematics and engineering in applying and analyzing novel numerical and semi analytical methods for obtaining solutions of integral equations of the first kind. Among these are the homotopy analysis method [1], operational Tau method [2], homotopy perturbation method [3], Adomian decomposition [3], quadrature rule [4], and automatic augmented Galerkin algorithms [5].

In this study, we develop the optimal homotopy asymptotic method (OHAM), which was proposed by Marinca et al. [6, 7], for solving the linear Fredholm integral equations of the first kind. This method is characterized by it is convergence criteria which are more flexible than other methods.

The general form of the linear Fredholm integral equations of the first kind is

$$
\begin{equation*}
f(s)=\int_{a}^{b} K(s, t) g(t) d t \tag{1}
\end{equation*}
$$

where $a$ and $b$ are constant and the functions $k(s, t)$ and $f(s)$ are known.

It should be noted that OHAM has been applied to the nonlinear Fredholm integral equations of the second kind by [8].

## 2. Application of OHAM to the Linear Fredholm Integral Equations of the First Kind

In this section, we formulate the optimal homotopy asymptotic method (OHAM) for solving the linear Fredholm integral equations of the first kind following the procedure as outlined in $[6,7]$ and other papers. Let us consider a form of the linear Fredholm integral equation of the first kind:

$$
\begin{equation*}
f(s)-\int_{a}^{b} K(x, t) g(t) d t=0 \tag{2}
\end{equation*}
$$

Using OHAM, we can obtain a family of equations as follows:

$$
\begin{align*}
(1-p) & {[L(g(s, p))+f(s)] } \\
& =H(p)[L(g(s, p))+f(s)+N(g(s, p))] \tag{3}
\end{align*}
$$

where $p \in[0,1]$ is an embedding parameter, $g(s, p)$ is unknown function, and $H(p)$ is an (nonzero) auxiliary function for $p \neq 0$ and $H(0)=0$ and given as $H(p)=\sum_{j=1}^{m} c_{j} p^{j}$ where $c_{j}, j=1,2, \ldots$, are auxiliary constants, and when $p=0$ and $p=1$ it holds that

$$
\begin{equation*}
g(s, 0)=g_{0}(s), \quad g(s, 1)=g(s) \tag{4}
\end{equation*}
$$

respectively. For obtaining the approximate solution, we use Taylor's series expansion about $p$ as follows:

$$
\begin{equation*}
g\left(s, p, c_{j}\right)=g_{0}(s)+\sum_{m=1}^{\infty} g_{m}\left(s, c_{j}\right) p^{m}, \quad j=1,2, \ldots . \tag{5}
\end{equation*}
$$

If the series (5) convergence occurs when $p=1$, one has

$$
\begin{equation*}
g\left(s, 1, c_{j}\right)=g_{0}(s)+\sum_{m=1}^{\infty} g_{m}\left(s, c_{j}\right), \quad j=1,2, \ldots \tag{6}
\end{equation*}
$$

Substituting (5) in (3) and equating the coefficients of like powers of $p$, we get as follows:

$$
\begin{gather*}
O\left(p^{0}\right): g_{0}(s)=-f(s), \\
O\left(p^{1}\right): g_{1}(s)=-c_{1} \int_{a}^{b} K(s, t) g_{0}(t) d t \\
O\left(p^{2}\right): g_{2}(s)=\left(1+c_{1}\right) g_{1}(s)-c_{1} \int_{a}^{b} K(s, t) g_{1}(t) d t \\
-c_{2} \int_{a}^{b} K(s, t) g_{0}(t) d t \\
O\left(p^{i}\right): g_{i}(s)=\left(1+c_{1}\right) g_{i-1}(s)+\sum_{j=2}^{i-1} c_{j} g_{i-j}(s) \\
-\sum_{k=1}^{i} c_{k} \int_{a}^{b} K(s, t) g_{i-k}(t) d t \tag{7}
\end{gather*}
$$

For finding the constants $c_{1}, c_{2}, c_{3}, \ldots$, we can get the result of the $m$ th-order approximations as follows:

$$
\begin{equation*}
g^{m}\left(s, c_{j}\right)=g_{0}(s)+\sum_{k=1}^{m} g_{k}\left(s, c_{j}\right), \quad j=1,2, \ldots, m \tag{8}
\end{equation*}
$$

If we substitute (8) into (1) we obtain the residual equation

$$
\begin{equation*}
R\left(s, c_{j}\right)=L\left(g^{m}\left(s, c_{j}\right)\right)+f(s)-\int_{a}^{b} K(s, t) g^{m}\left(t, c_{j}\right) d t \tag{9}
\end{equation*}
$$

If $R\left(s, c_{j}\right)=0$, then $g^{m}\left(s, c_{j}\right)$ will be the exact solution. The least squares method can be used to determine $c_{1}, c_{2}, c_{3}, \ldots$. At first we consider the functional

$$
\begin{equation*}
J\left(c_{j}\right)=\int_{a}^{b} R^{2}\left(s, c_{j}\right) d s \tag{10}
\end{equation*}
$$

By using Galerkin's method we get the following system:

$$
\begin{equation*}
\frac{\partial J}{\partial c_{j}}=2 \int_{a}^{b} R\left(s, c_{j}\right) \frac{\partial R}{\partial c_{j}} d s \tag{11}
\end{equation*}
$$

and then minimizing it to obtain the values of $c_{1}, c_{2}, \ldots, m$, we have

$$
\begin{equation*}
\frac{\partial J}{\partial c_{1}}=\frac{\partial J}{\partial c_{2}}=\cdots=\frac{\partial J}{\partial c_{m}}=0 \tag{12}
\end{equation*}
$$

With these constants, the approximate solution is determined.

## 3. Numerical Examples and Discussion

In this section, three examples of the linear Fredholm integral equations of the first kind were solved to show the efficiency of the present method. Maple software with long format and double accuracy was used to carry out the computations.

Example 1. We consider the following equation [9]:

$$
\begin{equation*}
\frac{1}{2} \sin (s)=\int_{0}^{\pi / 2} \frac{2}{\pi} \sin (s) \sin (t) g(t) d t \tag{13}
\end{equation*}
$$

for which the exact solution is $g(s)=\sin (s)$. Applying OHAM to the linear Fredholm integral equation of first kind yields

$$
\begin{gather*}
L(g(s, p))=g(s) \\
N(g(s, p))=-\int_{0}^{\pi / 2} \frac{2}{\pi} \sin (s) \sin (t) g(t) d t  \tag{14}\\
f(s)=\frac{1}{2} \sin (s)
\end{gather*}
$$

which satisfies

$$
\begin{align*}
& (1-p)\left[\left(g_{0}(s)+p g_{1}(s)+p^{2} g_{2}(s)+\cdots\right)+\frac{1}{2} \sin (s)\right] \\
& =\left(p c_{1}+p^{2} c_{2}+p^{3} c_{3}+\cdots\right) \\
& \quad \times\left[\left(g_{0}(s)+p g_{1}(s)+p^{2} g_{2}(s)+\cdots\right)+\frac{1}{2} \sin (s)\right. \\
& \quad-\int_{0}^{\pi / 2} \frac{2}{\pi} \sin (s) \sin (t)\left(g_{0}(t)+p g_{1}(t)\right. \\
& \left.\left.\quad+p^{2} g_{2}(t)+\cdots\right) d t\right] . \tag{15}
\end{align*}
$$

Now we use (7) to obtain a series of problems:

$$
\begin{align*}
& O\left(p^{0}\right): g_{0}(s)=-\frac{1}{2} \sin (s) \\
& O\left(p^{1}\right): g_{1}(s)=-c_{1} \int_{0}^{\pi / 2} \frac{2}{\pi} \sin (s) \sin (t) g_{0}(t) d t \\
& O\left(p^{2}\right): g_{2}(s)=\left(1+c_{1}\right) g_{1}(s) \\
&-c_{1} \int_{0}^{\pi / 2} \frac{2}{\pi} \sin (s) \sin (t) g_{1}(t) d t \\
&-c_{2} \int_{0}^{\pi / 2} \frac{2}{\pi} \sin (s) \sin (t) g_{0}(t) d t \tag{16}
\end{align*}
$$

Hence the solutions are

$$
\begin{gather*}
O\left(p^{0}\right): g_{0}(s)=-\frac{1}{2} \sin (s) \\
O\left(p^{1}\right): g_{1}(s)=\frac{1}{4} c_{1} \sin (s)  \tag{17}\\
O\left(p^{2}\right): g_{2}(s)= \\
=\frac{1}{4}\left(1+c_{1}\right) c_{1} \sin (s) \\
\\
\\
-\frac{1}{8} c_{1}^{2} \sin (s)+\frac{1}{4} c_{2} \sin (s)
\end{gather*}
$$

By substituting $g_{0}(s), g_{1}(s)$, and $g_{3}(s)$ solutions in (6), we obtain

$$
\begin{align*}
g(s)= & -\frac{1}{2} \sin (s)+\frac{1}{4} c_{1} \sin (s) \\
& +\frac{1}{4}\left(1+c_{1}\right) c_{1} \sin (s)  \tag{18}\\
& -\frac{1}{8} c_{1}^{2} \sin (s)+\frac{1}{4} c_{2} \sin (s)
\end{align*}
$$

For the calculations of the constants $c_{1}$ and $c_{2}$, the use of the technique mentioned in (8)-(12) yields

$$
\begin{equation*}
c_{1}=6.000000004, \quad c_{2}=-24.00000002 \tag{19}
\end{equation*}
$$

Substituting values in (18), the final solution becomes

$$
\begin{equation*}
g(s)=\sin (s) \tag{20}
\end{equation*}
$$

This is the exact solution.
Table 1 shows some numerical results of these solutions calculated according to the present method.

The exact solution, OHAM solution and absolute error of this example are shown in Figure 1.

Example 2. We consider the following equation [10]:

$$
\begin{equation*}
\frac{1}{4} s^{2}=\int_{0}^{1} \frac{5}{2} s^{2} t^{2} g(t) d t \tag{21}
\end{equation*}
$$

for which the exact solution is $g(s)=(1 / 2) s^{2}$. Applying OHAM to the linear Fredholm integral equation of first kind yields

$$
\begin{gather*}
L(g(s, p))=g(s) \\
N(g(s, p))=-\int_{0}^{1} \frac{5}{2} s^{2} t^{2} g(t) d t  \tag{22}\\
f(s)=\frac{1}{4} s^{2}
\end{gather*}
$$

which satisfies

$$
\begin{align*}
(1-p) & {\left[\left(g_{0}(s)+p g_{1}(s)+p^{2} g_{2}(s)+\cdots\right)+\frac{1}{4} s^{2}\right] } \\
= & \left(p c_{1}+p^{2} c_{2}+p^{3} c_{3}+\cdots\right) \\
& \times\left[\left(g_{0}(s)+p g_{1}(s)+p^{2} g_{2}(s)+\cdots\right)+\frac{1}{4} s^{2}\right. \\
& \left.\quad-\int_{0}^{1} \frac{5}{2} s^{2} t^{2}\left(g_{0}(t)+p g_{1}(t)+p^{2} g_{2}(t)+\cdots\right) d t\right] \tag{23}
\end{align*}
$$

Table 1: Numerical results of Example 1.

| $s$ | $g_{\text {exact }}$ | $g_{\text {OHAM }}$ | $\left\|g_{\text {exact }}-g_{\text {OHAM }}\right\|$ |
| :--- | :---: | :---: | :---: |
| 0 | 0 | 0 | 0 |
| 0.1 | 0.09983341665 | 0.09983341665 | 0 |
| 0.2 | 0.1986693308 | 0.1986693308 | 0 |
| 0.3 | 0.2955202067 | 0.2955202067 | 0 |
| 0.4 | 0.3894183423 | 0.3894183423 | 0 |
| 0.5 | 0.4794255386 | 0.4794255386 | 0 |
| 0.6 | 0.5646424734 | 0.5646424734 | 0 |
| 0.7 | 0.6442176872 | 0.6442176872 | 0 |
| 0.8 | 0.7173560909 | 0.7173560909 | 0 |
| 0.9 | 0.7833269096 | 0.7833269096 | 0 |
| 1.0 | 0.8414709848 | 0.8414709848 | 0 |

Now we use (7) to obtain a series of problems:

$$
\begin{gather*}
O\left(p^{0}\right): g_{0}(s)=-\frac{1}{4} s^{2} \\
O\left(p^{1}\right): g_{1}(s)=-c_{1} \int_{0}^{1} \frac{5}{2} s^{2} t^{2} g_{0}(t) d t \\
O\left(p^{2}\right): g_{2}(s)=\left(1+c_{1}\right) g_{1}(s) \\
-c_{1} \int_{0}^{1} \frac{5}{2} s^{2} t^{2} g_{1}(t) d t-c_{2} \int_{0}^{1} \frac{5}{2} s^{2} t^{2} g_{0}(t) d t \tag{24}
\end{gather*}
$$

Hence the solutions are

$$
\begin{gather*}
O\left(p^{0}\right): g_{0}(s)=-\frac{1}{4} s^{2} \\
O\left(p^{1}\right): g_{1}(s)=\frac{1}{8} c_{1} s^{2}  \tag{25}\\
O\left(p^{2}\right): g_{2}(s)=\frac{1}{8}\left(1+c_{1}\right) c_{1} s^{2}-\frac{1}{16} c_{1}^{2} s^{2}+\frac{1}{8} c_{2} s^{2} .
\end{gather*}
$$

By substituting $g_{0}(s), g_{1}(s)$, and $g_{3}(s)$ solutions in (6), we obtain

$$
\begin{equation*}
g(s)=-\frac{1}{4} s^{2}+\frac{1}{8} c_{1} s^{2}+\frac{1}{8}\left(1+c_{1}\right) c_{1} s^{2}-\frac{1}{16} c_{1}^{2} s^{2}+\frac{1}{8} c_{2} s^{2} . \tag{26}
\end{equation*}
$$

For the calculations of the constants $c_{1}$ and $c_{2}$, the use of the technique mentioned in (8)-(12) yields

$$
\begin{equation*}
c_{1}=6, \quad c_{2}=-24 \tag{27}
\end{equation*}
$$

Substituting values in (26), the final solution becomes

$$
\begin{equation*}
g(s)=\frac{1}{2} s^{2} \tag{28}
\end{equation*}
$$

This is the exact solution.
Table 2 shows some numerical results of these solutions calculated according to the present method.

The exact solution, OHAM solution and absolute error of this example are shown in Figure 2.

(a) Results for Example 1


- Absolute error for Example 1
(b) Absolute error for Example 1
Figure 1

$$
0.4-2
$$

(a) Results for Example 2
(b) Absolute error for Example 2

Figure 2

Table 2: Numerical results of Example 2.

| $s$ | $g_{\text {exact }}$ | $g_{\text {OHAM }}$ | $\left\|g_{\text {exact }}-g_{\text {OHAM }}\right\|$ |
| :--- | :---: | :---: | :---: |
| 0 | 0 | 0 | 0 |
| 0.1 | 0.005 | 0.005 | 0 |
| 0.2 | 0.02 | 0.02 | 0 |
| 0.3 | 0.045 | 0.045 | 0 |
| 0.4 | 0.08 | 0.08 | 0 |
| 0.5 | 0.125 | 0.125 | 0 |
| 0.6 | 0.18 | 0.18 | 0 |
| 0.7 | 0.245 | 0.245 | 0 |
| 0.8 | 0.32 | 0.32 | 0 |
| 0.9 | 0.405 | 0.405 | 0 |
| 1.0 | 0.5 | 0.5 | 0 |

Example 3. We consider the following equation [9]:

$$
\begin{equation*}
\frac{1}{2} s^{2}=\int_{0}^{1} 2 s^{2} t g(t) d t \tag{29}
\end{equation*}
$$

for which the exact solution is $g(s)=(1 / 2) s^{2}$. Applying OHAM to the linear Fredholm integral equation of first kind yields

$$
\begin{gather*}
L(g(s, p))=g(s), \\
N(g(s, p))=-\int_{0}^{1} 2 s^{2} t g(t) d t,  \tag{30}\\
f(s)=\frac{1}{2} s^{2}
\end{gather*}
$$

which satisfies

$$
\begin{align*}
& (1-p)\left[\left(g_{0}(s)+p g_{1}(s)+p^{2} g_{2}(s)+\cdots\right)+\frac{1}{2} s^{2}\right] \\
& =\left(p c_{1}+p^{2} c_{2}+p^{3} c_{3}+\cdots\right) \\
& \quad \times\left[\left(g_{0}(s)+p g_{1}(s)+p^{2} g_{2}(s)+\cdots\right)+\frac{1}{2} s^{2}\right. \\
& \left.\quad-\int_{0}^{1} 2 s^{2} t\left(g_{0}(t)+p g_{1}(t)+p^{2} g_{2}(t)+\cdots\right) d t\right] . \tag{31}
\end{align*}
$$



Figure 3

Now we use (7) to obtain a series of problems:

$$
\begin{gather*}
O\left(p^{0}\right): g_{0}(s)=-\frac{1}{2} s^{2} \\
O\left(p^{1}\right): g_{1}(s)=-c_{1} \int_{0}^{1} 2 s^{2} t g_{0}(t) d t \\
O\left(p^{2}\right): g_{2}(s)  \tag{32}\\
=\left(1+c_{1}\right) g_{1}(s)-c_{1} \int_{0}^{1} 2 s^{2} t g_{1}(t) d t \\
-c_{2} \int_{0}^{1} 2 s^{2} t g_{0}(t) d t
\end{gather*}
$$

Hence the solutions are

$$
\begin{gather*}
O\left(p^{0}\right): g_{0}(s)=-\frac{1}{2} s^{2}, \\
O\left(p^{1}\right): g_{1}(s)=\frac{1}{4} c_{1} s^{2},  \tag{33}\\
O\left(p^{2}\right): g_{2}(s)=\frac{1}{4}\left(1+c_{1}\right) c_{1} s^{2}-\frac{1}{8} c_{1}^{2} s^{2}+\frac{1}{4} c_{2} s^{2} .
\end{gather*}
$$

By substituting $g_{0}(s), g_{1}(s)$, and $g_{3}(s)$ solutions in (6), we obtain

$$
\begin{align*}
g(s)= & -\frac{1}{2} s^{2}+\frac{1}{4} c_{1} s^{2}+\frac{1}{4}\left(1+c_{1}\right) c_{1} s^{2} \\
& -\frac{1}{8} c_{1}^{2} s^{2}+\frac{1}{4} c_{2} s^{2} . \tag{34}
\end{align*}
$$

For the calculations of the constants $c_{1}$ and $c_{2}$, the use of the technique mentioned in (8)-(12) yields

$$
\begin{equation*}
c_{1}=6, \quad c_{2}=-24 \tag{35}
\end{equation*}
$$

Substituting values in (34), the final solution becomes

$$
\begin{equation*}
g(s)=s^{2} . \tag{36}
\end{equation*}
$$

This is the exact solution.

Table 3: Numerical results of Example 3.

| $s$ | $g_{\text {exact }}$ | $g_{\text {OHAM }}$ | $\left\|g_{\text {exact }}-g_{\text {OHAM }}\right\|$ |
| :--- | :---: | :---: | :---: |
| 0 | 0 | 0 | 0 |
| 0.1 | 0.01 | 0.01 | 0 |
| 0.2 | 0.04 | 0.04 | 0 |
| 0.3 | 0.09 | 0.09 | 0 |
| 0.4 | 0.16 | 0.16 | 0 |
| 0.5 | 0.25 | 0.25 | 0 |
| 0.6 | 0.36 | 0.36 | 0 |
| 0.7 | 0.49 | 0.49 | 0 |
| 0.8 | 0.64 | 0.64 | 0 |
| 0.9 | 0.81 | 0.81 | 0 |
| 1.0 | 1.0 | 1.0 | 0 |

Table 3 shows some numerical results of these solutions calculated according to the present method.

The exact solution, OHAM solution and absolute error of this example are shown in Figure 3.

## 4. Conclusions

In this paper, we presented the application of the OHAM in solving the linear Fredholm integral equations of the first kind. This method was tested on three different examples. This method proved to be an accurate and efficient technique for finding approximate solutions for the linear Fredholm integral equations of the first kind.

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## Research Article

# Numerical Study of Two-Dimensional Volterra Integral Equations by RDTM and Comparison with DTM 

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Received 17 April 2013; Accepted 10 June 2013
Academic Editor: Santanu Saha Ray
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#### Abstract

The two-dimensional Volterra integral equations are solved using more recent semianalytic method, the reduced differential transform method (the so-called RDTM), and compared with the differential transform method (DTM). The concepts of DTM and RDTM are briefly explained, and their application to the two-dimensional Volterra integral equations is studied. The results obtained by DTM and RDTM together are compared with exact solution. As an important result, it is depicted that the RDTM results are more accurate in comparison with those obtained by DTM applied to the same Volterra integral equations. The numerical results reveal that the RDTM is very effective, convenient, and quite accurate compared to the other kind of nonlinear integral equations. It is predicted that the RDTM can be found widely applicable in engineering sciences.


## 1. Introduction

Mathematical modeling of many problems in science, engineering, physics, and other disciplines leads to linear and nonlinear integrodifferential equations (IDE). The great use of mathematical models including integrodifferential equations is one of the main reasons obtaining the solutions of this kind of problems (see, e.g., [1-3] and the references therein). So, it is very important to get some information about the analytical solutions of these problems because these solutions give significant information about the character of the modeled event. But, in some cases, it is more difficult to obtain analytical solutions of these models. These are usually difficult to solve analytically, and in many cases the solution must be approximated. To approximate the solutions of these models, in recent years several numerical approaches have been proposed.

In this paper, we consider the following Volterra type of integral equation $[4,5]$ :

$$
\begin{equation*}
u(x, t)-\int_{0}^{t} \int_{0}^{x} K(x, t, y, z, u(y, z)) d y d z=f(x, t) \tag{1}
\end{equation*}
$$

where $K$ and $f$ are continuous functions and $K$ has the following form:

$$
\begin{equation*}
K(x, t, y, z, u(y, z))=\sum_{i=0}^{m} p_{i}(x, t) q_{i}(y, z, u(y, z)) \tag{2}
\end{equation*}
$$

The one-dimensional Volterra type of integral equation has been solved by many numerical methods, such as collocation methods [1], Taylor-series expansion methods [2], Gausstype quadratures method [3], spectral methods [6], Chebyshev polynomial method [7], Tau method [8], sine-cosine wavelets method [9], Monte Carlo method [10], and Haar functions method [11].

But in two-dimensional cases, a small amount of work has been done (see, e.g., [12-14]). Very recently, Tari et al. in [4] employed the classic differential transform method for solving two-dimensional Volterra type of integral equations (1), and Jang in [5] improved the proofs of the presented theorems by Tari et al. in [4]. They derived fundamental properties of the differential transforms of some kernel functions $K$ in Volterra integral equations.

However, the classic differential transform method, introduced by Zhou [15], is based on the definition of the differential transform, which is a Taylor series. Thus, it requires a cumbersome calculation to obtain the basic properties of the differential transforms. Some of DTM applications are mentioned in [16-21].

Recently, Keskin and Oturanç introduced a reduced form of DTM as reduced DTM (RDTM) and applied it to approximate some PDE [22] and factional PDEs [23]. More recently, Abazari and Ganji [24] extended RDTM to study the partial differential equation with proportional delay in $t$ and shrinking in $x$ and showed that, as a special advantage of RDTM rather than DTM, the reduced differential transform recursive equations produce exactly all the Poisson series coefficients of solutions based on the initial condition as weighted function, whereas the differential transform recursive equations produce exactly all the Taylor series coefficients of solutions.

Here, we suggest the RDTM, for the approximating of the solutions of the two-dimensional Volterra integral equations (1) with the same kernel functions in [4, 5]. In order to demonstrate the effectiveness of the RDTM, the illustrative examples for the same kernel function of references [5] are presented. These examples show that the RDTM produces exactly all the Poisson series coefficients (see Remark 5) of the exact solutions, whereas, the classic DTM produces exactly all the Taylor series coefficients of the exact solutions. As an important result, notwithstanding the simplicity and robustness of RDTM, it is depicted that the RDTM results are more accurate in comparison with those obtained by classic DTM.

## 2. Basic Definitions

With reference to the articles [16-21], the basic definitions of two-dimensional differential transform method (DTM) and their reduced form (RDTM) are introduced in the following two subsections, respectively.
2.1. Two-Dimensional DTM. Consider a function of two variables $w(x, t)$, and suppose that it can be represented as a product of two single-variable functions, that is, $w(x, t)=$ $f(x) g(t)$. On the basis of the properties of the one-dimensional differential transform, the function $w(x, t)$ can be represented as

$$
\begin{equation*}
w(x, t)=\sum_{i=0}^{\infty} F(i) x^{i} \sum_{j=0}^{\infty} G(j) t^{j}=\sum_{i=0}^{\infty} \sum_{j=0}^{\infty} W(i, j) x^{i} t^{j}, \tag{3}
\end{equation*}
$$

where $W(i, j)=F(i) G(j)$ is called the spectrum of $w(x, t)$.
The basic definitions and operations for two-dimensional differential transform are introduced as follows.

Definition 1. If $w(x, t)$ is analytic and continuously differentiable with respect to time $t$ in the domain of interest, then

$$
\begin{equation*}
W(m, n)=\frac{1}{m!n!}\left[\frac{\partial^{m+n}}{\partial x^{m} \partial t^{n}} w(x, t)\right]_{\substack{x=x_{0} \\ t=t_{0}}}, \tag{4}
\end{equation*}
$$

where the spectrum function $W(m, n)$ is the transformed function, which is also called $T$-function in brief.

The differential inverse transform of $W(k, h)$ is defined as

$$
\begin{equation*}
w(x, t)=\sum_{m=0}^{\infty} \sum_{n=0}^{\infty} W(m, n)\left(x-x_{0}\right)^{m}\left(t-t_{0}\right)^{n} . \tag{5}
\end{equation*}
$$

Combining (4) and (5), it can be obtained that

$$
\begin{align*}
w(x, t)= & \sum_{m=0}^{\infty} \sum_{n=0}^{\infty} \frac{1}{m!n!}\left[\frac{\partial^{m+n}}{\partial x^{m} \partial t^{n}} w(x, t)\right]_{\substack{x=x_{0} \\
t=t_{0}}}  \tag{6}\\
& \times\left(x-x_{0}\right)^{m}\left(t-t_{0}\right)^{n}
\end{align*}
$$

When $\left(x_{0}, t_{0}\right)$ are taken as $(0,0)$, then (5) can be expressed as

$$
\begin{equation*}
w(x, t)=\sum_{m=0}^{\infty} \sum_{n=0}^{\infty} W(m, n) x^{m} t^{n} \tag{7}
\end{equation*}
$$

In real applications, the function $w(x, t)$ is represented by a finite series of (7) that can be written as

$$
\begin{equation*}
w_{M, N}(x, t)=\sum_{m=0}^{M} \sum_{n=0}^{N} W(m, n) x^{m} t^{n}+R_{M, N}(x, t) \tag{8}
\end{equation*}
$$

and (7) implies that $R_{M, N}(x, t)=\sum_{m=M+1}^{\infty} \sum_{n=N+1}^{\infty} W(m$, $n) x^{m} t^{n}$ is negligibly small. Usually, the values of $M$ and $N$ are decided by convergency of the series coefficients.

From the above definitions, it can be found that the concept of the two-dimensional differential transform is derived from the two-dimensional Taylor series expansion. With (4) and (5), the fundamental mathematical operations performed using the two-dimensional differential transform may be readily obtained, and these are listed in Table 1. (See $[4,5,15,16]$.)

Recently, Jang [5] extended the two-dimensional DTM on (1) as follows.

Theorem 2. Assume that $U(m, n), V(m, n), H(m, n)$, and $G(m, n)$ are the differential transforms of the functions $u(x, t)$, $v(x, t), h(x, t)$, and $g(x, t)$, respectively; then we have the following:
(a) if $g(x, t)=\int_{0}^{t} \int_{0}^{x} u(y, z) d y d z$, then

$$
\begin{gather*}
G(m, 0)=G(0, n)=0, \quad m, n=0,1, \ldots \\
G(m, n)=\frac{1}{m n} U(m-1, n-1), \quad m, n=1,2, \ldots, \tag{9}
\end{gather*}
$$

Table 1: The fundamental operations of two-dimensional differential transform method.

| Original function | Transformed function |
| :--- | ---: |
| $w(x, t)=u(x, t) \pm v(x, t)$ | $W(m, n)=U(m, n) \pm V(m, n)$ |
| $w(x, t)=c u(x, t)$ | $W(m, n)=c U(m, n)$ |
| $w(x, t)=\frac{\partial}{\partial x} u(x, t)$ | $W(m, n)=(m+1) U(m+1, n)$ |
| $w(x, t)=\frac{\partial}{\partial t} u(x, t)$ | $W(m, n)=(m+1) U(m, n+1)$ |
| $w(x, t)=\frac{\partial^{r+s}}{\partial x^{r} \partial t^{s}} u(x, t)$ | $W(m, n)=\frac{(m+r)!(n+s)!}{m!n!} U(m+r, n+s)$ |
| $w(x, t)=u(x, t) v(x, t)$ | $W(m, n)=\sum_{r=0}^{m} \sum_{s=0}^{n} U(r, n-s) V(m-r, s)$ |
| $w(x, t)=x^{\alpha} t^{\beta}$ | $W(m, n)=\delta(m-\alpha, n-\beta)= \begin{cases}1 & m=\alpha, n=\beta \\ 0 & \text { otherwise }\end{cases}$ |

(b) if $g(x, t)=\int_{0}^{t} \int_{0}^{x} u(y, z) v(y, z) d y d z$, then

$$
\begin{gather*}
G(m, 0)=G(0, n)=0, \quad m, n=0,1, \ldots, \\
G(m, n)=\frac{1}{m n} \sum_{\ell=0}^{n-1} \sum_{k=0}^{m-1} U(k, \ell) V(m-k-1, n-\ell-1), \\
m, n=1,2, \ldots, \tag{10}
\end{gather*}
$$

(c) if $g(x, t)=h(x, t) \int_{0}^{t} \int_{0}^{x} u(y, z) d y d z$, then

$$
\begin{gather*}
G(m, 0)=G(0, n)=0, \quad m, n=0,1, \ldots \\
G(m, n)=\frac{1}{m n} \sum_{\ell=0}^{n-1} \sum_{k=0}^{m-1} H(k, \ell) \frac{V(m-k-1, n-\ell-1)}{(m-k)(n-\ell)}, \\
m, n=1,2, \ldots \tag{11}
\end{gather*}
$$

Proof. See [5].
Theorem 3. Assume that $U(m, n), V(m, n)$, and $G(m, n)$ are the differential transforms of the functions $u(x, t), v(x, t)$, and $g(x, t)$, respectively; then we have the following:
(a) if $g(x, t)=\int_{0}^{t} \int_{0}^{x}(u(y, z) / v(y, z)) d y d z$, then

$$
G(m, 0)=G(0, n)=0, \quad m, n=0,1, \ldots
$$

$$
\begin{align*}
& \sum_{\ell=0}^{n-1} \sum_{k=0}^{m-1}(m-k+1)(n-\ell+1) V(k, \ell) G(m-k+1, n-\ell+1) \\
& \quad=U(m, n) \tag{12}
\end{align*}
$$

(b) if $g(x, t)=(1 / v(x, t)) \int_{0}^{t} \int_{0}^{x} u(y, z) d y d z$, then

$$
\begin{gather*}
G(m, 0)=G(0, n)=0, \quad m, n=0,1, \ldots \\
\sum_{\ell=0}^{n-1} \sum_{k=0}^{m-1} V(k, \ell) G(m-k+1, n-\ell+1)  \tag{13}\\
=\frac{1}{(m+1)(n+1)} U(m, n)
\end{gather*}
$$

Proof. See [5].
2.2. Two-Dimensional Reduced DTM (RDTM). Consider a function of two variables $w(x, t)$, and suppose that it can be represented as a product of two single-variable functions, that is, $w(x, t)=f(x) g(t)$. Based on the properties of onedimensional differential transform, the function $w(x, t)$ can be represented as

$$
\begin{equation*}
w(x, t)=\sum_{i=0}^{\infty} F(i) x^{i} \sum_{j=0}^{\infty} G(j) t^{j}=\sum_{i=0}^{\infty} \sum_{j=0}^{\infty} W(i, j) x^{i} t^{j}, \tag{14}
\end{equation*}
$$

where $W(i, j)=F(i) G(j)$ is called the spectrum of $w(x, t)$.
Remark 4. The poisson function series generates a multivariate Taylor series expansion of the input expression $w$, with respect to the variables $X$, to order $n$, using the variable weights $W$.

Remark 5. The relationship introduced in (14) is the poisson series form of the input expression $w(x, t)$, with respect to the variables $x$ and $t$, to order $N$, using the variable weights $W_{k}(x)$.

Similarly on previous section, the basic definitions of twodifferential reduced differential transformation are introduced as follows.

Definition 6. If $w(x, t)$ is analytical function in the domain of interest, then the spectrum function

$$
\begin{equation*}
W_{k}(x)=\frac{1}{k!}\left[\frac{\partial^{k}}{\partial t^{k}} w(x, t)\right]_{t=t_{0}} \tag{15}
\end{equation*}
$$

is the reduced transformed function of $w(x, t)$.
Similarly on previous sections, the lowercase $w(x, t)$ respects the original function while the uppercase $W_{k}(x)$ stands for the reduced transformed function. The differential inverse transform of $W_{k}(x)$ is defined as

$$
\begin{equation*}
w(x, t)=\sum_{k=0}^{\infty} W_{k}(x)\left(t-t_{0}\right)^{k} \tag{16}
\end{equation*}
$$

Combining (15) and (16), it can be obtained that

$$
\begin{equation*}
w(x, t)=\sum_{k=0}^{\infty} \frac{1}{k!}\left[\frac{\partial^{k}}{\partial t^{k}} w(x, t)\right]_{t=t_{0}}\left(t-t_{0}\right)^{k} \tag{17}
\end{equation*}
$$

In real applications, the function $w(x, t)$ is represented by a finite series of (16), around $t_{0}=0$, and can be written as

$$
\begin{equation*}
w_{n}(x, t)=\sum_{k=0}^{n} W_{k}(x) t^{k}+R_{n}(x, t) \tag{18}
\end{equation*}
$$

and (18) implies that $R_{n}(x, t)=\sum_{k=n+1}^{\infty} W_{k}(x) t^{k}$ is negligibly small. Usually, the values of $n$ and $m$ are decided by convergency of the series coefficients. From the above proposition, it can be found that the concept of the reduced two-dimensional differential transform is derived from the two-dimensional differential transform method. With (15) and (16), the fundamental mathematical operations performed by reduced twodimensional differential transform can readily be obtained and listed in Table 2.

Similarly on previous subsection, we can extend the RDTM on Volterra integral equations (1) as follow.

Theorem 7. Assume that $U_{k}(x), V_{k}(x), H_{k}(x)$, and $W_{k}(x)$ are the reduced differential transforms of the functions $u(x, t)$, $v(x, t), h(x, t)$, and $w(x, t)$, respectively; then we have the following:

$$
\begin{gather*}
\text { (a) if } w(x, t)=\int_{t_{0}}^{t} \int_{x_{0}}^{x} u(y, z) v(y, z) d y d z \text {, then } \\
W_{k}(x)=\frac{1}{k} \int_{x_{0}}^{x}\left(\sum_{r=0}^{k-1} U_{r}(y) V_{k-r-1}(y)\right) d y, \quad k=1,2, \ldots, \tag{19}
\end{gather*}
$$

(b) if $w(x, t)=h(x, t) \int_{t_{0}}^{t} \int_{x_{0}}^{x} u(y, z) d y d z$, then

$$
\begin{equation*}
W_{k}(x)=\sum_{r=0}^{k} \frac{1}{k-r} H_{r}(x) \int_{x_{0}}^{x} U_{k-r-1}(y) d y, \quad k=1,2, \ldots \tag{20}
\end{equation*}
$$

Proof. (a) According to the fundamental operations of twodimensional RDTM listed in Table 2 and from Leibnitz formula, we get

$$
\begin{align*}
\frac{\partial^{k}}{\partial t^{k}} w(x, t)= & \frac{\partial^{k}}{\partial t^{k}}\left(\int_{t_{0}}^{t} \int_{x_{0}}^{x} u(y, z) v(y, z) d y d z\right) \\
= & \int_{x_{0}}^{x} \frac{\partial^{k-1}}{\partial t^{k-1}}\{u(y, t) v(y, t)\} d y \\
= & \int_{x_{0}}^{x}\left\{\begin{array}{c}
\sum_{r=0}^{k-1}\binom{k-1}{r} \frac{\partial^{r}}{\partial t^{r}} u(y, t) \\
\\
\end{array} \quad \frac{\partial^{k-r-1}}{\partial t^{k-r-1}} v(y, t)\right\} d y \tag{21}
\end{align*}
$$

therefore

$$
\begin{align*}
{\left[\frac{\partial^{k}}{\partial t^{k}} w(x, t)\right]_{t=t_{0}}=} & \int_{x_{0}}^{x}\left\{\begin{array}{c}
k-1 \\
r=0 \\
k-1 \\
r
\end{array}\right) r!(k-r-1)! \\
& \left.\quad \times U_{r}(y) V_{k-r-1}(y)\right\} d y \\
= & (k-1)!\int_{x_{0}}^{x}\left\{\sum_{r=0}^{k-1} U_{r}(y) V_{k-r-1}(y)\right\} d y \tag{22}
\end{align*}
$$

and then, from using (15), for $k=1,2, \ldots$, we get

$$
\begin{equation*}
W_{k}(x)=\frac{1}{k} \int_{x_{0}}^{x}\left\{\sum_{r=0}^{k-1} U_{r}(y) V_{k-r-1}(y)\right\} d y \tag{23}
\end{equation*}
$$

(b) Analogous to part (a), we get

$$
\begin{align*}
\frac{\partial^{k}}{\partial t^{k}} w(x, t) & =\frac{\partial^{k}}{\partial t^{k}}\left(h(x, t) \int_{t_{0}}^{t} \int_{x_{0}}^{x} u(y, z) d y d z\right) \\
& =\sum_{r=0}^{k}\binom{k}{r} \frac{\partial^{r}}{\partial t^{r}} h(x, t) \int_{x_{0}}^{x} \frac{\partial^{k-r-1}}{\partial t^{k-r-1}} u(y, t) d y \tag{24}
\end{align*}
$$

Table 2: The fundamental operations of two-dimensional RDTM.

| Original function | Reduced transformed function |
| :--- | ---: |
| $w(x, t)=u(x, t) \pm v(x, t)$ | $W_{k}(x)=U_{k}(x) \pm V_{k}(x)$ |
| $w(x, t)=\frac{\partial}{\partial x} u(x, t)$ | $W_{k}(x)=\frac{\partial}{\partial x} U_{k}(x)$ |
| $w(x, t)=\frac{\partial}{\partial t} u(x, t)$ | $W_{k}(x)=(k+1) U_{k+1}(x)$ |
| $w(x, t)=\frac{\partial^{r+s}}{\partial x^{r} \partial t^{s}} u(x, t)$ | $W_{k}(x)=\frac{(k+s)!}{k!} \frac{\partial^{r}}{\partial x^{r}} U_{k+s}(x)$ |
| $w(x, t)=u(x, t) v(x, t)$ | $W_{k}(x)=\sum_{r=0}^{k} U_{r}(x) V_{k-r}(x)$ |
| $w(x, t)=x^{m} t^{n}$ | $W_{k}(x)=x^{m} \delta(k-n)= \begin{cases}x^{m} & k=n \\ 0 & \text { otherwise }\end{cases}$ |

therefore

$$
\begin{align*}
{\left[\frac{\partial^{k}}{\partial t^{k}} w(x, t)\right]_{t=t_{0}}=} & \sum_{r=0}^{k}\binom{k}{r} r!(k-r-1)!H_{r}(x) \\
& \times \int_{x_{0}}^{x} U_{k-r-1}(y) d y \\
= & \sum_{r=0}^{k} \frac{k!}{k-r} H_{r}(x) \int_{x_{0}}^{x} U_{k-r-1}(y) d y \tag{25}
\end{align*}
$$

and then from using (15), for $k=1,2, \ldots$, we get

$$
\begin{equation*}
W_{k}(x)=\sum_{r=0}^{k} \frac{1}{k-r} H_{r}(x) \int_{x_{0}}^{x} U_{k-r-1}(y) d y . \tag{26}
\end{equation*}
$$

Theorem 8. Assume that $U_{k}(x), V_{k}(x)$, and $W_{k}(x)$ are the reduced differential transforms of the functions $u(x, t), v(x, t)$, and $w(x, t)$, respectively; then we have the following:
(a) if $w(x, t)=\int_{t_{0}}^{t} \int_{x_{0}}^{x}(u(y, z) / v(y, z)) d y d z$, then

$$
\begin{equation*}
U_{k}(x)=\sum_{r=0}^{k}(r+1) \frac{\partial W_{r+1}(x)}{\partial x} V_{k-r}(y), \quad k=0,1,2, \ldots, \tag{27}
\end{equation*}
$$

(b) if $w(x, t)=(1 / v(x, t)) \int_{t_{0}}^{t} \int_{x_{0}}^{x} u(y, z) d y d z$, then

$$
\begin{equation*}
k \sum_{r=0}^{k} W_{r}(x) V_{k-r}(x)=\int_{x_{0}}^{x} U_{k-1}(y) d y, \quad k=1,2, \ldots \tag{28}
\end{equation*}
$$

Proof. (a) By following the same manner as in the Theorem 7, we get

$$
\begin{equation*}
u(x, t)=\frac{\partial^{2} w(x, t)}{\partial x \partial t} v(x, t) \tag{29}
\end{equation*}
$$

then

$$
\begin{align*}
\frac{\partial^{k}}{\partial t^{k}} u(x, t) & =\frac{\partial^{k}}{\partial t^{k}}\left(\frac{\partial^{2} w(x, t)}{\partial x \partial t} v(x, t)\right) \\
& =\sum_{r=0}^{k}\binom{k}{r} \frac{\partial^{r+2}}{\partial x \partial t^{r+1}} w(x, t) \frac{\partial^{k-r}}{\partial t^{k-r}} v(y, t) \tag{30}
\end{align*}
$$

therefore

$$
\begin{equation*}
k!U_{k}(x)=\sum_{r=0}^{k}\binom{k}{r}(r+1)!(k-r)!\frac{\partial W_{r+1}}{\partial x}(x) V_{k-r}(x) \tag{31}
\end{equation*}
$$

and then from using (15), for $k=0,1,2, \ldots$, we get

$$
\begin{equation*}
U_{k}(x)=\sum_{r=0}^{k}(r+1) \frac{\partial W_{r+1}}{\partial x}(x) V_{k-r}(x) \tag{32}
\end{equation*}
$$

(b) Analogous to part (a), we get

$$
\begin{equation*}
w(x, t) v(x, t)=\int_{t_{0}}^{t} \int_{x_{0}}^{x} u(y, z) d y d z \tag{33}
\end{equation*}
$$

then

$$
\begin{equation*}
\frac{\partial^{k}}{\partial t^{k}}\{w(x, t) v(x, t)\}=\int_{x_{0}}^{x} \frac{\partial^{k-1}}{\partial t^{k-1}} u(y, t) d y \tag{34}
\end{equation*}
$$

therefore

$$
\begin{equation*}
\sum_{r=0}^{k}\binom{k}{r} \frac{\partial^{r}}{\partial t^{r}} w(x, t) \frac{\partial^{k-r}}{\partial t^{k-r}} v(x, t)=\int_{x_{0}}^{x} \frac{\partial^{k-1}}{\partial t^{k-1}} u(y, t) d y \tag{35}
\end{equation*}
$$

and then from using (15), for $k=1,2, \ldots$, we get

$$
\begin{equation*}
k!\sum_{r=0}^{k} W_{r}(x) V_{k-r}(x)=(k-1)!\int_{x_{0}}^{x} U_{k-1}(y) d y \tag{36}
\end{equation*}
$$

and therefore

$$
\begin{equation*}
k \sum_{r=0}^{k} W_{r}(x) V_{k-r}(x)=\int_{x_{0}}^{x} U_{k-1}(y) d y \tag{37}
\end{equation*}
$$

## 3. Numerical Results of DTM and RDTM

In this section, the reduced differential transform technique is described to solve a class of Volterra integral equations (1) with kernel functions of (2). In order to demonstrate the effectiveness of the RDTM, the illustrative examples for the same kernel function of [5] are presented. In each example, the numerical results of DTM, RDTM, and their comparisons with exact solution are given in separate tables. The results of the test examples show that the RDTM results are more powerful than DTM results.

Example 9. In the first example, consider the following twodimensional Volterra integral equation [5]:

$$
\begin{align*}
u(x, t) & -\int_{0}^{t} \int_{0}^{x} \frac{u(y, z)}{2+\sin (y+z)} d y d z  \tag{38}\\
& =(x-t)(4+2 \sin (x+t)-x t) .
\end{align*}
$$

(a) DTM: Jang [5] solved this equation by using DTM and obtained the following five-term DTM solution:

$$
\begin{align*}
u_{5,5}(x, t)= & \left(4 x+2 x^{2}-\frac{x^{4}}{3}\right)+\left(-4-\frac{2 x^{3}}{3}+\frac{x^{5}}{15}\right) t \\
& +\left(-2+\frac{x^{4}}{12}\right) t^{2}+\left(\frac{2 x}{3}-\frac{x^{5}}{180}\right) t^{3}  \tag{39}\\
& +\left(\frac{1}{3}-\frac{x^{2}}{12}\right) t^{4}+\left(-\frac{x}{15}+\frac{x^{3}}{180}\right) t^{5}
\end{align*}
$$

(b) RDTM: from Volterra integral equation (38), it is easy to see that the $u(x, 0)=x(4+2 \sin (x))$, and therefore RDTM version is

$$
\begin{equation*}
U_{0}(x)=x(4+2 \sin (x)) . \tag{40}
\end{equation*}
$$

By applying the RDTM properties listed in Theorem 8, on Volterra integral equation (38), for $k=0,1,2, \ldots$, we get

$$
\begin{align*}
& U_{k}(x) \\
& \begin{aligned}
=\sum_{r=0}^{k}(r & +1) \frac{d}{d x} \\
& \times\left\{U_{r+1}(x)\right. \\
& \quad-\sum_{\ell=0}^{r+1}\left(x \delta_{\ell, 0}-\delta_{\ell, 1}\right) \\
& \times\left[4 \delta_{r+1-\ell, 0}+\frac{2 \sin (x+((r+1-\ell) \pi / 2))}{(r+1-\ell)!}\right. \\
& \left.\left.-x \delta_{r+1-\ell, 1}\right]\right\} \\
& \times\left\{2 \delta_{k-r, 0}+\sin \left(x+\frac{(k-r) \pi}{2}\right)\right\},
\end{aligned}
\end{align*}
$$

where $U_{i}(x)$ is the reduced differential transform of $u(x, t)$. After expanding the RDTM recurrence equations (41), with initial value of (40), for $k=0,1,2,3,4$, the first five terms of $U_{k}(x)$ are obtained as follows:

$$
\begin{align*}
& U_{1}(x)=2 x \cos (x)-4-2 \sin (x) \\
& U_{2}(x)=-x \sin (x)-2 \cos (x) \\
& U_{3}(x)=\sin (x)-\frac{1}{3} x \cos (x)  \tag{42}\\
& U_{4}(x)=\frac{1}{3} \cos (x)+\frac{1}{12} x \sin (x) \\
& U_{5}(x)=\frac{1}{60} x \cos (x)-\frac{1}{12} \sin (x)
\end{align*}
$$

In the same manner, the rest of the components can be obtained by using the recursive equations (41). Substituting the quantities (41) in (18), the approximation solution of Volterra integral equation (38) in the Poisson series form is

$$
\begin{align*}
U_{5}(x, t)= & 2 x(2+\sin (x))+(2 x \cos (x)-4-2 \sin (x)) t \\
& +(-x \sin (x)-2 \cos (x)) t^{2}+\left(\sin (x)-\frac{x \cos (x)}{3}\right) t^{3} \\
& +\left(\frac{\cos (x)}{3}+\frac{x \sin (x)}{12}\right) t^{4} \\
& +\left(\frac{x \cos (x)}{60}-\frac{\sin (x)}{12}\right) t^{5}, \tag{43}
\end{align*}
$$

which is the same as the first five terms of the Poisson series of the exact solution $u(x, t)=2(x-t)(2+\sin (x+t))$.

Table 3: Comparisons of the exact solution $u(x, t)=2(x-t)(2+\sin (x+t))$, with $U_{5,5}(x, t)$ obtained by classic DTM [5] and $U_{5}(x, t)$ obtained by RDTM at some test points ( $x, t$ ) in Example 9 .

| $x$ | $t$ | $u(x, t)$ | Classic DTM [5] |  | Reduced DTM |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  |  |  | $U_{5,5}(x, t)$ | $\left\|u(x, t)-U_{5,5}(x, t)\right\|$ | $U_{5}(x, t)$ | $\left\|u(x, t)-U_{5}(x, t)\right\|$ |
| 0.2 | 0.1 | +0.4591040413 | +0.4591115320 | $7.4906677321 e-06$ | $+0.4591040577$ | $1.6393507773 e-08$ |
|  | 0.4 | -1.0258569894 | -1.0257575253 | $9.9464024681 e-05$ | -1.0257906614 | $6.6327927673 e-05$ |
|  | 0.7 | -2.7833269096 | -2.7813944067 | $1.9325029608 e-03$ | -2.7814532650 | $1.8736446239 e-03$ |
|  | 1 | -4.6912625375 | -4.6755840000 | $1.5678537548 e-02$ | -4.6756686352 | $1.5593902330 e-02$ |
| 0.5 | 0.1 | +2.0517139787 | +2.0516661667 | $4.7812049361 e-05$ | +2.0517139939 | $1.5172903822 e-08$ |
|  | 0.4 | +0.5566653819 | +0.5573213333 | $6.5595140784 e-04$ | +0.5567258924 | $6.0510490391 e-05$ |
|  | 0.7 | -1.1728156344 | -1.1698785000 | $2.9371343869 e-03$ | -1.1711316158 | $1.6840186272 e-03$ |
|  | 1 | -2.9974949866 | -2.9817708333 | $1.5724153271 e-02$ | -2.9836944470 | $1.3800539634 e-02$ |
| 0.8 | 0.1 | +3.8966576735 | +3.8929643413 | $3.6933321451 e-03$ | +3.8966576865 | $1.3025739598 e-08$ |
|  | 0.4 | +2.3456312688 | +2.3439851520 | $1.6461167738 e-03$ | +2.3456822669 | $5.0998086549 e-05$ |
|  | 0.7 | +0.5994989973 | +0.6014245867 | $1.9255893459 e-03$ | +0.6008906123 | $1.3916149372 e-03$ |
|  | 1 | -1.1895390524 | -1.1754026667 | $1.4136385685 e-02$ | -1.1783733000 | $1.1165752332 e-02$ |
| 1 | 0.1 | +5.2041732481 | +5.1881855000 | 1.5987748111e-02 | +5.2041732592 | $1.1127633925 e-08$ |
|  | 0.4 | +3.5825396760 | +3.5680853333 | $1.4454342653 e-02$ | +3.5825824990 | $4.2822997450 e-05$ |
|  | 0.7 | +1.7949988863 | +1.7840151667 | $1.0983719605 e-02$ | +1.7961453619 | $1.1464756153 e-03$ |
|  | 1 | +0.0000000000 | -0.0000000000 | -0.0000000000 | $+0.0090050384$ | $9.0050384311 e-03$ |

The numerical results obtained with RDTM are presented in Table 3, in comparison with the classic DTM solution of [5] and the exact solution $u(x, t)=2(x-t)(2+\sin (x+t))$, for some points of the intervals $0 \leq x \leq 1$ and $0 \leq t \leq 1$.

Example 10. In the second example, consider the following two-dimensional Volterra integral equation [5]:

$$
\begin{align*}
u(x, t) & -2 \int_{0}^{t} \int_{0}^{x} e^{y-z} u(y, z) d y d z \\
& =\sin (x+t)\left(e^{x-t}+1\right)-e^{-t} \sin (t)-e^{x} \sin (x) \tag{44}
\end{align*}
$$

(a) DTM: the approximation solution of this equation is also obtained by DTM in [5] as follows:

$$
\begin{align*}
u_{5,5}(x, t)= & \left(x-\frac{x^{3}}{6}+\frac{x^{5}}{120}\right)+\left(1-\frac{x^{2}}{2}+\frac{x^{4}}{24}\right) t \\
& +\left(-\frac{x}{2}+\frac{x^{3}}{24}-\frac{x^{5}}{240}\right) t^{2}+\left(-\frac{1}{6}+\frac{x^{2}}{12}-\frac{x^{4}}{144}\right) t^{3} \\
& +\left(\frac{x}{24}-\frac{x^{3}}{144}+\frac{x^{5}}{2880}\right) t^{4} \\
& +\left(\frac{1}{120}-\frac{x^{2}}{240}+\frac{x^{4}}{2880}\right) t^{5} \tag{45}
\end{align*}
$$

(b) RDTM: it is easy to see that the $u(x, 0)=\sin (x)$, and therefore RDTM version is

$$
\begin{equation*}
U_{0}(x)=\sin (x) \tag{46}
\end{equation*}
$$

By applying the RDTM on nonlinear Volterra integral equation (44), for $k=1,2, \ldots$, we get

$$
\begin{align*}
U_{k}(x)- & \left\{\sum_{r=0}^{k} \frac{\sin (x+r \pi / 2)}{r!}\left(\frac{(-1)^{k-r} e^{x}}{(k-r)!}+\delta_{k-r, 0}\right)\right. \\
& \left.-\sum_{r=0}^{k} \frac{(-1)^{r} \sin ((k-r) \pi / 2)}{r!(k-r)!}-\delta_{k, 0} e^{x} \sin (x)\right\} \\
= & \frac{2}{k} \int_{0}^{x}\left\{\sum_{r=0}^{k-1} e^{y} \frac{(-1)^{r}}{r!} U_{k-r-1}(y)\right\} d y \tag{47}
\end{align*}
$$

where $U_{i}(x)$ is the reduced differential transform of $u(x, t)$. After expanding the RDTM recurrence equations (47), with initial value of (46), for $k=1,2,3,4,5$, the first five terms of $U_{k}(x)$ are obtained as follows:

$$
\begin{aligned}
& U_{1}(x)=\cos (x) \\
& U_{2}(x)=-\frac{1}{2} \sin (x)
\end{aligned}
$$

TABLE 4: Comparisons of the exact solution $u(x, t)=\sin (x+t)$, with $U_{5,5}(x, t)$ obtained by classic DTM [5] and $U_{5}(x, t)$ obtained by reduced DTM at some test points $(x, t)$ in Example 10.

| $x$ | $t$ | $u(x, t)$ | Classic DTM [5] |  | Reduced DTM |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  |  |  | $U_{5,5}(x, t)$ | $\left\|u(x, t)-U_{5,5}(x, t)\right\|$ | $U_{5}(x, t)$ | $\left\|u(x, t)-U_{5}(x, t)\right\|$ |
| 0.2 | 0.1 | 0.2955202067 | 0.2955202184 | $1.1688660428 e-08$ | 0.2955202070 | $2.9532337686 e-10$ |
|  | 0.4 | 0.5646424734 | 0.5646439552 | $1.4818049646 e-06$ | 0.5646439183 | $1.4448767682 e-06$ |
|  | 0.7 | 0.7833269096 | 0.7833750550 | $4.8145422517 e-05$ | 0.7833749959 | $4.8086256376 e-05$ |
|  | 1 | 0.9320390860 | 0.9325020000 | $4.6291403277 e-04$ | 0.9325019239 | $4.6283789648 e-04$ |
| 0.5 | 0.1 | 0.5646424734 | 0.5646461680 | $3.6945737146 e-06$ | 0.5646424741 | $6.8315986201 e-10$ |
|  | 0.4 | 0.7833269096 | 0.7833397500 | $1.2840372517 e-05$ | 0.7833299139 | $3.0042708277 e-06$ |
|  | 0.7 | 0.9320390860 | 0.9321460859 | $1.0699997027 e-04$ | 0.9321309857 | $9.1899716402 e-05$ |
|  | 1 | 0.9974949866 | 0.9983398437 | $8.4485714595 e-04$ | 0.9983208230 | $8.2583639762 e-04$ |
| 0.8 | 0.1 | 0.7833269096 | 0.7834038828 | $7.6973172517 e-05$ | 0.7833269106 | $1.0099717729 e-09$ |
|  | 0.4 | 0.9320390860 | 0.9322215424 | $1.8245643277 e-04$ | 0.9320433813 | $4.2953023213 e-06$ |
|  | 0.7 | 0.9974949866 | 0.9978859380 | $3.9095139595 e-04$ | 0.9976224907 | $1.2750404846 e-04$ |
|  | 1 | 0.9738476309 | 0.9752880000 | $1.4403691218 e-03$ | 0.9749626963 | $1.1150653929 e-03$ |
| 1 | 0.1 | 0.8912073601 | 0.8915382743 | $3.3091424412 e-04$ | 0.8912073612 | $1.1792198329 e-09$ |
|  | 0.4 | 0.9854497300 | 0.9861662222 | $7.1649223376 e-04$ | 0.9854546786 | $4.9486333631 e-06$ |
|  | 0.7 | 0.9916648105 | 0.9928385451 | $1.1737346864 e-03$ | 0.9918098802 | $1.4506974814 e-04$ |
|  | 1 | 0.9092974268 | 0.9118055556 | 2.5081287299 - 03 | 0.9105512242 | $1.2537973843 e-03$ |

$$
\begin{align*}
& U_{3}(x)=-\frac{1}{6} \cos (x), \\
& U_{4}(x)=\frac{1}{24} \sin (x), \\
& U_{5}(x)=-\frac{1}{120} \cos (x) . \tag{48}
\end{align*}
$$

In the same manner, the rest of the components were obtained by using the recursive equations (47). Substituting the quantities (48) in (18), the approximation solution of Volterra integral equation (44) in the Poisson series form is

$$
\begin{align*}
U_{5}(x, t)= & \sin (x)+\cos (x) t-\frac{\sin (x)}{2} t^{2} \\
& -\frac{\cos (x)}{6} t^{3}+\frac{\sin (x)}{24} t^{4}-\frac{\cos (x)}{120} t^{5}, \tag{49}
\end{align*}
$$

which is the same as the first five terms of the Poisson series of the exact solution $u(x, t)=\sin (x+t)$. The numerical results obtained with reduced DTM are presented in Table 4, in comparison with the classic DTM solution of [5] and the exact solution $u(x, t)=\sin (x+t)$, for some points of the intervals $0 \leq x \leq 1$ and $0 \leq t \leq 1$.

Example 11. In the third example, consider the following twodimensional Volterra integral equation [5]:

$$
\begin{align*}
u(x, t) & -e^{t-x} \int_{0}^{t} \int_{0}^{x} u(y, z) d y d z \\
& =\sinh (x+t)\left(e^{t-x}+1\right)-e^{t-x}(\sinh (x)-\sinh (t)) \tag{50}
\end{align*}
$$

(a) DTM: the approximation solution of this equation is also obtained by DTM in [5] as follows

$$
\begin{align*}
u_{5,5}(x, t)= & \left(-x-\frac{x^{3}}{6}-\frac{x^{5}}{120}\right)+\left(1+\frac{x^{2}}{2}+\frac{x^{4}}{24}\right) t \\
& +\left(-\frac{x}{2}-\frac{x^{3}}{12}-\frac{x^{5}}{240}\right) t^{2}+\left(\frac{1}{6}+\frac{x^{2}}{12}+\frac{x^{4}}{144}\right) t^{3} \\
& +\left(-\frac{x}{24}-\frac{x^{3}}{144}-\frac{x^{5}}{2880}\right) t^{4} \\
& +\left(\frac{1}{120}+\frac{x^{2}}{240}+\frac{x^{4}}{2880}\right) t^{5} . \tag{51}
\end{align*}
$$

(b) RDTM: it is easy to see that the $u(x, 0)=-\sinh (x)$, and therefore RDTM version is

$$
\begin{equation*}
U_{0}(x)=-\sinh (x) . \tag{52}
\end{equation*}
$$

Table 5: Comparisons of the exact solution $u(x, t)=\sinh (t-x)$, with $U_{5,5}(x, t)$ obtained by classic DTM [5] and $U_{5}(x, t)$ obtained by reduced DTM at some test points ( $x, t$ ) in Example 11.

| $x$ | $t$ | $u(x, t)$ | Classic DTM [5] |  | Reduced DTM |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  |  |  | $U_{5,5}(x, t)$ | $\left\|u(x, t)-U_{5,5}(x, t)\right\|$ | $U_{5}(x, t)$ | $\left\|u(x, t)-U_{5}(x, t)\right\|$ |
| 0.2 | 0.1 | -0.1001667500 | -0.1001667561 | $6.0968226301 e-09$ | -0.1001667498 | $2.5944103810 e-10$ |
|  | 0.4 | +0.2013360025 | +0.2013367851 | $7.8252557270 e-07$ | +0.2013368189 | $8.1631586063 e-07$ |
|  | 0.7 | +0.5210953055 | +0.5211116472 | $1.6341756253 e-05$ | +0.5211117115 | $1.6406043554 e-05$ |
|  | 1 | +0.8881059822 | $+0.8881853333$ | $7.9351145710 e-05$ | $+0.8881854339$ | $7.9451747271 e-05$ |
| 0.5 | 0.1 | -0.4107523258 | -0.4107529453 | $6.1950968455 e-07$ | -0.4107523251 | $7.0149602793 e-10$ |
|  | 0.4 | -0.1001667500 | -0.1001714167 | $4.6666468226 e-06$ | -0.1001641445 | $2.6055545215 e-06$ |
|  | 0.7 | +0.2013360025 | +0.2013887643 | $5.2761781823 e-05$ | +0.2014033478 | $1.7345289013 e-05$ |
|  | 1 | +0.5210953055 | +0.5215820313 | $4.8672575625 e-04$ | +0.5216052465 | $5.0994098756 e-04$ |
| 0.8 | 0.1 | -0.7585837018 | -0.7585783977 | $5.3041062001 e-06$ | -0.7585837006 | $1.2071608158 e-09$ |
|  | 0.4 | -0.4107523258 | -0.4108535808 | $1.0125499718 e-04$ | -0.4107476947 | $4.6310571234 e-06$ |
|  | 0.7 | -0.1001667500 | -0.1002690359 | $1.4228584682 e-04$ | -0.1000423588 | $1.2439120515 e-04$ |
|  | 1 | +0.2013360025 | +0.2019546667 | $6.1866412557 e-04$ | +0.2023226727 | $9.8667016118 e-04$ |
| 1 | 0.1 | -1.0265167257 | -1.0264561563 | $6.0569458175 e-05$ | -1.0265167241 | $1.6018943949 e-09$ |
|  | 0.4 | -0.6366535821 | -0.6370106667 | $3.5708451843 e-04$ | -0.6366473802 | $6.2019846800 e-06$ |
|  | 0.7 | -0.3045202934 | -0.3051720521 | $6.5175863619 e-04$ | -0.3043519610 | $1.6833243259 e-04$ |
|  | 1 | +0.0000000000 | -0.0000000000 | -0.0000000000 | $+0.0013512390$ | $1.3512390404 e-03$ |

By applying the RDTM on nonlinear Volterra integral equation (50), for $k=1,2, \ldots$, we get

$$
\begin{align*}
\sum_{r=0}^{k}\{ & \left\{U_{r}(x)+\frac{\delta_{r, 0}\left(1+e^{-x}\right)}{2}\right. \\
& \left.+\frac{1}{2 r!}\left(1+(-1)^{r} e^{x}-\left(1+2^{r}\right)\left(e^{-x}+e^{-2 x}\right)\right)\right\}\left\{\frac{(-1)^{k-r}}{(k-r)!} e^{x}\right\} \\
= & \frac{1}{k} \int_{0}^{x} U_{k-r}(y) d y \tag{53}
\end{align*}
$$

where $U_{i}(x)$ is the reduced differential transform of $u(x, t)$. After expanding the RDTM recurrence equations (53), with initial value of (52), for $k=1,2,3,4,5$, the first five terms of $U_{k}(x)$ are obtain as follows:

$$
\begin{aligned}
& U_{1}(x)=\cosh (x), \\
& U_{2}(x)=-\frac{1}{2} \sinh (x), \\
& U_{3}(x)=\frac{1}{6} \cosh (x), \\
& U_{4}(x)=\frac{1}{24} \sinh (x), \\
& U_{5}(x)=\frac{1}{120} \cosh (x) .
\end{aligned}
$$

In the same manner, the rest of the components were obtained by using the recursive equations (47). Substituting the quantities (48) in (18), the approximation solution of Volterra integral equation (44) in the Poisson series form is

$$
\begin{align*}
U_{5}(x, t)= & -\sinh (x)+\cosh (x) t-\frac{\sinh (x)}{2} t^{2} \\
& +\frac{\cosh (x)}{6} t^{3}-\frac{\sinh (x)}{24} t^{4}+\frac{\cosh (x)}{120} t^{5} \tag{55}
\end{align*}
$$

which is same as the first five terms of the Poisson series of the exact solution $u(x, t)=\sinh (t-x)$. The numerical results obtained with reduced DTM are presented in Table 5, in comparison with the classic DTM solution of [5] and the exact solution $u(x, t)=\sinh (t-x)$, for some points of the intervals $0 \leq x \leq 1$ and $0 \leq t \leq 1$.

## 4. Conclusions

In this study, we presented the definition and operation of both two-dimensional differential transformation method (DTM) and their reduced form, the so-called reduced-DTM (RDTM) for finding the solutions of a class of Volterra integral equations. For illustration purposes, we consider three different examples. It is worth pointing out that both DTM and RDTM have convergence for the solutions; actually, the accuracy of the series solution increases when the number of terms in the series solution is increased. From the computational process of DTM and RDTM, we find that the RDTM is easier to apply. In other words, it is obvious that DTM has very complicated computational process rather than RDTM.

The RDTM reduces the computational difficulties of the DTM and all the calculations can be made with simple manipulations MATLAB. Actually, as a special advantage of RDTM rather than DTM, the reduced differential transform recursive equations produce exactly all the Poisson series coefficients of solutions, whereas the differential transform recursive equations produce exactly all the Taylor series coefficients of solutions. The reliability of the RDTM and the reduction in the size of computational domain give this method a wider applicability. For small value of $x, t$, in Tables 3,4 , and 5 , we find that the RDTM has a smaller error than DTM. Also, for large values of $x, t$, we may increase the accuracy of the series solution by computing more terms, which is quite easy using MATLAB.

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# A Pressure-Stabilized Lagrange-Galerkin Method in a Parallel Domain Decomposition System 

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Received 30 April 2013; Accepted 14 June 2013
Academic Editor: Santanu Saha Ray
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#### Abstract

A pressure-stabilized Lagrange-Galerkin method is implemented in a parallel domain decomposition system in this work, and the new stabilization strategy is proved to be effective for large Reynolds number and Rayleigh number simulations. The symmetry of the stiffness matrix enables the interface problems of the linear system to be solved by the preconditioned conjugate method, and an incomplete balanced domain preconditioner is applied to the flow-thermal coupled problems. The methodology shows good parallel efficiency and high numerical scalability, and the new solver is validated by comparing with exact solutions and available benchmark results. It occupies less memory than classical product-type solvers; furthermore, it is capable of solving problems of over 30 million degrees of freedom within one day on a PC cluster of 80 cores.


## 1. Introduction

The Lagrange-Galerkin method raises wide concern about the finite-element simulation of fluid dynamics. Based on the approximation of the material derivative along the trajectory of fluid particle, the method is natural in the simulation to physical phenomena, and it is demonstrated to be unconditionally stable for a wide class of problems [1-5]. A number of researches about the Lagrange-Galerkin method were done in the case of single processor element (PE) (cf. [6-8]); the symmetry of the matrices and good stability of the scheme were reported; using a numerical integration based on a division of each element, Rui and Tabata [9] developed a second scheme for convection-diffusion problem; Massarotti et al. [10] used a second-order characteristic curve method, and a special iteration was used to keep the symmetry of the stiffness matrix. The Lagrange-Galerkin method uses an implicit time discretization, and therefore an element searching algorithm is necessary to implement it. The element searching may become very expensive when the geometry is complicated or the mesh size is very small. Due to its doubtable efficiency and feasibility for complex simulations in the case of single PE , rare research has been done to implement it in parallel, by which the enormous computation power enables us to solve more challenging simulation problems.

The present study is concentrated on improving the solvability of the Lagrange-Galerkin method on large scale and complex problems by domain decompositions. Piecewise linear interpolations are thus employed for velocity, pressure, and temperature; therefore, the so-called inf-sup condition [11] should be satisfied, which is the first difficulty to be overcome in this work. Stabilization methods for incompressible flow problems were reported by many researchers (cf. [12-15]). Park and Sung proposed a stabilization for Rayleigh-Bénard convection by using feedback control [16]; for consistently stabilized finite element methods, Barth et al. classified the stabilization techniques and studied influence of the stabilization parameter in convergence [17]; Bochev et al. stated the requirements on choice of stabilization parameter if time step and mesh are allowed to vary independently [18]. As far as we know, it may not be enough to investigate what stabilization techniques are efficient for nonsteady and nonlinear flow problems approximated by LagrangeGalerkin methods in a domain decomposition system, where the interface problem can be solved by preconditioned conjugate gradient (PCG) method. In this paper, a pressurestabilization method, which keeps the symmetry of the linear system and is effective for high Reynolds number and Rayleigh number simulations, is introduced to implement
the Lagrange-Galerkin method in a domain decomposition system.

The element searching algorithm in a domain decomposition system using unstructured grids is the second difficulty to implement the Lagrange-Galerkin method in a domain decomposition system (cf. [5, 19]). Minev et al. reported an optimized binary searching algorithm for single PE by storing the necessary data structures in a way similar to the CSR compact storage format; however, the element information data is stored distributedly in the domain decomposition system by the skyline format, and a different way needs to be found to overcome the extra difficulty caused by the parallel computing algorithm. This step is critical, in the sense that it can be very computationally expensive and can thus make the entire algorithm impractical.

The remainder of this paper is organized as follows: in Section 2, the formula of the governing equation and the pressure-stabilization Lagrange-Galerkin method is described; Section 3 focuses on the parallel implementation of this scheme. Numerical results and comparisons with classical asymmetric product type methods in [20] are shown in Section 4. Conclusions are drawn in Section 5.

## 2. Formulation

2.1. The Governing Equations. Let $\Omega$ be a three-dimensional polyhedral domain with the boundary $\partial \Omega$. The conservation equations of mass and momentum are governed by

$$
\begin{gather*}
\frac{\partial u}{\partial t}+(u \cdot \nabla) u-2 \nu \nabla \cdot D(u)+\nabla p=f^{\text {buoyancy }} \\
\text { in } \Omega \times(0, \bar{t}), \\
\nabla \cdot u=0 \quad \text { in } \Omega \times(0, \bar{t}), \\
u=\widehat{u} \quad \text { on } \Gamma_{1} \times(0, \bar{t}),  \tag{1}\\
\sum_{j=0}^{3} \sigma_{i j} n_{j}=0 \quad \text { on } \frac{\partial \Omega}{\Gamma_{1} \times(0, \bar{t})} \\
u=u_{0} \quad \text { in } \Omega, \text { at } t=0
\end{gather*}
$$

where $\Gamma_{1} \subset \partial \Omega$ and

$$
\begin{equation*}
f^{\text {buoyancy }}=\beta\left(T_{r}-T\right) g \tag{2}
\end{equation*}
$$

is the gravity force per unit mass derived on the basis of Boussinesq approximation. $g$ is the gravity $\left[\mathrm{m} / \mathrm{s}^{2}\right], \beta, T$, and $T_{r}$ are the thermal expansion coefficient $[1 / K]$, the temperature $[K]$, and the reference temperature $[K]$, and $u, t, v$, and $p$ are velocity vector [ $\mathrm{m} / \mathrm{s}$ ], time [ s ], kinematic viscosity coefficient $\left[\mathrm{m}^{2} / \mathrm{s}\right]$, and kinematic pressure $\left[\mathrm{m}^{2} / \mathrm{s}^{2}\right]$, respectively. $\sigma_{i j}$ is the stress tensor $\left[\mathrm{N} / \mathrm{m}^{2}\right]$ defined by

$$
\begin{gather*}
\sigma_{i j}(u, p) \equiv-p \delta_{i j}+2 v D_{i j}(u) \\
D_{i j}(u) \equiv \frac{1}{2}\left(\frac{\partial u_{i}}{\partial x_{j}}+\frac{\partial u_{j}}{\partial x_{i}}\right), \quad i, j=1,2,3 \tag{3}
\end{gather*}
$$

with the Kronecker delta $\delta_{i j}$.

The fluid is assumed to be incompressible according to Boussinesq approximation, and the density is assumed to be constant except in the gravity force term where it depends on temperature according to the indicated linear law; see (2). The energy equation is

$$
\begin{gather*}
\frac{\partial T}{\partial t}+u \cdot \nabla T-a \Delta T=S \quad \text { in } \Omega \times(0, \bar{t}) \\
T=\widehat{T} \quad \text { on } \Gamma_{2} \times(0, \bar{t}) \\
a \frac{\partial T}{\partial n}=0 \quad \text { on } \frac{\partial \Omega}{\Gamma_{2} \times(0, \bar{t})}  \tag{4}\\
T=T_{0} \quad \text { in } \Omega, \text { at } t=0
\end{gather*}
$$

where $\Gamma_{2} \subset \partial \Omega, a$ is the thermal diffusion coefficient $\left[\mathrm{m}^{2} / \mathrm{s}\right]$, and $S$ is the source term with the unit of $[K / s]$.
2.2. The Lagrange-Galerkin Finite-Element Method. Some preliminaries are arranged for the derivation of a finite element scheme of (1) and (4). Let the subscript $h$ denote the representative length of the triangulation, and let $\mathfrak{\Im}_{h} \equiv$ $\{K\}$ denote a triangulation of $\Omega$ consisting of tetrahedral elements. Given that $g$ is a vector valued function on $\Gamma_{1}$, the finite element spaces are as follows:

$$
\begin{align*}
& X_{h} \equiv\left\{v_{h} \in C^{0}(\bar{\Omega})^{3} ;\left.v_{h}\right|_{K} \in P_{1}(K)^{3}, \forall K \in \Im_{h}\right\} \\
& M_{h} \equiv\left\{q_{h} \in C^{0}(\bar{\Omega}) ;\left.q_{h}\right|_{K} \in P_{1}(K), \forall K \in \Im_{h}\right\} \\
& V_{h}(g) \equiv\left\{v_{h} \in X_{h} ; v_{h}(P)=g(P), \forall P \in \Gamma_{1}\right\}  \tag{5}\\
& \Theta_{h}(b) \equiv\left\{\theta_{h} \in M_{h} ; \theta_{h}(P)=b(P), \forall P \in \Gamma_{2}\right\} \\
& V_{h} \equiv V_{h}(0), \Theta_{h} \equiv \Theta_{h}(0), Q_{h}=M_{h}
\end{align*}
$$

Let $(\cdot, \cdot)$ defines the $L_{2}$ inner product; the continuous bilinear forms $a$ and $b$ are introduced by

$$
\begin{align*}
& a(u, v) \equiv 2 v(D(u), D(v)), \\
& b(u, v) \equiv-(\nabla \cdot u, q) \tag{6}
\end{align*}
$$

respectively.
Let $\Delta t$ be the time increment, and let $N_{\bar{t}} \equiv[\bar{t} / \Delta t]$ be the total step number. Let the superscript $n$ denote the time step; a finite element approximation of (1) is described as follows: find $\left\{\left(u_{h}^{n}, p_{h}^{n}\right)\right\}_{n=1}^{N_{\bar{t}}} \in V_{h}(g) \times Q_{h}$, such that for $\left(v_{h}, q_{h}\right) \in V_{h} \times Q_{h}$,

$$
\begin{gather*}
\left(\frac{u_{h}^{n}-u_{h}^{n-1} \circ X_{1}\left(u_{h}^{n-1}, \Delta t\right)}{\Delta t}, v_{h}\right)+a\left(u_{h}^{n}, v_{h}\right) \\
+b\left(v_{h}, p_{h}^{n}\right)=\left(f^{n}, v_{h}\right),  \tag{7}\\
b\left(u_{h}^{n}, q_{h}\right)=0,
\end{gather*}
$$

where $X_{1}(\cdot, \cdot)$ denotes a first-order approximation of a particle's position [5], and the notation $\circ$ denotes the composition of functions.

For the purpose of large scale computation, a piecewise equal-order interpolation for velocity and pressure is used, as can be seen from (5). Pressure stabilization is thus needed to keep the necessary link between $V_{h}$ and $Q_{h}$. A penalty Galerkin least-squares (GLS) stabilization method for pressure is proved in [12] to hold the same asymptotic error estimates as the method of Hughes et al. [21] and it is computationally cheap. For P1/P1 elements, the stabilization is reduced to

$$
\begin{equation*}
\sum_{K \in \Im_{h}} \delta_{K} h_{K}^{2}\left(\nabla p_{h}^{n},-\nabla q_{h}\right)_{K} \tag{8}
\end{equation*}
$$

which does no modification to the momentum equation because of vanishing of the second-order derivate term. Here, $h_{K}$ denotes the maximum diameter of an element $K$. Unlike $[6,12]$, where a constant $\delta(>0)$ is used as the stabilization parameter, an element-wise stabilization parameter

$$
\delta_{K}=\left\{\begin{array}{l}
\alpha,  \tag{9}\\
\text { for } \log _{10}\left[\operatorname{Max}\left\{\left\|\nabla p_{h}^{n-1}\right\|_{2}\right\}_{i=1}^{4}\right] \leq 1, \\
\alpha \times \log _{10}\left[\operatorname{Max}\left\{\left\|\nabla p_{h}^{n-1}\right\|_{2}\right\}_{i=1}^{4}\right] \\
\text { otherwise }
\end{array}\right.
$$

is used in this work, where $\nabla p_{h}^{n-1}$ is gradient of the FEM approximated pressure at $t^{n-1}$ and $i$ is the number of the nodal point in a tetrahedral element. Since $\alpha$ is very important to balance the accuracy and convergence of the scheme, it is discussed in Section 4.1. The localized stabilization parameter is designed to be adaptive to the pressure gradient, and thus it has a better control on the pressure field.

By adding (8) to (7), a pressure-stabilized FEM scheme for Navier-Stokes problems is achieved. The nonsteady iteration loops for solving (1) and (4) and then reads the following.

Step 1. Compute the particle's coordinates by

$$
\begin{equation*}
X_{1}\left(u_{h}^{n-1}, \Delta t\right) \equiv x-u_{h}^{n-1} \Delta t \tag{10}
\end{equation*}
$$

and search the element holding the particle at $t^{n-1}$.
Step 2. Find $T_{h}^{n}$ by

$$
\begin{equation*}
\left(\frac{T_{h}^{n}-T_{h}^{n-1} \circ X_{1}\left(u_{h}^{n-1}, \Delta t\right)}{\Delta t}, \theta_{h}\right)+\left(a \nabla T_{h}^{n}, \nabla \theta_{h}\right)=\left(S^{n}, \theta_{h}\right) \tag{11}
\end{equation*}
$$

Step 3. Find $\left(u_{h}^{n}, p_{h}^{n}\right)$ by

$$
\begin{aligned}
&\left(\frac{u_{h}^{n}-}{} u_{h}^{n-1} \circ X_{1}\left(u_{h}^{n-1}, \Delta t\right)\right. \\
& \Delta t\left.v_{h}\right)+a_{0}\left(u_{h}^{n}, v_{h}\right) \\
&+b\left(v_{h}, p_{h}^{n}\right)+b\left(u_{h}^{n}, q_{h}\right) \\
&+\sum_{K \in \mathfrak{\Im}_{h}} \delta_{K} h_{K}^{2}\left(\nabla p_{h}^{n},-\nabla q_{h}\right)_{K} \\
&=\left(f^{n}, v_{h}\right)+\left(\beta\left(T_{r}-T_{h}^{n}\right) g, v_{h}\right)
\end{aligned}
$$

Step 4. Compute the relative error by a $H^{1} \times L^{2} \times H^{1}$ norm defined by

$$
\begin{align*}
\|(u, p, T)\|_{H^{1} \times L^{2} \times H^{1}} \equiv & \frac{1}{\sqrt{\operatorname{Re}}}\|u\|_{H^{1}(\Omega)^{3}}  \tag{13}\\
& +\|p\|_{L^{2}}+\|T\|_{H^{1}(\Omega)}
\end{align*}
$$

where Re denotes the Reynolds number, and set

$$
\begin{align*}
\operatorname{diff} & =\frac{\left\|\left(u^{n}, p^{n}, T^{n}\right)-\left(u^{n-1}, p^{n-1}, T^{n-1}\right)\right\|_{H^{1} \times L^{2} \times H^{1}}}{\left\|\left(u^{n-1}, p^{n-1}, T^{n-1}\right)\right\|_{H^{1} \times L^{2} \times H^{1}}}  \tag{14}\\
& \leq \operatorname{Err}_{\mathrm{NS}}
\end{align*}
$$

as the steady-state criterion; if (14) is satisfied or the number of loops reaches the maximum, then stop the iteration; otherwise, repeat Steps 1-3.

As can be seen from Steps 2 and 3, both (1) and (4) are approximated by the Lagrange-Galerkin method, and the searching algorithm only needs to be performed once in a nonsteady loop. It can also be seen that the solver is also flexible, and it can solve pure Navier-Stokes problems by setting the body force in (2) to external force and omitting Step 2.

## 3. Implementation

3.1. A Parallel Domain Decomposition System. To begin with the parallel domain decomposition method, the domain decomposition is introduced briefly as follows. The whole domain is decomposed into a number of subdomains without overlapping, and the solution of each subdomain is superimposed on the equation of the inner boundary of the subdomains. By static condensation, the linear system

$$
\begin{equation*}
K \bar{u}=\bar{f} \tag{15}
\end{equation*}
$$

is written as

$$
\begin{align*}
& {\left[\begin{array}{ccccc}
K_{I I}^{(1)} & 0 & \cdots & 0 & K_{I B}^{(1)} R_{B}^{(1)} \\
0 & \ddots & & \vdots & \vdots \\
\vdots & & \ddots & \vdots & \vdots \\
0 & \cdots & \cdots & K_{I I}^{(N)} & K_{I B}^{(N)} R_{B}^{(N)} \\
R_{B}^{(1) T} K_{I B}^{(1) T} & \cdots & \cdots & R_{B}^{(N) T} K_{I B}^{(N) T} & K_{B B}
\end{array}\right]\left[\begin{array}{c}
\bar{u}_{I}^{(1)} \\
\vdots \\
\vdots \\
\bar{u}_{I}^{(N)} \\
\bar{u}_{B}
\end{array}\right]} \\
& \quad=\left[\begin{array}{c}
\bar{f}_{I}^{(1)} \\
\vdots \\
\vdots \\
\bar{f}_{I}^{(N)} \\
\bar{f}_{B}
\end{array}\right], \tag{16}
\end{align*}
$$

where $K$ is the stiffness matrix, $\bar{u}$ denotes the unknowns (u and $p$ ), and $\bar{f}$ is the force vector. $R$ is the restriction operator consists of $0-1$ matrix. The superscripts $(N)$ means the
$N$ th subdomain, and subscript $I$ and $B$ relate to the element of the inner boundary, and interface boundary respectively.

From (16), it can be observed that the interface problems

$$
\begin{align*}
\sum_{i=1}^{N} R_{B}^{(i)^{T}} & \left(K_{B B}^{(i)}-K_{I B}^{(i)^{T}} K_{I I}^{(i)^{-1}} K_{I B}^{(i)}\right) R_{B}^{(i)} \bar{u}_{B}  \tag{17}\\
& =\sum_{i=1}^{N} R_{B}^{(i)^{T}}\left(\bar{f}_{B}^{(i)}-K_{I B}^{(i)^{T}} K_{I I}^{(i)^{-1}} \bar{f}_{I}^{(i)}\right)
\end{align*}
$$

and the inner problems

$$
\begin{equation*}
\bar{u}_{I}^{(i)}=K_{I I}^{(i)^{-1}}\left(\bar{f}_{I}^{(i)}-K_{I B}^{(i)^{T}} R_{B}^{(i)} \bar{u}_{B}\right), \quad i=1, \ldots, N \tag{18}
\end{equation*}
$$

can be solved separately [22]. In this work, the interface problems are solved first iteratively, and the inner problems are then solved by substituting $\bar{u}_{B}$ in to (18).

The Lagrange-Galerkin method keeps the symmetry of the stiffness matrix, and the GLS pressure-stabilization term in (8) also produces a symmetric matrix; therefore, $K$ is symmetric in (15), and a PCG method is employed to get the $u_{I}$ in (18), and to avoid drawback of the classical domain decomposition method, such as Neumann-Neumann and diagonalscaling, a balanced domain decomposition preconditioner is used to prevent the growing of condition number with the number of subdomains. An identity matrix is chosen as the coarse matrix, and the coarse problem is solved incompletely by omitting the fill-ins in some sensitive places during the Cholesky factorization. By using this inexact balanced domain decomposition preconditioning, the coarse matrix is sparser and thus easier to be solved; therefore, the new solver is expected to have better solvability on large scale computation models.
3.2. The Lagrange-Galerkin Method in Parallel. The element searching algorithm requires a global-wise element information to determine the position of one particle in the previous time step. However, in the parallel domain decomposition system, the whole domain is split into several parts one processor element (PE) works only on the current part, and it does not contain any element information of other parts. Each part is further divided into many subdomains, and the domain decomposition is performed by the PE in charge of the part. This parallelity causes a computational difficulty: for each time step, the particle is not limited within one part; therefore, exchanging the data between different PEs is necessary, which demands the PEs to communicate in the subdomain-wise computation.

In order to know the position of a particle at $t^{n-1}$, a neighbour elements list is created at the beginning of the analysis. Based on the information of neighbour elements and the coordinates calculated by (10), it is possible to find the element holding this particle at $t^{n-1}$. A 2 -dimensional searching algorithm is present as follows ( $\lambda_{i}$ is the barycentric coordinates, and $n e\left(\lambda_{i}\right)$ is the neighbour element; see Figure 1):


Figure 1: A searching algorithm.
(1) initialize: $e_{0}=e_{\text {current }}$;
(2) iterate $i=0,1, \ldots$, Maxloops;

If $\lambda_{1}, \lambda_{2}, \lambda_{3}>0$, return $e_{i}$;
else if $n e\left(\operatorname{Min}\left\{\lambda_{1}, \lambda_{2}, \lambda_{3}\right\}\right) \neq$ boundary

$$
e_{i+1}=n e\left(\operatorname{Min}\left\{\lambda_{1}, \lambda_{2}, \lambda_{3}\right\}\right) ;
$$

else break;
(3) return $e_{i}$.

The request of the old solutions, which is the $u_{h}^{n-1}$ in (10), is relatively trivial when using single PEs or simply solving the problem parallel using symmetric multiprocessing; however, in the domain decomposition system, the particle is not limited within one part; it may pass the interface of different parts, as can be seen from Figure 1. Because one PE only has the elements information that belongs to the current part, communications between PEs are necessary. However, the number of total elements in one subdomain may be different, which means that some point to point communication techniques, such as MPI_Send/MPI_Recv or MPI_Sendrecv in MPICH, cannot be used in element wise computation. In the previous research [23], a global variable to store all the old solutions is constructed. This method maintains a high computation speed but costs a huge memory usage. To reduce the memory consumption, a request-response system is used in this work. In the computation, the searching algorithm is performed first, and the element that contains the current particle in the previous time step is thus known; therefore, the PE to get $u_{h}^{n-1}$ from is also known. However, as the sender does not know which PE requires message from itself, the receiver has to send its request to the sender first; after the request is detected, the sender sends the message to the receiver. The procedure is as follows:
(1) by scanning all the particles in the current subdomain, an array including all the data that is needed by the current PE is sent to all the other PEs.
(2) All PEs check if there is any request to itself. If it exists, PEs will prepare an array of the needed data and send it.
(3) The current PE receives the data sent by other PEs.

Data transferred by MPI communication should be packaged properly to avoid the overflow of MPI buffer in case


Figure 2: Convergence of (a) different constant $\delta$ at $\mathrm{Re}=10^{3}$; (b) different Re for $\delta=0.005$ and localized $\delta_{K}$.


Figure 3: Numerical scalability.
of large-scale computation. Nonblocking communication is employed, and as the 3 steps are performed subsequently, thus the computation time and communication time will be overlapped.

## 4. Numerical Results and Discussion

The parallel efficiency of new solver is firstly evaluated in this section, and to validate the scheme, exact solutions and available benchmark results classical computational models are compared. The CG convergence is judged by Euclidian norm with a tolerance of $10^{-6}$, and for nonsteady iteration, $\operatorname{Err}_{\mathrm{NS}}=10^{-4}$ is set as the criterion, using the $H^{1} \times L^{2} \times H^{1}$ norm defined in (13). For pure Navier-Stokes problems, a similar $H^{1} \times L^{2}$ norm, which is related to velocity and pressure, is employed to judge the steady state.
4.1. Efficiency Test. The BDD serious preconditioners were employed in this work; they are very efficient, and their iteration numbers are about lâĄĎ10 of the normal domain decomposition preconditioners (cf. [23]). The inexact preconditioner mentioned in Section 3 also shows good convergence and is more suitable for large scale computations [24].

It was set as the default preconditioner for all the following computations of this research.

The penalty methods are not consistent since the substitution of an exact solution into the discrete equations (12) leaves a residual that is proportional to the penalty parameter (cf. [17]); therefore, $\delta_{K}$ should be determined carefully. Numerical experiments of a lid-driven cavity flow were tested, and the mesh size was $62 \times 62 \times 62$. The total degrees of freedom (DOF) are $1,000,188$, and the results are given by Figure 2. For the purpose of higher accuracy, $\delta_{K}$ is expected to be small; however, the convergence turns worse when $\delta_{K}$ goes small, as can be seen from Figure 2(a). In Figure 2(b), a constant $\delta=0.005$ is used for different Reynolds numbers, and no convergence is achieved within 10000 PCG loops for $\mathrm{Re}=10^{6}$; and the comparison shows that the $\delta_{K}$ performs much better than a constant $\delta=0.005$ when $\alpha$ is set to 0.005 .

The parallel efficiency is assessed firstly by freezing the mesh size of test problem and refining the domain decomposition by decreasing the subdomain size and therefore increasing the number of subdomains; the comparison of the numerical scalability of the current scheme with and without the preconditioner is assessed by Figure 3. It can be seen that with the preconditioning technique, the iterative procedure


$$
\begin{aligned}
& \rightarrow \text { BiCGSTAB (ADV_sFlow 0.5) } \\
& \square \text { BiCGSTAB2 (ADV_sFlow 0.5) } \\
& \rightarrow \text { GPBiCG (ADV_sFlow 0.5) } \\
& \rightarrow \text { Current }
\end{aligned}
$$


(b)

Figure 4: Time and memory usages.



Figure 5: Numerical scalability.
of current scheme converges rapidly, and the convergence is independent of the number of subdomains.

Based on the paralyzed Lagrange-Galerkin method, the new solver makes a symmetric stiffness matrix, therefore only the lower/upper triangular matrix needs to be saved. Moreover, nonblocking MPI communication is used instead of constructing global arrays to keep the old solutions, and the current solver is expected to reduce the memory consumption without sacrificing the computation speed. The usage of time and memory of solving the thermal driven cavity problem by different solvers is compared, and the results are given by Figure 4.

The test problem was solved by the new solver and the ADV_sFlow 0.5 [25], which contains some nonsymmetric product-type solvers like GPBiCG, BiCGSTAB, and BiCGSTAB2 [20]. The ADV_sFlow 0.5 employed a domain decomposition system similar to the work; however, no precondition technique is used because of the non-symmetry of the stiffness matrix in (15). The comparisons of elapsed time and memory occupation of the new solver and that of product-type solvers in ADV_sFlow 0.5 are show in Figures 4(a) and 4(b). As can be seen, the current scheme reduces the demand of computation time and memory consumption remarkably, and it is more suitable for large scale problems than product-type solvers.

The parallel scalability of the searching algorithm is also a concern for us, as it characterizes the ability of an algorithm to deliver larger speed-up using a larger number of PEs. To know this, the number of subdomains in one part is fixed, and computations on the test problem of various mesh sizes are performed by the new scheme. The speedup is shown in Figure 5. Three models were tested by the searching algorithm. With an increase in the mesh size of the computation model, the parallel scalability of the searching algorithm tends to be better. An explanation to this is that when the DOF increase, the number of elements in one subdomain is also increasing; therefore, the searching algorithm is accelerated more efficiently. However, too many elements in one subdomain will occupy more memory, and a trade-off strategy is necessary for parameterization.
4.2. Validation Tests. In this section, a variety of test problems have been presented in order to prove the capability of the parallel Lagrange-Galerkin algorithm. Benchmarks test of


Figure 6: A plane Couette flow model.


Figure 7: Numerical results versus exact solutions.

Navier-Stokes problems are in Sections 4.2.1, 4.2.2, and 4.2.3, and flow-thermal coupled problems are in Sections 4.2.4 and 4.2.5.
4.2.1. A Plane Couette Flow. The solver for Navier-Stokes equations in (1) was first tested with a 3D plane Couette flow. Under ideal conditions, the model is of infinite length; therefore, 4 times of the height is used as the length of the model see Figure 6. A constant velocity ( $\widehat{\mathcal{u}}, 0,0$ ) is applied on the upper horizontal face, and no-slip conditions are set on the lower horizontal face. A pressure gradient is imposed along $x_{1}$ for all the faces as essential boundary conditions.

An unstructured 3D mesh was generated by ADVENTURE_TetMesh [25], and the local density around the plane of $x_{1}=2$, where the data was picked from, was enriched. The total DOF is around $1,024,000$. The so-called Brinkman number [26] is believed to be the dominating parameter of the flow, and a serious of numerical experiments is done at various Brinkman number. To simulate the infinity length better, the exact solution is enforced on both the left face $\left(x_{1}=0\right)$ and the right face $\left(x_{1}=4\right)$ as Dirichlet boundary conditions. The comparisons between the computation results and exact solutions are given by Figure 7. Dotted


Figure 8: A lid-driven cavity model.
lines are used to present the results, and they are named as "Num_Res_l(B)," where $B$ stands for the Brinkman number. Crossed lines in Figure 7 present the computation results with no exact solutions setting on the left and right faces, and they are named as "Num_Res_0(B)."

It can be seen form Figure 7 that both of these two sets of computation results show good agreement with the exact solution, and dotted lines are closer to the exact solution, representing a better simulation to the ideal condition (cf. [27, 28]).
4.2.2. A Lid-Driven Cavity. The Navier-Stokes problems solver was then verified by a lid-driven cavity flow. The ideal gas flows over the upper face of the cube, and no-slip conditions are applied to all other faces, as in Figure 8.

All the faces of the cube were set with Dirichlet boundary conditions, and a zero reference pressure was at the centre of the cube to keep the simulation stable. The pressure profiles of the scheme using localized stabilization parameter in (9) and the scheme using constant $(\delta=1)$ parameter are compared, and the results are show in Figure 9.

Figure 9(a) shows the pressure counters of the scheme with the localized stabilization parameter in (9) and the Figure 9(b), shows scheme with a constant parameter. The model was run at $\operatorname{Re}=10^{4}$, and oscillations are viewed in Figure 9(b); however, the isolines in Figure 9(a) is quite smooth, showing that the pressure-stabilization term has a better control on the pressure field at high Reynolds number.

The model was run at different Reynolds numbers with a $128 \times 128 \times 128$ mesh to test the solvability of the new scheme. As shown in Figure 10, when the Reynolds number increases, the eddy at right bottom of plane $x_{1} x_{3}$ vanishes, while the eddy at the left bottom appears due to the increasing in the speed, and the flow goes more likely around the wall. The primary eddy goes lower and lower when Reynolds number becomes higher, and the particle is no longer limited to a single side of the cavity; it can pass from one side to the other, and back again violating the mirror symmetry, as is seen from other planes of Figure 10. Similar 3D results for high Reynolds number were reported by [29], and the solvability of the new solver for high Reynolds number was confirmed.


Figure 9: Pressure counters $\left(\operatorname{Re}=10^{4}\right)$.


Figure 10: Velocity and pressure profiles for different Reynolds number: $\operatorname{Re}=1,000$ (top), $\operatorname{Re}=3,200$ (middle), and $\operatorname{Re}=12,000$ (bottom) along different middle planes: plane $x_{1} x_{2}$ (left), plane $x_{1} x_{3}$ (middle), and plane $x_{2} x_{3}$ (right).
4.2.3. Backward Facing Step. The solver for Navier-Stokes equations was then tested with backward facing step, the fluid considered was air. The problem definition is shown in Figure 11, and the height of the step $h$ is the characteristic length. An unstructured 3D mesh was generated with 419,415
nodal points and 2,417,575 tetrahedral elements, and the local density of mesh was increased around the step.

A laminar flow is considered to enter the domain at inlet section, the inlet velocity profile is parabolic, and the Reynolds number is based on the average velocity at the inlet.


Figure 11: Backward facing steps.

(b)

Figure 12: Pressure counters (a) and velocity vectors (b) $(\mathrm{Re}=200)$.


Figure 13: Primary reattachment lengths.

The total length of the domain is 30 times the step height, so that the zero pressure is set at the outlet. A full 3D simulation of the step geometry for $100 \leq \operatorname{Re} \leq 800$ is present in this paper, and the primary reattachment lengths are predicted.

To determine the reattachment length, the position of the zero-mean-velocity line was measured. The points of detachment and reattachment were taken as the extrapolated zero-velocity line down the wall. The pressure contour in Figure 12(a) confirms the success of the pressure-stabilization method; velocity vectors and the primary attachment are demonstrated in Figure 12(b); similar results have been documented by many, like in $[10,30]$.


Figure 14: The model of infinite plates.

The comparison of primary reattachment length between current results and other available benchmark results are show in Figure 13. It is seen that the agreement is excellent at different Rayleigh numbers (cf. [31, 32]).
4.2.4. Natural Convection of Flat Plates. In order to test the coupled solver of Navier-Stokes equations and the convection-diffusion equation, the third application model was the natural convection between two infinite flat plates. The geometry is given in 3-dimensional by Figure 14. Noslip boundary conditions applied on the left and right vertical walls. The temperature on the left wall is assumed to be lower and set at $5[K]$; the right wall is set at $6[K]$. An unstructured 3D mesh about 1 million tetrahedral elements was generated, and the local grid density around the mid-plane was enriched.

The model was run at the size of $20 \times 20 \times 80$ to get the numerical solutions, and it was compared with the exact solutions in Figure 15. To simulate the infinity length of the plate better, the exact solution is enforced on both the upper face $\left(x_{3}=4\right)$ and the lower face $\left(x_{3}=0\right)$ as what is done in Section 4.2.1, and the results are present by a dotted line ("Num_Res_l") in Figure 12. And the model without exact solution set as boundary is named as "Num_Res_0" in Figure 15.


Figure 15: Numerical results versus exact solutions.


Figure 16: A thermal-driven cavity.

With the parameter setting of $\nu=0.5, T_{r}=5.5, \beta=1.0$, and $a=1$, the numerical experiment was performed. Results of the profile on $u_{3}(\cdot, 0,2)$, which are believed by many to be very sensitive, are shown in Figure 15. Both "Num_Res_0" and "Num_Res_l" are in great agreement with the exact solutions, and "Num_Res_l" is closer to the exact solution, producing a better simulation to the ideal condition. Similar results have been documented in [33].
4.2.5. Thermal-Driven Cavity. The new solver is also applied to a 3-dimensional nonlinear thermal driven cavity flow problem, which is cavity full of ideal gas; see Figure 16.

No-slip boundary conditions are assumed to prevail on all the walls of the cavity. Both the horizontal walls are assumed to be thermally insulated, and the left and right sides are kept at different temperatures. The cube is divided into $120 \times 120 \times$ 120 small cubes, and each small cube contains six tetrahedral elements. The time step is set to 0.01 s , with $\mathrm{Pr}=0.71$ and $\mathrm{Ra}=10^{4}$; the steady state is achieved after 0.39 s , as in Figure 17.

Figures 17(a) and 17(b) show the contour of vorticity and the velocity vectors at the steady stage, respectively, from the front view. The temperature contour is shown in Figure 17(c), and pressure profiles are show in Figure 17(d). The previous results convince us of the success in solving flow-thermal
coupled problems described by (1) and (4). Similar threedimensional results can also be found in [33,34]. The pressure profile in Figure 17(d) is smooth and symmetric, implying that the stabilization item in (8) works well.

In order to further validate the new solver, a comparison of temperature and velocity profiles of the current solver and other benchmark results was made. The centreline velocity results $w(\cdot, 0.5,0.5)$ and the temperature results $T(\cdot, 0.5,0.5)$, which are believed to be very sensitive in this simulation, are present in diagrams (a) and (b) of Figure 18, respectively. The velocity results share close resemblance to that of the ADV_sFlow 0.5, and they both show the more end-wall effects compared with the results of 2D case. The three temperature results show good agreement with each other, and the line representing the current results is the smoothest, as the mesh is the finest among the three. Similar results have been reported by other researchers (cf. [10, 33, 34]).

Thermal convection problems are believed to be dominated by two dimensionless numbers by many researchers, the Prandtl number and the Rayleigh number. To acquaint ourselves with the solvability of the new solver and to challenge applications of higher difficulty, a wide range of Rayleigh numbers from $10^{3}$ to $10^{7}$ is studied with $\operatorname{Pr}=0.71$, and the results for the steady-state solution are presented in Figure 19. Dimensionless length is used and the variation of Rayleigh number is determined by changing the characteristic length of the model.

The local Nusselt number $\left(\mathrm{Nu}=\partial T / \partial x_{1}\right)$ is a concern of many researchers, as they are sensitive to the mesh size. In Figure 19, the diagram (a) and the diagram (b) represent the local Nusselt number at the hot wall and the cold wall, respectively. Similar results can also be found in [10, 30, 35, 36]. The capability of the solver based on domain decomposed Lagrange-Galerkin scheme for high Rayleigh number is also confirmed by this figure.

The new solver enables the simulation of large scale problems, thus models of Rayleigh number up to $10^{7}$ can be run on small PC cluster. In this simulation, an unstructured mesh of $30,099,775$ DOF is generated, the time step, is 0.01 s and it takes about 24 hours to finish, using the a small Linux cluster of 64 PEs ( 64 cores@2.66 GHz).

## 5. Conclusions

A pressure-stabilized Lagrange-Galerkin method is implemented in a domain decomposition system in this research. By using localized stabilization parameter, the new scheme shows better control in the pressure field than constant stabilization parameter; therefore it has good solvability at high Reynolds number and high Rayleigh number. The reliability and accuracy of the present numerical results are validated by comparing with the exact solutions and recognized numerical results. Based on a domain decomposition method, the element searching algorithm shows good numerical scalability and parallel efficiency. The new solver reduces the memory consumption and is faster than classical product-type solvers. It is able to solve large scale problems of over 30 million degrees of freedom within one day by a small PC cluster.

(c)

Figure 17: Steady state of the thermal-driven cavity $\left(\mathrm{Ra}=10^{4}\right)$.


FIGURE 18: Centerline temperature velocity profiles of the symmetry plane $\left(\mathrm{Ra}=10^{4}\right)$.


Figure 19: Local Nusselt number along the hot wall (a) and the cold wall (b).

## Acknowledgments

This work was supported by the National Science Foundation of China (NSFC), Grants 11202248, 91230114, and 11072272; the China Postdoctoral Science Foundation, Grant 2012M521646, and the Guangdong National Science Foundation, Grant S2012040007687.

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# A Note on the Triple Laplace Transform and Its Applications to Some Kind of Third-Order Differential Equation 

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Received 25 March 2013; Accepted 20 May 2013
Academic Editor: R. K. Bera
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We introduced a relatively new operator called the triple Laplace transform. We presented some properties and theorems about the relatively new operator. We examine the triple Laplace transform of some function of three variables. We make use of the operator to solve some kind of third-order differential equation called "Mboctara equations."

## 1. Introduction

The topic of partial differential equations is one of the most important subjects in mathematics and other sciences. The behaviour of the solution very much depends essentially on the classification of PDEs therefore the problem of classification for partial differential equations is very natural and well known since the classification governs the sufficient number and the type of the conditions in order to determine whether the problem is well posed and has a unique solution. The Laplace transform has been intensively used to solve nonlinear and linear equations [1-7]. The Laplace transform is used frequently in engineering and physics; the output of a linear time invariant system can be calculated by convolving its unit impulse response with the input signal. Performing this calculation in Laplace space turns the convolution into a multiplication; the latter is easier to solve because of its algebraic form. The Laplace transform can also be used to solve differential equations and is used extensively in electrical engineering [1-7]. The Laplace transform reduces a linear differential equation to an algebraic equation, which can then be solved by the formal rules of algebra. The original differential equation can then be solved by applying the inverse Laplace transform. The English electrical engineer Oliver Heaviside first proposed a similar scheme, although without using the Laplace transform, and the resulting operational calculus is credited as the Heaviside calculus. Recently Kılıçman et al. [8-11] extended the Laplace transform to
the concept of double Laplace transform. This new operator has been intensively used to solve some kind of differential equation [11] and fractional differential equations. The aim of this work is to extend the Laplace transform to the triple Laplace transform. We will start with the definition of the triple Laplace transform.

## 2. Definitions and Theorems

Definition 1. Let $f$ be a continuous function of three variables; then, the triple Laplace transform of $f(x, y, t)$ is defined by

$$
\begin{align*}
& L_{x, y, t} {[f(x, y, t)] } \\
&=F(p, s, k) \iiint_{0}^{\infty} \exp [-p x] \exp [-s y]  \tag{1}\\
& \times \exp [-k t] f(x, y, t) d x d y d t
\end{align*}
$$

where, $x, y, t>0$ and $p, s, k$ are Laplace variables, and

$$
\begin{aligned}
& f(x, y, t) \\
& =\frac{1}{2 \pi i} \int_{\alpha-i \infty}^{\alpha+i \infty} e^{p x} \\
& \quad \times\left[\frac{1}{2 \pi i} \int_{\beta-i \infty}^{\beta+i \infty} e^{s y}\right.
\end{aligned}
$$

$$
\begin{align*}
& \times\left[\frac{1}{2 \pi i} \int_{\mu-i \infty}^{\mu+i \infty} e^{k t}\right. \\
& \quad \times F(p, s, k) d k] d s] d p \tag{2}
\end{align*}
$$

is the inverse triple Laplace transform.
Property 2. Assuming that the continuous function $f(x, y, t)$ is triple Laplace transformable, then,

$$
\begin{align*}
L_{t, y, x}[ & \left.\frac{\partial^{3} f(x, y, t)}{\partial x \partial y \partial t}\right] \\
= & p s k F(p, s, k)-p s F(p, s, 0)-p s F(p, 0, k) \\
& +p F(p, 0,0)-s k F(0, s, k)+s F(0, s, 0) \\
& +k F(0,0, k)-F(0,0,0), \\
L_{x, x, t}[ & \left.\frac{\partial^{3} f(x, y, t)}{\partial t \partial x^{2}}\right] \\
= & k p^{2} F(p, y, k)-p k F(0, y, k)-\frac{\partial F(0, y, k)}{\partial x}  \tag{3}\\
L_{x x x}[ & \left.\frac{\partial^{3} f(x, y, t)}{\partial x^{3}}\right] \\
= & p^{3} F(p, y, t)-p^{2} F(0, y, t) \\
& -p \frac{\partial F(0, y, t)}{\partial x}-\frac{\partial^{2} F(0, y, t)}{\partial x^{2}} .
\end{align*}
$$

## 3. Uniqueness and Existence of the Triple Laplace Transform

In this section, we will study the uniqueness and existence of triple Laplace transform. First of all, let $f(x, y, t)$ be a continuous function on the interval $[0, \infty)$ which is of exponential order, that is, for some $a, b, c \in R$. Consider

$$
\begin{equation*}
\sup _{x, y, t>0}\left|\frac{f(x, y, t)}{\exp [a x+b y+c t]}\right|<0 \tag{4}
\end{equation*}
$$

Under the previous condition, the triple Laplace transform,

$$
\begin{align*}
& F(p, s, k)=\iiint_{0}^{\infty} \exp [-p x] \exp [-s y]  \tag{5}\\
& \times \exp [-k t] f(x, y, t) d x d y d t
\end{align*}
$$

exists for all $p>a, s>b$, and $k>c$ and is in actuality infinitely differentiable with respect to $p>a, s>b$ and $k>c$. All functions in this study are assumed to be of exponential order. The following theorem shows that $f(x, y, t)$ can be uniquely obtained from $F(p, s, t)$.

Theorem 3. Let $f(x, y, t)$ and $g(x, y, t)$ be continuous functions defined for $x, y, t \geq 0$ and having Laplace transforms, $F(p, s, k)$ and $G(p, s, k)$, respectively. If $F(p, s, k)=G(p, s, k)$, then $f(x, y, t)=g(x, y, t)$.

Proof. From the definition of the inverse Laplace transform, if $\alpha, \beta$, and $\mu$ are sufficiently large, then the integral expression, by

$$
\begin{align*}
& f(x, y, t) \\
& =\frac{1}{2 \pi i} \int_{\alpha-i \infty}^{\alpha+i \infty} e^{p x} \\
& \qquad \quad \times\left[\frac{1}{2 \pi i} \int_{\beta-i \infty}^{\beta+i \infty} e^{s y}\right. \\
& \quad \times\left[\frac{1}{2 \pi i} \int_{\mu-i \infty}^{\mu+i \infty} e^{k t}\right. \\
& \quad \times F(p, s, k) d k] d s] d p \tag{6}
\end{align*}
$$

for the triple inverse Laplace transform, can be used to obtain

$$
\begin{align*}
& f(x, y, t) \\
& \begin{array}{l}
=\frac{1}{2 \pi i} \int_{\alpha-i \infty}^{\alpha+i \infty} e^{p x} \\
\\
\quad \times\left[\frac{1}{2 \pi i} \int_{\beta-i \infty}^{\beta+i \infty} e^{s y}\right. \\
\\
\quad \times\left[\frac{1}{2 \pi i} \int_{\mu-i \infty}^{\mu+i \infty} e^{k t}\right. \\
\\
\quad \times F(p, s, k) d k] d s] d p
\end{array}
\end{align*}
$$

By hypothesis, we have that $F(p, s, k)=G(p, s, k)$. then replacing this in the previous expression, we have the following:

$$
\begin{align*}
& f(x, y, t) \\
& \begin{array}{l}
=\frac{1}{2 \pi i} \int_{\alpha-i \infty}^{\alpha+i \infty} e^{p x} \\
\\
\times\left[\frac{1}{2 \pi i} \int_{\beta-i \infty}^{\beta+i \infty} e^{s y}\right. \\
\\
\times\left[\frac{1}{2 \pi i} \int_{\mu-i \infty}^{\mu+i \infty} e^{k t}\right. \\
\\
\quad \times G(p, s, k) d k] d s] d p
\end{array}
\end{align*}
$$

which boil down to

$$
\begin{aligned}
& f(x, y, t) \\
& =\frac{1}{2 \pi i} \int_{\alpha-i \infty}^{\alpha+i \infty} e^{p x} \\
& \times\left[\frac{1}{2 \pi i} \int_{\beta-i \infty}^{\beta+i \infty} e^{s y}\right. \\
& \times\left[\frac{1}{2 \pi i} \int_{\mu-i \infty}^{\mu+i \infty} e^{k t}\right. \\
& \quad \times G(p, s, k) d k] d s] d p
\end{aligned}
$$

$$
\begin{equation*}
=g(x, y, t) \tag{9}
\end{equation*}
$$

and this proves the uniqueness of the triple Laplace transform.

Theorem 4. If, at the point $(p, s, k)$, the integrals

$$
\begin{align*}
F_{1}(p, s, k)=\iiint_{0}^{\infty} & \exp [-p x] \exp [-s y] \\
& \times \exp [-k t] f_{1}(x, y, t) d x d y d t  \tag{10}\\
F_{2}(p, s, k)=\iiint_{0}^{\infty} & \exp [-p x] \exp [-s y] \\
& \times \exp [-k t] f_{2}(x, y, t) d x d y d t
\end{align*}
$$

are convergent and in addition if

$$
\begin{align*}
F_{3}(p, s, k)=\iiint_{0}^{\infty} & \exp [-p x] \exp [-s y]  \tag{11}\\
& \times \exp [-k t] f_{3}(x, y, t) d x d y d t
\end{align*}
$$

is absolutely convergent, then, the following expression:

$$
\begin{equation*}
F(p, s, k)=F_{1}(p, s, k) F_{2}(p, s, k) F_{3}(p, s, k) \tag{12}
\end{equation*}
$$

is the Laplace transform of the function

$$
\begin{align*}
& f(x, y, t) \\
& \qquad \begin{array}{l}
=\int_{0}^{t} \int_{0}^{y} \int_{0}^{x} f_{3}\left(x-\left(x_{1}+\rho\right), y-\left(y_{1}+\sigma\right)\right. \\
\\
\left.t-\left(t_{1}+\tau\right)\right) f_{2}\left(x_{1}-\rho, y_{1}-\sigma, t_{1}-\tau\right) \\
\\
\times f_{1}(\rho, \sigma, \tau) d \rho d \sigma d \tau
\end{array}
\end{align*}
$$

and the integral

$$
\begin{align*}
& F(p, s, k)=\iiint_{0}^{\infty} \exp [-p x] \exp [-s y]  \tag{14}\\
& \times \exp [-k t] f(x, y, t) d x d y d t
\end{align*}
$$

is convergent at the point $(p, s, k)$; for the readers who are interested, they can see the proof in [11, 12].

Theorem 5. A function $f(x, y, t)$ which is continuous on $[0, \infty)$ and satisfies the growth condition (4) can be recovered from only $F(p, s, k)$ as

$$
\begin{align*}
f(x, y, t)= & \lim _{\substack{n_{1} \rightarrow \infty \\
n_{2} \rightarrow \infty \\
n_{3} \rightarrow \infty}} \frac{(-1)^{n_{1}+n_{2}+n_{3}}}{n_{1}!n_{2}!n_{3}!}\left(\frac{n_{1}}{x}\right)^{n_{1}+1}\left(\frac{n_{2}}{y}\right)^{n_{2}+1}  \tag{15}\\
& \times\left(\frac{n_{3}}{t}\right)^{n_{3}+1} \mathrm{X}^{n_{1}+n_{2}+n_{3}}\left[\frac{n_{1}}{x}, \frac{n_{2}}{y}, \frac{n_{3}}{t}\right]
\end{align*}
$$

Evidently, the main difficulty in using Theorem 5 for computing the inverse Laplace transform is the repeated symbolic differentiation of $F(p, s, k)$.

Let us see how Theorem 5 can be applicable. Let us consider the following functions:

$$
\begin{equation*}
f(x, y, t)=\exp [-a x-b y-c t] \tag{16}
\end{equation*}
$$

Naturally the triple Laplace transform of the previous function is given later as

$$
\begin{equation*}
F(p, s, k)=\frac{1}{(p-a)(s-b)(k-c)} . \tag{17}
\end{equation*}
$$

Now applying the high-order mixed derivative to the previous expression, we obtain the following:

$$
\begin{align*}
\frac{\partial^{n_{1}+n_{2}+n_{3}}[F(p, s, k)]}{\partial p^{n_{1}} \partial s^{n_{2}} \partial k^{n_{3}}}= & n_{1}!n_{2}!n_{3}!(-1)^{n_{1}+n_{2}+n_{3}} \\
& \times(a+P)^{-1-n_{1}}(s+b)^{-1-n_{2}}(c+k)^{-1-n_{3}} \tag{18}
\end{align*}
$$

Applying Theorem 5 in the previous expression, we obtain the following result:

$$
\begin{align*}
f(x, y, t)= & \lim _{\substack{n_{1} \rightarrow \infty \\
n_{2} \rightarrow \infty \\
n_{3} \rightarrow \infty}} \frac{n_{1}{ }^{1+n_{1}} n_{2}{ }^{1+n_{2}} n_{3}{ }^{1+n_{3}}}{x^{n_{1}+1} y^{n_{2}+1} t^{n_{3}+1}}\left(a+\frac{n_{1}}{x}\right)^{-n_{1}-1}  \tag{19}\\
& \times\left(b+\frac{n_{2}}{y}\right)^{-n_{2}-1}\left(c+\frac{n_{3}}{t}\right)^{-n_{3}-1} .
\end{align*}
$$

Making a change of variable in the previous expression, we obtain the following simplified result:

$$
\begin{align*}
f(x, y, t)= & \lim _{\substack{n_{1} \rightarrow \infty \\
n_{2} \rightarrow \infty \\
n_{3} \rightarrow \infty}}\left(1+\frac{a n_{1}}{x}\right)^{-n_{1}-1}\left(1+\frac{b n_{2}}{y}\right)^{-n_{2}-1}  \tag{20}\\
& \times\left(1+\frac{c n_{3}}{t}\right)^{-n_{3}-1} .
\end{align*}
$$

Using together, the application of logarithm and the L'Hôpital's rule on the previous expression, we arrive at the following result:

$$
\begin{align*}
\ln (f(x, y, t)) & =-a x-b y-c t \Longrightarrow f(x, y, t) \\
& =\exp [-a x-b y-c t] \tag{21}
\end{align*}
$$

## 4. Some Properties of Triple Laplace Transform

In this section, we present some properties of the triple Laplace transform. Note that these properties follow from those of the double Laplace transform introduced by Kılıçman and Eltayeb [8]. The properties of the triple Laplace transform will enable us to find further transform pairs $\{f(x, y, t), F(p, s, k)\}$ :

$$
\text { (i) } \begin{align*}
F(p & +a, s+b, k+d) \\
& =L_{x, y, t}\left[e^{-a x-y b-c t} f(x, y, t)\right](p, s, k) \tag{22}
\end{align*}
$$

We will present the proof

$$
\begin{align*}
& L_{x, y, t}\left[e^{-a x-y b-c t} f(x, y, t)\right](p, s, k) \\
& =\iiint_{0}^{\infty} \exp [-p x] \exp [-s y] \exp [-k t] \exp [-a x] \\
& \quad \times \exp [-b y] \exp [-c t] f(x, y, t) d x d y d t \\
& \int_{0}^{\infty} \exp [-p x] \exp [-a x] \\
& \times\left(\iint_{0}^{\infty} \exp [-s y] \exp [-k t] \exp [-b y]\right. \\
& \quad \times \exp [-c t] f(x, y, t) d t d y) d t \tag{23}
\end{align*}
$$

Note that the integral inside the bracket satisfies the properties of the double Laplace transform and is given as [11]

$$
\begin{align*}
& \left(\iint_{0}^{\infty} \exp [-s y] \exp [-k t] \exp [-b y] \exp [-c t]\right.  \tag{24}\\
& \quad \times f(x, y, t) d t d y)=F(x, s+b, k+d)
\end{align*}
$$

Thus

$$
\begin{align*}
& \int_{0}^{\infty} \exp [-p x] \exp [-a x] F(x, s+b, k+d) d t  \tag{25}\\
& \quad=F(p+a, s+b, k+d)
\end{align*}
$$

and this completes the proof.
(ii) The following can also be observed:

$$
\begin{equation*}
\frac{1}{\alpha \beta \gamma} F\left(\frac{p}{\alpha}, \frac{s}{\beta}, \frac{k}{\gamma}\right)=L_{x, y, t}[f(\alpha x, \beta y, \gamma t)](p, s, k) . \tag{26}
\end{equation*}
$$

We will present the proof

$$
\begin{align*}
& L_{x, y, t}[f(\alpha x, \beta y, \gamma t)](p, s, k) \\
& =\iiint_{0}^{\infty} \exp [-p x] \exp [-s y] \exp [-k t] \\
& \quad \times f(\alpha x, \beta y, \gamma t) d x d y d t  \tag{27}\\
& \quad \begin{aligned}
& \int_{0}^{\infty} \exp [-p x]\left(\iint_{0}^{\infty} \exp [-s y] \exp [-k t]\right. \\
&\quad \times f(\alpha x, \beta y, \gamma t) d y d t) d x
\end{aligned}
\end{align*}
$$

Note that the double integral inside the bracket satisfies the property of the double Laplace transform as [11]

$$
\begin{align*}
& \left(\iint_{0}^{\infty} \exp [-s y] \exp [-k t] f(\alpha x, \beta y, \gamma t) d y d t\right) \\
& \quad=\frac{1}{\beta \gamma} F\left(\alpha x, \frac{s}{\beta}, \frac{k}{\gamma}\right) . \tag{28}
\end{align*}
$$

Thus

$$
\begin{align*}
L_{x, y, t} & {[f(\alpha x, \beta y, \gamma t)](p, s, k) } \\
& =\int_{0}^{\infty} \exp [-p x] \frac{1}{\beta \gamma} F\left(\alpha x, \frac{s}{\beta}, \frac{k}{\gamma}\right) d x  \tag{29}\\
& =\frac{1}{\alpha \beta \gamma} F\left(\frac{p}{\alpha}, \frac{s}{\beta}, \frac{k}{\gamma}\right),
\end{align*}
$$

and this completes the proof.
(iii) The following property can also be observed:

$$
\begin{align*}
& \frac{\partial^{n+m+v}[F(p, s, k)]}{\partial p^{n} \partial s^{n} \partial k^{v}}  \tag{30}\\
& \quad=L_{x, y, t}\left[(-1)^{n+m+v} x^{n} y^{m} t^{v} f(x, y, t)\right](p, s, k)
\end{align*}
$$

We will present the proof

$$
\begin{align*}
F(p, s, k)=\iiint_{0}^{\infty} & \exp [-p x] \exp [-s y] \exp [-k t]  \tag{31}\\
& \times f(x, y, t) d x d y d t
\end{align*}
$$

Then,

$$
\begin{align*}
& \frac{\partial^{n+m+v}[F(p, s, k)]}{\partial p^{n} \partial s^{n} \partial k^{v}} \\
& =\frac{\partial^{n+m+v}}{\partial p^{n} \partial s^{n} \partial k^{v}}\left(\iiint_{0}^{\infty} \exp [-p x] \exp [-s y]\right. \\
&  \tag{32}\\
& \quad \times \exp [-k t] f(x, y, t) d x d y d t)
\end{align*}
$$

Now making use of the convergence properties of the improper integral involved, we can interchange the operation of differentiation and integration and differentiate with
respect to $p, s$, and $k$ under the integral sign. Thus, we arrive at the following expression:

$$
\begin{align*}
& \frac{\partial^{n+m+v}[F(p, s, k)]}{\partial p^{n} \partial s^{n} \partial k^{v}} \\
& =\frac{\partial^{n}}{\partial p^{n}} \int_{0}^{\infty} \exp [-p x] \\
&  \tag{33}\\
& \quad \times\left(\frac{\partial^{m+v}}{\partial s^{n} \partial k^{v}} \iint_{0}^{\infty} \exp [-s y] \exp [-k t]\right. \\
& \\
& \quad \times f(x, y, t) d y d t) d x
\end{align*}
$$

Note that the expression in the bracket satisfies the property of the double Laplace transform as [11]

$$
\begin{gather*}
\frac{\partial^{m+v}}{\partial s^{n} \partial k^{v}} \iint_{0}^{\infty} \exp [-s y] \exp [-k t] f(x, y, t) d y d t  \tag{34}\\
=L_{y, t}\left[(-1)^{m+v} y^{m} t^{v} f(x, y, t)\right](s, k)
\end{gather*}
$$

Thus

$$
\begin{align*}
& \frac{\partial^{n+m+v}[ }{\partial p^{n} \partial s^{n} \partial k^{v}} \\
& =\frac{\partial^{n}}{\partial p^{n}} \int_{0}^{\infty} \exp [-p x] \\
& \quad \times\left(L_{y, t}\left[(-1)^{m+v} y^{m} t^{v} f(x, y, t)\right](s, k)\right) d x . \tag{35}
\end{align*}
$$

And finally, we obtain

$$
\begin{align*}
& \frac{\partial^{n+m+v}[F(p, s, k)]}{\partial p^{n} \partial s^{n} \partial k^{v}}  \tag{36}\\
& \quad=L_{x, y, t}\left[(-1)^{n+m+v} x^{n} y^{m} t^{v} f(x, y, t)\right](p, s, k),
\end{align*}
$$

and this completes the proof.
Now using the previous three properties, we will show the proof of Theorem 5 .

Proof of Theorem 5. Let us define the set of functions depending on parameters $m, n$, and $v$ as

$$
\begin{equation*}
h_{m, n, v}(x, y, t)=\frac{m^{m+1} n^{n+1} v^{v+1}}{m!n!v!} x^{m} y^{n} t^{v} e^{-m x-n y-v t} \tag{37}
\end{equation*}
$$

It worth noting that the previous function is a kind of threedimensional density of probability, and it therefore follows that

$$
\begin{equation*}
\iiint_{0}^{\infty} h_{m, n, v}(x, y, t) d x d y d t=1 \tag{38}
\end{equation*}
$$

In addition of this, we will have that

$$
\begin{equation*}
\lim _{\substack{m \rightarrow \infty \\ n \rightarrow \infty \\ v \rightarrow \infty}} \iiint_{0}^{\infty} h_{m, n, v}(x, y, t) \psi(x, y, t) d x d y d t=\psi(1,1,1) \tag{39}
\end{equation*}
$$

where $\psi(x, y, t)$ is any continuous function. Let $\Psi(p, s, k)$ denote the triple Laplace transform of the continuous function $\psi(x, y, t)$. However, if one defines the function $M(x, y, t)=f(x \alpha, y \beta, t \gamma)$, making use of the second property established in (29), we arrive at the following:

$$
\begin{equation*}
\frac{1}{\alpha \beta \gamma} F\left(\frac{p}{\alpha}, \frac{s}{\beta}, \frac{k}{\gamma}\right)=L_{x, y, t}[f(\alpha x, \beta y, \gamma t)](p, s, k) . \tag{40}
\end{equation*}
$$

Here if one applies the third property, in particular by replacing $p=m / x, s=n / y, k=v / t$ as follows:

$$
\begin{gather*}
L_{x y t}(M(x, y, t))=\frac{1}{\alpha \beta \gamma} F\left(\frac{p}{\alpha}, \frac{s}{\beta}, \frac{k}{\gamma}\right),  \tag{41}\\
\frac{\partial^{n+m+v}\left[L_{x y t}(M(x, y, t))\right]}{\partial p^{n} \partial s^{n} \partial k^{v}} \\
=\frac{\partial^{n+m+v}[(1 / \alpha \beta \gamma) F(p / \alpha, s / \beta, k / \gamma)]}{\partial p^{n} \partial s^{n} \partial k^{v}}  \tag{42}\\
=\frac{1}{\alpha^{m+1} \beta^{n+1} \gamma^{v+1}} \\
\quad \times \frac{\partial^{n+m+v}[F(p / \alpha, s / \beta, k / \gamma)]}{\partial p^{n} \partial s^{n} \partial k^{v}} .
\end{gather*}
$$

Now let us put $\psi(x, y, t)=e^{-p x-s y-k t} M(x, y, t)$. Now if we make use of (38), we obtain the following

$$
\begin{align*}
& \psi(1,1,1)=e^{-p-s-k} M(1,1,1)=e^{-p-s-k} f(\alpha, \beta, \gamma) \\
&=\lim _{\substack{m \rightarrow \infty \\
n \rightarrow \infty \\
v \rightarrow \infty}} \frac{m^{m+1} n^{n+1} v^{v+1}}{m!n!v!} \iiint_{0}^{\infty} x^{m} y^{n} t^{v} e^{-p x-s y-k t} \\
& \quad \times e^{-m x-n y-v t} \Psi(x, y, t) d x d y d t \\
&=\lim _{\substack{m \rightarrow \infty \\
n \rightarrow \infty \\
v \rightarrow \infty}} \frac{m^{m+1} n^{n+1} v^{v+1}}{m!n!v!} L_{x y t}\left[x^{m} y^{n} t^{v} e^{-m x-n y-v t} \Psi(x, y, t)\right] . \tag{43}
\end{align*}
$$

Now taking into account properties (i) and (ii), (42) together with the function $M(x, y, t)$, we arrive at the following:

$$
\begin{aligned}
& L_{x y t} {\left[x^{m} y^{n} t^{v} e^{-m x-n y-v t} \Psi(x, y, t)\right] } \\
&=(-1)^{m+n+v} \frac{\partial^{n+m+v}\left[L_{x y t}\left(e^{-m x-n y-k t} \Psi(x, y, t)\right)(p, s, k)\right]}{\partial p^{n} \partial s^{n} \partial k^{v}} \\
&=(-1)^{m+n+v} \frac{1}{\alpha^{m} \beta^{n} \gamma^{v}} \\
& \times \frac{\partial^{n+m+v}\left[L_{x y t}(\Psi(x, y, t))(p+m, s+n, k+v)\right]}{\partial p^{n} \partial s^{n} \partial k^{v}} \\
&=(-1)^{m+n+v} \frac{1}{\alpha^{m} \beta^{n} \gamma^{v}} \\
& \times\left(\left(\partial ^ { n + m + v } \left[L_{x y t}(f(\alpha x, \beta y, \gamma t))\right.\right.\right. \\
&\left.\left.\left.\times\left(\frac{p+m}{\alpha}, \frac{s+n}{\beta}, \frac{k+v}{\gamma}\right)\right]\right)\left(\partial p^{n} \partial s^{n} \partial k^{v}\right)^{-1}\right)
\end{aligned}
$$

$$
\begin{align*}
= & (-1)^{m+n+v} \\
& \times \frac{1}{\alpha^{m} \beta^{n} \gamma^{v}} \frac{\partial^{n+m+v}[F((p+m) / \alpha,(s+n) / \beta,(k+v) / \gamma)]}{\partial p^{n} \partial s^{n} \partial k^{v}} . \tag{44}
\end{align*}
$$

Now observe that from (44) with the fact that $f(\alpha, \beta, t)=$ $\psi(1,1,1) e^{p+s+k}$, we arrive at the following:

$$
\begin{align*}
& f(\alpha, \beta, t) \\
&= e^{p+s+k} \lim _{\substack{m \rightarrow \infty \\
n \rightarrow \infty \\
v \rightarrow \infty}} \frac{m^{m+1} n^{n+1} v^{v+1}}{m!n!v!} \\
& \times\left(\frac{m}{\alpha}\right)^{m+1}\left(\frac{n}{\beta}\right)^{n+1}\left(\frac{v}{\gamma}\right)^{v+1} \\
& \times \frac{\partial^{n+m+v}[F((p+m) / \alpha,(s+n) / \beta,(k+v) / \gamma)]}{\partial p^{n} \partial s^{n} \partial k^{v}} . \tag{45}
\end{align*}
$$

The previously mentioned is true for any $p, s, k$ in the complete space, in particular, for $p=0, s=0, k=0$, and in this case Theorem 5 is covered.

## 5. Application to Third-Order Partial Differential Equation

In this section, we present the application of this operator for solving some kind of third-order partial differential equations.

Example 1. consider the following third-order partial differential equation:

$$
\begin{equation*}
\partial_{x y t} u(x, y, t)+u(x, y, t)=0 \tag{46}
\end{equation*}
$$

The previous equation is called the Mboctara equation and is subjected to the following boundaries and initial conditions:

$$
\begin{array}{ll}
u(x, y, 0)=e^{x+y}, & u(x, 0, t)=e^{x-t} \\
u(0, y, t)=e^{y-t}, & u(x, y, 1)=e^{x+y-1} \tag{47}
\end{array}
$$

Now applying the triple Laplace transform on both sides of (46), we obtain the following:

$$
\begin{equation*}
p s k U(p, s, k)+U(p, s, k)=G(p, s, k) . \tag{48}
\end{equation*}
$$

Here

$$
\begin{align*}
G(p, s, k)= & p s U(p, s, 0)+p s U(p, 0, k)-p U(p, 0,0) \\
& +s k U(0, s, k)-s U(0, s, 0)-k U(0,0, k) \\
& +U(0,0,0) . \tag{49}
\end{align*}
$$

Factorising the right side of equation (49), we obtain the following:

$$
\begin{equation*}
U(p, s, k)=\frac{G(p, s, k)}{1+p s k} \tag{50}
\end{equation*}
$$

Now applying the inverse triple Laplace transform on the previous equation we obtain the following solution:

$$
\begin{equation*}
u(x, y, t)=L_{x y t}^{-1}\left[\frac{G(p, s, k)}{1+p s k}\right]=e^{x+y-t} \tag{51}
\end{equation*}
$$

This is the exact solution for Mboctara equation.
Example 2. Let us consider the following nonhomogeneous Mboctara equation

$$
\begin{equation*}
\partial_{x y t} u(x, y, t)+u(x, y, t)=-e^{x-2 y+t} \tag{52}
\end{equation*}
$$

subjected to the following initial and boundaries conditions:

$$
\begin{gather*}
u(x, 0,0)=e^{x}, \quad \partial_{t} u(x, 0, t)=e^{x+t}, \quad \partial_{x} u(x, 0, t)=e^{x+t}, \\
u(0,0,0)=1, \quad u(x, 0.5, t)=e^{x+t-1} \tag{53}
\end{gather*}
$$

Now applying the triple Laplace transform on both sides of (52), we obtain the following:

$$
\begin{align*}
p s k U & (p, s, k)+U(p, s, k) \\
& =G(p, s, k)-\frac{1}{(1+p)(2+s)(1+k)} \tag{54}
\end{align*}
$$

Factorising the right side of (54), we obtain the following:

$$
\begin{equation*}
U(p, s, k)=\frac{G(p, s, k)-1 /(1+p)(2+s)(1+k)}{1+p s k} . \tag{55}
\end{equation*}
$$

Now applying the inverse triple Laplace transform on the previous equation, we obtain the following solution

$$
\begin{align*}
u(x, y, t) & =L_{x y t}^{-1}\left[\frac{G(p, s, k)-1 /(1+p)(2+s)(1+k)}{1+p s k}\right] \\
& =e^{x-2 y+t} \tag{56}
\end{align*}
$$

This is the exact solution for nonhomogeneous Mboctara equation.

Example 3. Let us consider the following nonhomogeneous Mboctara equation

$$
\begin{align*}
\partial_{x y t} u(x, y, t)+u(x, y, t)= & \cos (x) \cos (y) \cos (-t) \\
& -\sin (x) \sin (y) \sin (-t) \tag{57}
\end{align*}
$$

subjected to the following initial and boundaries conditions:

$$
\begin{gather*}
u(x, y, 0)=\cos (x) \cos (y) \\
\partial_{t} u(x, y, 0)=\partial_{x} u(0, y, t)=\partial_{y} u(x, 0, t)=0  \tag{58}\\
u\left(x, \frac{\pi}{2}, t\right)=u\left(x, y, \frac{\pi}{2}\right)=u\left(\frac{\pi}{2}, y, t\right)=0
\end{gather*}
$$

Table 1: Table of triple Laplace transform for some function of three variables.

| Functions $f(x, y, t)$ | Triple laplace transform $F(p, s, k)$ |
| :---: | :---: |
| $a b c$ | $\frac{a b c}{p s k}$ |
| xyt | $\frac{1}{p^{2} s^{2} k^{2}}$ |
| $x^{n} y^{m} t^{v}, n, m, v$ are natural numbers | $k^{-1-v} s^{-1-m} p^{-n-1} \Gamma(1+n) \Gamma(1+m) \Gamma(1+v)$ |
| $x^{n} y^{m} t^{v} e^{-a x-b y-c t}$ | $(k+c)^{-1-v}(s+b)^{-1-m}(p+a)^{-n-1} \Gamma(1+n) \Gamma(1+m) \Gamma(1+v)$ |
| $e^{-a x-b y-c t}$ | $\frac{1}{(a+p)(b+s)(c+k)}$ |
| $\cos (x) \cos (y) \cos (t)$ | $\frac{k s p}{\left(1+p^{2}\right)\left(1+s^{2}\right)\left(1+k^{2}\right)}$ |
| $\sin (x) \sin (y) \sin (t)$ | $\frac{1}{\left(1+p^{2}\right)\left(1+s^{2}\right)\left(1+k^{2}\right)}$ |
| $\sin (x+y+t)$ | $\frac{-1+p s+k(p+s)}{\left(1+p^{2}\right)\left(1+s^{2}\right)\left(1+k^{2}\right)}$ |
| $\cos (x+y+t)$ | $-\frac{k+p+s-k p s}{\left(1+p^{2}\right)\left(1+s^{2}\right)\left(1+k^{2}\right)}$ |
| $\sqrt{x y t}$ | $\frac{\pi \sqrt{\pi}}{8^{3} \sqrt{k s p}}$ |
| $e^{a x+y b+c t} \sinh (a x) \sinh (b y) \sinh (c t)$ | $\frac{(b)(c)(a)}{\left(-2 a p+p^{2}\right)\left(-2 b s+s^{2}\right)\left(-2 c k+k^{2}\right)}$ |
| $e^{a x+y b+c t} \cosh (a x) \cosh (b y) \cosh (c t)$ | $\frac{(b-s)(c-k)(a-p)}{\left(-2 a p+p^{2}\right)\left(-2 b s+s^{2}\right)\left(-2 c k+k^{2}\right)}$ |
| $\operatorname{Erf}\left[\frac{a}{2 \sqrt{x}}\right] \operatorname{Erf}\left[\frac{b}{2 \sqrt{y}}\right] \operatorname{Erf}\left[\frac{c}{2 \sqrt{t}}\right]$ | $\frac{e^{-\sqrt{c^{2} k}-\sqrt{b^{2} s}}}{k p s}\left(-1+e^{-\sqrt{c^{2} k}}\right)\left(1-e^{-\sqrt{a^{2} p}}\right)\left(-1+e^{-\sqrt{b^{2} s}}\right)$ |
| $\frac{\sin (a x)}{x} \frac{\sin (b y)}{y} \frac{\sin (c t)}{t}$ | $\arctan \left(\frac{\sqrt{a^{2}}}{p}\right) \arctan \left(\frac{\sqrt{b^{2}}}{s}\right) \arctan \left(\frac{\sqrt{c^{2}}}{k}\right)$ |
| $\frac{\cos (a x)}{x^{n}} \frac{\cos (b y)}{y^{m}} \frac{\cos (c t)}{t^{v}}$ | $\begin{aligned} & k^{-1+v}\left(1+\frac{b^{2}}{s^{2}}\right)^{1 / 2(-1+m)} s^{-1+m} \cos [c t] \cos \left((-1+m) \arctan \left[\frac{\sqrt{b^{2}}}{s}\right]\right) \Gamma(1-m) \Gamma(1-v) \\ & \times\left(1+\frac{a^{2}}{p^{2}}\right)^{1 / 2(-1+n)} p^{-1+n} \cos \left((n-1) \arctan \left(\frac{\|a\|}{p}\right)\right) \Gamma(1-n) \end{aligned}$ |
| $\frac{\sin (a x)}{x^{n}} \frac{\sin (b y)}{y^{m}} \frac{\sin (c t)}{t^{v}}$ | $\begin{aligned} & k^{-1+v}\left(1+\frac{b^{2}}{s^{2}}\right)^{1 / 2(-1+m)} s^{-1+m} \Gamma(1-m) \Gamma(1-v)\left(1+\frac{a^{2}}{p^{2}}\right)^{1 / 2(-1+n)} \\ & \times p^{-1+n} \Gamma(1-n) \operatorname{sign}(a) \sin \left((n-1) \arctan \left(\frac{\|a\|}{p}\right)\right) \Gamma(1-n) \end{aligned}$ |
| $J_{n}(x) J_{n}(y) J_{n}(t)$ | $\begin{gathered} 8^{-n}(k s p)^{-1-n} \text { Hypergeometric } 2 F 1\left(\frac{1+n}{2}, \frac{2+n}{2},-\frac{1}{k^{2}}\right) \\ \text { Hypergeometric } 2 F 1\left(\frac{1+n}{2}, \frac{2+n}{2},-\frac{1}{s^{2}}\right) \\ {\left[\text { Hypergeometric } 2 F 1\left(\frac{1+n}{2}, \frac{2+n}{2},-\frac{1}{p^{2}}\right)\right.} \end{gathered}$ |
| $I_{n}(x) I_{n}(y) I_{n}(t)$ | $\begin{gathered} 8^{-n}(k s p)^{-1-n} \text { Hypergeometric } 2 F 1\left(\frac{1+n}{2}, \frac{2+n}{2}, 1+n, \frac{1}{k^{2}}\right) \\ \text { Hypergeometric } 2 F 1\left(\frac{1+n}{2}, \frac{2+n}{2}, 1+n, \frac{1}{s^{2}}\right) \\ \text { Hypergeometric } 2 F 1\left(\frac{1+n}{2}, \frac{2+n}{2}, 1+n, \frac{1}{p^{2}}\right) \end{gathered}$ |

Exact solution of the nonhomogeneous Mboctara equation (4.1)


Figure 1: Numerical simulation of the exact solutions of the Homogeneous and non-homogeneous Mboctara equations.

Now applying the triple Laplace transform on both sides of (57), we obtain the following:

Factorising the right side of (59), we obtain the following:

$$
\begin{align*}
p s k U & (p, s, k)+U(p, s, k) \\
= & G(p, s, k)+\frac{k s p}{\left(1+p^{2}\right)\left(1+s^{2}\right)\left(-1+k^{2}\right)}  \tag{59}\\
& -\frac{1}{\left(1+p^{2}\right)\left(1+s^{2}\right)\left(-1+k^{2}\right)} . \tag{60}
\end{align*}
$$

$$
\begin{aligned}
U(p, s, k)= & \frac{G(p, s, k)}{1+p s k}+\frac{k s p /\left(1+p^{2}\right)\left(1+s^{2}\right)\left(-1+k^{2}\right)}{1+p s k} \\
& -\frac{1 /\left(1+p^{2}\right)\left(1+s^{2}\right)\left(-1+k^{2}\right)}{1+p s k}
\end{aligned}
$$

Now applying the inverse triple Laplace transform on the previous equation, we obtain the following solution:

$$
\begin{align*}
& u(x, y, t) \\
& =L_{x y t}^{-1}\left[\frac{G(p, s, k)}{1+p s k}+\frac{k s p /\left(1+p^{2}\right)\left(1+s^{2}\right)\left(-1+k^{2}\right)}{1+p s k}\right. \\
& \left.-\frac{1 /\left(1+p^{2}\right)\left(1+s^{2}\right)\left(-1+k^{2}\right)}{1+p s k}\right] \\
& =\cos (x) \cos (y) \cos (-t) . \tag{61}
\end{align*}
$$

This is the exact solution for nonhomogeneous Mboctara equation.

Example 4. consider the following nonlinear nonhomogeneous with variable coefficient Mboctara equation:

$$
\begin{align*}
& e^{x+y+t} \partial_{x y t} u(x, y, t)-3 u^{2}(x, y, t)+e^{x+y+t} u(x, y, t) \\
& \quad=e^{2 x+2 y+2 t},  \tag{62}\\
& u_{x}(x, y, 0)=e^{x+y}, \quad u(0,0,0)=1, \\
& u(1,0,0)=e, \quad \partial_{x y t} u(0,0,0)=1 .
\end{align*}
$$

Now applying the triple Laplace transform on both sides of (62) and then using the properties of the triple Laplace transform and after factorising as in the previous examples and taking the inverse triple Laplace transform, we obtain the following as an exact solution of this type of Mboctara equation:

$$
\begin{equation*}
u(x, y, t)=e^{x+y+t} \tag{63}
\end{equation*}
$$

The numerical simulations of the exact solutions of the Mboctara equation are depicted in Figure 1(a) (4.1), Figure 1(b) (4.6), Figure 1(c) (4.6) and Figure 1(d) (4.11), respectively.

## 6. Triple Laplace Transform of Some Functions of Three Variables

In this section, we examine the triple Laplace transform of some functions in Table 1:

$$
\begin{aligned}
& L_{x y t}\left(Y_{n}(x) Y_{n}(y) Y_{n}(t)\right) \\
& =8^{-n}(k s p)^{-1-n} \operatorname{Csc}^{3}[n \pi] \\
& \quad \times\left(-\left(4 k^{2}\right)^{n} \text { Hypergeometric } 2 F 1\right. \\
& \\
& \quad \times\left(\frac{1+n}{2}, \frac{2+n}{2}, 1-n,-\frac{1}{k^{2}}\right) \\
& \\
& +\cos (n \pi) \text { Hypergeometric } 2 F 1 \\
& \\
& \left.\quad \times\left(\frac{1+n}{2}, \frac{2+n}{2}, 1+n,-\frac{1}{k^{2}}\right)\right)
\end{aligned}
$$

$$
\begin{align*}
& \times( -\left(4 s^{2}\right)^{n} \text { Hypergeometric2F1 } \\
& \times\left(\frac{1+n}{2}, \frac{2+n}{2}, 1-n,-\frac{1}{s^{2}}\right) \\
&+\cos (n \pi) \text { Hypergeometric } 2 F 1 \\
&\left.\times\left(\frac{1+n}{2}, \frac{2+n}{2}, 1+n,-\frac{1}{s^{2}}\right)\right) \\
& \times\left(-\left(4 p^{2}\right)^{n} \text { Hypergeometric } 2 F 1\right. \\
& \quad \times\left(\frac{1+n}{2}, \frac{2+n}{2}, 1-n,-\frac{1}{p^{2}}\right) \\
& \quad+\cos (n \pi) \text { Hypergeometric } 2 F 1 \\
&\left.\times\left(\frac{1+n}{2}, \frac{2+n}{2}, 1+n,-\frac{1}{p^{2}}\right)\right) . \tag{64}
\end{align*}
$$

## 7. Conclusion

This work presents the definition of the triple Laplace transform. Some triple Laplace transform is presented in Table 1. Some theorems and properties of this new relatively new operator are presented. Applications of the new operator, for solving some kind of third-order partial differential equations called Mboctara equation, are presented. Numerical solutions of the Mboctara equation are given.

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## Research Article

# Approximate Solution of Tuberculosis Disease Population Dynamics Model 

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Received 22 March 2013; Accepted 2 June 2013
Academic Editor: R. K. Bera
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#### Abstract

We examine possible approximate solutions of both integer and noninteger systems of nonlinear differential equations describing tuberculosis disease population dynamics. The approximate solutions are obtained via the relatively new analytical technique, the homotopy decomposition method (HDM). The technique is described and illustrated with numerical example. The numerical simulations show that the approximate solutions are continuous functions of the noninteger-order derivative. The technique used for solving these problems is friendly, very easy, and less time consuming.


## 1. Introduction

Tuberculosis, MTB, or TB (short for tubercle bacillus) is a common and in many cases lethal, infectious disease caused by various strains of Mycobacterium, usually Mycobacterium Tuberculosis [1]. Tuberculosis typically attacks the lungs, but can also affect other parts of the body. It is spread through the air when people who have an active TB infection cough, sneeze, or otherwise transmit their saliva through the air [2]. Most infections are asymptomatic and latent, but about one in ten latent infections eventually progresses to active disease which, if left untreated, kills more than $50 \%$ of those so infected. Interested reader can find more about this model in [3-7].

Based on the standard SIRS model, the model population was compartmentalised into the susceptible ( $S$ ) and the infected ( $I$ ) which is further broken down into latently infected $\left(I_{L}\right)$ and actively infected $\left(I_{A}\right)$ while the recovered subpopulation is ploughed back into the susceptible group due to the possibility of reinfection after successful treatment of the earlier infection. The model monitors the temporary dynamics in the population of susceptible people $(t)$, TB latently infected people $I_{L}(t)$, and TB actively infected people
$I_{A}(t)$ as captured in the model system of ordinary differential equations that follows.

$$
\begin{align*}
& \frac{d S(t)}{d t}=v f N-\alpha I_{A} S(t)+\delta S(t)+T_{A} I_{A}(t)+T_{L} I_{L}(t) \\
& \frac{d I_{L}(t)}{d t}=(1-P) \alpha I_{A} S(t)-\beta_{A} I_{L}(t)-T_{L} I_{L}(t)-\delta I_{L}(t) \\
& \frac{d I_{A}(t)}{d t}=P \alpha I_{A} S(t)+\beta_{A} I_{L}(t)-T_{A} I_{A}(t)-\delta I_{A}(t)-\varepsilon I_{A}(t) \tag{1}
\end{align*}
$$

subject to the initial conditions

$$
\begin{equation*}
S(0)=N, \quad I_{L}(0) \geq 0, \quad I_{A}(0) \geq 0, \tag{2}
\end{equation*}
$$

where $N$ is the total number of new people in the location of interest; $S$ is the number of susceptible people in the location; $I_{L}$ is the number of TB latently infected people; $I_{A}$ is the number of TB actively infected people; $v$ is the probability that a susceptible person is not vaccinated; $f$ is the efficient rate of vaccines; $T_{L}$ is the success rate of latent $T_{B}$ therapy; $T_{A}$ is the active TB treatment cure rate; $\alpha$ is the TB instantaneous
incidence rate per susceptible; $\delta$ is humans natural death rate; $P$ is the proportion of infection instantaneously degenerating into active TB; $\varepsilon$ is the TB-induced death rate; and $\beta_{A}$ is the breakdown rate from latent to active TB. The equilibrium analysis of the model was studied in [8]. Equation (1) together with (2) does not have an exact solution and is usually solved numerically.

The purpose of this paper is to derive approximate analytical solutions for the standard form as well as the fractional version of (1) together with (2) using the relatively new analytical technique, the homotopy decomposition method (HDM).

The paper is structured as follows. In Section 2, we present the basic ideal of the homotopy decomposition method for solving partial differential equations. We present the application of the HDM for system Tuberculosis disease population dynamics model in Section 3. In Section 4, we present the application of the HDM for system of fractional Tuberculosis disease population dynamics model. The conclusions are then given finally in Section 5.

## 2. Fundamental Information about Homotopy Decomposition Method

To demonstrate the elementary notion of this technique, we consider a universal nonlinear nonhomogeneous partial differential equation with initial conditions of the following form [9-13].

$$
\begin{array}{r}
\frac{\partial^{m} U(x, t)}{\partial t^{m}}=L(U(x, t))+N(U(x, t))+f(x, t)  \tag{3}\\
m=1,2,3 \ldots
\end{array}
$$

focused on the primary condition

$$
\begin{array}{r}
\frac{\partial^{i} U(x, 0)}{\partial t^{i}}=y_{i}(x), \quad \frac{\partial^{m-1} U(x, 0)}{\partial t^{m-1}}=0  \tag{4}\\
i=0,1,2, \ldots, m-2
\end{array}
$$

where $m$ is the order of the derivative, where $f$ is an identified function, $N$ is the common nonlinear differential operator, $L$ denotes a linear differential operator, and $m$ is the order of the derivative. The procedures first stage here is to apply the inverse operator $\partial^{m} / \partial t^{m}$ on both sides of (3) to obtain

$$
\begin{align*}
U(x, t)= & \sum_{k=0}^{m-1} \frac{t^{k}}{k!} \frac{d^{k} u(x, 0)}{d t^{k}} \\
& +\int_{0}^{t} \int_{0}^{t_{1}} \cdots \int_{0}^{t_{m-1}} L(U(x, \tau)) \\
& +N(U(x, \tau))+f(x, \tau) d \tau \cdots d t \tag{5}
\end{align*}
$$

The multi-integral in (3) can be transformed to

$$
\begin{align*}
& \int_{0}^{t} \int_{0}^{t_{1}} \cdots \int_{0}^{t_{m-1}} L(U(x, \tau)) \\
& \quad+N(U(x, \tau))+f(x, \tau) d \tau \cdots d t  \tag{6}\\
& =\frac{1}{(m-1)!} \int_{0}^{t}(t-\tau)^{m-1} L(U(x, \tau)) \\
& \quad+N(U(x, \tau))+f(x, \tau) d \tau
\end{align*}
$$

so that (3) can be reformulated as

$$
\begin{align*}
& U(x, t)=\sum_{k=0}^{m-1} \frac{t^{k}}{k!} y_{i}(x) \\
& \\
& \quad+\frac{1}{(m-1)!} \int_{0}^{t}(t-\tau)^{m-1} L(U(x, \tau))  \tag{7}\\
& \\
& \quad+N(U(x, \tau))+f(x, \tau) d \tau
\end{align*}
$$

Using the homotopy scheme, the solution of the aforementioned integral equation is given in series form as

$$
\begin{align*}
& U(x, t, p)=\sum_{n=0}^{\infty} p^{n} U_{n}(x, t)  \tag{8}\\
& U(x, t)=\lim _{p \rightarrow 1} U(x, t, p)
\end{align*}
$$

and the nonlinear term can be decomposed as

$$
\begin{equation*}
N U(r, t)=\sum_{n=1}^{\infty} p^{n} \mathscr{H}_{n}(U) \tag{9}
\end{equation*}
$$

where $p \in(0,1]$ is an implanting parameter. $\mathscr{H}_{n}(U)$ is the polynomials that can be engendered by

$$
\begin{equation*}
\mathscr{H}_{n}\left(U_{0}, \ldots, U_{n}\right)=\frac{1}{n!} \frac{\partial^{n}}{\partial p^{n}}\left[N\left(\sum_{j=0}^{n} p^{j} U_{j}(x, t)\right)\right] \tag{10}
\end{equation*}
$$

$$
n=0,1,2 \ldots
$$

The homotopy decomposition method is obtained by the combination of decomposition method with Abel integral and is given by

$$
\begin{align*}
& \sum_{n=0}^{\infty} p^{n} U_{n}(x, t) \\
& \quad=T(x, t)+p \frac{1}{(m-1)!} \\
& \quad \times \int_{0}^{t}(t-\tau)^{m-1}\left[f(x, \tau)+L\left(\sum_{n=0}^{\infty} p^{n} U_{n}(x, \tau)\right)\right.  \tag{11}\\
& \left.\quad+\sum_{n=0}^{\infty} p^{n} \mathscr{H}_{n}(U)\right] d \tau
\end{align*}
$$

with

$$
\begin{equation*}
T(x, t)=\sum_{k=0}^{m-1} \frac{t^{k}}{k!} y_{i}(x) \tag{12}
\end{equation*}
$$

Relating the terms of same powers of $p$, this gives solutions of various orders. The initial guess of the approximation is $T(x, t)$ that is actually the Taylor series of the exact solution of order $m$. Note that this initial guess insures the uniqueness of the series decompositions [9].

## 3. Application of the HDM to the Model with Integer-Order Derivative

In this section, we employ this method for deriving the set of the mathematical equations describing the tuberculosis disease population dynamics model.

Resulting from the steps involved in the HDM method, we reach at the following integral equations that are very simple to solve:

$$
\begin{aligned}
& p^{0}: S_{0}(t)=S(0), \\
& p^{0}: I_{L 0}(t)=I_{L}(0), \\
& p^{0}: I_{A 0}(t)=I_{A}(0), \\
& p^{1}: S_{1}(t) \\
& =\int_{0}^{t}\left(v f N-\alpha I_{A 0} S_{0}(\tau)+\delta S_{0}(\tau)\right. \\
& \left.+T_{A} I_{A 0}(\tau)+T_{L} I_{L 0}(\tau)\right) d \tau, \quad S_{1}(0)=0, \\
& p^{1}: I_{L 1}(t) \\
& =\int_{0}^{t}\left((1-P) \alpha I_{A 0} S_{0}(\tau)-\beta_{A} I_{L 0}(\tau)\right. \\
& \left.-T_{L} I_{L} 0(\tau)-\delta I_{L 0}(\tau)\right) d \tau, \quad I_{L 1}(0)=0, \\
& p^{1}: I_{A 1}(t) \\
& =\int_{0}^{t}\left(P \alpha I_{A 0} S_{0}(\tau)+\beta_{A} I_{L 0}(\tau)\right. \\
& \left.-T_{A} I_{A 0}(\tau)-\delta I_{A 0}(\tau)-\varepsilon I_{A 0}(\tau)\right) d \tau, \quad I_{A 1}(0)=0, \\
& p^{n}: S_{n}(t) \\
& =\int_{0}^{t}\left(v f N-\alpha \sum_{j=0}^{n-1} I_{A j} S_{n-j-1}(\tau)+\delta S_{n-1}(\tau)\right. \\
& \left.+T_{A} I_{A(n-1)}(\tau)+T_{L} I_{L(n-1)}(\tau)\right) d \tau, \\
& S_{n-1}(0)=0, \\
& p^{n}: I_{L n}(t) \\
& =\int_{0}^{t}\left((1-P) \alpha \sum_{j=0}^{n-1} I_{A j} S_{n-j-1}(\tau)-\beta_{A} I_{L(n-1)}(\tau)\right.
\end{aligned}
$$

$$
\begin{array}{r}
\left.-T_{L} I_{L(n-1)}(\tau)-\delta I_{L(n-1)}(\tau)\right) d \tau \\
I_{L n}(0)=0
\end{array}
$$

$$
\begin{align*}
& p^{n}: I_{A n}(t) \\
&=\int_{0}^{t}\left(P \alpha \sum_{j=0}^{n-1} I_{A j} S_{n-j-1}(\tau)+\beta_{A} I_{L(n-1)}(\tau)-T_{A} I_{A(n-1)}(\tau)\right. \\
&\left.\quad-\delta I_{A(n-1)}(\tau)-\varepsilon I_{A(n-1)}(\tau)\right) d \tau, \quad I_{A n}(0)=0 . \tag{13}
\end{align*}
$$

Integrating the previous, we obtain the following components:

$$
\begin{gather*}
S_{0}(t)=S(0) ; \quad I_{L 0}(t)=I_{L}(0) \\
I_{A 0}(t)=I_{A}(0), \\
S_{1}(t)=\left(v f N-\alpha I_{A 0} S_{0}+\delta S_{0}+T_{A} I_{A 0}+T_{L} I_{L 0}\right) t \\
I_{L 1}(t)=\left((1-P) \alpha I_{A 0} S_{0}-\beta_{A} I_{L 0}-T_{L} I_{L} 0-\delta I_{L 0}\right) t \\
I_{A 1}(t)=\left(P \alpha I_{A 0} S_{0}+\beta_{A} I_{L 0}-T_{A} I_{A 0}-\delta I_{A 0}-\varepsilon I_{A 0}\right) t \tag{14}
\end{gather*}
$$

For simplicity, let us put

$$
\begin{gather*}
a=\left(v f N-\alpha I_{A 0} S_{0}+\delta S_{0}+T_{A} I_{A 0}+T_{L} I_{L 0}\right), \\
b=\left((1-P) \alpha I_{A 0} S_{0}-\beta_{A} I_{L 0}-T_{L} I_{L} 0-\delta I_{L 0}\right), \\
c=\left(P \alpha I_{A 0} S_{0}+\beta_{A} I_{L 0}-T_{A} I_{A 0}-\delta I_{A 0}-\varepsilon I_{A 0}\right), \\
S_{2}(t)=\frac{1}{2} t^{2}\left(b T_{A}+c T_{L}-a I_{A 0} \alpha-b S_{0} \alpha+a \delta\right) \\
=\frac{t^{2}}{2} a_{1},  \tag{15}\\
I_{L 2}(t)=\frac{1}{2} t^{2}\left(-c T_{L}+a I_{A 0} \alpha-a I_{A 0} P \alpha\right. \\
\left.\quad+b S_{0} \alpha-c \beta_{A}-c \delta\right)=\frac{t^{2}}{2} b_{1}, \\
I_{A 2}(t)=\frac{1}{2} t^{2}\left(a I_{A 0} P \alpha+b P \alpha S_{0}-b T_{A}\right. \\
\left.\quad+c \beta_{A}-b \delta-b \varepsilon\right)=c_{1} \frac{t^{2}}{2} .
\end{gather*}
$$

In general, we obtain the following recursive formulas:

$$
\begin{align*}
& S_{n}(t)=\frac{t^{n}}{n!} a_{n} \\
& I_{L n}(t)=\frac{t^{n}}{n!} b_{n}  \tag{16}\\
& I_{A n}(t)=c_{n} \frac{t^{n}}{n!}
\end{align*}
$$

where $a_{n}, b_{n}$, and $c_{n}$ depend on the fixed set of empirical parameters. It therefore follows that the approximate solution of the system (1) is given as

$$
\begin{align*}
& S_{N}(t)=\sum_{n=0}^{N} \frac{t^{n}}{n!} a_{n}, \\
& I_{L N}(t)=\sum_{n=0}^{N} \frac{t^{n}}{n!} b_{n},  \tag{17}\\
& I_{A N}(t)=\sum_{n=0}^{N} \frac{t^{n}}{n!} c_{n} .
\end{align*}
$$

If for instance one supposes that the total number of new people in the location of interest is $N=100$; the initial number of susceptible people in the location is $S(0)=96$; the initial number of TB latently infected people is $I_{L}(0)=3$; the initial number of TB actively infected people is $I_{A}(0)=1$; the probability that a susceptible person is not vaccinated is $v=0.5$; the efficient rate of vaccines is $f=0.5$; the success rate of latent TB therapy is $T_{L}=0.8$; the active TB treatment cure rate is $T_{A}=0.74$; the TB instantaneous incidence rate per susceptible is $\alpha=0.41$; humans natural death rate is $\delta=1 /(366 \times 70)$; the proportion of infection instantaneously degenerating into active TB is $P=0.0197$; the TB-induced death rate is $\varepsilon=0.0735$; and the breakdown rate from latent to active TB is $\beta_{A}=0.01$, then the following approximate solution is obtained as a result of the first 8 terms of the series decomposition:

$$
\begin{align*}
& S(t)= 96-11.2162 t+62.1069 t^{2}-29.5924 t^{3}-149.2 t^{4} \\
&+48.3455 t^{5}-20.6378 t^{6}+15.5857 t^{7}+\cdots \\
& I_{L}(t)=3+36.8527 t-62.9161 t^{2}-797.302 t^{3}+151.174 t^{4} \\
&-48.8926 t^{5}+20.7629 t^{6}-15.6036 t^{7}+\cdots \\
& I_{A}(t)=1-0.706394 t+0.252053 t^{2}-0.252832 t^{3} \\
&-1.96203 t^{4}+0.573666 t^{5}-0.131459 t^{6} \\
&+0.0190148 t^{7}+\cdots . \tag{18}
\end{align*}
$$

If in addition we assume that no new person migrates or is born in this area, we obtain the following figures. The approximate solutions of the main problem are depicted in Figures 1, 2, and 3, respectively.

Figure 1 shows that, if there is migration or newborn in the location of interest, the number of susceptible people will vanish as time goes, because of the natural death rate and due to TB. Note that any person that is latently infected is removed from the set of susceptible. Figure 2 indicates that the number of people that are latently infected will increase up to a certain time and then vanish as time goes. The number of susceptible people, will become latently infected since some are not vaccinated against the TB and finally will vanish due to. Figure 3 indicates that the number of TB actively infected


Figure 1: Approximate solution for the number of susceptible people in the location.


Figure 2: Approximate solution for the number of TB latently infected people.
people will also vanish because of the natural death rate and the death due to TB.

## 4. Application of the HDM to the Model with Noninteger-Order Derivative

Fractional calculus has been used to model physical and engineering processes, which are found to be best described by fractional differential equations. It is worth noting that the standard mathematical models of integer-order derivatives, including nonlinear models, do not work adequately in many cases. In the recent years, fractional calculus has played a very important role in various fields such as mechanics, electricity, chemistry, biology, economics, notably control theory, and signal and image processing. Major topics


Figure 3: Approximate solution for the number of TB actively infected people.
include anomalous diffusion; vibration and control; continuous time random walk; Levy statistics, fractional Brownian motion; fractional neutron point kinetic model; power law; Riesz potential; fractional derivative and fractals; computational fractional derivative equations; nonlocal phenomena; history-dependent process; porous media; fractional filters; biomedical engineering; fractional phase-locked loops, and groundwater problem (see [14-21]).

### 4.1. Properties and Definitions

Definition 1. A real function $f(x), x>0$, is said to be in the space $C_{\mu}, \mu \in \mathbb{R}$, if there exists a real number $p>\mu$, such that $f(x)=x^{p} h(x)$, where $h(x) \in C[0, \infty)$, and it is said to be in space $C_{\mu}^{m}$ if $f^{(m)} \in C_{\mu}, m \in \mathbb{N}$.

Definition 2. The Riemann-Liouville fractional integral operator of order $\alpha \geq 0$, of a function $f \in C_{\mu}, \mu \geq-1$, is defined as

$$
\begin{gather*}
J^{\alpha} f(x)=\frac{1}{\Gamma(\alpha)} \int_{0}^{x}(x-t)^{\alpha-1} f(t) d t, \quad \alpha>0, x>0  \tag{19}\\
J^{0} f(x)=f(x)
\end{gather*}
$$

Properties of the operator can be found in [14-16]. We mention only the following: for $f \in C_{\mu}, \mu \geq-1, \alpha, \beta \geq 0$, and $\gamma>-1$,

$$
\begin{gather*}
J^{\alpha} J^{\beta} f(x)=J^{\alpha+\beta} f(x) \\
J^{\alpha} J^{\beta} f(x)=J^{\beta} J^{\alpha} f(x) J^{\alpha} x^{\gamma}=\frac{\Gamma(\gamma+1)}{\Gamma(\alpha+\gamma+1)} x^{\alpha+\gamma} \tag{20}
\end{gather*}
$$

Lemma 3. If $m-1<\alpha \leq m, m \in \mathbb{N}$, and $f \in C_{\mu}^{m}, \mu \geq-1$, then

$$
\begin{gather*}
D^{\alpha} J^{\alpha} f(x)=f(x) \\
J^{\alpha} D_{0}^{\alpha} f(x)=f(x)-\sum_{k=0}^{m-1} f^{(k)}\left(0^{+}\right) \frac{x^{k}}{k!}, \quad x>0 \tag{21}
\end{gather*}
$$

Definition 4 (partial derivatives of fractional order). Assume now that $f(\mathbf{x})$ is a function of $n$ variables $x_{i} i=1, \ldots, n$ also of class $C$ on $D \in \mathbb{R}_{n}$. We define partial derivative of order $\alpha$ for $f$ respect to $x_{i}$ the function

$$
\begin{equation*}
a \partial_{\underline{\mathbf{x}}}^{\alpha} f=\left.\frac{1}{\Gamma(m-\alpha)} \int_{a}^{x_{i}}\left(x_{i}-t\right)^{m-\alpha-1} \partial_{x_{i}}^{m} f\left(x_{j}\right)\right|_{x_{j}=t} d t \tag{22}
\end{equation*}
$$

where $\partial_{x_{i}}^{m}$ is the usual partial derivative of integer-order $m$.
4.2. Approximate Solution of Fractional Version. The system of equations under investigation here is given as

$$
\begin{align*}
\frac{d^{\mu} S(t)}{d t^{\mu}}= & v f N-\alpha I_{A} S(t)+\delta S(t) \\
& +T_{A} I_{A}(t)+T_{L} I_{L}(t), \quad 0<\mu \leq 1, \\
\frac{d^{\eta} I_{L}(t)}{d t^{\eta}}= & (1-P) \alpha I_{A} S(t)-\beta_{A} I_{L}(t)  \tag{23}\\
& -T_{L} I_{L}(t)-\delta I_{L}(t), \quad 0<\eta \leq 1 \\
\frac{d^{v} I_{A}(t)}{d t^{v}}= & P \alpha I_{A} S(t)+\beta_{A} I_{L}(t)-T_{A} I_{A}(t) \\
& -\delta I_{A}(t)-\varepsilon I_{A}(t), \quad 0<v \leq 1
\end{align*}
$$

Following the discussion presented earlier, we arrive at the following equations:

$$
\begin{aligned}
& p^{0}: S_{0}(t)=S(0), \\
& p^{0}: I_{L 0}(t)=I_{L}(0), \\
& p^{0}: I_{A 0}(t)=I_{A}(0), \\
& p^{1}: S_{1}(t) \\
& =\frac{1}{\Gamma(\mu)} \int_{0}^{t}(t-\tau)^{\mu-1} \\
& \times\left(v f N-\alpha I_{A 0} S_{0}(\tau)+\delta S_{0}(\tau)\right. \\
& \left.+T_{A} I_{A 0}(\tau)+T_{L} I_{L 0}(\tau)\right) d \tau, \\
& S_{1}(0)=0, \\
& p^{1}: I_{L 1}(t) \\
& =\frac{1}{\Gamma(\eta)} \int_{0}^{t}(t-\tau)^{\eta-1} \\
& \times\left((1-P) \alpha I_{A 0} S_{0}(\tau)-\beta_{A} I_{L 0}(\tau)\right. \\
& \left.-T_{L} I_{L} 0(\tau)-\delta I_{L 0}(\tau)\right) d \tau, \\
& I_{L 1}(0)=0 \text {, } \\
& p^{1}: I_{A 1}(t) \\
& =\frac{1}{\Gamma(v)} \int_{0}^{t}(t-\tau)^{v-1} \\
& \times\left(P \alpha I_{A 0} S_{0}(\tau)+\beta_{A} I_{L 0}(\tau)\right. \\
& \left.-T_{A} I_{A 0}(\tau)-\delta I_{A 0}(\tau)-\varepsilon I_{A 0}(\tau)\right) d \tau, \\
& I_{A 1}(0)=0,
\end{aligned}
$$

$$
\begin{align*}
& p^{n}: S_{n}(t) \\
& =\frac{1}{\Gamma(\mu)} \int_{0}^{t}(t-\tau)^{\mu-1} \\
& \times\left(v f N-\alpha \sum_{j=0}^{n-1} I_{A j} S_{n-j-1}(\tau)+\delta S_{n-1}(\tau)\right. \\
& \left.+T_{A} I_{A(n-1)}(\tau)+T_{L} I_{L(n-1)}(\tau)\right) d \tau, \\
& S_{n-1}(0)=0, \\
& p^{n}: I_{L n}(t) \\
& =\frac{1}{\Gamma(\eta)} \int_{0}^{t}(t-\tau)^{\eta-1} \\
& \times\left((1-P) \alpha \sum_{j=0}^{n-1} I_{A j} S_{n-j-1}(\tau)-\beta_{A} I_{L(n-1)}(\tau)\right. \\
& \left.-T_{L} I_{L(n-1)}(\tau)-\delta I_{L(n-1)}(\tau)\right) d \tau, \\
& I_{L n}(0)=0, \\
& p^{n}: I_{A n}(t) \\
& =\frac{1}{\Gamma(v)} \int_{0}^{t}(t-\tau)^{v-1} \\
& \times\left(P \alpha \sum_{j=0}^{n-1} I_{A j} S_{n-j-1}(\tau)+\beta_{A} I_{L(n-1)}(\tau)\right. \\
& -T_{A} I_{A(n-1)}(\tau)-\delta I_{A(n-1)}(\tau) \\
& \left.-\varepsilon I_{A(n-1)}(\tau)\right) d \tau, \\
& I_{A n}(0)=0 . \tag{24}
\end{align*}
$$

Integrating the previous, we obtain the following components:

$$
\begin{gathered}
S_{0}(t)=S(0) ; \quad I_{L 0}(t)=I_{L}(0) ; \\
I_{A 0}(t)=I_{A}(0), \\
S_{1}(t)=-\frac{11.2162 t^{\mu}}{\Gamma(1+\mu)} ; \\
I_{L 1}(t)=\frac{36.8527 t^{\eta}}{\Gamma(1+\eta)}, \\
I_{A 1}(t)=-\frac{0.706394 t^{v}}{\Gamma(1+v)}, \\
S_{2}(t)=t^{\mu}\left(\frac{29.4822 t^{\eta}}{\Gamma(1+\eta+\mu)}+\frac{4.59822 t^{\mu}}{\Gamma(1+2 \mu)}\right. \\
\left.+\frac{27.2809 t^{v}}{\Gamma(1+v+\mu)}\right),
\end{gathered}
$$

$$
\begin{gathered}
I_{L 2}(t)=-t^{\eta}\left(\frac{29.8522 t^{\eta}}{\Gamma(1+2 \eta)}+\frac{4.58965 t^{\mu}}{\Gamma(1+\eta+\mu)}\right. \\
\left.\quad+\frac{27.7492 t^{v}}{\Gamma(1+v+\mu)}\right) \\
I_{A 2}(t)=t^{v}\left(\frac{0.368527 t^{\eta}}{\Gamma(1+\eta+v)}-\frac{0.00901337 t^{\mu}}{\Gamma(1+v+\mu)}\right. \\
\left.+\frac{0.520184 t^{v}}{\Gamma(1+2 v)}\right)
\end{gathered}
$$

$$
S_{3}(t)=t^{\mu}\left(-\frac{3.24846 t^{\mu+v} \Gamma(1+\mu+v)}{\Gamma(1+\mu) \Gamma(1+v) \Gamma(1+2 \mu+v)}\right.
$$

$$
-\frac{0.298522 t^{2 \eta}}{\Gamma(1+2 \eta+v)}-\frac{15.7583 t^{\eta+\mu}}{\Gamma(1+\eta+2 \mu)}
$$

$$
-\frac{1.88509 t^{2 \mu}}{\Gamma(1+3 \mu)}-\frac{36.4319 t^{\eta+v}}{\Gamma(1+\mu+v+\eta)}
$$

$$
\left.-\frac{10.836 t^{\mu+v}}{\Gamma(1+2 \mu+v)}-\frac{20.0895 t^{2 v}}{\Gamma(1+\mu+2 v)}\right)
$$

$$
\begin{gather*}
I_{L 3}(t)=t^{\eta}\left(-\frac{1148.5 t^{2 \eta}}{\Gamma(1+3 \eta)}-\frac{164.513 t^{\eta+\mu}}{\Gamma(1+2 \eta+\mu)}+\frac{1.88158 t^{2 \mu}}{\Gamma(1+\eta+2 \mu)}\right. \\
-\frac{1067.59 t^{\eta}}{\Gamma(1+2 \eta+v)}+\frac{11.1633 t^{\mu+v}}{\Gamma(1+\eta+\mu+v)} \\
\left.+\frac{3.2421 t^{\mu+v} \Gamma(1+\mu+v)}{\Gamma(1+\mu) \Gamma(1+v) \Gamma(1+\eta+\mu+v)}\right) \\
I_{A 3}(t)=t^{v}\left(\frac{0.00636699 t^{\mu+v} \Gamma(1+\mu+v)}{\Gamma(1+\mu) \Gamma(1+v) \Gamma(1+2 v+\mu)}\right. \\
\quad-\frac{0.298522 t^{2 \eta}}{\Gamma(1+2 \eta+v)}-\frac{0.0222046 t^{\eta+\mu}}{\Gamma(1+\eta+2 \mu)} \\
+\frac{0.00369513 t^{2 \mu}}{\Gamma(1+3 \mu)}-\frac{0.548873 t^{\eta+v}}{\Gamma(1+\mu+2 v)} \\
\left.+\frac{0.0285603 t^{\mu+v}}{\Gamma(1+\mu+2 v)}-\frac{0.38306 t^{2 v}}{\Gamma(1+3 v)}\right) \tag{25}
\end{gather*}
$$

The remaining terms can be obtained in the same manner. But here only few terms of the series solutions are considered, and the asymptotic solution is given as

$$
\begin{gather*}
S(t)=S_{0}(t)+S_{1}(t)+S_{2}(t)+S_{3}(t)+\cdots, \\
I_{L}(t)=I_{L 0}(t)+I_{L 1}(x, t)+I_{L 2}(x, t)+I_{L 3}(x, t)+\cdots \\
I_{A}(t)=I_{A 0}(t)+I_{A 1}(x, t)+I_{A 2}(x, t)+I_{A 3}(x, t)+\cdots \tag{26}
\end{gather*}
$$

The following figures show the simulated solutions for different values of the fractional order derivatives. The approximate


Figure 4: Approximate for $\mu=0.45, \eta=0.7$, and $v=0.85$.


Figure 5: Approximate for $\mu=0.45, \eta=0.7$, and $v=0.85$.


Figure 6: Approximate for $\mu=0.45, \eta=0.7$, and $v=0.85$.


Figure 7: Approximate for $\mu=0.045, \eta=0.5$, and $v=0.085$.


Figure 8: Approximate for $\mu=0.045, \eta=0.5$, and $v=0.085$.
solutions of the main problem are depicted in Figures 4, 5, 6, 7,8 , and 9 , respectively.

The numerical simulations show that the approximate solutions are continuous functions of the noninteger-order derivative. It is worth noting that the standard mathematical models of integer-order derivatives, including nonlinear models, do not work adequately in many cases. It is therefore advisable to use the fractional model for describing this problem.

## 5. Conclusion

The tuberculosis model was examined for the case of integerand noninteger-order derivatives. Both systems of nonlinear equations were solved with an iterative analytical model


Figure 9: Approximate for $\mu=0.045, \eta=0.5$, and $v=0.085$.
called the homotopy decomposition model method. The basic characters of the relatively new technique are presented in detail. The approximate solutions of the noninteger case are increasing continuous functions of the fractional order derivative. The technique used for solving these problems is friendly, very easy, and less time consuming. The numerical solutions in both cases display the biological behaviour of the real world situation.

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# A Rectangular Mixed Finite Element Method with a Continuous Flux for an Elliptic Equation Modelling Darcy Flow 

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Received 26 March 2013; Accepted 29 May 2013
Academic Editor: Santanu Saha Ray
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#### Abstract

We introduce a mixed finite element method for an elliptic equation modelling Darcy flow in porous media. We use a staggered mesh where the two components of the velocity and the pressure are defined on three different sets of grid nodes. In the present mixed finite element, the approximate velocity is continuous and the conservation law still holds locally. The LBB consistent condition is established, while the $L 2$ error estimates are obtained for both the velocity and the pressure. Numerical examples are presented to confirm the theoretical analysis.


## 1. Introduction

We consider the discretization technique for the elliptic problem modelling the flow in saturated porous media, or the classical Darcy flow problem, including a system of mass conservation law and Darcy's law [1, 2]. The most popular numerical methods for this elliptic equation focus on mixed finite element methods, since by this kind of methods the original scalar variable, called pressure, and its vector flux, named Darcy velocity, can be approximated simultaneously and maintain the local conservation. The classical theory for the mixed finite element, which includes the LBB consistent condition, the existence and uniqueness of the approximate solution, and the error estimate, has been established. Some mixed finite element methods such as RT mixed finite element and BDM mixed finite element are introduced (as in [36]), which satisfy the consistent condition and have optimal order error estimate [7, 8]. Give some stabilized mixed finite methods by adding to the classical mixed formulation some least squares residual forms of the governing equations.

By using the abovementioned mixed finite element methods, the approximate velocity is continuous in the normal direction and discontinuous in the tangential direction on the edges of the element. This is reasonable for the case of heterogenous permeability, yet it is desirable that the flux be continuous in some applications [9]. In particular, when
we track the characteristic segment using the approximate velocity, the discontinuities of the velocity may introduce some difficulties when the characteristic line cross the edges of element. While applying mass-conservative characteristic finite element method to the coupled system of compressible miscible displacement in porous media, the continuous flux is crucial [10]. A brief description of this point will be found at the last part of this paper.

To overcome this disadvantage, Arbogast and Wheeler [11] introduced a mixed finite element method with an approximate velocity continuous in both the normal direction and the tangential direction, which was got by adding some freedom to the RT mixed finite element. In this paper, we introduced a mixed finite element method with an approximate velocity continuous in all directions. It is based on rectangular mesh and uses continuous piecewise bilinear functions to approximate the velocity components and uses piecewise constant functions to approximate the pressure. We obtain the element by improving a kind of element for Stokes equation and Navier-Stokes equation given by Han [12], Han and Wu [13], and Han and Yan [14]. By using this mixed finite element, we can get continuous velocity vector and maintain the local conservation. Comparing to the mixed finite element method in [11], we need less degrees of freedom for the same convergence rate. The LBB consistent condition and L2 error estimates of velocity and pressure are also established.

The outline of the rest of this paper is organized as follows. In Sections 2 and 3, we recall the model problem and weak formulation for the mixed finite element method and then establish the discrete inf-sup consistent condition and L2 error estimates for the velocity and the pressure in Section 4. In Section 5, we present some numerical examples which verify the efficiency of the proposed mixed finite element method. A valuable application of this method to mass-conservative characteristic (MCC) scheme for the coupled compressible miscible displacement in porous media closes the paper in Section 6.

## 2. The Mixed Finite Formulation for Darcy Equation

The mathematical model for viscous flow in porous media includes Darcy's law and conservation law of mass, written as follows:

$$
\begin{array}{ll}
u=-\frac{\kappa}{\mu} \nabla p & \text { on } \Omega \text { (Darcy's law) } \\
\operatorname{div} u=\phi & \text { on } \Omega \text { (mass conservation) }  \tag{1}\\
u \cdot n=0 & \text { on } \Gamma
\end{array}
$$

where $\kappa>0$ is the permeability, $\mu>0$ is the viscosity, and $\phi$ is the volumetric flow rate source or sink. $\Gamma$ is the boundary of $\Omega$, and $n$ is the unit outward normal vector to $\Gamma$. The variable $u=\left(u_{1}, u_{2}\right)$ is the Darcy velocity vector, and $p$ is the pressure. The source $\phi$ must satisfy the consistency constraint

$$
\begin{equation*}
\int_{\Omega} \phi d \Omega=0 \tag{2}
\end{equation*}
$$

Let $L^{2}(\Omega)$ be the space of square integrable function in $\Omega$ with inner product $(\cdot, \cdot)$ and norm $\|\cdot\|$. We use the notation of the Hilbert space

$$
\begin{equation*}
H(\operatorname{div}, \Omega)=\left\{u \in\left[L^{2}(\Omega)\right]^{2} ; \operatorname{div} u \in L^{2}(\Omega)\right\} \tag{3}
\end{equation*}
$$

with norm

$$
\begin{equation*}
\|u\|_{H(\operatorname{div}, \Omega)}=\left\{\|u\|^{2}+\|\operatorname{div} u\|^{2}\right\}^{1 / 2} \tag{4}
\end{equation*}
$$

Define the following subspaces of $H(\operatorname{div}, \Omega)$ and $L^{2}(\Omega)$ :

$$
\begin{gather*}
V=H_{0}(\operatorname{div}, \Omega)=\{u \in H(\operatorname{div}, \Omega): u \cdot n=0 \text { on } \Gamma\}, \\
S=\left\{q \mid q \in L^{2}(\Omega): \int_{\Omega} q d \Omega=0\right\} . \tag{5}
\end{gather*}
$$

The classical weak variational formulation of Problem (1) is as follows: find $(u, p) \in V \times S$, such that

$$
\begin{align*}
& a(u, v)-b(v, p)=0 \quad \forall v \in V \\
& b(u, q)=(\phi, q) \quad \forall q \in S \tag{6}
\end{align*}
$$

Here,

$$
\begin{equation*}
a(u, v)=\int_{\Omega} \frac{\mu}{\kappa} u \cdot v d x \quad b(v, q)=\int_{\Omega} q \operatorname{div} v d x \tag{7}
\end{equation*}
$$

The following discussion and discrete analysis are related to the weak form (6). Let $V_{0}$ be a closed subspace of $V$ via

$$
\begin{equation*}
V_{0}=\{v \in V: b(v, q)=0, \forall q \in S\} . \tag{8}
\end{equation*}
$$

For the bilinear forms $a(u, v)$ and $b(v, q)$, we have the standard result.

Lemma 1. The bilinear form $a(u, v)$ is bounded on $V \times V$ and coercive on $V_{0}$, and the bilinear form $b(v, q)$ is bound on $V \times S$. Namely,
(1) there exist two constants $C_{1}>0$ and $\alpha>0$ such that

$$
\begin{gather*}
|a(u, v)| \leq C_{1}\|u\|_{H(\operatorname{div}, \Omega)}\|v\|_{H(\operatorname{div}, \Omega)} \quad \forall u, v \in V, \\
a(u, u) \geq \alpha\|u\|_{H(\operatorname{div}, \Omega)}^{2} \quad \forall u \in V_{0} \tag{9}
\end{gather*}
$$

(2) there is a constant $C_{2}>0$ such that

$$
\begin{equation*}
|b(v, q)| \leq C_{2}\|q\|_{0, \Omega}\|v\|_{H(\operatorname{div}, \Omega)} \quad \forall q \in S, v \in V \tag{10}
\end{equation*}
$$

For the space $V$ and $S$, the Ladyzhenskaya-Babus̆kaBrezzi( $L-B-B$ ) condition holds; see [15, 16], for example.

Lemma 2. There is a constant $\beta>0$ such that

$$
\begin{equation*}
\sup _{v \in V} \frac{b(v, q)}{\|v\|_{H(\operatorname{div}, \Omega)}} \geq \beta\|q\|_{0, \Omega}, \quad \forall q \in S . \tag{11}
\end{equation*}
$$

It is clear that there exists a unique solution $(u, p) \in V \times S$ to the Problem (6).

## 3. Finite Element Discretization

In this section, we present the mixed finite element based on rectangular mesh for the Darcy flow problem.

In [13], Han and Wu introduced a mixed finite element for Stokes problem and then extended to solve the Navier-Stokes problem [14]. Based on this element, we introduced the new mixed finite element with a continuous flux approximation for Darcy flow problem.

For simplicity, we suppose that the domain $\Omega$ is a unit square, and the mixed finite element discussed here can be easily generalized to the case when the domain $\Omega$ is a rectangular.

Let $N$ be a given integer and $h=1 / N$. We construct the finite-dimensional subspaces of $S$ and $V$ by introducing three different quadrangulations $\tau_{h}, \tau_{h}^{1}, \tau_{h}^{2}$ of $\Omega$.

First, we divide $\Omega$ into uniform squares

$$
\begin{array}{r}
T_{i, j}=\left\{(x, y): x_{i-1} \leq x \leq x_{i}, y_{j-1} \leq y \leq y_{j}\right\} \\
i, j=1, \ldots, N \tag{12}
\end{array}
$$

where $x_{i}=i h$ and $y_{j}=j h$. The corresponding quadrangulation is denoted by $\tau_{h}$. See Figure 1(a).

$$
\begin{equation*}
\tau_{i, j}=\left\{T_{i, j}: i, j=1, \ldots, N\right\} . \tag{13}
\end{equation*}
$$



Figure 1: Quadrangulations: (a) $\tau_{h}$, (b) $\tau_{h}^{1}$, and (c) $\tau_{h}^{2}$.

Then, for all $T_{i, j} \in \tau_{h}$, we connect all the neighbor midpoints of the vertical sides of $T_{i, j}$ by straight segments if the neighbor midpoints have the same vertical coordinate. Then, $\Omega$ is divided into squares and rectangles. The corresponding quadrangulation is denoted by $\tau_{h}^{1}$ (see Figure 1(b)). Similarly, for all $T_{i, j} \in \tau_{h}$, we connect all the neighbor midpoints of the horizontal sides of $T_{i, j}$ by straight line segments if the neighbor midpoints have the same horizontal coordinate. Then, we obtained the third quadrangulation of $\Omega$, which is denoted by $\tau_{h}^{2}$ (see Figure 1(c)).

Based on the quadrangulation $\tau_{h}$, we define the piecewise constant functional space used to approximate the pressure

$$
\begin{equation*}
S_{h}:=\left\{q_{h}:\left.q_{h}\right|_{T}=\text { constant }, \forall T \in \tau_{h} ; \int_{\Omega} q_{h} d x=0\right\} \tag{14}
\end{equation*}
$$

$S_{h}$ is a subspace of $S$.
Furthermore, using the quadrangulations $\tau_{h}^{1}$ and $\tau_{h}^{2}$, we construct a subspace of $V$. Denote by $\Gamma_{1}, \Gamma_{2}, \Gamma_{3}$, and $\Gamma_{4}$ the south, right, north, and left sides on the boundary of $\Omega$. Set

$$
\begin{aligned}
V_{h}^{1}= & \left\{v_{h} \in C^{(0)}(\bar{\Omega}):\left.v_{h}\right|_{T^{1}} \in Q_{1,1}\left(T^{1}\right) \forall T^{1} \in \tau_{h}^{1},\right. \\
& \left.v_{h}=0 \text { on } \Gamma_{2} \cup \Gamma_{4}\right\}, \\
V_{h}^{2}= & \left\{v_{h} \in C^{(0)}(\bar{\Omega}):\left.v_{h}\right|_{T^{2}} \in Q_{1,1}\left(T^{2}\right) \forall T^{2} \in \tau_{h}^{2},\right. \\
& \left.v_{h}=0 \text { on } \Gamma_{1} \cup \Gamma_{3}\right\},
\end{aligned}
$$

where $Q_{1,1}$ denotes the piecewise bilinear polynomial space with respect to the variables $x$ and $y$. Let

$$
\begin{equation*}
V_{h}=V_{h}^{1} \times V_{h}^{2} \tag{16}
\end{equation*}
$$

Obviously, $V_{h}$ is a subspace of $V$.
Using the subspaces $V_{h}$ and $S_{h}$ instead of $V$ and $S$ in the variational Problem (6), we obtain the discrete problem: find $\left(u_{h}, p_{h}\right) \in V_{h} \times S_{h}$, such that

$$
\begin{align*}
& a\left(u_{h}, v_{h}\right)-b\left(v_{h}, p_{h}\right)=0 \quad \forall v_{h} \in V_{h},  \tag{17}\\
& b\left(u_{h}, q_{h}\right)=\left(\phi, q_{h}\right) \quad \forall q_{h} \in S_{h} .
\end{align*}
$$

## 4. Convergence Analysis and Error Estimate

In this section, we give the corresponding convergence analysis and error estimate. Firstly, we define an interpolating for the following analysis.

For the quadrangulation $\tau_{h}$, we divided the edges of all squares into two sets. The first one denoted by $L_{V}$ contains all vertical edges, and the second one denoted by $L_{H}$ contains all horizontal edges. We define the interpolation operator
$\Pi: V \rightarrow V_{h}$ by $\Pi u=\left(\Pi_{h}^{1} u_{1}, \Pi_{h}^{2} u_{2}\right) \in V_{h}^{1} \times V_{h}^{2}$, which satisfy the following:

$$
\begin{array}{ll}
\int_{l} \Pi_{h}^{1} u_{1} d s=\int_{l} u_{1} d s & \forall l \in L_{V^{\prime}} \\
\int_{l} \Pi_{h}^{2} u_{2} d s=\int_{l} u_{2} d s & \forall l \in L_{H^{\prime}} \tag{18}
\end{array}
$$

where $L_{V^{\prime}}$ is a set of edges of elements got by bisecting the most bottom element edges and the most top element edges of $L_{V}$ and $L_{H^{\prime}}$ are got by bisecting the most left element edges and the most right element edges of $L_{H}$. See Figures 2 and 3.

Lemma 3. For any $u \in V$, the interpolating $\Pi u \in V_{h}$ is unique$l y$ determined by (18).

Proof. It is easy to see that (18) is equivalent to an equation of $A X=B$, where $A$ is a matrix and $X, B$ are vectors. Direct calculation shows that

$$
\begin{equation*}
A=h * \operatorname{diag}\left\{A_{1}, A_{1}, \ldots\right\} \tag{19}
\end{equation*}
$$

and the form of submatrix $A_{1}$ is as follows

$$
\left(\begin{array}{cccccccccc}
\frac{1}{4} & \frac{1}{4} & 0 & 0 & 0 & \cdots & 0 & 0 & 0 & 0  \tag{20}\\
0 & \frac{3}{8} & \frac{1}{8} & 0 & 0 & \cdots & 0 & 0 & 0 & 0 \\
0 & \frac{1}{8} & \frac{3}{4} & \frac{1}{8} & 0 & \cdots & 0 & 0 & 0 & 0 \\
0 & 0 & \frac{1}{8} & \frac{3}{4} & \frac{1}{8} & \cdots & 0 & 0 & 0 & 0 \\
\cdots & \cdots & \cdots & \cdots & \cdots & \cdots & \cdots & \cdots & \cdots & \\
0 & 0 & 0 & 0 & 0 & \cdots & \frac{1}{8} & \frac{3}{4} & \frac{1}{8} & 0 \\
0 & 0 & 0 & 0 & 0 & \cdots & 0 & \frac{1}{8} & \frac{3}{8} & 0 \\
0 & 0 & 0 & 0 & 0 & \cdots & 0 & 0 & \frac{1}{4} & \frac{1}{4}
\end{array}\right) .
$$

We can see that the matrix is invertible and the equation is solvable, and therefore $X$ can be uniquely determined.

Assume that the solution ( $u, p$ ) of Problem (6) has the following smoothness properties:

$$
\begin{equation*}
u \in V^{\prime}:=V \bigcap H^{2}((\Omega))^{2}, \quad p \in S \bigcap H^{1}(\Omega) \tag{21}
\end{equation*}
$$

Then, we should give the following lemma about the properties of the interpolations defined in (18).

Lemma 4. (i) There exist two constants $C_{3}$ and $C_{4}$ independent of $h$, such that

$$
\begin{gather*}
|u-\Pi u|_{i, 2, \Omega} \leq C_{3} h^{j-i}|u|_{j, 2, \Omega}, \quad i=0,1, \quad i \leq j \leq 2  \tag{22}\\
\inf _{q_{h} \in S_{h}}\left\|p-q_{h}\right\| \leq C_{4} h|p|_{1, \Omega} \tag{23}
\end{gather*}
$$

(ii) There exists a constant $C_{5}$ independent of $h$ such that

$$
\begin{equation*}
\|\Pi u\|_{H(\operatorname{div}, \Omega)} \leq C_{5}\|u\|_{1, \Omega} \quad \forall u \in V . \tag{24}
\end{equation*}
$$



Figure 2: Some edges on $L_{V}$ and corresponding edges on $L_{V^{\prime}}$.
(iii) For any $u \in V$, we have that

$$
\begin{equation*}
\int_{\Omega} q_{h} \operatorname{div}(u-\Pi u) d x=0, \quad \forall q_{h} \in S_{h} . \tag{25}
\end{equation*}
$$

Proof. The estimates (22), (23), and (24) follow from Definition (18) and the approximation theory; see [1], for example.

For (25), based on Green formulation, we know that

$$
\begin{align*}
\int_{\Omega} q_{h} \operatorname{div}(u-\Pi u) d x= & \sum_{T \in \tau_{h}} \int_{T} q_{h} \operatorname{div}(u-\Pi u) d x \\
= & \sum_{T \in \tau_{h}} \int_{\partial T} q_{h}(u-\Pi u) \cdot \vec{n} d s \\
& -\sum_{T \in \tau_{h}} \int_{T} \nabla q_{h} \cdot(u-\Pi u) d x \\
= & \sum_{l \in L_{V}} \int_{l} q_{h}\left(u_{1}-\Pi_{h}^{1} u_{1}\right) n_{1} d s \\
& +\sum_{l \in L_{H}} \int_{l} q_{h}\left(u_{2}-\Pi_{h}^{2} u_{2}\right) n_{2} d s \\
= & \sum_{l \in L_{V^{\prime}}} \int_{l} q_{h}\left(u_{1}-\Pi_{h}^{1} u_{1}\right) n_{1} d s \\
& +\sum_{l \in L_{H^{\prime}}} \int_{l} q_{h}\left(u_{2}-\Pi_{h}^{2} u_{2}\right) n_{2} d s \\
= & 0 . \tag{26}
\end{align*}
$$

Here, $\vec{n}=\left(n_{1}, n_{2}\right)$, and we use (18) for the last step. The proof is completed.


Figure 3: Some edges on $L_{H}$ and corresponding edges on $L_{H^{\prime}}$.

Theorem 5. The discrete Inf-sup condition is valid; namely, there is a constant $\beta \geq 0$, such that

$$
\begin{equation*}
\sup _{v_{h} \in V_{h}} \frac{b\left(v_{h}, q_{h}\right)}{\left\|v_{h}\right\|_{H(\operatorname{div}, \Omega)}} \geq \beta\left\|q_{h}\right\|, \quad \forall q_{h} \in S_{h} . \tag{27}
\end{equation*}
$$

Proof. From the process above, we obtain that $b\left(v, q_{h}\right)=$ $b\left(\Pi v, q_{h}\right)$, any $v \in V, q_{h} \in S_{h}$. For any $p_{h} \in S_{h}$, there exists $v \in\left(H_{0}^{1}(\Omega)\right)^{2}$, such that

$$
\begin{equation*}
\nabla \cdot v=q_{h}, \quad\|v\|_{1, \Omega} \leq C_{6}\left\|q_{h}\right\|, \tag{28}
\end{equation*}
$$

where $C_{6}$ is a constant independent of $q_{h}$; then we obtain

$$
\begin{align*}
\sup _{v_{h} \in V_{h}} \frac{b\left(v_{h}, q_{h}\right)}{\left\|v_{h}\right\|_{H(\operatorname{div}, \Omega)}} & \geq \frac{b\left(\Pi v, q_{h}\right)}{\|\Pi v\|_{H(\operatorname{div}, \Omega)}} \\
& =\frac{b\left(v, q_{h}\right)}{\|\Pi v\|_{H(\operatorname{div}, \Omega)}}  \tag{29}\\
& =\frac{\left\|q_{h}\right\|_{0}^{2}}{\|\Pi v\|_{H(\operatorname{div}, \Omega)}} .
\end{align*}
$$

Using Lemma 4, we have that

$$
\begin{equation*}
\sup _{v_{h} \in V_{h}} \frac{b\left(v_{h}, q_{h}\right)}{\left\|v_{h}\right\|_{H(\operatorname{div}, \Omega)}} \geq \frac{1}{C_{5}} \frac{\left\|q_{h}\right\|_{0}^{2}}{\|v\|_{1, \Omega}} \geq \frac{1}{C_{5} C_{6}}\left\|q_{h}\right\| . \tag{30}
\end{equation*}
$$

Taking $\beta=1 / C_{5} C_{6}$, we complete the proof of (27).
With the analysis technique presented by Arbogast and Wheeler [11], we consider the numerical analysis of the mixed finite element presented in this paper. Recall the $R T_{0}$ mixed element spaces $V_{h}^{\prime} \times S_{h}^{\prime}[3,5,6]$ based on the partition $\tau_{h}$

$$
\begin{equation*}
V_{h}^{\prime}=Q_{1,0}\left(\tau_{h}\right) \times Q_{0,1}\left(\tau_{h}\right), \quad S_{h}^{\prime}=S_{h} \tag{31}
\end{equation*}
$$

Define the interpolation operator $\Pi^{\prime}: V \rightarrow V_{h}^{\prime}$ by the following equations:

$$
\begin{array}{ll}
\int_{l} \Pi^{\prime} u_{1} d s=\int_{l} u_{1} d s & \forall l \in L_{V}  \tag{32}\\
\int_{l} \Pi^{\prime} u_{2} d s=\int_{l} u_{2} d s & \forall l \in L_{H}
\end{array}
$$

Denote by $P_{S}: S \rightarrow S_{h}$ the $L^{2}$ projection operator and by $P_{V^{\prime}}: V \rightarrow V_{h}^{\prime}$ the $\left(L^{2}(\Omega)\right)^{2}$ vector projection operator. The following properties of the projections hold:

$$
\begin{align*}
& \left\|p-P_{S} p\right\|_{0} \leq C h|p|_{1}  \tag{33}\\
& \left\|u-P_{V^{\prime}} u\right\|_{0} \leq C h\|u\|_{1} .
\end{align*}
$$

Then, we have an important property about the operator $\Pi^{\prime}$.

Lemma 6. For any $u \in Q_{1,1}\left(\tau_{h}^{1}\right) \times Q_{1,1}\left(\tau_{h}^{2}\right)$, there holds the equivalence $\Pi^{\prime} u=P_{V^{\prime}} u$; namely,

$$
\begin{equation*}
\left(u-\Pi^{\prime} u, v\right)=0, \quad \forall v \in V_{h}^{\prime} \tag{34}
\end{equation*}
$$

Proof. As the definition of $V_{h}^{\prime}$ is based on each element $T$, we focus our discussion on arbitrary element $e \subset \tau_{h}, e=\left[x_{0}, x_{0}+\right.$ $h] \times\left[y_{0}, y_{0}+h\right]$. Firstly, we consider the $x$-component (see Figure 4). The analysis for $y$-component is similar.

For a function $U_{1} \in V_{h}^{1}$, on an element $e$, it is uniquely given by its node values $u_{i}, i=1, \ldots, 6$. As $U_{1}$ is a continuous bilinear function on each of the two parts as shown in Figure 4. Then, from (32), we know that $\Pi^{\prime} u_{1}=a+b x$ is given by

$$
\begin{align*}
& \int_{l_{1}}(a+b x) d s=\left(a+b x_{0}\right) * h=\left(u_{1}+2 u_{3}+u_{5}\right) * \frac{h}{4} \\
& \int_{l_{2}}(a+b x) d s=\left(a+b\left(x_{0}+h\right)\right) * h=\left(u_{2}+2 u_{4}+u_{6}\right) * \frac{h}{4} \tag{35}
\end{align*}
$$

We deduce that

$$
\begin{gather*}
a=\frac{u_{1}+2 u_{3}+u_{5}-4 b x_{0}}{4} \\
b=\left(\left(u_{2}-u_{1}\right)+2\left(u_{4}-u_{3}\right)+\left(u_{6}-u_{5}\right)\right) * \frac{1}{4 h} . \tag{36}
\end{gather*}
$$

It is clear that we just need to verify (34) for both $v=1$ and $v=x$.

We first consider $v=1$. Denote by $\varphi_{i}$ the node basis function at the point $i$, which implies that $\varphi_{i}\left(x_{j}\right)=\delta_{i, j}$, which has the value 1 if and only if $i=j$; otherwise, it is zero. By direct calculation, we can get the basis, for example,

$$
\begin{equation*}
\varphi_{1}=\frac{1}{4}\left(2-\frac{2}{h} x+\frac{2}{h} x_{0}\right)\left(2-\frac{4}{h} y+\frac{4}{h} y_{0}\right) \tag{37}
\end{equation*}
$$

so

$$
\begin{align*}
& \int_{e} U_{1} d x d y \\
&= \int_{x_{0}}^{x_{0}+h} \int_{y_{0}}^{y_{0}+\frac{h}{2}}\left(u_{1} \varphi_{1}+u_{2} \varphi_{2}+u_{3} \varphi_{3}+u_{4} \varphi_{4}\right) d x d y \\
&+\int_{x_{0}}^{x_{0}+h} \int_{y_{0}+h / 2}^{y_{0}+h}\left(u_{3} \varphi_{3}+u_{4} \varphi_{4}+u_{5} \varphi_{5}+u_{6} \varphi_{6}\right) d x d y \\
&= \frac{h^{2}}{8}\left(u_{1}+u_{2}+u_{3}+u_{4}\right)+\frac{h^{2}}{8}\left(u_{3}+u_{4}+u_{5}+u_{6}\right) \\
&= \frac{h^{2}}{8}\left(u_{1}+u_{2}+2 u_{3}+2 u_{4}+u_{5}+u_{6}\right) . \tag{38}
\end{align*}
$$



Figure 4: An element on $\tau_{h}$ and its corresponding portion on $\tau_{h}^{1}$.

By direct computation, we can easily see that $\int_{e} \Pi^{\prime} U_{1} d x d y$ has the same value, so

$$
\begin{equation*}
\int_{e} \Pi^{\prime} U_{1} d x d y=\int_{e} U_{1} d x d y \tag{39}
\end{equation*}
$$

When $v=x$, we have that

$$
\begin{align*}
\int_{e}^{\Pi^{\prime} U_{1} * x d x d y=} & \int_{x_{0}}^{x_{0}+h} \int_{y_{0}}^{y_{0}+h} a x+b x^{2} d x d y \\
= & a\left(x_{0} h^{2}+\frac{h^{3}}{2}\right)  \tag{40}\\
& +b\left(x_{0}^{2} h^{2}+x_{0} h^{3}+\frac{1}{3} h^{4}\right)
\end{align*}
$$

where $a, b$ are defined in (36). Next, we compare the coefficients of $u_{i}$ in (40) with the coefficients in $\int_{e} U_{1} * x d x d y$,

$$
\begin{align*}
\int_{e} \varphi_{1} * x d x d y & =\int_{y_{0}}^{y_{0}+h / 2} \int_{x_{0}}^{x_{0}+h} \varphi * x d x d y \\
& =\frac{h}{2} \int_{x_{0}}^{x_{0}+h} \frac{1}{4}\left(2 x-\frac{2}{h} x^{2}+\frac{2}{h} x_{0} x\right)  \tag{41}\\
& =\frac{1}{24} h^{3}+\frac{1}{8} x_{0} h^{2}=k_{1}
\end{align*}
$$

which determine $k_{1}$ as the coefficient of $u_{1}$. With similar computation, we obtain that

$$
\begin{align*}
& k_{5}=k_{1}, \quad k_{2}=k_{6}=\frac{x_{0} h^{2}}{8}+\frac{h^{3}}{12}, \\
& k_{3}=\frac{x_{0} h^{2}}{4}+\frac{h^{3}}{12}, \quad k_{4}=\frac{x_{0} h^{2}}{4}+\frac{h^{3}}{6} . \tag{42}
\end{align*}
$$

Comparing with (40), we can find that (34) is true with $v=x$. So, we certify the lemma.

Theorem 7. If $(u, p)$ satisfy (6) and ( $u_{h}, p_{h}$ ) satisfy (17), then there exists a positive constant $C$ independent of $h$ such that the following error estimates hold:

$$
\begin{gather*}
\left\|u-u_{h}\right\|_{0} \leq C h\|u\|_{1}  \tag{43}\\
\left\|p-p_{h}\right\|_{0} \leq C h\left(\|u\|_{1}+\|p\|_{1}\right) .
\end{gather*}
$$

Proof. First, we focus on the error $u-u_{h}$. From (6), (17), (18), and (32), we derive that
$b\left(u, q_{h}\right)=b\left(u_{h}, q_{h}\right)=b\left(\Pi u, q_{h}\right)=b\left(\Pi^{\prime} u, q_{h}\right), \quad \forall q_{h} \in S_{h}$.

Let $v=\Pi^{\prime} v_{h}$ in (6); then

$$
\begin{equation*}
a\left(u, \Pi^{\prime} v_{h}\right)-b\left(\Pi^{\prime} v_{h}, p\right)=0 \tag{45}
\end{equation*}
$$

Namely,

$$
\begin{equation*}
a\left(P_{v^{\prime}} u, v_{h}\right)-\left(P_{s} \nabla \cdot v_{h}, p\right)=a\left(P_{v^{\prime}} u, v_{h}\right)-b\left(v_{h}, P_{s} p\right)=0 \tag{46}
\end{equation*}
$$

Here, we used the property $\nabla \cdot \Pi^{\prime} v=P_{s} \nabla \cdot v$. Subtracting from (17), we get that

$$
\begin{equation*}
a\left(P_{v^{\prime}} u-u_{h}, v_{h}\right)-b\left(v_{h}, P_{s} p-p_{h}\right)=0 . \tag{47}
\end{equation*}
$$

Take

$$
\begin{equation*}
v_{h}=\Pi u-u_{h}, \quad q_{h}=P_{s} p-p_{h} \tag{48}
\end{equation*}
$$

Then
$a\left(P_{v^{\prime}} u-u_{h}, \Pi u-u_{h}\right)-b\left(\Pi u-u_{h}, P_{s} p-p_{h}\right)=0$.
Due to (44), we find that

$$
\begin{equation*}
b\left(\Pi u-u_{h}, P_{s} p-p_{h}\right)=0 . \tag{50}
\end{equation*}
$$

Now, we analyze the error $u-u_{h}$ based on the equations above

$$
\begin{align*}
a(u- & \left.u_{h}, u-u_{h}\right) \\
= & a\left(u-u_{h}, u-\Pi u\right)+a\left(u-u_{h}, \Pi u-u_{h}\right) \\
= & a\left(u-u_{h}, u-\Pi u\right)+a\left(u-P_{v^{\prime}} u, \Pi u-u_{h}\right) \\
& +a\left(P_{v^{\prime}} u-u_{h}, \Pi u-u_{h}\right) \\
\leq & \epsilon_{1}\left\|u-u_{h}\right\|_{0}^{2}+\frac{1}{\epsilon_{1}}\|u-\Pi u\|_{0}^{2}  \tag{51}\\
& +\epsilon_{2}\|\Pi u-u\|_{0}^{2}+\frac{1}{\epsilon_{2}}\left\|u-P_{v^{\prime}} u\right\|_{0}^{2} \\
& +\epsilon_{3}\left\|u-u_{h}\right\|_{0}^{2}+\frac{1}{\epsilon_{3}}\left\|u-P_{v^{\prime}} u\right\|_{0}^{2}
\end{align*}
$$

where $\epsilon_{i}>0, i=1,2,3$ are positive constants. Take the value of $\epsilon_{1}=\epsilon_{3}=\mu / 4 \kappa, \epsilon_{2}=1$, and combining with (22) and (33), we conclude that

$$
\begin{equation*}
\left\|u-u_{h}\right\|_{0} \leq C h\|u\|_{1} . \tag{52}
\end{equation*}
$$

We also can obtain a higher order error estimate for $\left\|P_{s} p-p_{h}\right\|$. Consider the classical duality argument. Let $\phi$ be the solution of the following elliptical problem:

$$
\begin{equation*}
\Delta \phi=P_{s} p-p_{h}, \quad \frac{\partial \phi}{\partial n}=0 . \tag{53}
\end{equation*}
$$

By the elliptic regularity, the estimate holds: $|\phi|_{H^{2}} \leq$ $C\left\|P_{s} p-p_{h}\right\|_{0}$. So

$$
\begin{align*}
&\left\|P_{s} p-p_{h}\right\|_{0}^{2} \\
&=\left(P_{s} p-p_{h}, \nabla \cdot \nabla \phi\right) \\
&=\left(P_{s} p-p_{h}, \nabla \cdot \Pi \nabla \phi\right) \\
&= a\left(P_{v^{\prime}} u-u_{h}, \Pi \nabla \phi\right) \\
&= a\left(P_{v^{\prime}} u-u_{h}, \Pi \nabla \phi-P_{v^{\prime}} \nabla \phi\right)+a\left(P_{v^{\prime}} u-u_{h}, P_{v^{\prime}} \nabla \phi\right) \\
&= a\left(P_{v^{\prime}} u-u_{h}, \Pi \nabla \phi-P_{v^{\prime}} \nabla \phi\right)+a\left(P_{v^{\prime}} u-u, P_{v^{\prime}} \nabla \phi\right) \\
&+a\left(u-u_{h}, P_{v^{\prime}} \nabla \phi\right) \\
&= a\left(P_{v^{\prime}} u-u_{h}, \Pi \nabla \phi-P_{v^{\prime}} \nabla \phi\right)+a\left(u-u_{h}, P_{v^{\prime}} \nabla \phi-\nabla \phi\right) \\
&+a\left(u-u_{h}, \nabla \phi\right) . \tag{54}
\end{align*}
$$

Now, we estimate the right hand terms of the above inequality. From (33), (22), and (52), we have

$$
\begin{align*}
a\left(P_{v^{\prime}} u-u_{h}, \Pi \nabla \phi-P_{v^{\prime}} \nabla \phi\right)= & a\left(P_{v^{\prime}} u-u, \Pi \nabla \phi-\nabla \phi\right) \\
& +a\left(u-u_{h}, \Pi \nabla \phi-\nabla \phi\right) \\
& +a\left(P_{v^{\prime}} u-u, \nabla \phi-P_{v^{\prime}} \nabla \phi\right) \\
& +a\left(u-u_{h}, \nabla \phi-P_{v^{\prime}} \nabla \phi\right) \\
\leq & C h^{2}\|u\|_{1}|\phi|_{H^{2}} \\
\leq & C h^{2}\|u\|_{1} \mid\left\|P_{s} p-p_{h}\right\|_{0} . \tag{55}
\end{align*}
$$

It is easy to see that

$$
\begin{align*}
& a\left(u-u_{h}, P_{v^{\prime}} \nabla \phi-\nabla \phi\right) \leq C h^{2}\|u\|_{1}|\phi|_{H^{2}} \\
& \quad \leq C h^{2}\|u\|_{1}\left\|P_{s} p-p_{h}\right\|_{0^{\prime}} \\
& \begin{aligned}
& a\left(u-u_{h}, \nabla \phi\right) \\
&=a(u-\Pi u, \nabla \phi)+a\left(\Pi u-u_{h}, \nabla \phi-P_{v^{\prime}} \nabla \phi\right) \\
&+a\left(\Pi u-u_{h}, P_{v^{\prime}} \nabla \phi\right) \\
& \leq C\left(h^{2}|u|_{2}\|\phi\|_{2}+h^{2}|u|_{1}\|\phi\|_{2}\right) \\
& \leq C h^{2}\|u\|_{2}\left\|P_{s} p-p_{h}\right\|_{0} .
\end{aligned}
\end{align*}
$$

Here, we used the fact that $a\left(\Pi u-u_{h}, P_{v^{\prime}} \nabla \phi\right)=0$ which is got from the Green formulation and (44).

Combining the above inequalities, we conclude that

$$
\begin{align*}
\left\|p-p_{h}\right\|_{0} & \leq\left\|p-P_{s} p\right\|_{0}+\left\|P_{s} p-p_{h}\right\|_{0} \\
& \leq \operatorname{Ch}\left(\|u\|_{1}+\|p\|_{0}\right) \tag{57}
\end{align*}
$$

We complete the proof.
It is worth mentioning that we analyze this mixed finite element method in a direct way as it is not straightforward to apply the classical inf-sup theory. We just have the coercivity property for $a\left(u_{h}, v_{h}\right)$ on the normal $L_{2}$ space, not in the subspace of $v_{0 h}=\left\{v_{h} \in V_{h}: b\left(v_{h}, q_{h}\right)=0, \forall q_{h} \in S_{h}\right\}$, and the same issue also occurs in [11]. The problem is that testing $(\nabla \cdot v, w)$ by $w \in W_{h}$ does not control the full divergence of $V$, and it does not occur when this method is applied to Stokes or Navier-Stokes equations (as in [13, 14]). As a result, we just obtain a convergence rate of $\left\|u-u_{h}\right\|_{0}$. Failing to obtain convergence rate of $\left\|u-u_{u}\right\|_{H(\operatorname{div}, \Omega)}$ is a weak point of this proposed mixed formulation compared to the classical Raviart-Thomas mixed method. But the significance of continuous flux applied to mass conservation can be found in Section 6.

## 5. Numerical Examples

In this section, we present some numerical results for the model Problem (1). For simplicity, we assume that the domain

Table 1: Three numerical test cases.

| Case | Coefficient $\mu / \kappa$ | True solution $u$ | True solution $p$ |
| :--- | :---: | :---: | :---: |
| 1 | 1 | $\binom{x^{2} y-x^{4} y}{x y^{4}-x y^{2}}$ | $(x-1 / 2)(y-1 / 2)$ |
| 2 | $\left(\begin{array}{cc}e^{2 x y^{2}} & 0 \\ 0 & \frac{1}{x+y}\end{array}\right)$ | $\binom{x^{2} y-x^{4} y}{x y^{4}-x y^{2}}$ | $(x-1 / 2)(y-1 / 2)$ |
| 3 | $\left(\begin{array}{cc}e^{2 x y^{2}} & 0 \\ 0 & \frac{1}{x+y}\end{array}\right)$ | $\binom{e^{-x y}}{x^{2} \cos y}$ | $y e^{x}$ |

is a unit square $\Omega=[0,1] \times[0,1]$ and the test cases are summarized in Table 1. We can choose the boundary conditions and the right hand terms according to the analytical solutions.

We compare our method to the formulation constructed by Arbogast and Wheeler [11]. Its corresponding discrete finite element spaces are

$$
\begin{gather*}
\bar{V}_{h}=\left\{v_{h} \in\left(C^{(0)}(\bar{\Omega})\right)^{2}:\left.v_{h}\right|_{T} \in Q_{1,2}(T)\right. \\
\left.\times Q_{2,1}(T), \forall T \in \tau_{h}\right\}, \\
\bar{S}_{h}=\left\{q_{h}:\left.q_{h}\right|_{T}=\right.\text { constant }  \tag{58}\\
\left.\forall T \in \tau_{h} ; \int_{\Omega} q_{h} d x=0\right\}
\end{gather*}
$$

The results of the error estimate with various norms are listed in Table 2, while the corresponding convergence rates of the presented method are shown in Table 3.

Close results of numerical errors for both formulations are shown in Table 2. From Table 3, we can see that $p$ converges to $p_{h}$ as $O(h)$ and $P_{s} p-p_{h}$ as $O\left(h^{2}\right)$ for our formulation, which both agree with the theorem. From the examples, we can observe that $u_{h}$ converges to $u$ somewhat better than expected, and it appears that on the uniform grid we attain $O\left(h^{3 / 2}\right)$ superconvergence in the $L^{2}$ norm which is similar to the tests of Arbogast's formulation [11]. Yet, the degrees of freedom of our method are less than Arbogast's scheme. As in the case of $64 * 64$, the degrees of freedom of Arbogast's scheme are 20866 and 12676 for our formulation. The convergence rate of $\left\|u-u_{h}\right\|_{H(\operatorname{div}, \Omega)}$ is first order, but here we cannot give the corresponding analysis.

## 6. A Valuable Application

In this section, we briefly show an application of the proposed mixed finite element method to the miscible displacement of one incompressible fluid by another in porous media. The model is as follows:

$$
\begin{gather*}
\mu(C) K^{-1} u+\nabla p=\gamma(C) \nabla d, \quad(x, t) \in \Omega \times J, \\
\phi \frac{\partial C}{\partial t}+\nabla \cdot(u C)-\nabla \cdot(D(u) \nabla C)=\widetilde{C} q, \quad(x, t) \in \Omega \times J, \\
\nabla \cdot u=g, \quad(x, t) \in \Omega \times J, \\
u \cdot n=g_{1}, \quad(x, t) \in \partial \Omega \times J, \\
C(x, 0)=C_{0}(x), \quad x \in \Omega \tag{59}
\end{gather*}
$$

where $\gamma(C)$ and $d$ are the gravity coefficient and vertical coordinate, $\phi(x)$ is the porosity of the rock, and $\widetilde{C} q$ represents a known source. $D(x, u)$ is the molecular diffusion and mechanical dispersion coefficient. For convenience, we denote that $f=\widetilde{C} q$ and $a(C)=\mu(C) K^{-1}$. Let $\chi:(0, T] \rightarrow R^{2}$ be the solution of the ordinary differential equation

$$
\begin{gather*}
\frac{d \chi}{d \tau}=\frac{\mathbf{u}(\chi(x, t ; \tau), \tau)}{\phi(x)},  \tag{60}\\
\chi(x, t ; t)=x .
\end{gather*}
$$

Let $V=H(\operatorname{div}, \Omega), S=L_{0}^{2}(\Omega), M=H^{1}(\Omega)$; then, we derive the entire weak formulation for the model: find $(\mathbf{u}, p, C) \in$ $V \times S \times M$, such that

$$
\begin{align*}
& (a(C) u, v)-(p, \nabla \cdot v)=(\gamma(C) \nabla d, \mathbf{v}), \quad \forall \mathbf{v} \in V \\
& \left(\phi(x) \frac{d C(\chi, \tau)}{d \tau}+g C, w\right)+(D \nabla C, \nabla w)=(f, w) \tag{61}
\end{align*}
$$

$\forall w \in M$,

$$
(\nabla \cdot u, \varphi)=(g, \varphi), \quad \forall \varphi \in S
$$

Let $\Delta t$ be the time step for both concentration and pressure; define

$$
\begin{equation*}
M_{h}=\left\{v_{h} \in C^{(0)}(\bar{\Omega}):\left.v_{h}\right|_{T} \in Q_{1,1}(T), \forall T \in \tau_{h}\right\} . \tag{62}
\end{equation*}
$$

Combing with the new characteristic finite element method which preserves the mass balance proposed by Rui and Tabata [10], the approximate characteristic line of $\chi$ is defined as

$$
\begin{equation*}
\chi^{n}(x)=x-\frac{u_{h}^{n}}{\phi(x)} \Delta t \tag{63}
\end{equation*}
$$

We obtain the corresponding full-discrete mass-conservative characteristic (MCC) scheme: find $\left(u_{h}, p_{h}, C_{h}\right) \in V_{h} \times S_{h} \times M_{h}$, such that

$$
\begin{gather*}
\left(a\left(C_{h}^{n}\right) u_{h}^{n}, v_{h}\right)-\left(p_{h}, \nabla \cdot v_{h}\right) \\
=\left(\gamma\left(C_{h}^{n}\right) \nabla d, v_{h}\right), \quad \forall v_{h} \in V_{h} \\
\left(\frac{\phi C_{h}^{n}-\left(\phi C_{h}^{n-1}\right) \circ \chi^{n} \gamma^{n}}{\Delta t}, \varphi_{h}\right)+\left(D\left(u_{h}^{n}\right) \nabla C_{h}^{n}, \nabla \varphi_{h}\right) \\
=\left(f, \varphi_{h}\right), \quad \forall \varphi_{h} \in M_{h} \\
\left(\nabla \cdot u_{h}^{n}, q_{h}\right)=\left(g, q_{h}\right), \quad \forall q_{h} \in S_{h} \\
C_{h}^{0}=\widetilde{C}^{0}, \tag{64}
\end{gather*}
$$

where

$$
\begin{align*}
\gamma^{n}= & \operatorname{det}\left(\frac{\partial \chi^{n}}{\partial x}\right) \\
= & 1-\frac{\nabla \cdot u_{h}^{n}}{\phi} \Delta t+u_{h}^{n} \frac{\nabla \phi}{\phi^{2}} \Delta t  \tag{65}\\
& +\nabla\left(\frac{u_{h, 1}^{n}}{\phi}\right) \cdot \operatorname{curl}\left(\frac{u_{h, 2}^{n}}{\phi}\right) \Delta t^{2} .
\end{align*}
$$

Table 2: The numerical error for fm. 1 (our formulation) and fm. 2 (Arbogast's formulation).

| Case | Mesh | $\left\\|u-u_{h}\right\\|$ |  | $\left\\|\nabla \cdot\left(u-u_{h}\right)\right\\|$ |  | $\left\\|p-p_{h}\right\\|$ |  | $\left\\|P_{s} p-p_{h}\right\\|$ |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  |  | fm. 1 | fm. 2 | fm. 1 | fm. 2 | fm. 1 | fm. 2 | fm. 1 | fm. 2 |
| 1 | 4 | $4.90 e-2$ | $5.67 e-2$ | $3.06 e-1$ | $3.24 e-1$ | $2.93 e-2$ | $2.93 e-2$ | $4.53 e-3$ | $4.17 e-3$ |
|  | 8 | 1.78 e-2 | $2.05 e-2$ | $1.53 e-1$ | $1.62 e-1$ | $1.47 e-2$ | $1.47 e-2$ | $1.24 e-3$ | $1.20 e-3$ |
|  | 16 | $6.45 e-3$ | $7.37 e-3$ | $7.67 e-2$ | $8.13 e-2$ | $7.37 e-3$ | $7.37 e-3$ | $3.18 e-4$ | $3.15 e-4$ |
|  | 32 | $2.31 e-3$ | $2.64 e-3$ | $3.84 e-2$ | $4.07 e-2$ | $3.68 e-3$ | 3.68 e-3 | $8.01 e-5$ | $7.98 e-5$ |
|  | 64 | $8.25 e-4$ | $9.38 e-4$ | $1.92 e-2$ | $2.03 e-2$ | $1.84 e-3$ | 1.84e-3 | $2.01 e-5$ | $2.01 e-5$ |
| 2 | 4 | $4.70 e-2$ | $5.47 e-2$ | $2.99 e-1$ | $3.22 e-1$ | $2.95 e-2$ | 2.94e-2 | $5.42 e-3$ | $4.97 e-3$ |
|  | 8 | $1.72 e-2$ | $1.99 e-2$ | $1.53 e-1$ | $1.63 e-1$ | $1.47 e-2$ | $1.47 e-2$ | $1.54 e-3$ | $1.48 e-3$ |
|  | 16 | $6.25 e-3$ | $7.19 e-3$ | 7.75 - - 2 | $8.27 e-2$ | $7.37 e-3$ | $7.37 e-3$ | $4.04 e-4$ | $3.98 e-4$ |
|  | 32 | $2.25 e-3$ | $2.58 e-3$ | $3.89 e-2$ | 4.15 - 2 | $3.68 e-3$ | 3.68 e-3 | $1.03 e-4$ | $1.02 e-4$ |
|  | 64 | $8.08 e-4$ | $9.21 e-4$ | $1.95 e-2$ | $2.08 e-2$ | $1.84 e-3$ | $1.84 e-3$ | $2.59 e-5$ | $2.58 e-5$ |
| 3 | 4 | 9.65 - 2 | $1.09 e-1$ | $4.14 e-1$ | $4.67 e-1$ | $1.49 e-1$ | $1.49 e-1$ | $7.39 e-3$ | $6.21 e-3$ |
|  | 8 | $3.79 e-2$ | $4.31 e-2$ | $2.16 e-1$ | $2.46 e-1$ | $7.44 e-2$ | $7.44 e-2$ | $2.14 e-3$ | $1.89 e-3$ |
|  | 16 | $1.42 e-2$ | $1.62 e-2$ | 1.11e-1 | $1.28 e-1$ | $3.72 e-2$ | $3.72 e-2$ | $5.72 e-4$ | $5.19 e-4$ |
|  | 32 | $5.19 e-3$ | $5.91 e-3$ | 5.63 - - 2 | $6.51 e-2$ | $1.86 e-2$ | $1.86 e-2$ | $1.47 e-4$ | 1.35 e-4 |
|  | 64 | $1.87 e-3$ | $2.13 e-3$ | $2.84 e-2$ | $3.28 e-2$ | $9.31 e-3$ | $9.31 e-3$ | $3.72 e-5$ | $3.44 e-5$ |

Table 3: The corresponding convergence rates of fm. 1 and fm. 2.

| Case | Mesh | $\left\\|u-u_{h}\right\\|$ |  | $\left\\|\nabla \cdot\left(u-u_{h}\right)\right\\|$ |  | $\left\\|p-p_{h}\right\\|$ |  | $\left\\|P_{s} p-p_{h}\right\\|$ |  |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
|  |  | fm. 1 | fm. 2 | fm. 1 | fm. 2 | fm. 1 | fm. 2 | fm. 1 | fm. 2 |
| 1 | 8 | 1.459 | 1.468 | 0.997 | 1.001 | 0.995 | 0.993 | 1.875 | 1.795 |
|  | 16 | 1.468 | 1.476 | 0.998 | 0.995 | 0.999 | 0.999 | 1.961 | 1.934 |
|  | 32 | 1.479 | 1.484 | 1.000 | 0.998 | 1.000 | 1.000 | 1.987 | 1.978 |
|  | 64 | 1.486 | 1.489 | 1.000 | 1.000 | 1.000 | 1.000 | 1.996 | 1.993 |
| 2 | 8 | 1.449 | 1.457 | 0.968 | 0.978 | 0.999 | 0.996 | 1.817 | 1.742 |
|  | 16 | 1.462 | 1.471 | 0.983 | 0.984 | 1.001 | 1.001 | 1.930 | 1.901 |
|  | 32 | 1.471 | 1.479 | 0.993 | 0.993 | 1.001 | 1.000 | 1.976 | 1.960 |
|  | 64 | 1.480 | 1.485 | 0.997 | 0.997 | 1.000 | 1.000 | 1.989 | 1.984 |
| 3 | 8 | 1.347 | 1.340 | 0.942 | 0.924 | 0.998 | 0.997 | 1.787 | 1.708 |
|  | 16 | 1.416 | 1.414 | 0.957 | 0.945 | 0.999 | 0.999 | 1.906 | 1.870 |
|  | 32 | 1.452 | 1.452 | 0.979 | 0.973 | 1.000 | 1.000 | 1.959 | 1.942 |
|  | 64 | 1.472 | 1.473 | 0.990 | 0.986 | 1.000 | 1.000 | 1.983 | 1.975 |

We can see that the continuous flux is indispensable for $\gamma^{n}$. Let $\varphi_{h}=1$ in (64), and summing it up from $n=1$ to $N$, we get the mass balance

$$
\begin{equation*}
\int_{\Omega} \phi C_{h}^{N} d x=\int_{\Omega} \phi C_{h}^{0} d x+\Delta t \sum_{n=1}^{N} \int_{\Omega} f^{n} d x . \tag{66}
\end{equation*}
$$

Here, we just give numerical example to show the feasibility of this application, and the theoretical analysis of stability, mass balance, and convergence of this discrete scheme will be discussed in the future. Firstly, we define compute mass error and relative mass error as follows:

$$
\begin{align*}
\text { compute mass error : } & \int_{\Omega} \phi C_{h}^{N} d x \\
& -\left(\int_{\Omega} \phi C_{h}^{0} d x+\Delta t \sum_{n=1}^{N} \int_{\Omega} f^{n} d x\right) \\
\text { relative mass error : } & \frac{\int_{\Omega} \phi C_{h}^{N} d x-\int_{\Omega} \phi C^{N} d x}{\int_{\Omega} \phi C^{N} d x} \tag{67}
\end{align*}
$$

Now, we select $\mu(C)=C$, and the following analytical solution of the problem is

$$
\begin{gather*}
u(x, y, t)=\left(e^{x}+t, e^{y}+t\right), \\
p(x, y, t)=e^{-t}\left(x^{2}+y^{2}\right),  \tag{68}\\
C(x, y, t)=e^{-t}\left(\left(x-\frac{1}{2}\right)^{2}+\left(y-\frac{1}{2}\right)^{2}\right) .
\end{gather*}
$$

The error results with different norms of this numerical simulation can be listed in Tables 4 and 5, and at last we give a mass error to check the mass conservation in Table 6.

As can be seen from Tables 4 and 5, we conjecture that almost all the convergence rates are true in general. From Table 6 we find that mass balance is right as computational mass error resulting from computer is inevitable and nearly invariable for different meshes, while the relative mass error decreases as was expected. The corresponding theoretical analysis about this system will be considered in the future work.

Table 4: Numerical error and convergence rate ( $\Delta t=C h$ ).

| Mesh | $5 \times 5$ |  |  | $10 \times 10$ |  |  | $20 \times 20$ |  |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Norm type | Error | Rate | Error | Rate | Error | Rate | Error | Rate |
| $\\|u\\|_{l^{2}\left(L^{2}\right)}$ | $1.83 e-4$ | - | $7.10 e-5$ | 1.36 | $3.38 e-5$ | 1.07 | $1.65 e-5$ | 1.03 |
| $\\|u\\|_{l^{\infty}\left(L^{2}\right)}$ | $1.29 e-2$ | - | $5.19 e-3$ | 1.31 | $2.64 e-3$ | 0.97 | $1.37 e-3$ | 0.95 |
| $\\|p\\|_{l^{2}\left(L^{2}\right)}$ | $1.33 e-3$ | - | $6.67 e-4$ | 1.00 | $3.33 e-4$ | 1.00 | $1.67 e-4$ | 1.00 |
| $\\|p\\|_{l^{\infty}\left(L^{2}\right)}$ | $9.43 e-2$ | - | $4.71 e-2$ | 1.00 | $2.35 e-2$ | 1.00 | $1.18 e-2$ | 1.00 |
| $\\|C\\|_{l^{2}\left(H^{1}\right)}$ | $2.32 e-3$ | - | $1.16 e-3$ | 1.01 | $5.78 e-4$ | 1.00 | $2.88 e-4$ | 1.00 |
| $\\|C\\|_{l^{\infty}\left(H^{1}\right)}$ | $1.63 e-1$ | - | $8.18 e-2$ | 1.00 | $4.11 e-2$ | 0.99 | $2.05 e-2$ | 0.99 |

Table 5: Numerical error and convergence rate $\left(\Delta t=C h^{2}\right)$.

| Mesh | $5 \times 5$ |  |  | $10 \times 10$ |  | $20 \times 20$ |  | $40 \times 40$ |  |
| :--- | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Norm type | Error | Rate | Error | Rate | Error | Rate | Error | Rate |  |
| $\\|C\\|_{l^{2}\left(L^{2}\right)}$ | $8.48 e-5$ | - | $2.13 e-5$ | 1.995 | $5.37 e-6$ | 1.986 | $1.36 e-6$ | 1.971 |  |
| $\\|C\\|_{l^{\infty}\left(L^{2}\right)}$ | $1.34 e-2$ | - | $3.37 e-3$ | 1.989 | $8.56 e-4$ | 1.978 | $2.21 e-4$ | 1.952 |  |

Table 6: Mass error for concentration C $(\Delta t=C h)$.

| Mesh | $5 \times 5$ | $10 \times 10$ | $20 \times 20$ | $40 \times 40$ |
| :--- | :---: | :---: | :---: | :---: |
| Compute mass error | $1.209 e-3$ | $1.243 e-3$ | $1.269 e-3$ | $1.284 e-3$ |
| Relative mass error | $2.068 e-2$ | $5.427 e-3$ | $1.487 e-3$ | $4.371 e-4$ |

## Acknowledgment

The work is supported by the National Natural Science Foundation of China Grant no. 11171190.

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